

Recent GW activity in pion photoproduction experiments and analysis within the SAID Framework

Presenter: William Briscoe

Co Authors: Igor Strakovsky*, Axel Schmidt, Olga Cortes,
Marshall Scott

GW Graduate Students: Phoebe Sharp, Sara Ratliff, Erin Seroka,
Izzy Illari, Peter Solazzo, Chan Kim, Giovanni Angelini.

* Thanks for the help in preparing slides

Hadron Spectroscopy

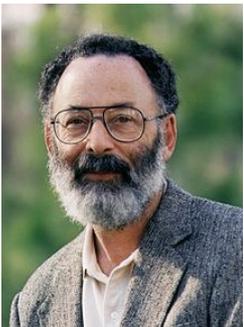


It is clear that we still need much more information about the existence and parameters of many baryon states, especially in the $N=2$ mass region, before this question of non-minimal $SU(6) \times O(3)$ super-multiplet can be settled.

Dick Dalitz, 1976

*The first problem is the notion of a resonance is not well defined. The ideal case is a narrow resonance far away from the thresholds, superimposed on slowly varying background. It can be described by a **Breit-Wigner** formula and is characterized by a pole in the analytic continuation of the partial wave amplitude into the low half of energy plane.*

Gerhard Höhler, 1987



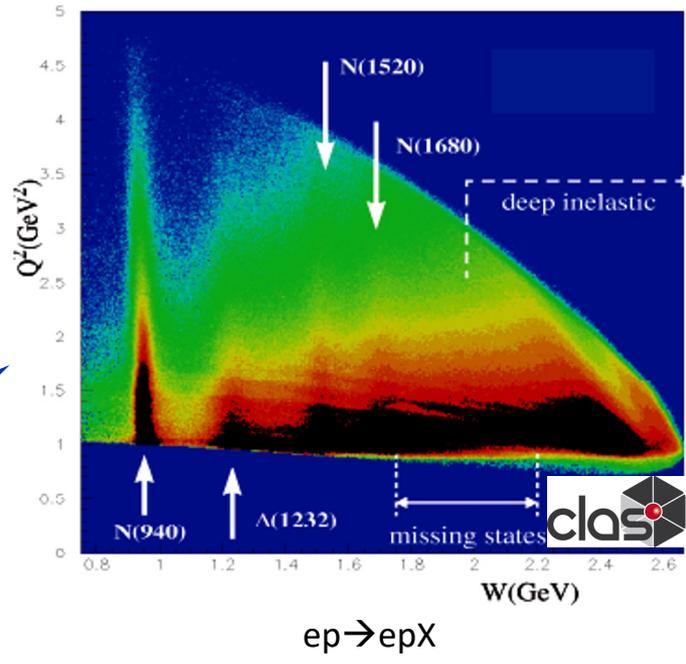
Why N^ s are important – The first is that nucleons are the stuff of which our world is made. My second reason is that they are simplest system in which the quintessentially non-**Abelian** character of QCD is manifest. The third reason is that history has taught us that, while relatively simple, Baryons are sufficiently complex to reveal physics hidden from us in the mesons.*

Nathan Isgur, 2000

There are Many Ways to Study N^*

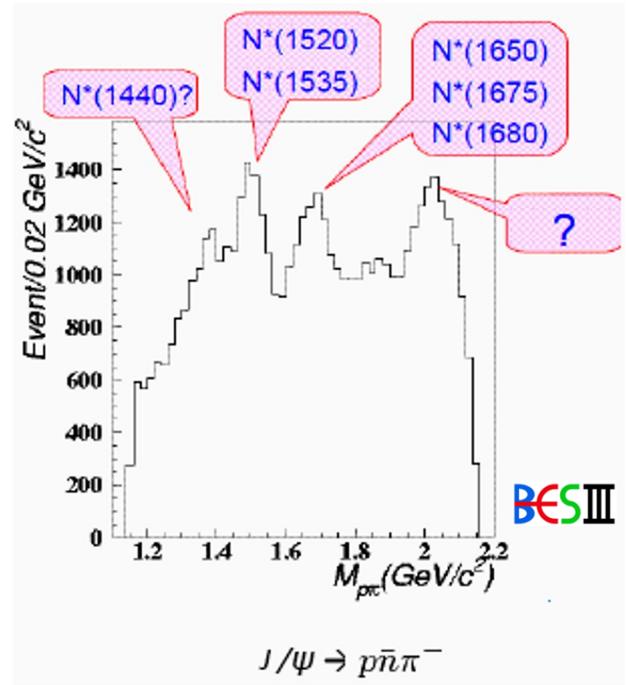
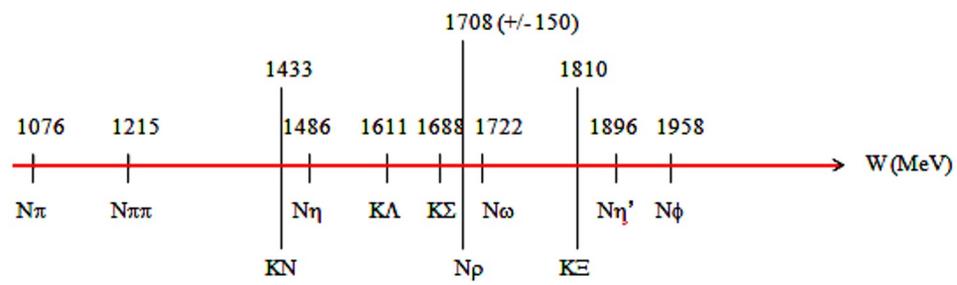
• Prolific source of N^* & Δ^* baryons is to measure many channels with different combinations of quantum numbers.

- $\pi N \rightarrow \pi N, \pi\pi N, \dots$
- $\gamma N \rightarrow \pi N, \pi\pi N, \dots$
- $\gamma^* N \rightarrow \pi N, \pi\pi N, \dots$
- $pp \rightarrow pp\pi^0, pp\pi\pi, \dots$
- $J/\psi \rightarrow p\bar{p}\pi^0, p\bar{n}\pi^-, \dots$



Most of **PDG Listings** info comes from these sources.

- πN elastic scattering is highly constrained.
- **Resonance** structure is correlated.
- Two-body final state, **fewer amplitudes**.



Where Are We Now...

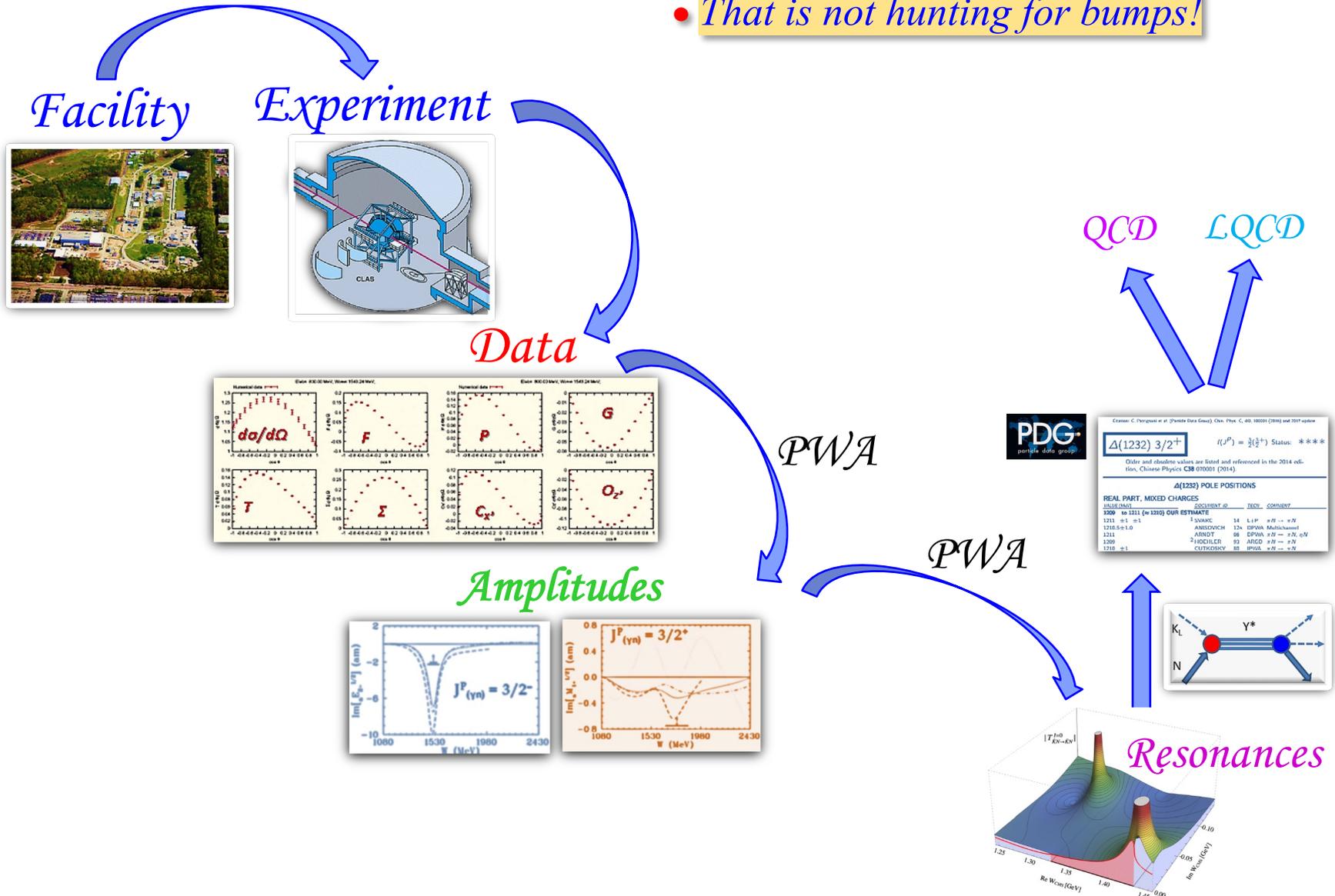
- Certain experiments provide unique info about resonance decay properties.
 - helicity couplings $A_{1/2}$ & $A_{3/2}$ for γp and γn decays come only from **pion** photo- & electro-production measurements.
- Helicity couplings normally extracted from full energy-dependent multipole amplitudes.
 - Until recently, the only available multipole amplitudes were for single pion photoproduction.
- Determination of $A_{1/2}$ & $A_{3/2}$ from meson photo- & electro-production requires knowledge of corresponding hadronic couplings.
 - Photo- & electro-production alone determine only the product of couplings to the γN and hadronic channels.

Every phenomenology group
[BnGa, EBAC, Gent, Giessen, JAW, Juelich, MAID, & SAID]
Uses the SAID πN results for constrain.

- Most modern **experimental efforts** focus on photo- or electro-production experiments
 - high-precision measurements with **hadron** beams (**pions** & **kaons**) are needed!
- PWAs are best way to determine **N*** properties
 - Multichannel approaches can help resolve inconsistencies.

Road Map to Baryon Spectroscopy

- That is not hunting for bumps!



