Lambda potential in dense nuclear matter from chiral EFT: Bridging heavy-ion collisions, hypernuclei, and neutron stars

Asanosuke JINNO (Kyoto Univ., Japan, D3)



in collaboration with Koichi Murase (Tokyo Metropolitan Univ.) Yasushi Nara (Akita International Univ.) Akira Ohnishi (YITP)

Y. Nara, <u>AJ</u>, K. Murase, and A. Ohnishi, Phys. Rev. C 106 (2022) 044902. <u>AJ</u>, K. Murase, Y. Nara, and A. Ohnishi, Phys. Rev. C 108 (2023) 065803. AJ, K. Murase, and Y. Nara, arXiv:2501.09881 (2025) (Proceeding for EXA/LEAP2024) + ongoing work

JLab Hypernuclear Collaboration Meeting 2025, 16 May, 2025

- Neutron star matter EOS study and Λ potential
- Evaluation of the Λ potential/YN+YNN forces from chiral EFT
 - **1.** Λ hypernuclear spectroscopy
 - 2. A directed flow v_1 of heavy-ion collisions



Nuclear matter EOS

(Nuclear matter) Equation Of State, EOS, $P(\epsilon, T, etc.)$:

Pressure as a function of the energy density, temperature, etc.

EOS of dense nuclear matter plays an important role in various physics!



Unified approach for neutron star EOS 4/25

A <u>unified approach</u> has provided <u>a strong constraint on EOS</u>, $P = P(\epsilon)$.

Lower density: based on NN + NNN int. from chiral EFT, or well constrained mean field model

Higher density: combining many observational/experimental information

Neutron star observation

e.g. E. Annala et al., Nature Phys. 16, 907 (2020); Y. Fujimoto, K. Fukushima, & K. Murase, PRD 101, 054016 (2020), L. Brandes, Weise, Kaiser, PRD 108, 094104 (2023).

Neutron star observation + heavy ion

(collective flow)

e.g. S. Huth et al., Nature 606 (2022) 276.; N. Rutherford et al.; Astrophys. J. Lett. 971 (2024) L19.





(D) HIC and Astro combined:



S. Huth et al., Nature 606 (2022) 276.

Hyperon composition is important! 5/25

Such an approach does <u>not</u> tell us the <u>detailed properties</u> of EOS.



Microscopic description would give more insight to hadron-quark phase transition and guide to construct finite temperature EOS.

Appearance of <u>hyperons significantly changes the EOS!</u>

cf. Hyperon puzzle of neutron stars, Demorest et al. Nature (2010).





Construction of the Λ potential



- Empirical value based on Λ hypernuclei $U_{\Lambda}(\rho_0) \approx -30$ MeV is fitted
- $\rho \gtrsim 3.5 \rho_0$ is extrapolated by using the anzats

 $U_{\Lambda}(\rho) = u_0 + u_1 \left(\frac{\rho}{\rho_0} - 1\right) + u_2 \left(\frac{\rho}{\rho_0} - 1\right)^2.$

7/25

<u>Validity of the A potential in $\rho < \rho_0$ and $\rho > \rho_0$ should be evaluated!</u>

Purpose of this research

8/25

Evaluating the validity of the strongly repulsive Λ potential based on chiral EFT by using the heavy-ion collision and hypernuclear data.



Unified approach for the Λ potential!

- Neutron star matter EOS study and Λ potential
- Evaluation of the Λ potential/YN+YNN forces from chiral EFT
 - **1.** <u>Λ hypernuclear spectroscopy</u>
 - 2. A directed flow v_1 of heavy-ion collisions



Motivations of using Λ binding energy ^{10/25}

$$\Lambda \text{ binding energy } B_{\Lambda} = \left(A^{-1}Z \right) - \left(A^{-1}Z \right)$$

Behaving differently from the empirical Λ potentials used in Skyrme-Hartree-Fock method. e.g. LY-IV: Lanskoy and Yamamoto (1997)



Motivations of using Λ binding energy 11/25

$$\Lambda \text{ binding energy } \boldsymbol{B}_{\Lambda} = \begin{pmatrix} A^{-1}\boldsymbol{Z} \end{pmatrix} - \begin{pmatrix} A \boldsymbol{Z} \end{pmatrix}$$

NLO result has been tested but does not explain the data well. Haidenbauer and Vidaña (2021) How about including ΛΝΝ int.?



Spherical Skyrme-Hartree-Fock for A hypernuclei <u>12/25</u>



(Note) For neutron and proton, we use SLy4 (Chabanat et al. (1998)). (Note2) CSB, pair correlation, and deformation are neglected.

Fitting of the Λ potential from chiral EFT 13/25



* The value of a_3^{Λ} is determined to reproduce the Λ binding energy of ${}^{13}_{\Lambda}C$ (11.88 MeV).

(:: Surface terms have a large effect. even-even nuclei)

Λ binding energies

<u>AJ</u>, K. Murase, Y. Nara, & A. Ohnishi, PRC 108, 065803 (2023).

 $(a_3^{\Lambda} \text{ in LY-IV model is also tuned to reproduce } {}^{13}_{\Lambda}C \text{ data})$



Chi2 (chiral YN) overbounds a few MeV for *s***-wave.**

14/25



Chi3 reproduces the data, at the same

level of accuracy as LY-IV.

Both the repulsive and attractive Λ potentials are consistent with the data.

Constraints on U_{Λ} from Λ hypernuclei 15/25

AJ, K. Murase, Y. Nara, & A. Ohnishi, PRC 108, 065803 (2023).

