

2024 EIC User Group Early Career Workshop

The Performance of the AstroPix Sensor for the Barrel Imaging Calorimeter in ePIC

BOBAE KIM (bobae.kim@anl.gov)

Argonne National Laboratory

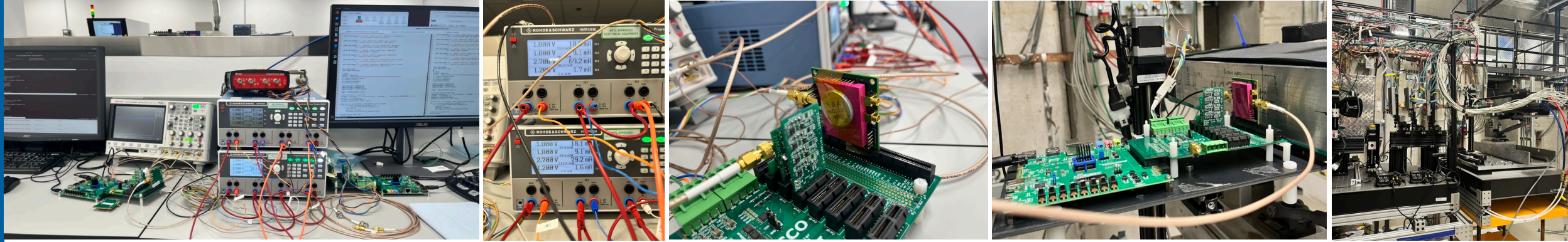


U.S. DEPARTMENT OF ENERGY

Argonne National Laboratory is a
U.S. Department of Energy laboratory
managed by UChicago Argonne, LLC.

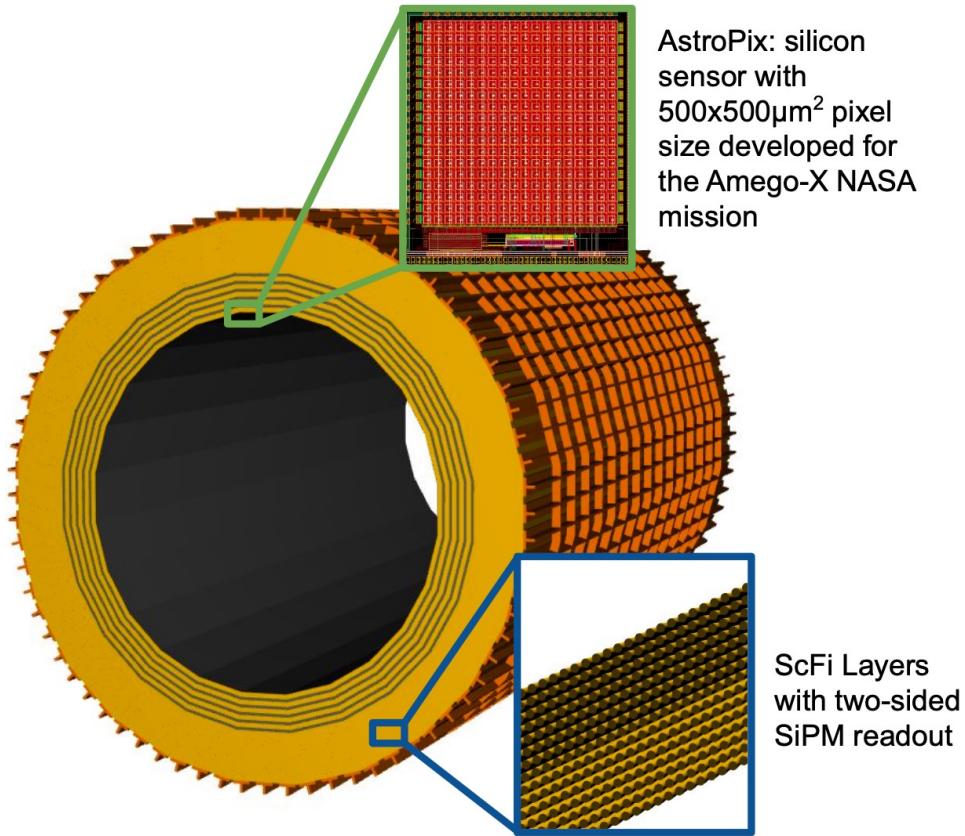


Contents



- Barrel Imaging Calorimeter in EIC detector
- Astropix Sensor: Imaging Part in Barrel Imaging Calorimeter
- Bench Test at ANL
 - Noise scan
 - Injection scan
 - Source test
- Beam Test at FANL: 120 GeV proton
 - Astropix v3 single chip
 - Astropix v3 single chip + Baby bcal
 - Double layer using Astropix v3 single chips
- Summary

