# JLab E12-14-012 experiment: 2014-2023

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## Outline

- 2014 PAC 42 talk
- E12-14-012 Motivations:
  - Neutrino Oscillation Experiments
  - Importance of cross sections in oscillation results
- Experimental setup
- (e,e') results on C, Ar, Ti and Al
- (e,e'p) results on Ar and Ti

### 2014: Benhar's talk at PAC-42

#### Motivation

- ★ The interpretation of the signals detected by most neutrino experiments require a *quantitative* understanding of the nuclear response to electroweak interaction
- ★ The results of numerical studies suggest that the impact of nuclear modelling on the determination of oscillation parameters may be of the order of 10%
- ★ Over the past decades electron scattering has provided a wealth of information (nucleon form factors, inclusive double differential cross sections, coincidence, semi-inclusive cross sections) that are now beginning to be exploited in the analysis of neutrino data
- ★ Currently unavailable electron scattering data will be needed for the analysis of future, high precision, neutrino experiments using liquid Argon detectors
- ★ The proposed measurement will provide information required for the simulation of neutrino interactions in argon *in all channels*

### E12-14-012: Goals and Motivations

- <u>Primary Goal</u>: Measurement of the spectral functions of Argon and Titanium through Ar-Ti (*e,e'p*) reactions
   Data Collected (Feb-March 2017):
  - Ar/Ti/C/Dummy/Optical (*e*,*e*′*p*) reactions for five different kinematic set-ups

- Ar/Ti/C/Dummy (*e*,*e*') reactions for one kinematic set-up

- <u>**Primary Motivation**</u>: To help improve the accuracy of the measurement of the neutrino-oscillation parameters, including the *CP violation in leptonic sector* (one of the top priority of the US particle physics community), in the future neutrino experiments, mainly DUNE, by:
- Measuring spectral function of argon (~ initial momentum and energy distributions of nucleons bound in argon) that can directly be used in the reconstruction of neutrino energies (currently the major source of uncertainty in neutrino experiments).
- Using measured argon spectral functions to further develop (extend) a fully consistent parameter-free theoretical (neutrino-nucleus) model that can be used in (every step of) the analysis of long baseline neutrino experiments.

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### Neutrino Oscillation Experiments



### Signature of Neutrino Oscillation

• Neutrino Spectrum Distortions

### Near to Far extrapolation

- Provides data-driven estimate of un-oscillated event rate at the Far detector.
- Neutrino spectrum distortions calculated from the ratio of neutrino spectrum at far detector to un-oscillated predicted event rate at far detector
- Influenced by uncertainties in the knowledge of flux and cross sections.

### **Observable Oscillation Parameters**

$$P(\nu_{\mu} \rightarrow \nu_{e}) = \sin^{2} 2\theta \sin^{2} \left( \frac{\Delta m^{2} L}{4E_{\nu}} \right)$$

Neutrino energy >1 GeV

Oscillation parameter determination depends on the reconstructed neutrino energy.



### **DUNE** Experiment



Experiments measure event rates which, for a given observable topology, can be naively computed as:

 $\frac{\text{Event Rate at near detector:}}{N_{\text{ND}}^{\alpha}(\boldsymbol{p}_{\text{reco}}) = \sum_{i} \phi_{\alpha}(E_{\text{true}}) \times \sigma_{\alpha}^{i}(\boldsymbol{p}_{\text{true}}) \times \epsilon_{\alpha}(\boldsymbol{p}_{\text{true}}) \times R_{i}(\boldsymbol{p}_{\text{true}}; \boldsymbol{p}_{\text{reco}})}$ 

$$\underbrace{ \frac{\text{Event Rate at far detector:}}{N_{\text{FD}}^{\alpha \to \beta}(\boldsymbol{p}_{\text{reco}}) = \sum_{i} \phi_{\alpha}(E_{\text{true}}) \times P_{\alpha\beta}(E_{\text{true}}) \times \sigma_{\beta}^{i}(\boldsymbol{p}_{\text{true}}) \times \epsilon_{\beta}(\boldsymbol{p}_{\text{true}}) \times R_{i}(\boldsymbol{p}_{\text{true}}; \boldsymbol{p}_{\text{reco}}) } }_{i}$$