QCD LQCD

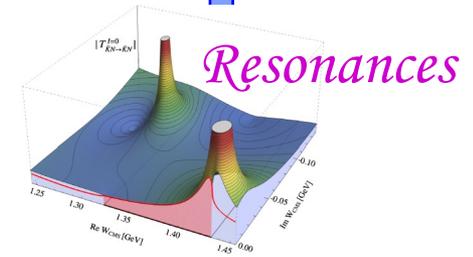
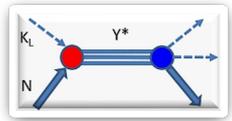


Ceasari, C. (arXiv:1404.0444), Chin. Phys. C, 40, 10001 (2016) and 2017 update

$\Delta(1232) 3/2^+$ $J^P = \frac{3}{2}^+$ Status: ***

$\Delta(1232)$ POLE POSITIONS

REAL PART, MIXED CHARGES	DOCUMENT ID	TEXT	COMMENT
1209	10.1211 (v1210) OUR ESTIMATE		
1211	n1 = 1	SWANEK	14 LIP $\neq N \rightarrow \neq N$
1218.5-1.0		ANDRUSOVICH	17M DPWA Multichannel
1211		ARNDT	98 DPWA $\neq N \rightarrow \neq N, \neq N$
1209		HÖCHLER	93 ARGD $\neq N \rightarrow \neq N$
1210	-1	CHUSHNYKH	89 IMPD $\neq N \rightarrow \neq N$



Amplitudes

PWA

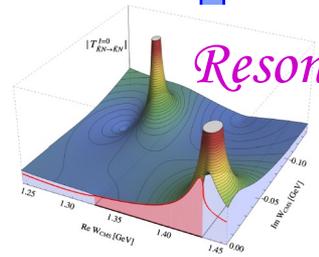
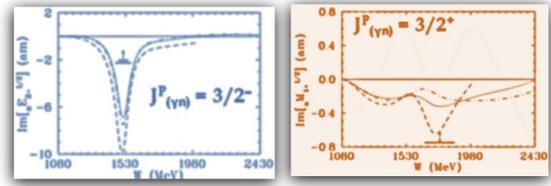
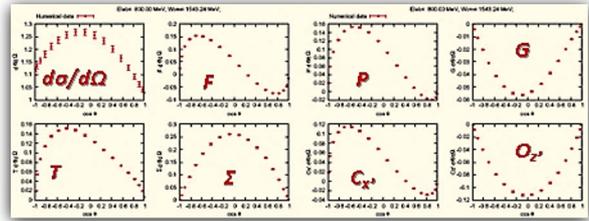
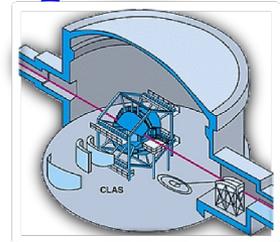
PWA

Resonances

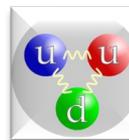
Facility

Experiment

Data



Baryon Sector @ PDG2025



PTEP

Progress of
Theoretical and
Experimental Physics



Review of Particle Physics

P.A. Zyla et al. Particle Data Group, *Prog. Theor. Exp. Phys.* 2020, 083C01 (2020)

PDG

particle data group



The Physical Society of Japan



OXFORD
UNIVERSITY PRESS



GW Contribution

S. Navas et al, Phys Rev D 110, 030001 (2024)



p	$1/2^+$	****	$\Delta(1232)$	$3/2^+$	****	Σ^+	$1/2^+$	****	Ξ^0	$1/2^+$	****	Λ_c^+	$1/2^+$	****
n	$1/2^+$	****	$\Delta(1600)$	$3/2^+$	***	Σ^0	$1/2^+$	****	Ξ^-	$1/2^+$	****	$\Lambda_c(2595)^+$	$1/2^-$	***
$N(1440)$	$1/2^+$	****	$\Delta(1620)$	$1/2^-$	****	Σ^-	$1/2^+$	****	$\Xi(1530)^0$	$3/2^+$	****	$\Lambda_c(2625)^+$	$3/2^-$	*
$N(1520)$	$3/2^-$	****	$\Delta(1700)$	$3/2^-$	****	$\Sigma(1305)$	$3/2^+$	****	$\Xi(1620)^0$	*		$\Lambda_c(2765)^+$	*	
$N(1535)$	$1/2^-$	****	$\Delta(1750)$	$1/2^+$	*	$\Sigma(1400)$	*		$\Xi(1690)^0$	***		$\Lambda_c(2890)^+$	$5/2^+$	***
$N(1650)$	$1/2^-$	****	$\Delta(1900)$	$1/2^-$	**	$\Sigma(1560)$	**		$\Xi(1820)^0$	$3/2^-$	***	$\Lambda_c(2940)^+$	***	
$N(1675)$	$5/2^-$	****	$\Delta(1905)$	$5/2^+$	****	$\Sigma(1580)$	$3/2^-$	**	$\Xi(1950)^0$	***		$\Sigma_c(2455)$	$1/2^+$	****
$N(1690)$	$5/2^+$	****	$\Delta(1910)$	$1/2^+$	****	$\Sigma(1620)$	$1/2^-$	**	$\Xi(2030)^0$	$\geq 3/2^+$	****	$\Sigma_c(2520)$	$3/2^+$	****
$N(1695)$	*		$\Delta(1920)$	$3/2^+$	**	$\Sigma(1660)$	$1/2^+$	***	$\Xi(2250)^0$	**		$\Sigma_c(2800)$	***	
$N(1700)$	$3/2^-$	***	$\Delta(1930)$	$5/2^-$	*	$\Sigma(1670)$	$3/2^-$	****	$\Xi(2370)^0$	**		Ξ_c^+	$1/2^+$	***
$N(1710)$	$1/2^+$	**	$\Delta(1940)$	$3/2^-$	**	$\Sigma(1690)$	**		$\Xi(2500)^0$	*		Ξ_c^0	$1/2^+$	***
$N(1720)$	$3/2^+$	****	$\Delta(1950)$	$7/2^+$	**	$\Sigma(1750)$	**		$\Xi(2645)^0$	**		Ξ_c^0	$1/2^+$	***
$N(1800)$	$5/2^+$	**	$\Delta(2000)$	$5/2^+$	*	$\Sigma(1770)$	$1/2^+$	*	$\Xi(2790)^0$	**		Ξ_c^0	$1/2^+$	***
$N(1810)$	$3/2^-$	**	$\Delta(2200)$	$7/2^-$	**	$\Sigma(1775)$	$1/2^-$	****	$\Xi(2930)^0$	**		Ξ_c^0	$3/2^+$	***
$N(1830)$	$1/2^+$	**	$\Delta(2300)$	$9/2^+$	**	$\Sigma(1840)$	$3/2^+$	***	$\Xi(2980)^0$	**		Ξ_c^0	$3/2^+$	***
$N(1850)$	$3/2^-$	**	$\Delta(2350)$	$5/2^-$	*	$\Sigma(1880)$	$1/2^+$	**	$\Xi(3055)^0$	**		Ξ_c^0	$1/2^-$	***
$N(1900)$	$7/2^-$	**	$\Delta(2390)$	$7/2^+$	*	$\Sigma(1915)^0$	$5/2^+$	****	$\Xi(3080)^0$	***		Ξ_c^0	$3/2^-$	***
$N(2000)$	$5/2^+$	**	$\Delta(2400)$	$9/2^-$	**	$\Sigma(1915)^+$	$5/2^+$	****	$\Xi(2930)^0$	*		Ξ_c^0	*	
$N(2040)$	$3/2^+$	*	$\Delta(2420)$	$11/2^+$	****	$\Sigma(2000)$	$1/2^-$	**	$\Xi(2980)^0$	***		Ξ_c^0	***	
$N(2060)$	$5/2^-$	**	$\Delta(2420)$	$11/2^+$	****	$\Sigma(2030)$	$7/2^+$	****	$\Xi(3055)^0$	**		Ξ_c^0	**	
$N(2100)$	$1/2^+$	*	$\Delta(2750)$	$13/2^-$	**	$\Sigma(2070)$	$5/2^+$	*	$\Xi(3123)^0$	*		Ξ_c^0	*	
$N(2120)$	$3/2^-$	**	$\Delta(2950)$	$15/2^+$	**	$\Sigma(2080)$	$3/2^+$	**	Ω_c^0	$1/2^+$	***	$\Omega_c(2770)^0$	$3/2^+$	***
$N(2190)$	$7/2^-$	****	Λ	$1/2^+$	****	$\Sigma(2100)$	$7/2^-$	**	Ω_c^0	$1/2^+$	***			
$N(2220)$	$9/2^+$	****	$\Lambda(1405)$	$1/2^-$	****	$\Sigma(2250)$	***		Ξ_c^+	*				
$N(2250)$	$9/2^-$	****	$\Lambda(1520)$	$3/2^-$	****	$\Sigma(2455)$	**		Ξ_c^+	*				
$N(2600)$	$11/2^-$	***	$\Lambda(1600)$	$1/2^+$	***	$\Sigma(2620)$	**		Λ_b^0	$1/2^+$	***			
$N(2700)$	$13/2^+$	**	$\Lambda(1670)$	$1/2^-$	****	$\Sigma(3000)$	*		Σ_b^0	$1/2^+$	***			
			$\Lambda(1690)$	$3/2^-$	****	$\Sigma(3170)$	*		Σ_b^0	$3/2^+$	***			
			$\Lambda(1800)$	$1/2^-$	***				Ξ_b^0	$1/2^+$	***			
			$\Lambda(1810)$	$1/2^-$	***				Ξ_b^0	$3/2^+$	***			
			$\Lambda(1820)$	$5/2^+$	****				Ξ_b^0	$1/2^+$	***			
			$\Lambda(1830)$	$5/2^-$	****				Ξ_b^0	$1/2^+$	***			
			$\Lambda(1890)$	$3/2^+$	***				Ξ_b^0	$1/2^+$	***			
			$\Lambda(2000)^0$	$7/2^+$	****									
			$\Lambda(2100)$	$7/2^-$	****									
			$\Lambda(2110)$	$5/2^+$	****									
			$\Lambda(2325)$	$3/2^-$	**									
			$\Lambda(2350)$	$9/2^+$	**									
			$\Lambda(2585)$	**										

• First hyperon was discovered in 1950.

• Pole position in complex energy plane for hyperons has been made only in 2010.



• PDG2024 has 133 Baryon Resonances (69 of them are 4* & 3*).

• In case of SU(6) x O(3), 434 states would be present if all revealed multiplets were fleshed out (three 70 & four 56).

• LQCD results are similar.

R. Koniuk & N. Isgur, Phys Rev Lett 44, 845 (1980)



Complete Experiment for Pion PhotoProduction

2002

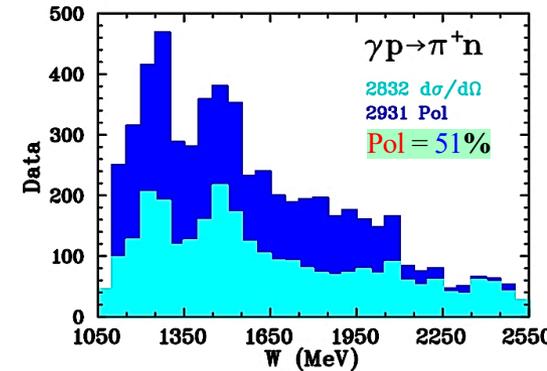
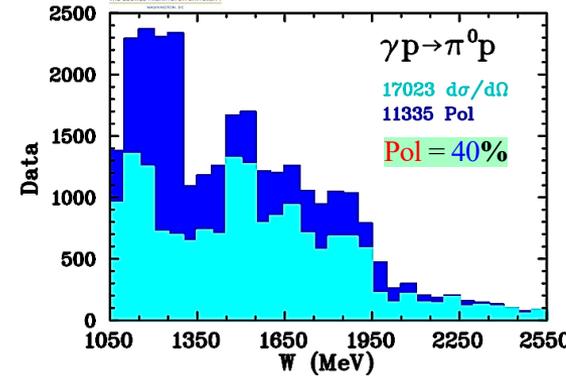