Barrel Imaging Calorimeter in EIC detector



- The electromagnetic calorimeter is the main detector for electron-pion separation in the barrel region.
 - Detection of electrons/photons to measure energy and position
- Requirements for Barrel EM Calorimeter
 - Require moderate **energy resolution** $(7 - 10)\%/\sqrt{E} \oplus (1 - 3)\%$
 - Require **electron-pion separation up to 10^4** at low momenta in combination with other detectors
 - Discriminate between **π^0 decays and single γ up to ~ 10 GeV**
 - **Low energy photon** reconstruction ~ 100 MeV
- 4(+2) layers of Astropix sensors interleaved with the first 5 Pb/SciFi layers
- Followed by a **large section of Pb/SciFi section**
- Total radiation thickness $\sim 17.1 X_0$ at $\eta=0$
- Sampling fraction $\sim 10\%$

R&D goals for FY24

- Commissioning and characterizing a small Pb/SciFi calorimeter prototype, termed Baby BCAL, using mixed electron/pion and proton beams at the Fermilab Test Beam Facility (FTBF)
- Testing the first integration of the AstroPix v3 single chip with the Pb/SciFi system and the double layer using AstroPix v3 single chips

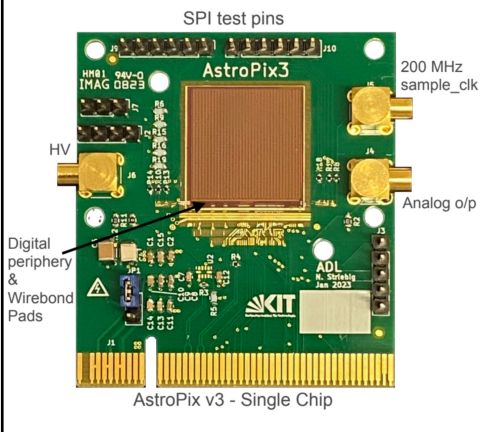

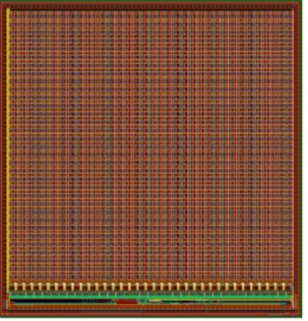
Astropix sensor: Introduction

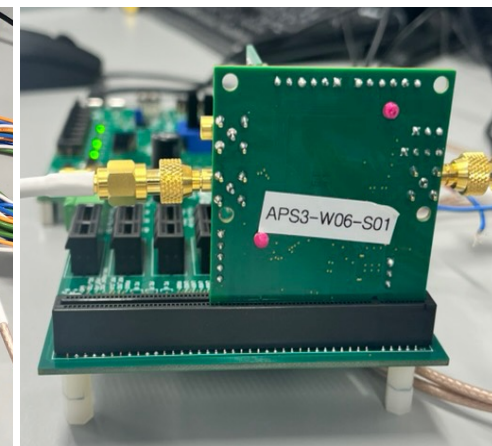
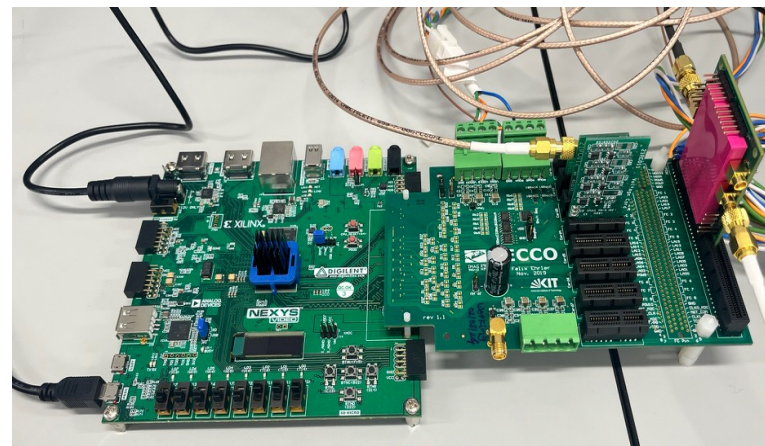
AstroPix R&D for BIC