### Neutrino Interactions



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### Deep Underground Neutrino Experiment

- DUNE is a future long-baseline neutrino oscillation and proton decay experiment
- 1300 km baseline, new broad-band (0.5–7 GeV) mega-watt scale  $v_{\mu}/\overline{v_{\mu}}$  beam from Fermilab to SURF
  - → 1.2 MW (2026) up to 2.4 MW (2032)
- Large international collaboration: 1100+ physicists, 178+ institutions, 32+ countries





### Appearance Probability as function of neutrino energy



Need energy to distinguish between different  $\delta_{CP}$ 

### Hall A at Jefferson Lab



### (e,e') and (e,e'p) processes

- (e, e'p) process (exclusive):
   Both outgoing electron and proton are detected
- (e,e') process (inclusive):
   Only scattered electron is detected

$$e + A \rightarrow e' + p + (A - 1)$$

$$e + A \rightarrow e' + X$$



### Why Titanium?

The shell model structure of the **protons in Ti** is nearly identical to that of the **neutrons in Ar** 



### **Inclusive analysis**



Ti(e,e') and C(e,e') inclusive cross sections published Phys. Rev. C 98, 014617 (2018) Ar(e,e') inclusive cross published in Phys. Rev. C 99, 054608 (2019) Al(e,e') inclusive cross section analysis published in Phys. Rev. C 100, 054606 (2019).

### (e,e'p) analysis: extraction of the spectral function



• Universal property of the nucleus, independent of the interaction.

### (e,e'p) cross section



## Missing momentum $\mathbf{p}_m$ and missing energy $E_m$

Without final-state interactions,

$$(E_e, \mathbf{k}_e) \xrightarrow{(E_e', \mathbf{k}_e')} (E_p', \mathbf{p}')$$

$$E_e + M - \underline{E_m} = E_e' + E_p'$$

known

$$\mathbf{k}_e + \mathbf{p}_m = \mathbf{k}_e \mathbf{,} + \mathbf{p'}$$

 $E_m - E_{\text{thr}}$  is the excitation energy  $p_m \equiv |\mathbf{p}_m|$  is the initial proton momentum

### Analysis procedure

1) Extract of the (*e,e'p*) cross section

2) Using  $\sigma_{cc1}$  of de Forest and nuclear transparency, obtain the reduced cross sections as a function of (a)  $p_m$  and (b)  $E_m$ .

3) Find the parameters of the spectral function (*i.e.*, spectroscopic factors) from the fits to the reduced cross sections as a function of  $p_m$ .

4) Using the priors from Step 3), find the parameters of the spectral function (*i.e.*, spectroscopic factors, peak positions, distribution widths) from the fits to the reduced cross sections as a function of *E*<sub>m</sub>. Correct for transparency.

## Exclusive analysis – kin1 - Ar - Missing energy and missing momentum

#### Effect of FSI



### Fit procedure and minimization

For each bin in the spectra of missing energy (100 bins between 1 and 100 MeV) and missing momentum (40 bins with momentum range changing between kinematics), we determined the product of the reduced MC cross section and the ratio of the data to simulation yield

$$\frac{d^2 \sigma_{cc1}^{\rm red}}{d\Omega dE'} = \Big(\frac{d^2 \sigma_{cc1}^{\rm red}}{d\Omega dE'}\Big)_{\rm MC} \times \frac{Y(E',\theta)}{Y_{\rm MC}(E',\theta)}$$

The fit performs a  $\chi^2$  minimization using the TMinuit package available in root.