Generating Λ potentials and choose ones well reproduces the experimental data. $U_{\Lambda}(\rho \leq \rho_0)$ and m^*_{Λ} are constrained to some extent, while $U_{\Lambda}(\rho > \rho_0)$ is scattered! We need other data to constrain $U_{\Lambda}(\rho > \rho_0)$!



- Neutron star matter EOS study and Λ potential
- Evaluation of the Λ potential/YN+YNN forces from chiral EFT
 - **1.** Λ hypernuclear spectroscopy
 - **2.** <u> Λ directed flow v_1 of heavy-ion collisions</u>



Directed flow v_1 ($\sqrt{s_{NN}} \approx 3-5$ GeV) 17/25

- The anisotropic collective flow v_n = (cosnφ) has been extensively investigated to extract the properties of dense matter equation of states.
 Recent review: A. Sorensen et al., Prog. Part. Nucl. Phys. 134 (2024) 104080.
- Directed flow: $v_1 = \langle \cos \phi \rangle = \langle p_x / p_T \rangle$ as a function of the rapidity $y = \tanh^{-1} \left(\frac{p_z}{E} \right)$ $(p_T^2 = p_x^2 + p_y^2)$



 v_1 has a non-trivial dependence on EOS.

Our previous result in Nara et al. (2022) 18/25



(Extension) Σ potential

 Previous work: Λ potentials for 300 () 200 MeV 100 all hyperons Y. Nara, AJ, K. Murase, & A. Ohnishi, PRC 106, 044902(2022) -50<u>∟</u> • Σ potential may affect the $\Lambda + \Sigma^0$ dynamics in heavy-ion collisions and can be constrained by data! AJ, K. Murase, and Y. Nara, arXiv:2501.09881 (2025) (Proceeding for EXA/LEAP2024)



19/25

Extensions in JAM2 after Nara et al. (2022) 20/25

Several (minor) extensions have been made after the Λ v1

paper: Y. Nara, AJ, K. Murase, & A. Ohnishi, PRC 106, 044902(2022).

New covariant RQMD

model (RQMDv2) Y. Nara, AJ, K. Murase, in preparation. QM2025 Poster

Covariant collision term

Y. Nara, <u>AJ</u>, T. Maruyama, K. Murase, and A. Ohnishi, Phys. Rev. C 108, 024910 (2023).

• YN cross section by chiral N2LO

Action for new RQMD

 Λ + Σ 0 v1 is still similar.

 $S = \int dt \langle \Phi | i\hbar \frac{\partial}{\partial t} - \hat{H} | \Phi \rangle$

Motivated by the time-dependent variational principle:

$$S = S_{\text{part}} + S_{\text{field}}$$
 $S_{field} = \int d^4 x \mathcal{L}_{\text{field}}$

$$S_{part} = \sum_{i=1}^{N} \int p_i \cdot dx_i - \int d^4x d^4p W(x, p) f(x, p)$$

Generalized potential: $W(x, p) = \frac{p^{*2}(x, p) - m^{*2}(x, p)}{2}$

Potential contribution: $p^{*\mu} = p^{\mu} - U^{\mu}, \quad m^* = m + U_s$ $f(x,p) = \sum_{i=1}^N f_i(x,p) = \sum_{i=1}^N \int ds \lambda_i(s) \delta((x-x_i(s)) \cdot \hat{a})) g(x,p;x_i(s),p_i(s))$

 $\lambda_i(s)$ is a Lagrange multiplier and g is a Gaussian profile.

Σ single-particle potentials

21/25

For fitting procedure, see Y. Nara, AJ, K. Murase, & A. Ohnishi, PRC 106, 044902(2022).



Can this difference be found in hyperon directed flows?

Λ+Σ0 v1 at $\sqrt{s_{NN}}$ =4.5 GeV



WIP projects for more quantitative modeling 23/25

- Calculating hyperon potentials with <u>chiral N2LO YN+YNN</u>
 YN: J. Haidenbauer, U.-G. Meißner, A. Nogga, & H. Le, Eur. Phys. J. A 59 (2023) 3, 63.
 YNN: working with Johann Haidenbauer.
- Implementing potentials of hyperon resonances (Y*) by employing the parity doublet model with Y. Nara and K. Murase

 To avoid the uncertainty in Y*, lower collision energy (HADES energy) may be preferred.

Interactions including resonance hyperon could affect the dynamics.



Σ potential is interesting.

• **S** potential has large uncertainty: $U_{\Sigma}(
ho_0) = 30 \pm 20 \text{ MeV}$

based on Σ^{-} atom data and (π^{+}, K^{+}) inclusive spectra

A. Gal, E. V. Hungerford, & D. J. Millener (2016).

- New nice measurements:
 - Differential cross section of Σp (J-PARC E40)
 Σ⁺p: T. Nanamura et al., PTEP 2022,

093D01 (2022).

Σ⁻*p*: K. Miwa et al., PRC 104, 045204 (2021).

 \rightarrow More attractive Σ potential

- Femtoscopy of Σ⁺p
 B. Heybeck (ALICE) talk at QM2025
- Formation of ${}_{\Sigma}^{5}$ He...? inferred from ${}_{\Lambda}^{5}$ He yield Yingjie Zhou (STAR) talk at QM 2025



24/25

Summary

NLO13(500) YN + YNN via decuplet saturation:

The strongly repulsive Λ potential is obtained and is found consistent with

 $2M_{\odot}$ NS (No Λ in NS)

Λ binding energy Λ directed flow





- We still need other strategy to distinguish between the repulsive and attractive Λ potentials.
- The $\Lambda + \Sigma 0$ flow is sensitive to the momentum dependence of the hyperon potentials and the density dependence of the Σ potential.
- Σ potential can be a key to pin down YN and YNN interaction.