- v3: Application in larger NASA payloads
 - Not final version for BIC
 - Single chip and quad chip: under test at ANL
- v4: Currently under test in NASA
- v5: Just submitted engineering run in mid June
 - For use in BIC
 - Similar to AstroPix v3

AstroPix v3 single chip

- 2x2 cm² -size with 35 × 35 pixel matrix, 500 μm × 500 μm pixel pitch, 720 μm thickness
- Very low power dissipation < 1.5 mW/cm²
- The good energy resolution (<10% @ 60 keV)
- Timestamp clock: 2 MHz (500 ns), 8-bit (0-255)
- ToT (Time over Threshold) clock: 200 MHz (5 ns), 12-bit (0-4,096)
- 10-byte data frame per hit
 - Header including Layer ID
 - 5 Bytes of hit information from sensor (chip ID, pixel location, time stamp, Time over threshold)
 - 4 Bytes of FPGA timestamp

v3	v4	v5
 <p>SPI test pins 200 MHz sample_clk Analog o/p Digital periphery & Wirebond Pads ADL KIT AstroPix v3 - Single Chip</p>	 <p>KIT AstroPix4 ADL KIT</p>	 <p>KIT Karlsruhe Institute of Technology</p>
<ul style="list-style-type: none"> Full size 2 x 2 cm² array 500 x 500 μm² pixel pitch 35 x 35 pixels 	<ul style="list-style-type: none"> 1 x 1 cm² array 500 x 500 μm² pixel pitch 16 x 13 pixels 	<ul style="list-style-type: none"> Full size 2 x 2 cm² array 500 x 500 μm² pixel pitch 35 x 33 pixels

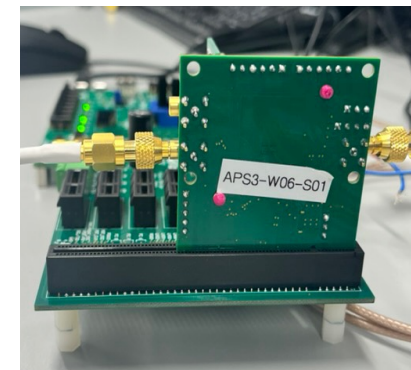
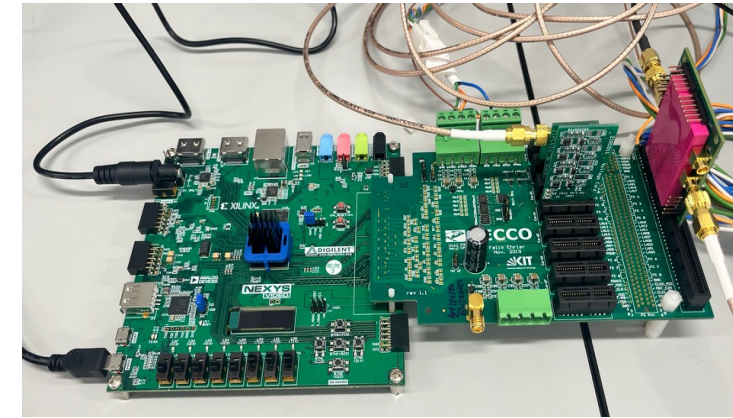
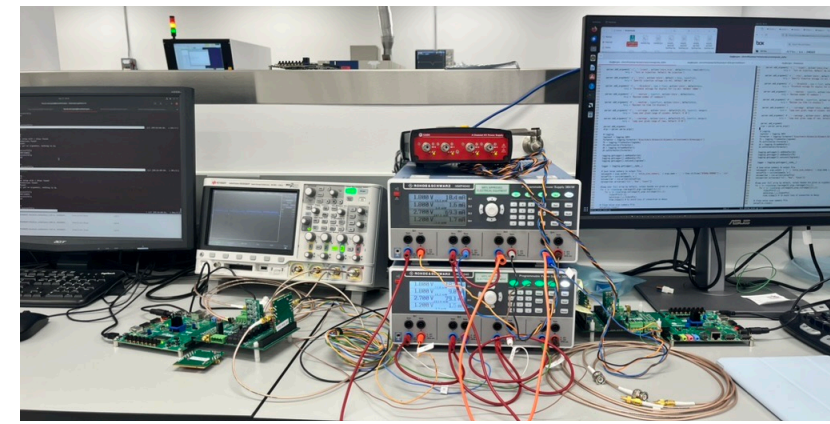