The index i labels the missing momentum(energy) bin,  $\alpha$  is the orbital index,  $f^{pred}_{\alpha}(i)$  is the parametrized prediction evaluated at bin i in the missing momentum spectra for orbital  $\alpha$ ,  $S_{\alpha}$  is the spectroscopic factor.

$$\chi^2 = \sum_i \chi_i^2 = \sum_i \left( \frac{\sigma_i^{\text{red, obs}} - \sum_\alpha S_\alpha f_\alpha^{\text{pred}}(i)}{\sigma_{\sigma_i^{\text{red, obs}}}} \right)^2$$

### *Ar* (*e*,*e*'*p*) – Phys. Rev. D 105, 112002, (2022)



		all priors	w/o $p_m$	w/o corr.
$\alpha$	$N_{lpha}$		$S_{lpha}$	
$1d_{3/2}$	2	$0.89\pm0.11$	$1.42\pm0.20$	$0.95\pm0.11$
$2s_{1/2}$	2	$1.72\pm0.15$	$1.22\pm0.12$	$1.80\pm0.16$
$1d_{5/2}$	6	$3.52\pm0.26$	$3.83\pm0.30$	$3.89\pm0.30$
$1p_{1/2}$	2	$1.53\pm0.21$	$2.01\pm0.22$	$1.83\pm0.21$
$1p_{3/2}$	4	$3.07\pm0.05$	$2.23\pm0.12$	$3.12\pm0.05$
$1s_{1/2}$	2	$2.51\pm0.05$	$2.05\pm0.23$	$2.52\pm0.05$
corr.	0	$3.77\pm0.28$	$3.85\pm0.25$	excluded
$\sum_{\alpha} S_{\alpha}$		$17.02\pm0.48$	$16.61\pm0.57$	$14.12\pm0.42$
d.o.f		206	231	232
$\chi^2/{ m d.o.f.}$		1.9	1.4	2.0