2020



1996-2018

Particle	$L_{2I,2J}$	Overall status	Status as seen in —					Status as seen in															
			$N\pi$	$N\eta$	AK	ΣK	$\Delta\pi$	$N\rho$	$N\gamma$	Particle	J^P	overall	$N\gamma$	$N\pi$	$\Delta\pi$	$N\sigma$	$N\eta$	AK	ΣK	$N\rho$	$N\omega$	$N\eta'$	
$N(939)$	P_{11}	****																					
$N(1440)$	P_{11}	****	****	*			***	*	***														
$N(1520)$	D_{13}	****	****	*			****	****	****														
$N(1535)$	S_{11}	****	****	****				*	**	***													
$N(1650)$	S_{11}	****	****	*		***	**	***	**	***													
$N(1675)$	D_{15}	****	****	*	*		****	*	****								*	*					
$N(1680)$	F_{15}	****	****				****	****	****								*	*					
$N(1700)$	D_{13}	***	***	*	**	*	**	*	**							*	*						
$N(1710)$	P_{11}	***	***	**	**	*	**	*	***							**	**	*	*	*			
$N(1720)$	P_{13}	****	****	*	**	*	*	**	**														
$N(1900)$	P_{13}	**	**						*														
$N(1990)$	F_{17}	**	**	*	*	*			*														*
$N(2000)$	F_{15}	**	**	*	*	*	*	*	**														
$\Delta(1232)$	P_{33}	****	****	F					****														****
$\Delta(1600)$	P_{33}	***	***	o			***	*	**														***
$\Delta(1620)$	S_{31}	****	****	r			****	****	****														****
$\Delta(1700)$	D_{33}	****	****	b		*	***	**	***														***
$\Delta(1750)$	P_{31}	*	*	i																			*
$\Delta(1900)$	S_{31}	**	**	d		*	*	**	*														*
$\Delta(1905)$	F_{35}	****	****	d		*	**	**	***														****
$\Delta(1910)$	P_{31}	****	****	e		*	*	*	*														****
$\Delta(1920)$	P_{33}	***	***	n		*	**	**	*														***
$\Delta(1930)$	D_{35}	***	***			*			**														***
$\Delta(1940)$	D_{33}	*	*	F					*														*
$\Delta(1950)$	F_{37}	****	****	o		*	****	*	****														****
$\Delta(2000)$	F_{37}	**	**	r					**														**
$\Delta(1232)$	$3/2^+$	****	****	****	****	****	****	****	****														****
$\Delta(1440)$	$1/2^+$	****	****	****	****	****	****	****	****														****
$\Delta(1520)$	$3/2^-$	****	****	****	****	****	****	****	****														****
$\Delta(1535)$	$1/2^-$	****	****	****	****	****	****	****	****														****
$\Delta(1650)$	$1/2^-$	****	****	****	****	****	****	****	****														****
$\Delta(1675)$	$5/2^-$	****	****	****	****	****	****	****	****														****
$\Delta(1680)$	$5/2^+$	****	****	****	****	****	****	****	****														****
$\Delta(1700)$	$3/2^-$	****	****	****	****	****	****	****	****														****
$\Delta(1710)$	$1/2^+$	****	****	****	****	****	****	****	****														****
$\Delta(1720)$	$3/2^+$	****	****	****	****	****	****	****	****														****
$\Delta(1860)$	$5/2^+$	**	**	**	**	**	**	**	**														**
$\Delta(1875)$	$3/2^-$	****	****	****	****	****	****	****	****														****
$\Delta(1880)$	$1/2^+$	****	****	****	****	****	****	****	****														****
$\Delta(1895)$	$1/2^-$	****	****	****	****	****	****	****	****														****
$\Delta(1900)$	$3/2^+$	****	****	****	****	****	****	****	****														****
$\Delta(1990)$	$7/2^+$	**	**	**	**	**	**	**	**														**
$\Delta(2000)$	$5/2^+$	**	**	**	**	**	**	**	**														**



$\pi^+n/\pi^0p = 20\%$

• Lack of $\gamma p \rightarrow \pi^+n$ vs $\gamma p \rightarrow \pi^0p$ data does not allow us to be as confident about determination of *neutron* couplings.

World Progress in Pion PhotoProduction

SAID: <http://gwdac.phys.gwu.edu/>

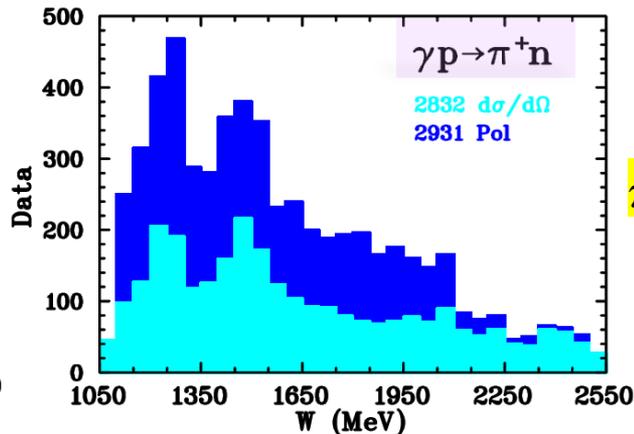
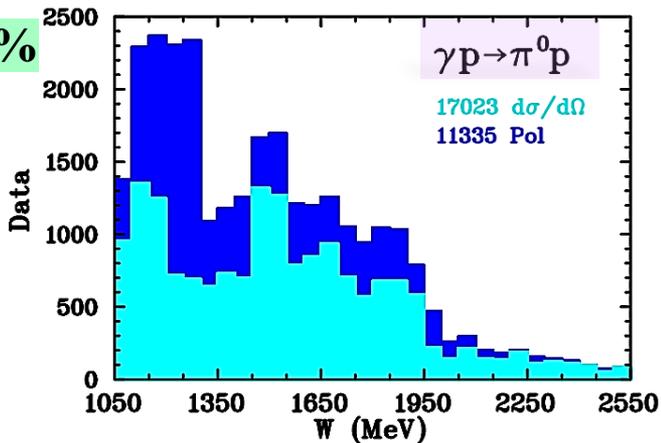
D. Ireland, E. Pasyuk, IIS, Prog Part Nucl Phys 111, 103752 (2020)

1996–2018

Pol = 40%



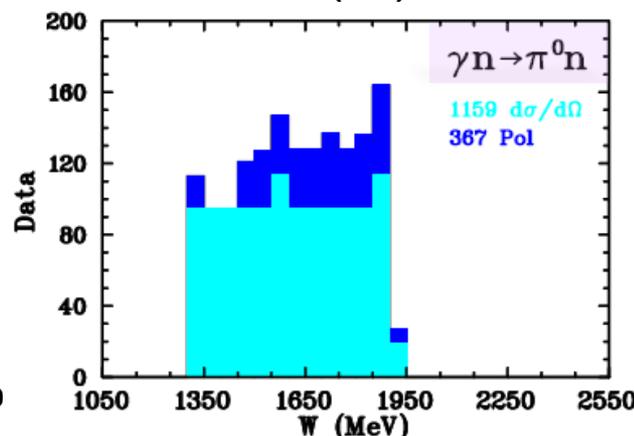
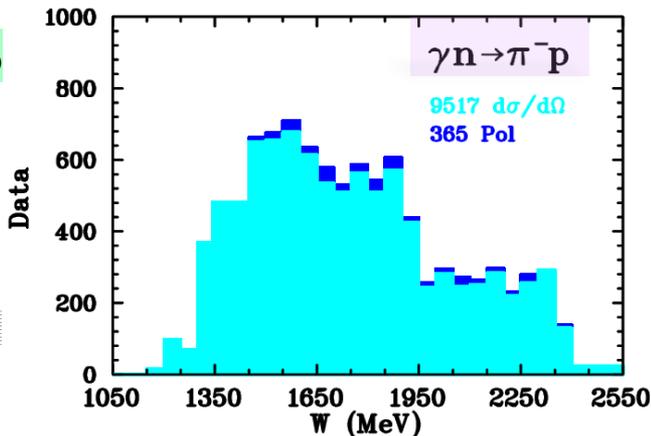
SPring-8



Pol = 51%

$\pi^+n/\pi^0p = 20\%$

Pol = 4%



Pol = 24%

$\pi^-p/\pi^0n = 15\%$



• There is disbalance between π^0p & π^+n data, 20%.

• Pion photoproduction on neutron much less known, 33%.

World Neutral & Charged PionPR Data

W.J. Briscoe, M. Doring, H. Habertzettl, M. Manley, M. Naruki, IS, E. Swanson, Eur Phys J A 51, 129 (2015)

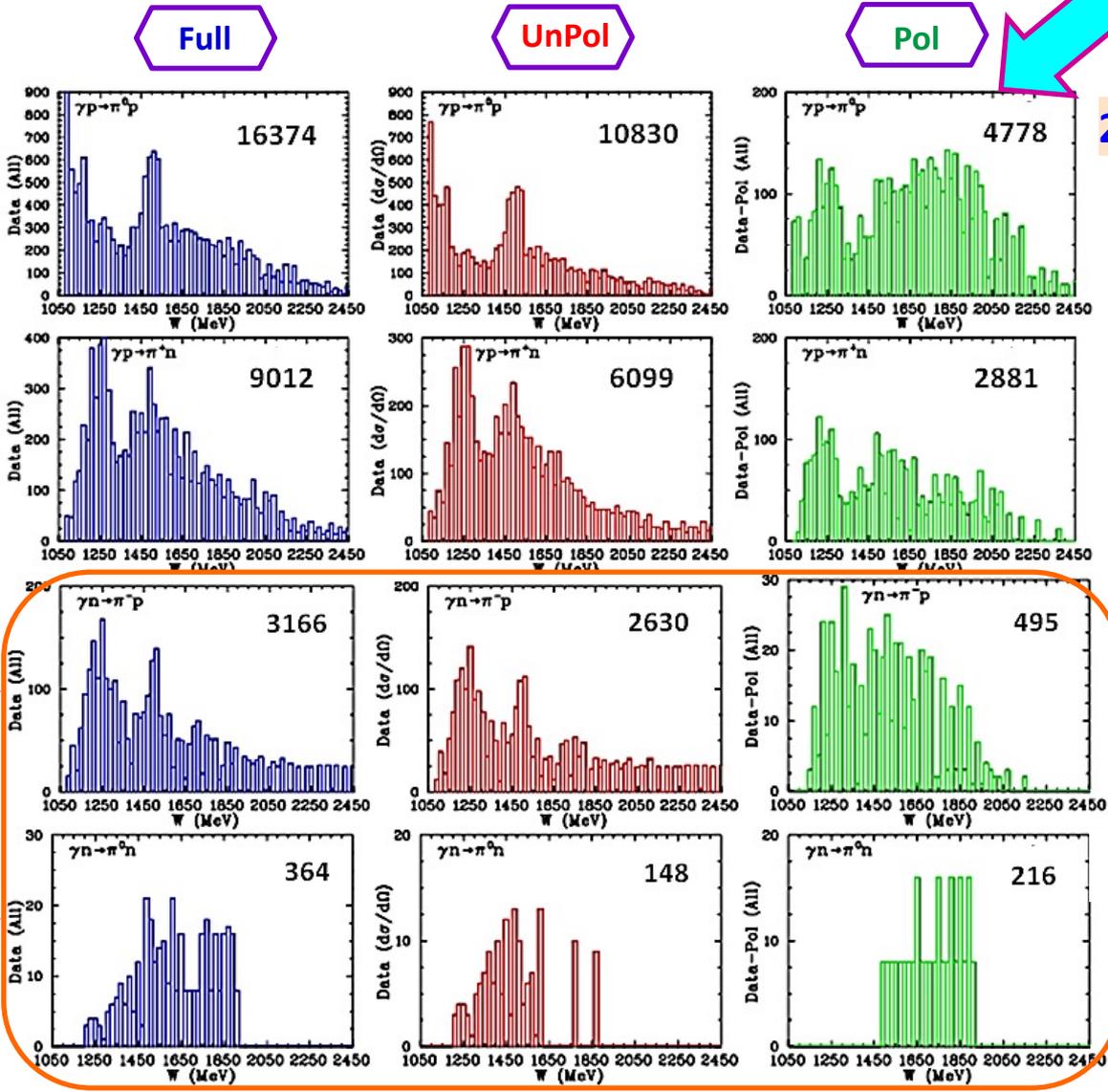
W < 2.5 GeV

$\gamma p \rightarrow \pi^0 p$

$\gamma p \rightarrow \pi^+ n$

$\gamma n \rightarrow \pi^+ p$

$\gamma n \rightarrow \pi^0 n$



2017

25858 — 19768 — 6090

9859 — 6078 — 3781

11856 — 11092 — 764

364 — 148 — 216

PWA for non-Strange Baryons

- **Originally** PWA arose as technology to determine amplitude of reaction via fitting scattering data.
That is **non-trivial mathematical problem** – looking for solution of **ill-posed** problem following to **Hadamard & Tikhonov**.



- **Resonances** appeared as **by-product**
[bound states objects with definite quantum numbers, mass, lifetime, & so on].

- **Standard PWA**
 - ⇒ Reveals only **wide** Resonances, but not too wide ($\Gamma < 500$ MeV) & possessing not too **small** BR ($BR > 4\%$).
 - ⇒ Tends (by construction) to **miss** narrow Res with $\Gamma < 20$ MeV.



Most of our current knowledge about bound states of **three light quarks** has come mainly from $\pi N \rightarrow \pi N$ PWA's:



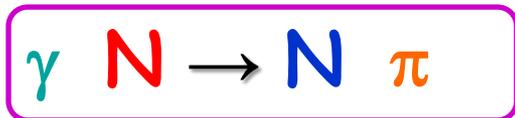
Karlsruhe-Helsinki,
Carnegie-Mellon-Berkeley,
& **GW.**



Main source of **EM** couplings is **GW**, **BnGa**, & **JuBoWa** analyses.