R&D Goal

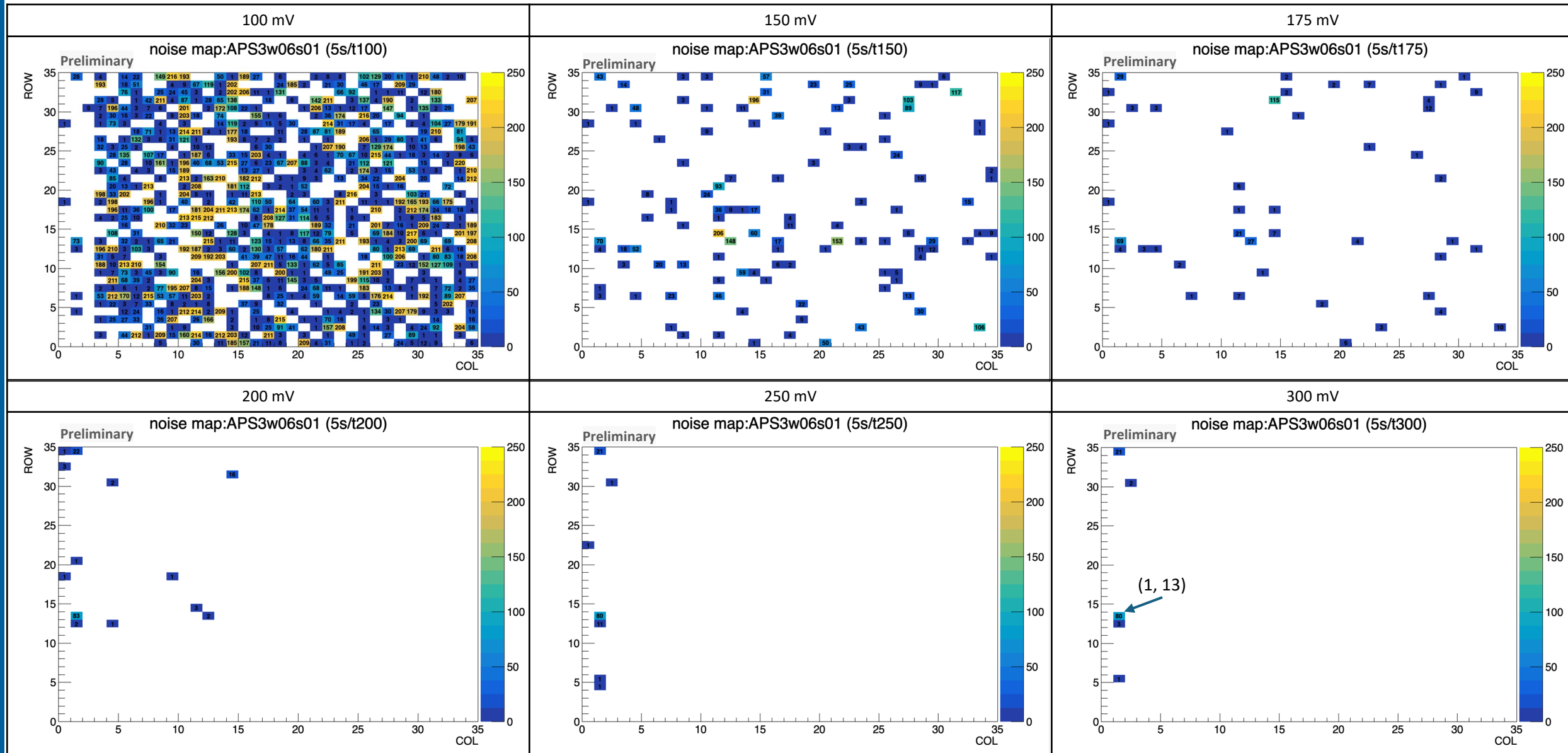
- M2: AstroPix chips prepared at the bench