	1.0		· · · · · · · ]	
() ()	$0.8$ $\frac{2s_{1/2}}{2s_{1/2}}$			
1/Me	0.6		-	
$\left( E_{m} ight)$ (	0.4	P3/2	-	
S(	0.2	corr $1s_{1/2}$		
	0.0 0 20	40 60	80 100	
		$E_m (MeV)$		
2	$E_{\alpha}$ (1	$\frac{E_m (\text{MeV})}{\text{MeV}}$	$\sigma_{\alpha}$ (1	MeV)
α	$E_{\alpha}$ (1) w/ priors	$\frac{E_m (MeV)}{MeV)}$ w/o priors	$\frac{\sigma_{\alpha}}{w/ \text{ priors}}$	MeV) w/o priors
$\frac{lpha}{1d_{3/2}}$	$\frac{E_{\alpha}}{\text{w/ priors}}$ $12.53 \pm 0.02$	$     \frac{E_m \text{ (MeV)}}{\text{MeV)}} \\     \frac{\text{w/o priors}}{10.90 \pm 0.12} $	$\frac{\sigma_{\alpha}}{\text{w/ priors}}$ $1.9 \pm 0.4$	$\frac{\text{MeV}}{\text{w/o priors}}$ $1.6 \pm 0.4$
$rac{lpha}{1d_{3/2}} \ 2s_{1/2}$	$     \frac{E_{\alpha}}{\text{w/ priors}} \\     \frac{12.53 \pm 0.02}{12.92 \pm 0.02} $	$\frac{E_m \text{ (MeV)}}{\text{MeV)}}$ $\frac{\text{W/o priors}}{10.90 \pm 0.12}$ $12.57 \pm 0.38$	$ \frac{\sigma_{\alpha} (1)}{\text{w/ priors}} $ $ \frac{1.9 \pm 0.4}{3.8 \pm 0.8} $	$\frac{\text{MeV})}{\text{w/o priors}}$ $\frac{1.6 \pm 0.4}{3.0 \pm 1.8}$
$rac{lpha}{1d_{3/2}} \ 2s_{1/2} \ 1d_{5/2}$	$     \frac{E_{\alpha}}{\text{w/ priors}}     12.53 \pm 0.02     12.92 \pm 0.02     18.23 \pm 0.02     $	$E_m (MeV)$ MeV) w/o priors 10.90 ± 0.12 12.57 ± 0.38 17.77 ± 0.80	$     \frac{\sigma_{\alpha} (1)}{\text{w/ priors}}     1.9 \pm 0.4     3.8 \pm 0.8     9.2 \pm 0.9   $	$\frac{\text{MeV})}{\text{w/o priors}} \\ 1.6 \pm 0.4 \\ 3.0 \pm 1.8 \\ 9.6 \pm 1.3 \\ \end{array}$
$rac{lpha}{1d_{3/2}}\ 2s_{1/2}\ 1d_{5/2}\ 1p_{1/2}$	$     \begin{array}{r} E_{\alpha} (1) \\ \hline W \text{ priors} \\ 12.53 \pm 0.02 \\ 12.92 \pm 0.02 \\ 18.23 \pm 0.02 \\ 28.8 \pm 0.7 \end{array} $	$\frac{E_m \text{ (MeV)}}{\text{MeV)}}$ $\frac{\text{MeV)}}{10.90 \pm 0.12}$ $12.57 \pm 0.38$ $17.77 \pm 0.80$ $28.7 \pm 0.7$	$ \frac{\sigma_{\alpha}}{\text{w/ priors}} \\ \frac{1.9 \pm 0.4}{3.8 \pm 0.8} \\ 9.2 \pm 0.9 \\ 12.1 \pm 1.0 $	$\begin{tabular}{ c c c c c c } \hline \hline MeV) \\ \hline $w/o$ priors \\ \hline $1.6 \pm 0.4$ \\ $3.0 \pm 1.8$ \\ $9.6 \pm 1.3$ \\ $12.0 \pm 3.6$ \end{tabular}$
$rac{lpha}{1d_{3/2}}\ 2s_{1/2}\ 1d_{5/2}\ 1p_{1/2}\ 1p_{3/2}$	$\begin{array}{r} E_{\alpha} (1) \\ \hline W/ \text{ priors} \\ 12.53 \pm 0.02 \\ 12.92 \pm 0.02 \\ 18.23 \pm 0.02 \\ 28.8 \pm 0.7 \\ 33.0 \pm 0.3 \end{array}$	$\frac{E_m \text{ (MeV)}}{\text{MeV}}$ $\frac{\text{MeV}}{10.90 \pm 0.12}$ $12.57 \pm 0.38$ $17.77 \pm 0.80$ $28.7 \pm 0.7$ $33.0 \pm 0.3$	$     \begin{array}{r} \sigma_{\alpha} (1) \\     \hline         w/ \text{ priors} \\         1.9 \pm 0.4 \\         3.8 \pm 0.8 \\         9.2 \pm 0.9 \\         12.1 \pm 1.0 \\         9.3 \pm 0.5 \\         \end{array} $	$\begin{tabular}{ c c c c c c } \hline MeV) \\ \hline $w/o$ priors \\ \hline $1.6 \pm 0.4$ \\ $3.0 \pm 1.8$ \\ $9.6 \pm 1.3$ \\ $12.0 \pm 3.6$ \\ $9.3 \pm 0.5$ \end{tabular}$
$lpha \ 1d_{3/2} \ 2s_{1/2} \ 1d_{5/2} \ 1p_{1/2} \ 1p_{3/2} \ 1s_{1/2} \ 1s_{1/2}$	$\begin{array}{r} E_{\alpha} (1) \\ \hline W/ \text{ priors} \\ 12.53 \pm 0.02 \\ 12.92 \pm 0.02 \\ 18.23 \pm 0.02 \\ 28.8 \pm 0.7 \\ 33.0 \pm 0.3 \\ 53.4 \pm 1.1 \end{array}$	$\frac{E_m \text{ (MeV)}}{\text{MeV)}}$ $\frac{\text{W/o priors}}{10.90 \pm 0.12}$ $12.57 \pm 0.38$ $17.77 \pm 0.80$ $28.7 \pm 0.7$ $33.0 \pm 0.3$ $53.4 \pm 1.0$	$ \frac{\sigma_{\alpha}}{\text{w/ priors}} \\ \frac{1.9 \pm 0.4}{3.8 \pm 0.8} \\ 9.2 \pm 0.9 \\ 12.1 \pm 1.0 \\ 9.3 \pm 0.5 \\ 28.3 \pm 2.2 \end{array} $	$\begin{tabular}{ c c c c c } \hline MeV) \\ \hline $ w/o$ priors \\ \hline $ 1.6 \pm 0.4 \\ $ 3.0 \pm 1.8 \\ $ 9.6 \pm 1.3 \\ $ 12.0 \pm 3.6 \\ $ 9.3 \pm 0.5 \\ $ 28.1 \pm 2.3 \\ \hline \end{tabular}$