Direct Amplitude Reconstruction in Pion Photo-Production



spin: $1 \quad \frac{1}{2} \rightarrow \frac{1}{2} \quad 0$

helicities: $2 \times 2 \times 2 / 2 = 4$

parity conservation \rightarrow

• In particle physics, **helicity** is projection of the spin \vec{S} onto direction of momentum, \hat{p} :

$$h = \vec{J} \cdot \hat{p} = \vec{L} \cdot \hat{p} + \vec{S} \cdot \hat{p} = \vec{S} \cdot \hat{p}$$

$$\hat{p} = \frac{\vec{p}}{|\vec{p}|}$$

Therefore, there are **4** independent invariant amplitudes

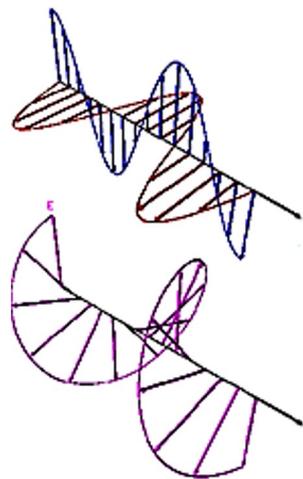
• In order to **determine** pion photoproduction amplitude [**4 modules** and **3 relative phases**], one has to carry out **7 independent** measurements at **fixed** (W, t) or (E, θ) .



8 • This extra observable is necessary to eliminate **sign ambiguity**.

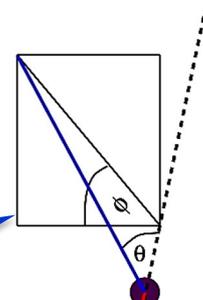
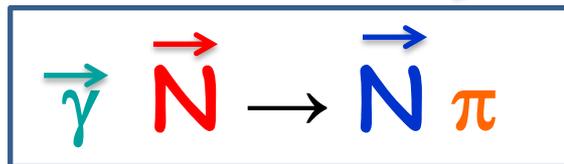
Complete Experiment for Pion PhotoProduction

- There are **16** non-redundant observables.
- They are **not completely independent** from each other.

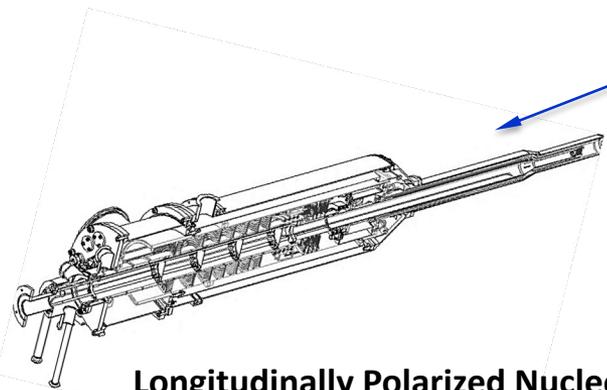


Linear
Polarized
Beam

Circular
Polarized
Beam



Nucleon Recoil
Polarization



Longitudinally Polarized Nucleon Target
Transverse Polarized Nucleon Target

- 1** un-pol measurement: $d\sigma/d\Omega$
- 3** single pol measurements: Σ, T, P
- 12** double pol measurements: $E, F, G, H, C_x, C_z, O_x, O_z, L_x, L_z, T_x, T_z$
- 18** triple polarization asymmetries
[9 for linear pol beam]
[9 for circular pol beam]
- 13** of them are non-vanishing



A. Sandorfi *et al.* AIP Conf. Proc. **1432**, 219 (2012)
K. Nakayama, private communication, 2014

Importance of Neutron Data

- **EM** interaction do not conserve **isospin**, so multipole amplitudes contain **isoscalar** & **isovector** contributions of **EM** current.

Proton

Neutron

$$\begin{aligned}
 A_{\pi^0 p} &= A^0 + \frac{1}{3} A^{1/2} + \frac{2}{3} A^{3/2} & A_{\pi^0 n} &= -A^0 + \frac{1}{3} A^{1/2} + \frac{2}{3} A^{3/2} \\
 A_{\pi^+ n} &= \sqrt{2} \left(A^0 + \frac{1}{3} A^{1/2} - \frac{1}{3} A^{3/2} \right) & A_{\pi^- p} &= \sqrt{2} \left(A^0 - \frac{1}{3} A^{1/2} + \frac{1}{3} A^{3/2} \right)
 \end{aligned}$$

- **Proton** data alone does not allow separation of **isoscalar** & **isovector** components.

Q: *Can we avoid?* A: **NO!**

- Need **data** on both **proton** & **neutron** !



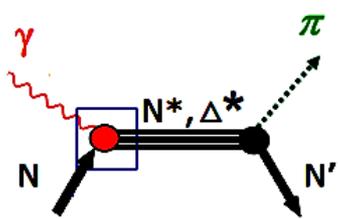


Photo-Decay Amplitudes in BW & Pole Forms

- Pole is main signature of resonance.

$$A_h^{BW} = C \sqrt{\frac{q_r}{k_r} \frac{\pi(2J+1)M_r \Gamma_r^2}{m_N \Gamma_{\pi,r}}} \tilde{A}_\alpha^h$$

Evaluated at Res Energy

$$A_h^{pole} = C \sqrt{\frac{q_p}{k_p} \frac{2\pi(2J+1)W_p}{m_N \text{Res}_{\pi/N}}} \text{Res } A_\alpha^h$$

Evaluated at Pole

TABLE I. Breit-Wigner and pole values for selected nucleon resonances. Masses, widths, and residues are given in units of MeV, the helicity 1/2 and 3/2 photo-decay amplitudes in units of $10^{-3}(\text{GeV})^{-1/2}$. Errors on the phases are generally 2–5 degrees. For isospin 1/2 resonances the values of the proton target are given.

Resonance	Breit-Wigner values				Pole values			
	(Mass, width)	$\Gamma_\pi/2$	$A_{1/2}$	$A_{3/2}$	(Re W_p , $-2 \text{Im } W_p$)	R_π	$A_{1/2}$	$A_{3/2}$
$\Delta(1232) 3/2^+$	(1233, 119)	60	-141 ± 3	-258 ± 5	(1211, 99)	52 $[-47^\circ]$	$-136 \pm 5 [-18^\circ]$	$-255 \pm 5 [-6^\circ]$
$N(1440) 1/2^+$	(1485, 284)	112	-60 ± 2		(1359, 162)	38 $[-98^\circ]$	$-66 \pm 5 [-38^\circ]$	
$N(1520) 3/2^-$	(1515, 104)	33	-19 ± 2	$+153 \pm 3$	(1515, 113)	38 $[-5^\circ]$	$-24 \pm 3 [-7^\circ]$	$+157 \pm 6 [+10^\circ]$
$N(1535) 1/2^-$	(1547, 188)	34	$+92 \pm 5$		(1502, 95)	16 $[-16^\circ]$	$+77 \pm 5 [+4^\circ]$	
$N(1650) 1/2^-$	(1635, 115)	58	$+35 \pm 5$		(1648, 80)	14 $[-69^\circ]$	$+35 \pm 3 [-16^\circ]$	



R.L. Workman *et al*, Phys Rev C **87**, 068201 (2013)

A. Svarc *et al*, Phys Rev C **89**, 065208 (2014)

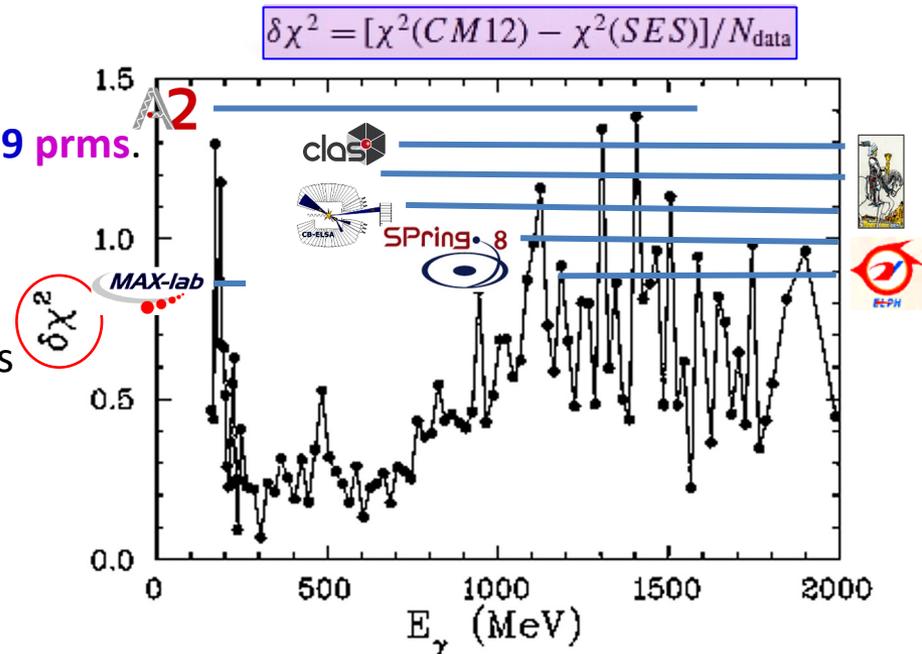


Single-Energy Solutions (SES)

R.L. Workman, M.W. Paris, W.J. Briscoe, IS, Phys Rev C **86**, 015202 (2012)

- **SAID** has employed both single-energy (**SES**) & energy-dependent (**Global**) solutions using **least-squares** technology over variety of energy ranges in order to estimate uncertainties.

- **SES**: based on bin of data spanning narrow **E** range [**5 – 75 MeV**] searches **2 to 29 prms.**
110 SES have been generated with central **E = 147 to 2650 MeV.**
of data in bin varies from **80 to 1100.**
- **Systematic deviation** between **SES** & **Global** fits is indication of
 - ⇒ Missing structure in **global** fit.
 - ⇒ Possible problems with particular **dataset.**



- **Diagonal Error Matrix** generated in **SES** fits.
 It can be used to estimate the overall uncertainties for **Global** solution.

Single Pion PhotoProduction on “Neutron” Target

- Accurate evaluation of **EM** couplings $N^* \rightarrow \gamma N$ & $\Delta^* \rightarrow \gamma N$ from **meson photoproduction** data remains paramount task in **hadron** physics.

- Only with good data on both **proton** & **neutron** targets, one can hope to disentangle **isoscalar** & **isovector** **EM** couplings of various N^* & Δ^* resonances,
as well as **isospin** properties of non-resonant **background amplitudes**.

K.M. Watson, Phys Rev 95, 228 (1954); R.L. Walker, Phys Rev 182, 1729 (1969)



- The lack of $\gamma n \rightarrow \pi^- p$ & $\gamma n \rightarrow \pi^0 n$ data does not allow us to be as confident about determination of **neutron** couplings relative to those of **proton**.

- **Radiative decay** width of **neutral baryons** may be extracted from π^- & π^0 photoproduction off **neutron**, which involves **bound neutron target** & needs use of **model-dependent nuclear (FSI) corrections**.



A.B. Migdal, JETP 1, 2 (1955); K.M. Watson, Phys Rev 95, 228 (1954)



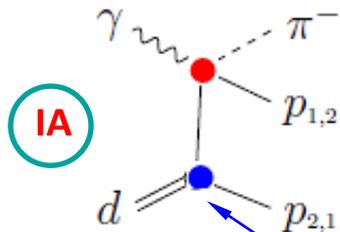
FSI for $\gamma d \rightarrow \pi p N$

$\gamma n \rightarrow \pi N$

V. Tarasov, A. Kudryavtsev, W. Briscoe, H. Gao, IS, Phys Rev C **84**, 035203 (2011)

V. Tarasov, A. Kudryavtsev, W. Briscoe, B. Krusche, IS, M. Ostrick, Phys At Nucl **79**, 216 (2016)

- FSI plays critical role in **state-of-the-art** analysis of $\gamma n \rightarrow \pi N$ data. It depends on (E, θ) .
- For $\gamma n \rightarrow \pi N$, effect is **5% – 60%**.