1-1. Noise Scan with respect to threshold

- Used Astropix v3 single chips: APS3w06s01
- Bias voltage: - 150 V
- LV supply
 - Digital and analog: 1.8 V
 - Gecco board: 2.7 V
 - Amplifier: 1.2 V
- Time windows: 5 secs
- Threshold: 100, 150, 175, 200, 250, and 300 mV



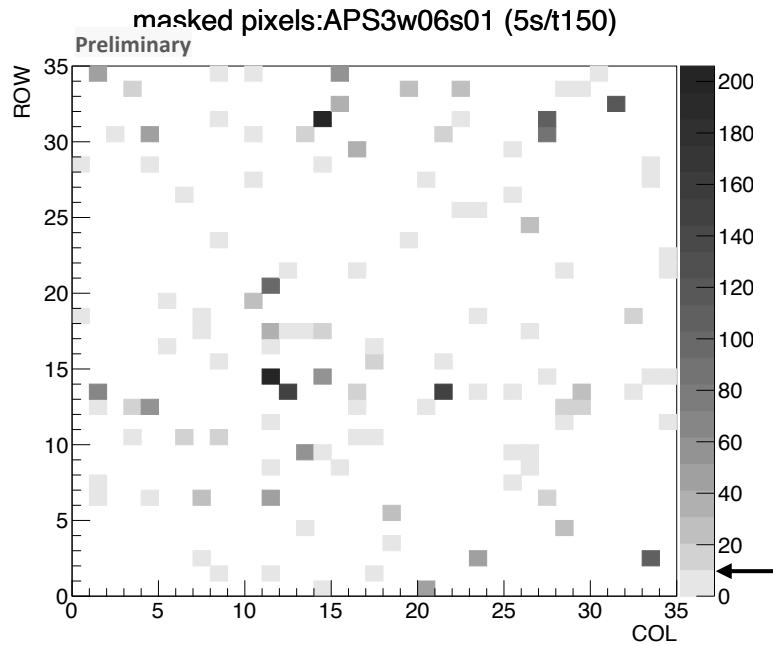
Noise Scan with respect to Threshold [100 mV – 300 mV]



Noise Scan: Masked Pixels

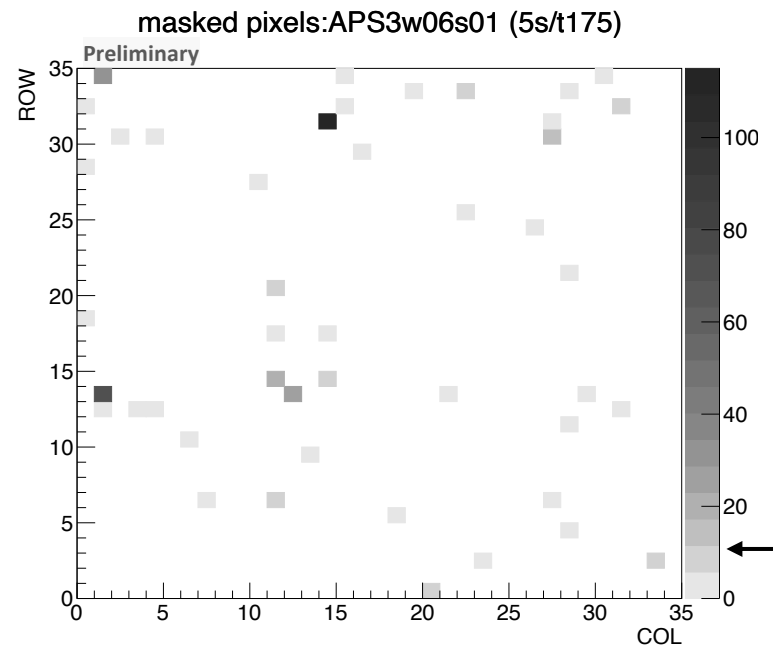
- Disable pixel : Noise > 10
- Used 200 mV threshold for source test and beam test