### *Ti* (*e*,*e*'*p*) – Phys. Rev. D 107, 012005, (2023)



50				A.
		all priors	w/o $p_m$	w/o corr.
α	$N_{lpha}$		$S_{lpha}$	
$1f_{7/2}$	2	$1.53\pm0.25$	$1.55\pm0.28$	$1.24\pm0.22$
$1d_{3/2}$	4	$2.79\pm0.37$	$3.15\pm0.54$	$3.21\pm0.37$
$2s_{1/2}$	2	$2.00\pm0.11$	$1.78\pm0.46$	$2.03\pm0.11$
$1d_{5/2}$	6	$2.25\pm0.16$	$2.34\pm0.19$	$3.57\pm0.29$
$1p_{1/2}$	2	$2.00\pm0.20$	$1.80\pm0.27$	$2.09\pm0.19$
$1p_{3/2}$	4	$2.90\pm0.20$	$2.92\pm0.20$	$4.07\pm0.15$
$1s_{1/2}$	2	$2.14\pm0.10$	$2.56\pm0.30$	$2.14\pm0.11$
corr.	0	$4.71\pm0.31$	$4.21\pm0.46$	excluded
$\sum_{\alpha} S_{\alpha}$		$20.32\pm0.65$	$20.30 \pm 1.03$	$18.33 \pm 0.59$
d.o.f		121	153	125
$\chi^2/{ m d.o.f.}$		0.95	0.71	1.23



	$E_{lpha}$ (1	MeV)	$\sigma_{lpha}~({ m MeV})$		
$\alpha$	w/ priors	w/o priors	w/ priors	w/o priors	
$1f_{7/2}$	$11.32\pm0.10$	$11.31\pm0.10$	$8.00 \pm 5.57$	$8.00\pm6.50$	
$1d_{3/2}$	$12.30\pm0.24$	$12.33\pm0.24$	$7.00\pm0.61$	$7.00\pm3.84$	
$2s_{1/2}$	$12.77\pm0.25$	$12.76\pm0.25$	$7.00\pm3.76$	$7.00\pm3.84$	
$1d_{5/2}$	$15.86\pm0.20$	$15.91\pm0.22$	$2.17\pm0.27$	$2.23\pm0.29$	
$1p_{1/2}$	$33.33\pm0.60$	$33.15\pm0.65$	$3.17\pm0.45$	$3.03\pm0.48$	
$1p_{3/2}$	$39.69 \pm 0.62$	$39.43 \pm 0.68$	$5.52\pm0.70$	$5.59\pm0.70$	
$1s_{1/2}$	$53.84 \pm 1.86$	$52.00 \pm 3.13$	$11.63 \pm 1.90$	$13.63 \pm 2.59$	
corr.	$25.20\pm0.02$	$25.00\pm0.29$	100 m	1	