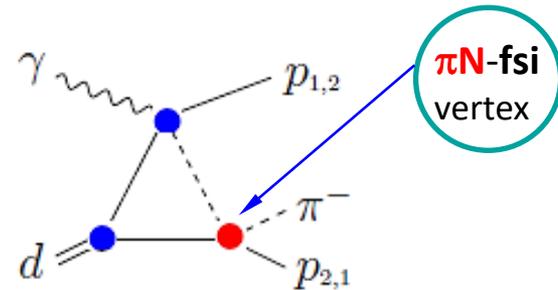
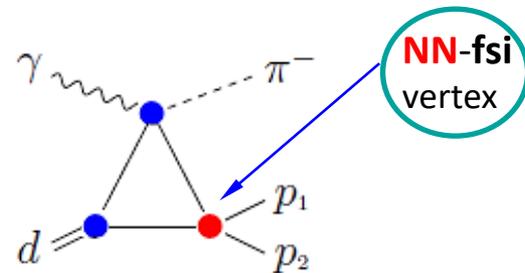


Fermi Smearing



Input: SAID: $\gamma N \rightarrow \pi N$, $\pi N \rightarrow \pi N$, $NN \rightarrow NN$ amplitudes for 3 leading terms

DWF: full Bonn NN Potential (there is no sensitivity to DWF).



$$R = (d\sigma/d\Omega_{\pi p}) / (d\sigma^{IA}/d\Omega_{\pi p})$$



$$\frac{d\sigma}{d\Omega}(\gamma n) = R^{-1} \frac{d\sigma}{d\Omega}(\gamma d)$$

FSI for $\gamma d \rightarrow \pi^- pp \rightarrow \gamma n \rightarrow \pi^- p$

V. Tarasov, A. Kudryavtsev, W. Briscoe, H. Gao, IS, Phys Rev C **84**, 035203 (2011)

$$R_{FSI} = (d\sigma/d\Omega_{\pi p}) / (d\sigma^{IA}/d\Omega_{\pi p})$$

Cuts:

$p_s < 200$ MeV/c
 $p_f < 200$ MeV/c

CLAS g10 & g13:

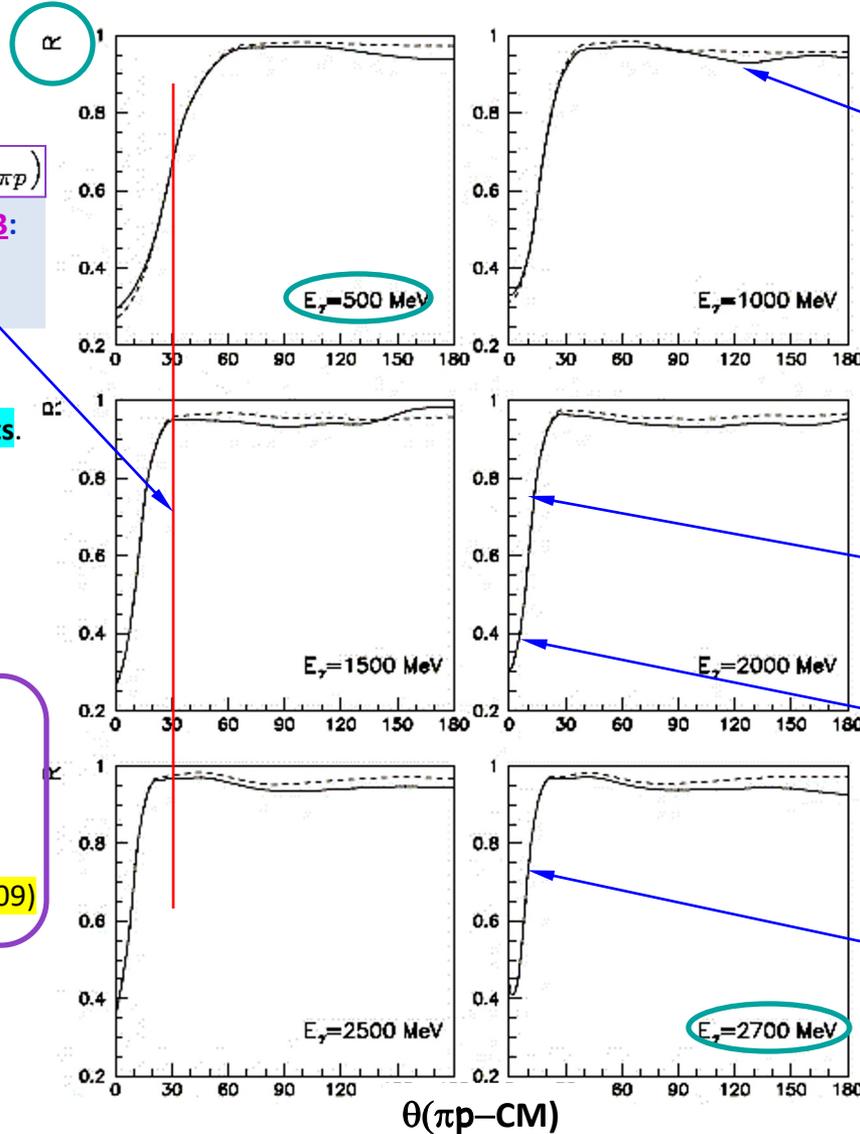
$E > 0.5$ GeV
 $\theta > 30$ deg

• There is no large sensitivity to cuts.

• Previous estimation of **Glauber FSI** gave order of **15–30%**.



W. Chen *et al*,
 Phys Rev Lett **103**, 012301 (2009)



- For **CLAS** data:
 - FSI correction factor **R < 1**.
 - Behavior is **smooth** vs. θ .
 - Effect: $\Delta\sigma/\sigma \leq 10\%$.

-- [IA + NN_{fsi}] / IA
 — [IA + (NN+ π N)_{fsi}] / IA

- There is **sizeable FSI** effect from **S-wave** part of **pp-FSI** at small angles.
- Region **narrows** as **E** increases.

Forward direction is Terra incognita

FSI for $\gamma d \rightarrow \pi^0 p n \rightarrow \gamma n \rightarrow \pi^0 n$

V.E. Tarasov, A.E. Kudryavtsev, WJB, B. Krusche, IIS, M. Ostrick, Phys At Nucl **79**, 216 (2016)

$$R_{FSI} = (d\sigma/d\Omega_{\pi p}) / (d\sigma^{IA}/d\Omega_{\pi p})$$

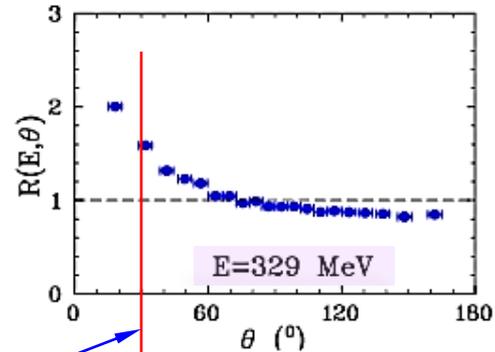
Cuts:

$p_s < 200$ MeV/c

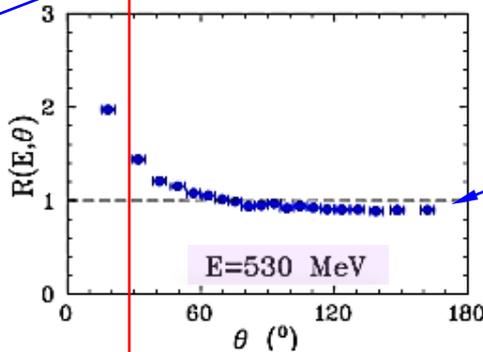
$p_f < 200$ MeV/c

- There is no large sensitivity to cuts.

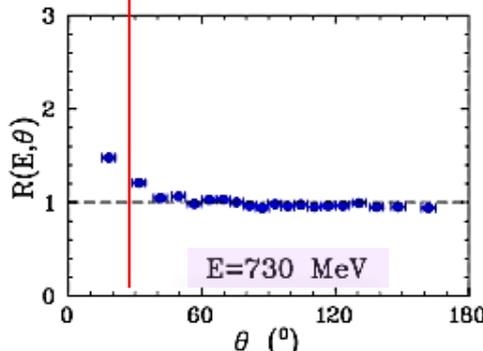
*Forward direction is
Terra incognita.
There are no data for
validation.*



- For $\Delta 2$ data:
 - Behavior is smooth vs. θ .
 - Effect: $\Delta\sigma/\sigma \leq 15\%$.



- There is sizeable FSI effect from S-wave part of pp-FSI at small angles.



*Assumption is
FSI for polarized
measurements
is small.*

FSI for $\gamma d \rightarrow \pi^0 np \rightarrow \gamma n \rightarrow \pi^0 n$ & $\gamma p \rightarrow \pi^0 p$

V. Tarasov, A. Kudryavtsev, W. Briscoe, B. Krusche, IS, M. Ostrick, Phys At Nucl **79**, 216 (2016)

- $\gamma n \rightarrow \pi^0 n$ case is much more complicated vs. $\gamma n \rightarrow \pi^- p$ because π^0 can come from both γn & γp initial interactions.

$$A(\gamma p \rightarrow \pi^0 p) = A_v + A_s$$

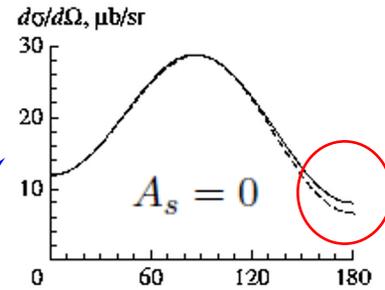
$$A(\gamma n \rightarrow \pi^0 n) = A_v - A_s$$

- The corrections for both target nucleons are **practically identical** for π^0 production in energy range of $\Delta(1232)3/2^+$ due to **isospin structure** of $\gamma N \rightarrow \pi N$ amplitude:

isoscalar *isovector*

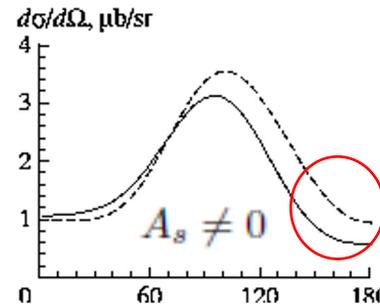
$$A_s = 0 \text{ or } A_v = 0$$

$$R_n = R_p$$

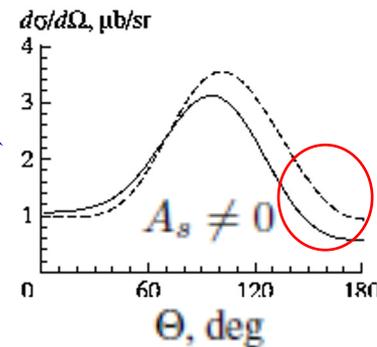


$\gamma p \rightarrow \pi^0 p$ (solid curves)
 $\gamma n \rightarrow \pi^0 n$ (dashed curves)

$\Delta(1232)3/2^+$



$N(1440)1/2^+$



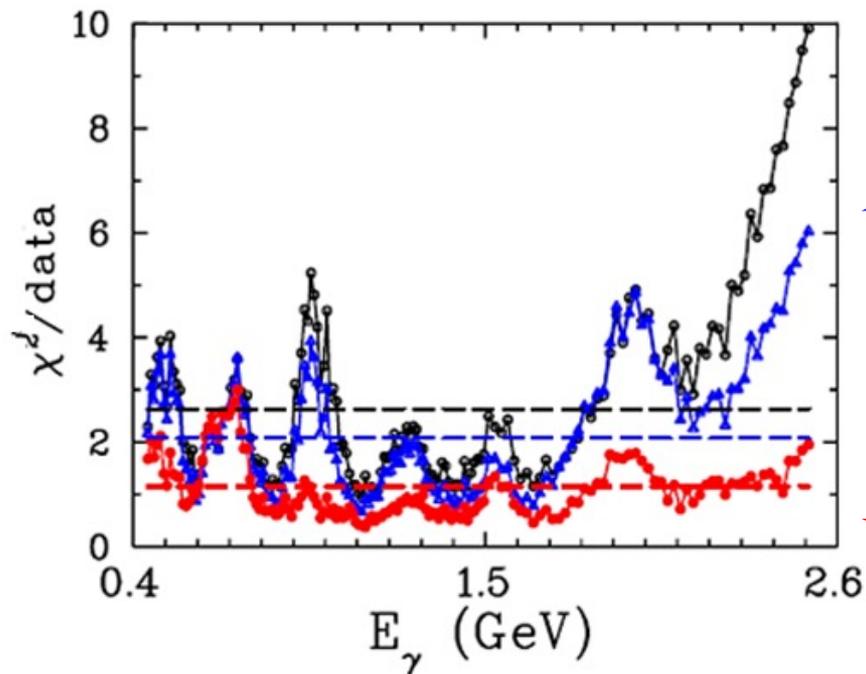
$N(1535)1/2^-$

- In general case,

$$R_n \neq R_p$$

Comparison of Previous & New SAID Fits for g_{13}

P. Mattione *et al*, Phys. Rev. C **96**, 035204 (2017)



• Recent SAID **PR15** applied to **g13** data **without & with FSI** corrections.