$V_{th} = 150$ mV



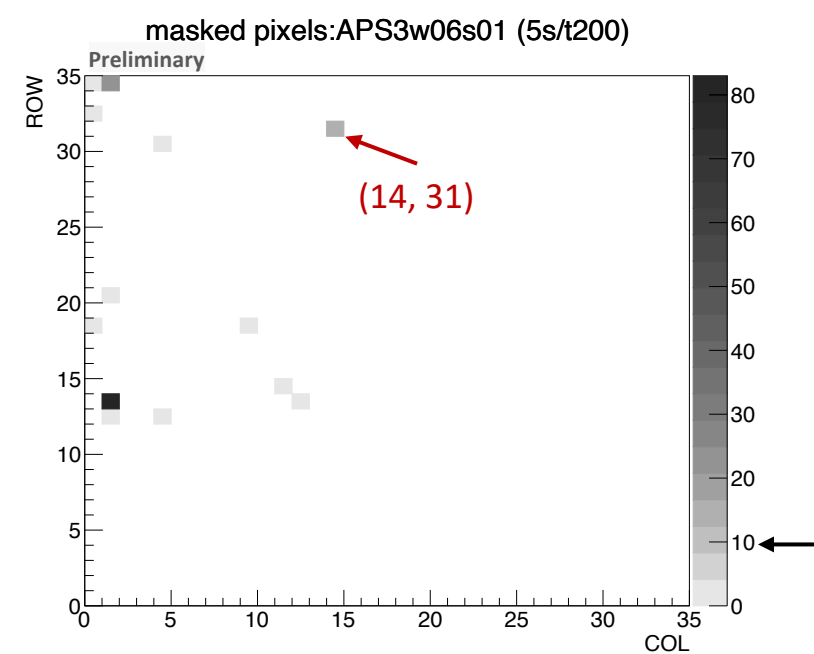
- Active pixels: ~ 96.5%
- 43/1225 pixels are disable.

$V_{th} = 175$ mV



- Active pixels: ~ 99.5%
- 6/1225 pixels are disable.

$V_{th} = 200$ mV



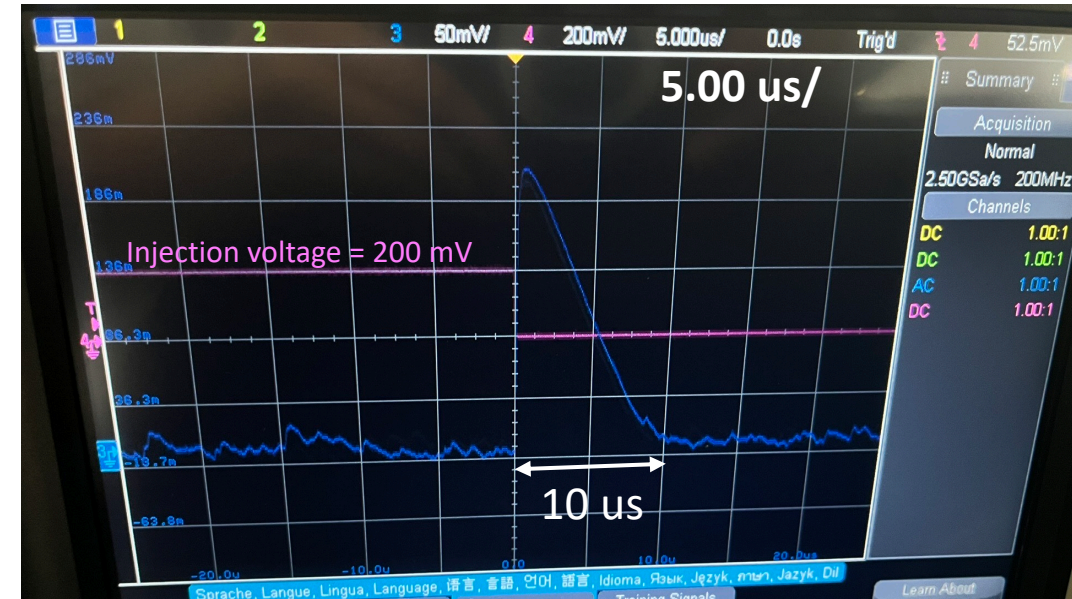
- Active pixels: ~ 99.8%
- 3/1225 pixels are disable.