### Summary

- We completed the data analysis for the full data set collected by the E12-14-012 experiment at Jefferson Lab in 2017.
- Data has been published for both inclusive and exclusive analysis 5 publications from 2017, 4 PhD thesis.
  - Inclusive cross sections for C and Ti, [Dai et al., PRC 98, 014617 (2018)]
  - Inclusive cross section for Ar, [Dai et al., PRC 99, 054608 (2019)]
  - Inclusive cross section for Al-7075, Ar, C and Ti of all (*e,e'*) data [Murphy *et al.*, PRC 100, 054606 (2019)]
  - Exclusive cross section for Ar, [Jiang et al., PRD 105, 112002 (2022)]
  - Exclusive cross section for Ti, [Jiang et al., PRD 107, 012005 (2023)]
  - H. Dai (VT), PhD thesis, 2019
  - M. Murphy (VT), PhD thesis, 2020
  - L. Gu (VT), PhD thesis, 2021
  - D. Abrams (UVA), PhD thesis, 2022
  - C. Lanham (VT), MS thesis, 2023

### **Summary**

- Separation of individual contributions requires improved analysis. Numerous theoretical developments are necessary. But we have data now.
- Interpretation of the Ti proton spectral function in terms of neutron spectral function on Ar is in progress, new theory development will be needed.

## Thank you

## Back up

## Test spectral function: 80% mean-field + 20% NN correlations



Independent-particle shell model

### Mean-field part of the spectral function



Relativistic MF calculations by C. Giusti

## Exclusive analysis - kin1 - Ti - Missing energy and missing momentum



## Ar missing energy distribution



Shapes – drawn in different colors represent the contribution of different orbitals

First three orbital shapes are estimates

Last three level are derived from data

## FSI analysis

- C. Giusti provided a relativistic code, tested against old data on <sup>16</sup>O and <sup>12</sup>C up to <sup>40</sup>Ca
- Compute reduced cross section both PWIA and DWIA for various wave functions, identify the energy for each orbital

α	$E_{\alpha}$	$\sigma_{lpha}$	$E^{lpha}_{ m low}$	$E^{\alpha}_{\mathrm{high}}$
			argon	
$1d_{3/2}$	12.53	2	8	14
$2s_{1/2}$	12.93	2	8	14
$1d_{5/2}$	18.23	4	14	20
$1p_{1/2}$	28.0	8	20	45
$1p_{3/2}$	33.0	8	20	45
$1s_{1/2}$	52.0	8	45	70
			titanium	
$1f_{7/2}$	11.45	2	8	14
$2s_{1/2}$	12.21	2	14	30
$1d_{3/2}$	12.84	2	14	30
$1d_{5/2}$	15.46	4	14	30
$1p_{1/2}$	35.0	8	30	54
$1p_{3/2}$	40.0	8	30	54
$1s_{1/2}$	62.0	8	53	80

### Boiling Study----Nathaly Santiesteban and H. Dai

- We calculated the normalized yield for different currents, and the change in yield represents change in target density
- The normalization is done with respect to the lowest current
- We fit the numbers with quadratic function and fix the I=0 point to 1
- When  $I = 9.67 \mu A$ , within 2% for all the runs, the boiling effect is 17.2%, with 0.7% uncertainty.

				1.00		– – init — best-fit
Current (µA)	Number of events	Yield (ev/μC)	Normalized Yield	1.00	-	Ar Best Fit A = 0.00039+/-0.00016
2.65 +/- 0.14	4898	1571.63 +/- 23.86	1 +/- 0.015	0.95	-	B = -0.02157 + 7 - 0.00387 C = 1.00000 + 7 - 0.02087 $\chi^2 / \text{ndf} = 1.764$
4.39+/-0.14	10283	1523.80 +/- 15.97	0.97 +/- 0.01	dield 20.90	-	
8.06 +/- 0.15	17460	1454.32 +/- 11.69	0.925 +/- 0.007	Vormalize		
11.81 +/- 0.17	26848	1352.62 +/- 8.77	0.860 +/- 0.005	∠ 0.85		
15.15 +/- 0.19	25764	1287.83 +/- 8.52	0.8194 +/- 0.0054	0.80	-	
18.08 +/- 0.21	26065	1263.59 +/- 8.31	0.804 +/- 0.0053	0.75	-	
					2 4 6 8 1	0 12 14 16 18