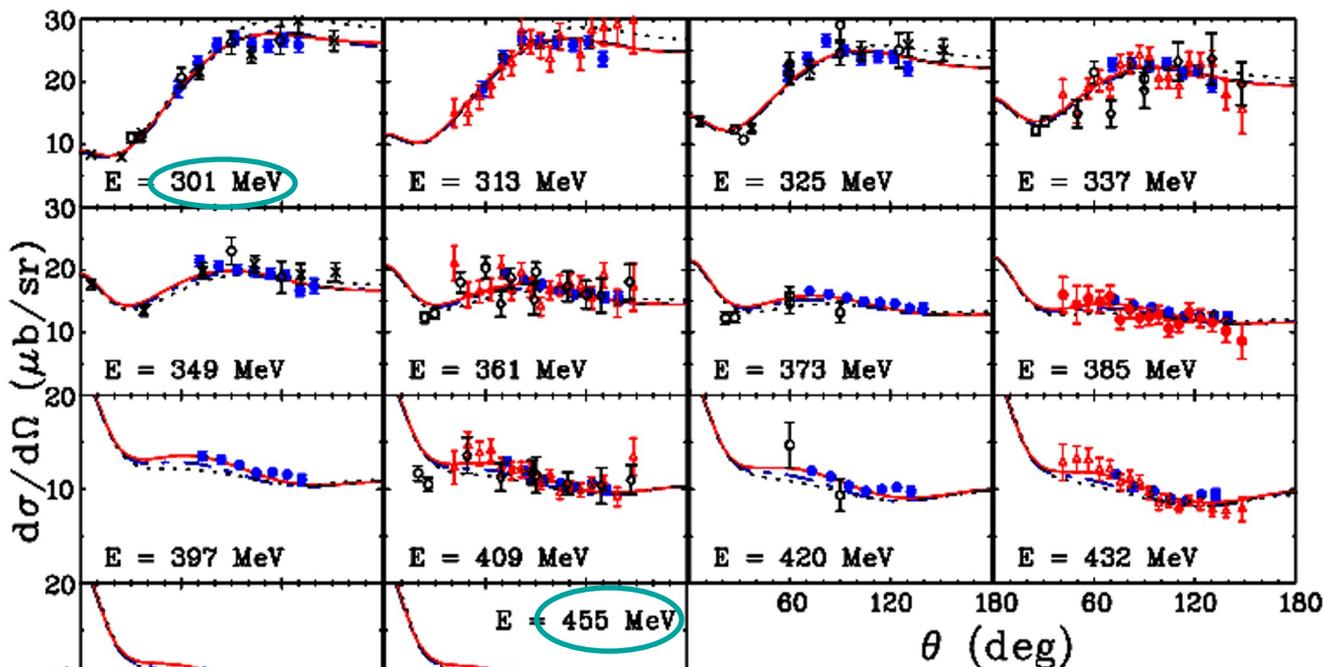
• New SAID **MA27** fit obtained after adding new **g13** data **with FSI** corrections.

• Obviously, **FSI** plays important role in $\gamma n \rightarrow \pi^- p$ $d\sigma/d\Omega$ determination.

MAMI-B for $\gamma n \rightarrow \pi^- p$ around Δ

W.J. Briscoe, A.E. Kudryavtsev, P. Pedroni, IS, V.E. Tarasov, R.L. Workman, Phys Rev C **86**, 065207 (2012)

- MAMI-B data for $\gamma n \rightarrow \pi^- p$ (including FSI corrections) & previous hadronic data for $\pi^- p \rightarrow n \gamma$ appear to agree well.



SAID-PE12
SAID-SN11
MAID07

FSI included

• T -invariance is good as 2×10^{-3}

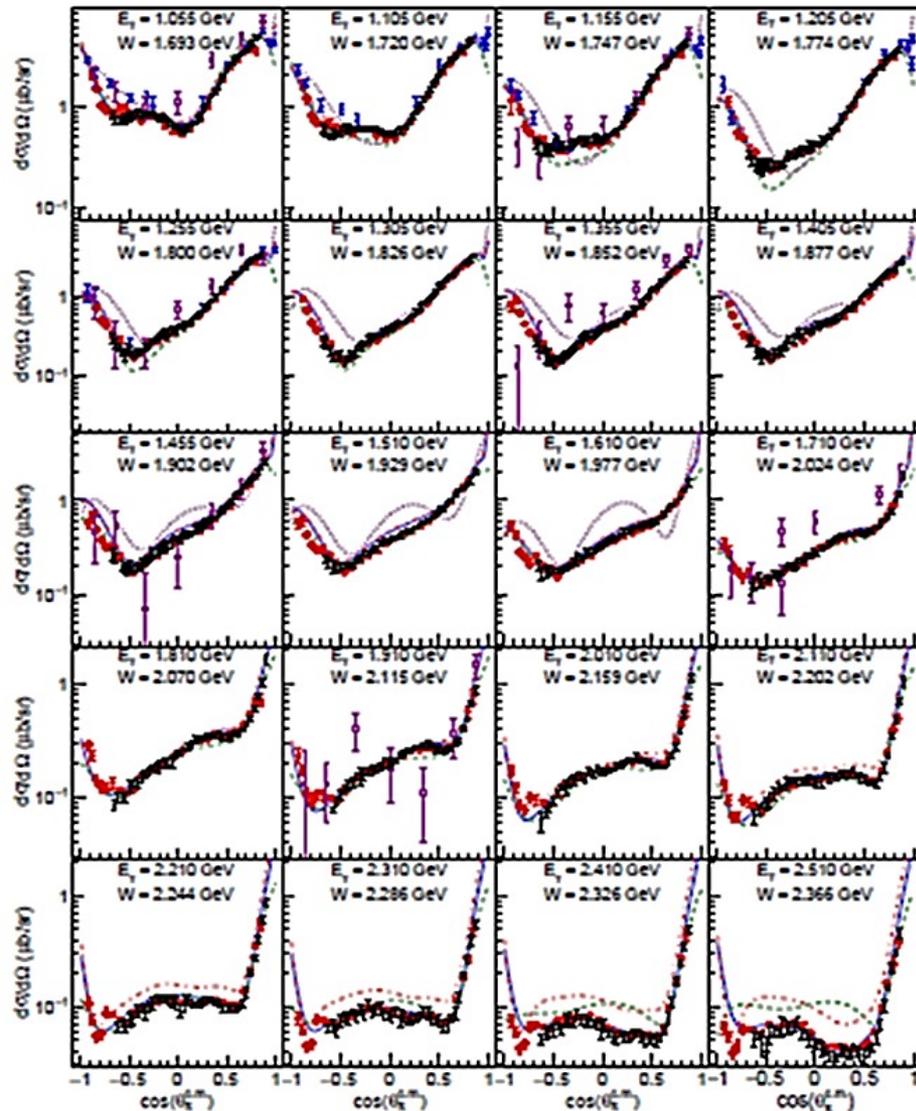
Data:

- – MAMI-B for $\gamma n \rightarrow \pi^- p$ sys=2%
J. Ahrens *et al*, Eur Phys J A **44**, 189 (2010)
- Δ – CB@BNL for $\pi^- p \rightarrow n \gamma$ sys=5%
A. Shafi *et al*, Phys Rev C **70**, 035204 (2004)
- o – TRIUMF, CERN, LBL, LAMPF for $\pi^- p \rightarrow n \gamma$



CLAS g13 for $\gamma n \rightarrow \pi^- p$ above 0.5 GeV

P. Mattione *et al*, Phys. Rev. C 96, 035204 (2017)

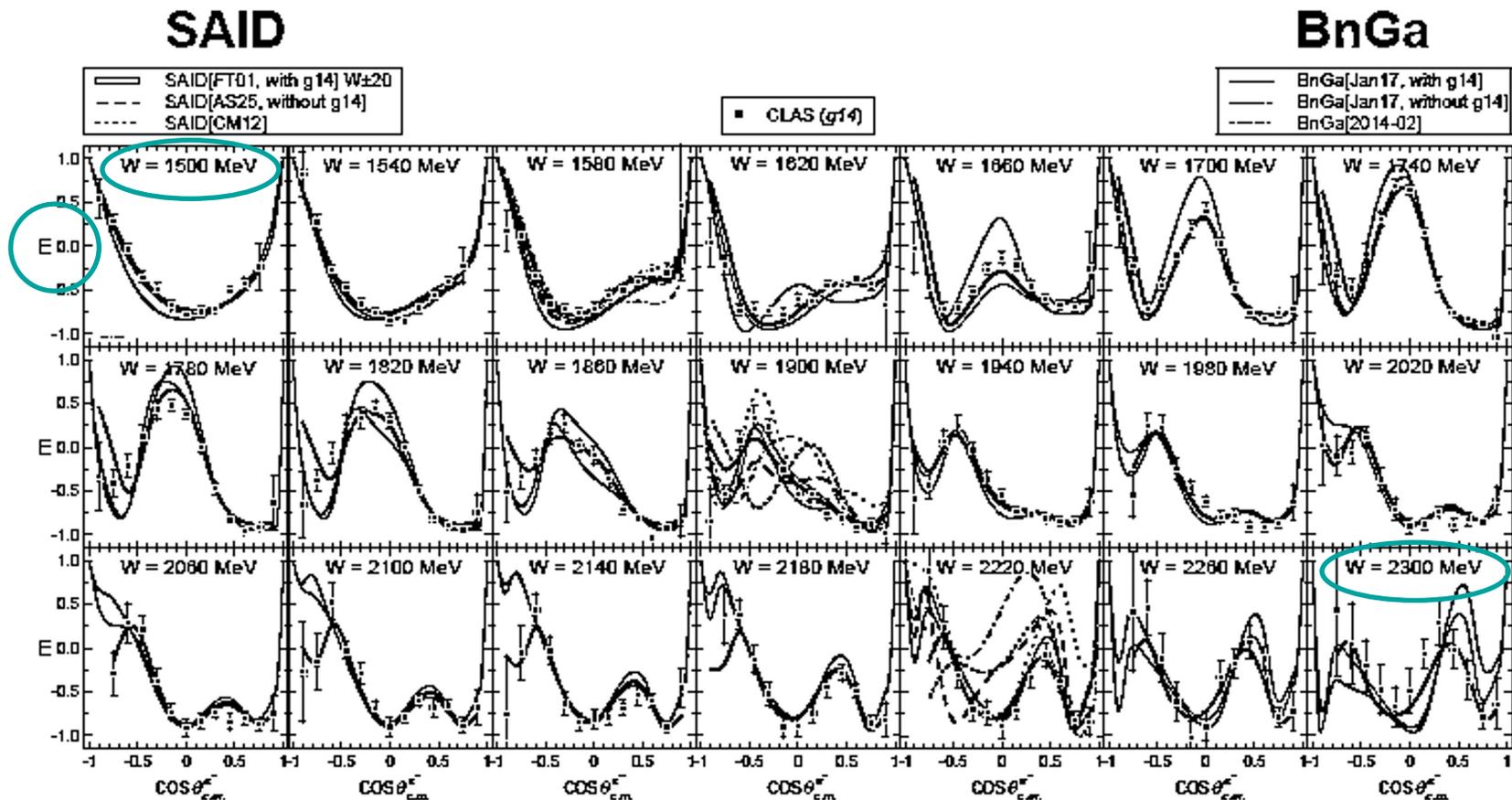


$E = 445\text{--}2510$ MeV
 $\pi^- p: 8428$ $d\sigma/d\Omega$

• These data a factor of nearly three increase in world statistics for this channel in this kinematic range.

FSI included





$E = 730\text{--}2345$ MeV
 π^-p : 266 E

No $\mathcal{F}SI$ included

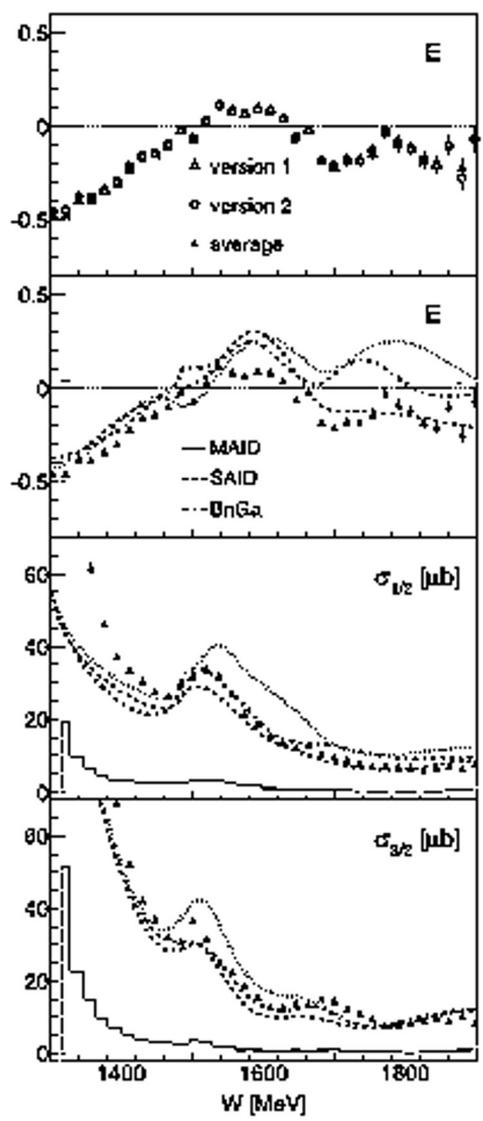
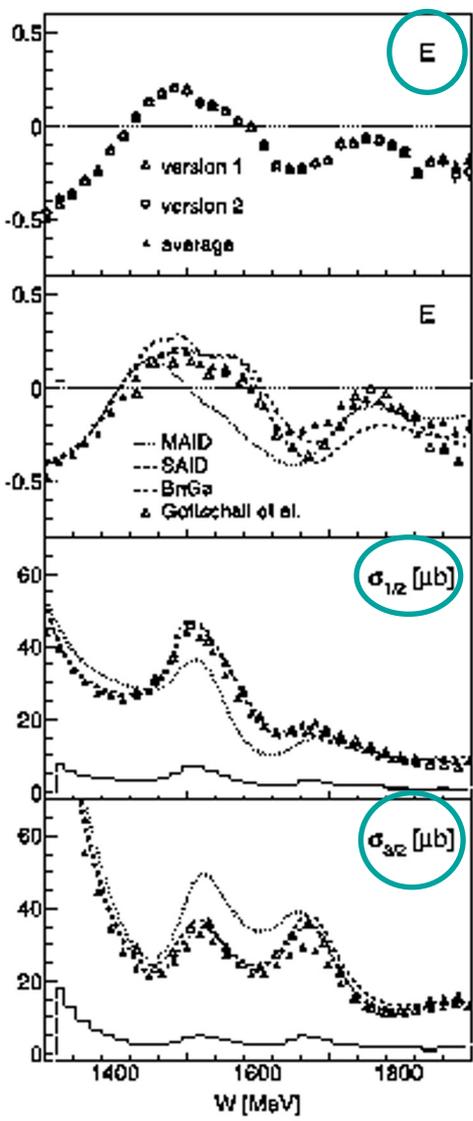


Recent CB@MAMI E for $\vec{\gamma}\vec{n} \rightarrow \pi^0 n$

M.Dieterle et al Phys Let B 523, 770 (2017)

“Proton”

“Neutron”



$[(\text{GeV})^{-1/2} \times 10^{-3}]$

$N(1680)5/2^+ \rightarrow N\gamma$

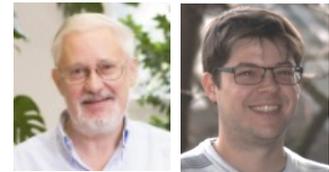
$pA^{3/2} = +133 \pm 12$ $pA^{1/2} = -15 \pm 6$
 $nA^{3/2} = -33 \pm 9$ $nA^{1/2} = +29 \pm 10$

• It couples weakly to **neutron**.

$N(1675)5/2^- \rightarrow N\gamma$

$pA^{3/2} = +20 \pm 5$ $pA^{1/2} = +19 \pm 8$
 $nA^{3/2} = -85 \pm 10$ $nA^{1/2} = -60 \pm 5$

• It couples strongly to **neutron**.

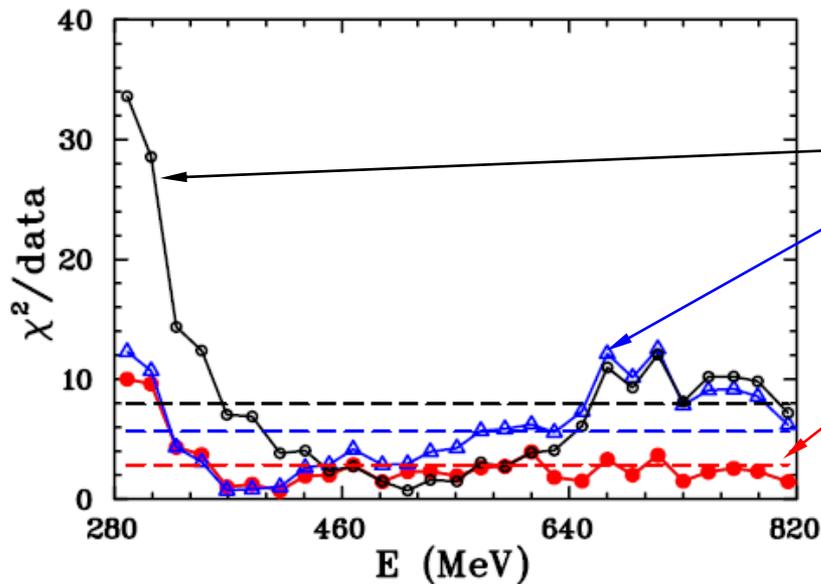


No FSI included

Comparison of Previous & New SAID Fits for $\gamma n \rightarrow \pi^0 n$



WJB et al, Phys Rev C **100**, 065205 (2019)



- Recent SAID MA27 applied to $\mathbf{A2}$ data without & with FSI corrections.
- New SAID MA19 fit obtained after adding new $\mathbf{A2}$ data with FSI corrections.

Obviously, FSI plays important role in $\gamma n \rightarrow \pi^0 n$ $d\sigma/d\Omega$ determination.

Reaction	MAID2007	SAID MA27	SAID MA19
Present $\gamma n \rightarrow \pi^0 n$	1855/492=3.77	2765/492=5.62	1405/492=2.86
Previous $\gamma p \rightarrow \pi^0 p$	156700/13988=11.20	25856/13988=1.85	23954/13988=1.71
$\gamma p \rightarrow \pi^+ n$	121150/5225=23.19	10785/5225=2.06	10371/5225=1.99
$\gamma n \rightarrow \pi^- p$	49471/4142=11.94	7087/4142=1.71	6530/4142=1.58
$\gamma n \rightarrow \pi^0 n$	27060/515=52.54	2958/515=5.74	2320/515=4.51
Total	354373/23870=14.85	46686/23870=1.96	43174/23870=1.81

FSI included

WJB et al, Phys Rev C **100**, 065205 (2019) • $\mathbf{A2}$ data included in SAID fits.

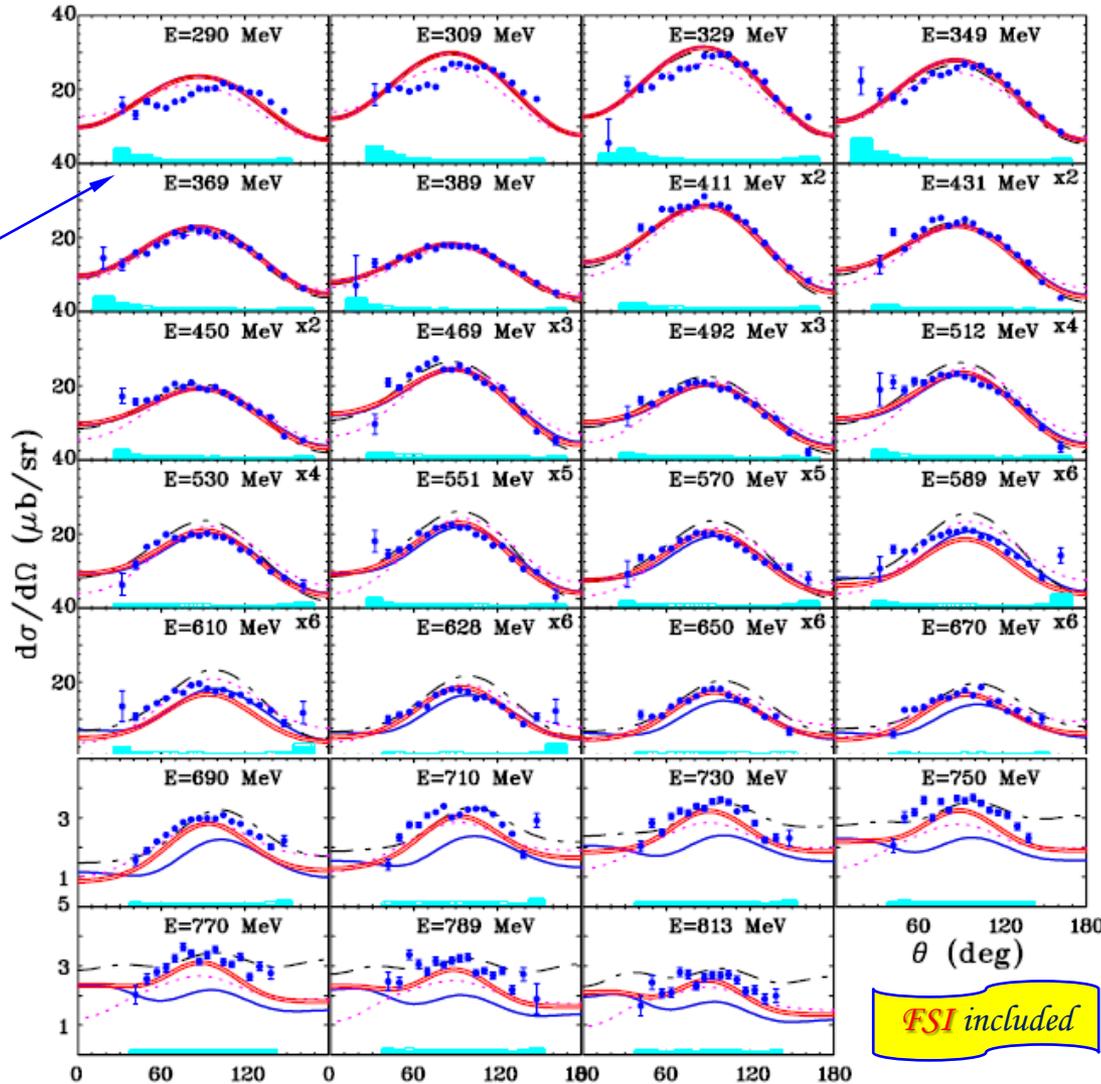
P. T. Mattione et al, Phys Rev C **96**, 035204 (2017)

D. Drechsel, S. S. Kamalov, & L. Tiator, Eur. Phys. J. A **34**, 69 (2007)

- Differential cross sections for $\gamma n \rightarrow \pi^0 n$.