curent  $\mu A$ 

### **Kinematic Setup**

	$E'_e$	$ heta_e$	$ \mathbf{p}' $	$\theta_{p'}$	$ \mathbf{q} $	$p_m$	$E_m$
	$(\mathrm{GeV})$	(deg)	(MeV)	(deg)	(MeV)	(MeV)	(MeV)
kin1	1.777	21.5	915	-50.0	865	50	73
kin2	1.716	20.0	1030	-44.0	846	184	50
kin3	1.799	17.5	915	-47.0	741	174	50
kin4	1.799	15.5	915	-44.5	685	230	50
kin5	1.716	15.5	1030	-39.0	730	300	50



kin1			kin3		
Collected Data	Hours	Events(k)	Collected Data	Hours	Events(k)
Ar Ti Dummy	29.6 12.5 0.75	43955 12755 955	Ar Ti Dummy	13.5 8.6 0.6	73176 28423 2948
kin2			kin4		
Collected Data	Hours Events(k)		Collected Data	Hours	Events(k)
Ar Ti Dummy Optics C	32.1 18.7 4.3 1.15 2.0	62981 21486 5075 1245 2318	Ar Ti Dummy Optics C	30.9 23.8 7.1 0.9 3.6	158682 113130 38591 4883 21922
kin5	kin5		kin5 - Inclus	ive	
Collected Data	Hours	Events (k)	Collected Data	Minute	es Events(k)
Ar Ti Dummy Optics	12.6 1.5 5.9 2.9	45338 61 16286 160	Ar Ti Dummy C	57 50 56 115	2928 2993 3235 3957

### The (e, e'p) cross section within PWIA

★ Factorization of the final state of the semi-exclusive process

 $e + A \rightarrow e' + p + (A - 1)_n$ 

leads to the simple expression

$$\frac{d\sigma_A}{dE_{e'}d\Omega_{e'}dE_pd\Omega_p} = \mathcal{K} \sigma_{ep} P(p_m, E_m) ,$$

with the missing momentum and missing energy defined by

 $\mathbf{p}_m = \mathbf{p} - \mathbf{q} ,$ 

 $\omega + M_A = \sqrt{(M_A - m + E_m)^2 + |\mathbf{p}_m|^2} + \sqrt{\mathbf{p}^2 + m^2} \rightarrow E_m \approx \omega - T_\mathbf{p} ,$ 

**q** and  $\omega$  being the momentum and energy transfer, respectively

★ Warning: while providing a clear interpretation of the reaction mechanism, PWIA fails to account for final state interactions of the outgoing nucleon

### Inclusion of Final State Interactions (FSI)

- Distorted Wave Impuse Approximation (DWIA). The plane-wave describing the outgoing nucleon is replaced by a *distorted wave*, obtained from a complex optical potential fitted to proton-nucleus scattering data
- ★ The momentum distributions of the shell model states are shifted by an amount  $\Delta_p$  and quenched by a factor  $\tilde{Z}$
- ★ For  $A \le 16$ , the accuracy of the optical potential approach has been tested comparing to the results of many-body calculations of the relevant overlaps



★ FSI effects beyond DWIA are minimized in *parallel kinematics* 

### Correlated part of the spectral function



- Ciofi degli Atti and Simula, PRC 53, 1689 (1996)
- Correlated nucleons form quasi-deuteron pairs, with the relative momentum distributed as in deuteron.
- NN pairs undergo CM motion (Gaussian distrib.)
- Excitation energy of the (A 1)-nucleons is their kinetic energy plus the pn knockout threshold

Ar(e, e'p) – Phys. Rev. D 105, 112002, (2022)



Ti(e, e'p) – Phys. Rev. D 107, 012005, (2023)