- New **A2** data cover broader angular rangy & focus on low energies.

systematics



$E = 290 - 813$ MeV
 $W = 1195 - 1533$ MeV
 $\theta = 18 - 162^\circ$
 $\pi^0 n: 492$ $d\sigma/d\Omega$

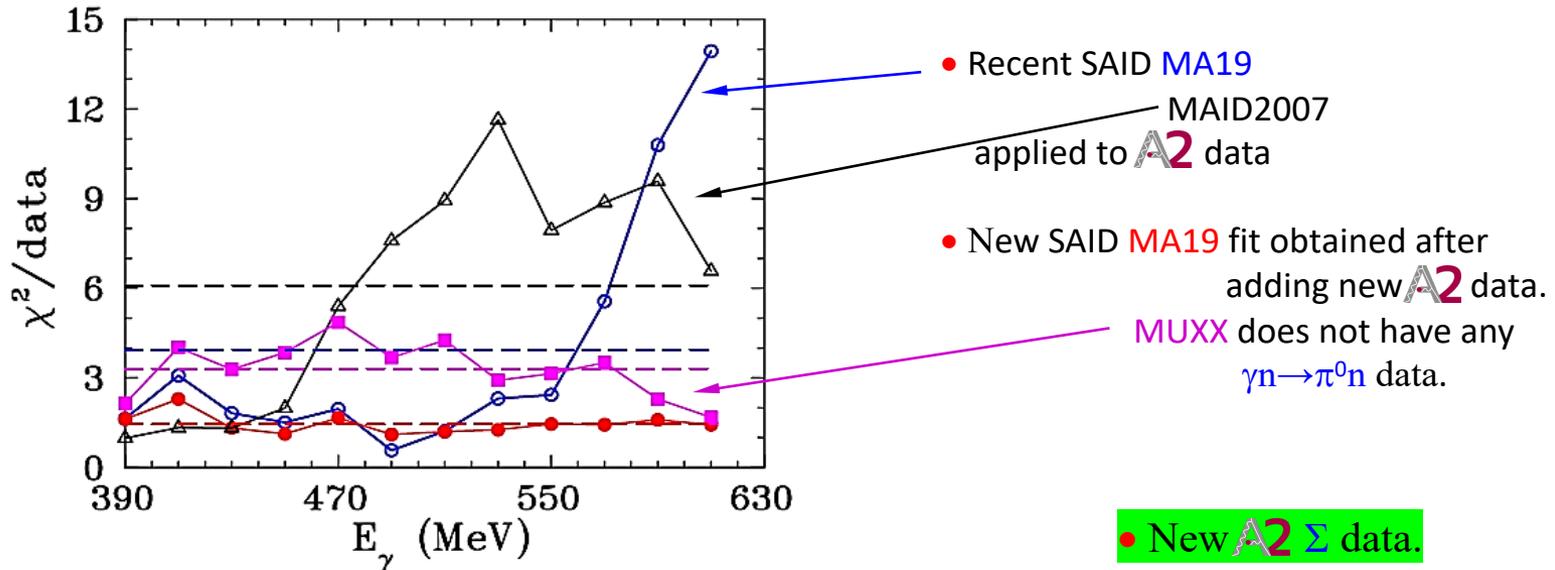
PWA:
 SAID MA19
 SAID MA27
 MAID2007
 BnGa2002

FSI included

Comparison of Previous & New SAID Fits for $\gamma n \rightarrow \pi^0 n$



C. Mullen *et al*, arXiv:2103.08400 [nucl-ex]



• $E_\gamma = 155$ to 1000 MeV

Solution	$\chi^2/(\pi^0 p \text{ data})$	$\chi^2/(\pi^+ n \text{ data})$
MU22	13274/9534=1.39	7454/4039=1.85
MUXX	13171/9534=1.38	7259/4039=1.80
MA19	12565/9534=1.32	7461/4039=1.85
MAID2007	73638/9534=7.72	14599/4039=3.61

Solution	$\chi^2/(\pi^0 n \text{ data})$	$\chi^2/(\pi^- p \text{ data})$
MU22	2345/ 798=2.94	5879/3456=1.70
MUXX	7639/ 798=9.57	5384/3456=1.56
MA19	2649/ 798=3.32	5999/3456=1.74
MAID2007	4846/ 798=6.07	15365/3456=4.45

Solution	$\chi^2/(\pi^0 n \text{ data})$
MU22	275/189=1.46
MUXX	624/189=3.30
MA19	743/189=3.93
MAID2007	1151/189=6.09

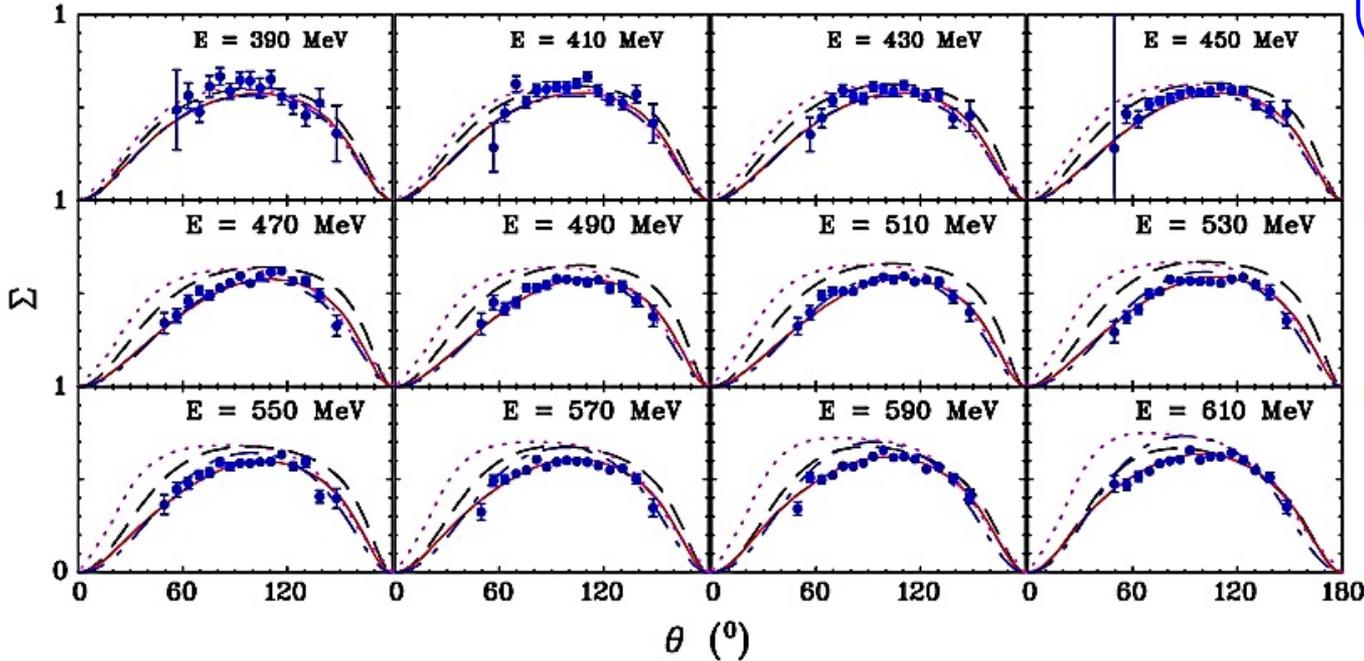
C. Mullen *et al*, arXiv:2103.08400 [nucl-ex]

WJB *et al*, Phys Rev C **100**, 065205 (2019)

D. Drechsel, S. S. Kamalov, & L. Tiator, Eur. Phys. J. A **34**, 69 (2007)

- Σ beam asymmetry for $\gamma n \rightarrow \pi^0 n$.

$E = 390 - 610$ MeV
 $W = 1271 - 1424$ MeV
 $\theta = 49 - 148^\circ$
 $\pi^0 n: 189 \Sigma$



- Systematic uncertainties for each bin have been added in quadrature.

- New **A2** data cover parts of Δ & **Roper** resonance regions.
- There are no significant changes in dominant multipoles below **1** GeV.

C. Mullen *et al*, arXiv:2103.08400 [nucl-ex]

WJB *et al*, Phys Rev C **100**, 065205 (2019)

D. Drechsel, S. S. Kamalov, & L. Tiator, Eur. Phys. J. A **34**, 69 (2007)

E. Gutz *et al*, Eur Phys J A **50**, 74 (2014)

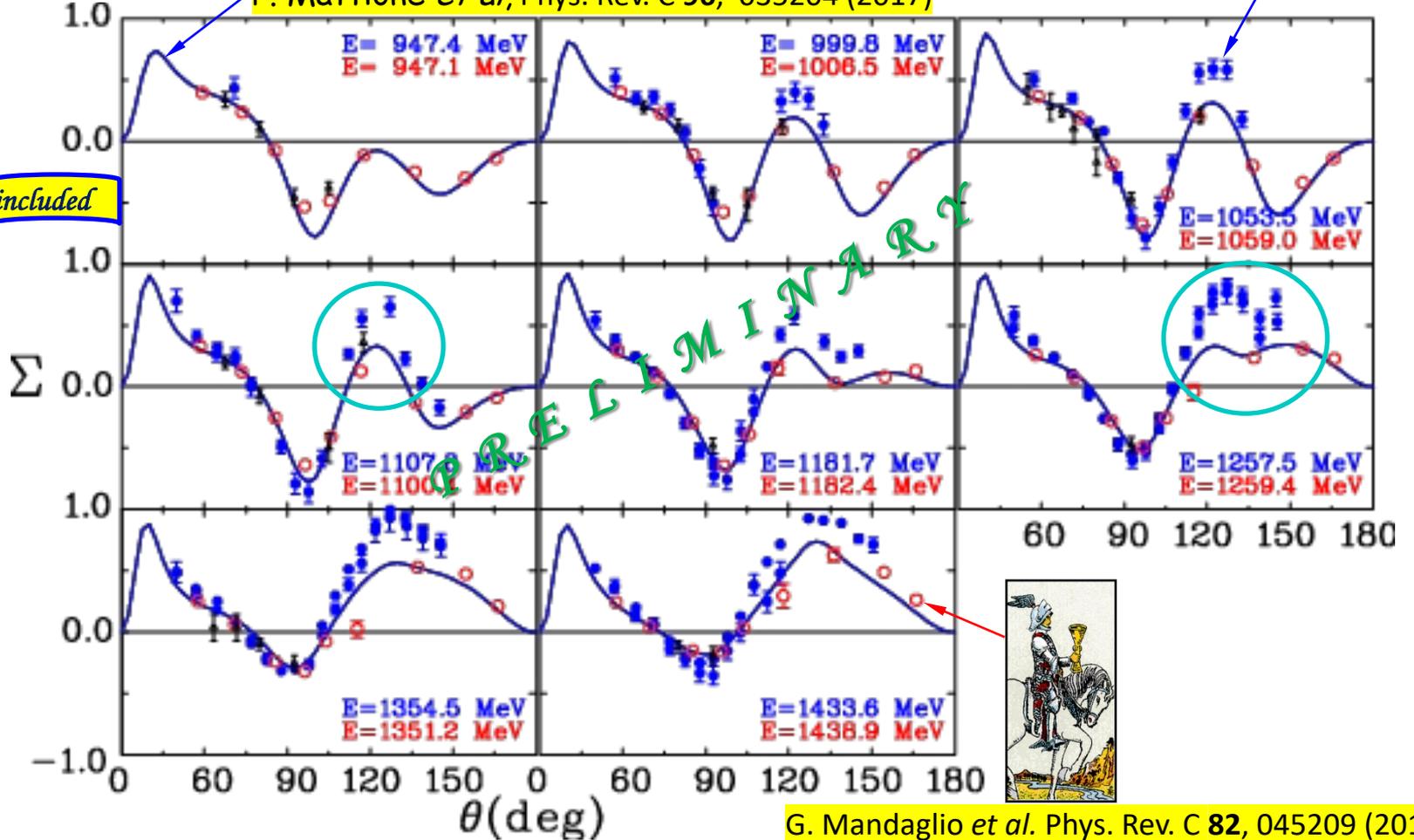
PWA:
 SAID MU22
 SAID MA19
 MAID2007
 BnGa2002

Σ for $\vec{\gamma}n \rightarrow \pi^- p$

D. Sokhan, 8th International Conference on Quarks and Nuclear Physics 2018, Tsukuba, Japan

P. Mattione *et al*, Phys. Rev. C 96, 035204 (2017)

No FSI included



G. Mandaglio *et al*. Phys. Rev. C 82, 045209 (2010)



Conclusions

We are all in it together!