→ Masked noisy pixel (14, 31)

R&D Goal

- M2: AstroPix chips prepared at the bench

1-2. Injection Scan

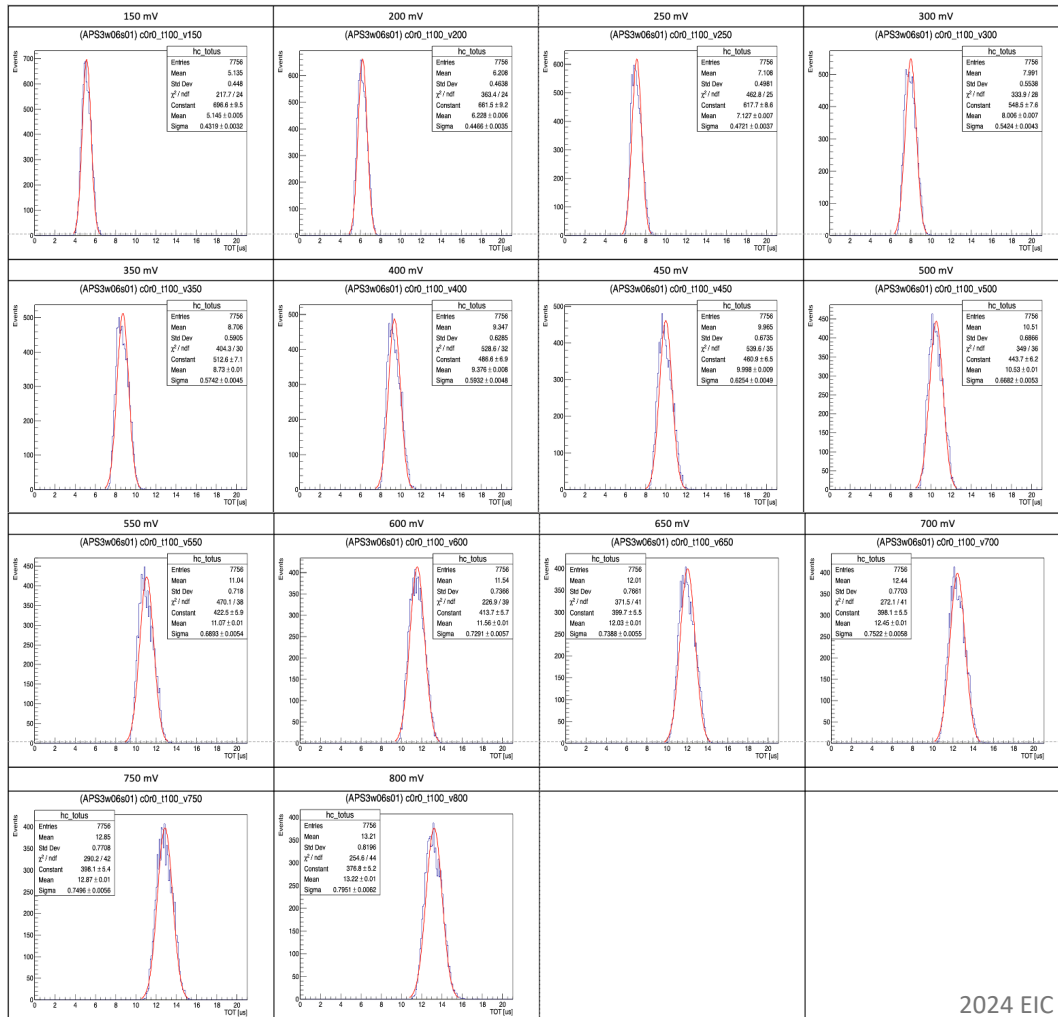


- Used Astropix v3 single chips: APS3w06s01
- Bias voltage: - 150 V
- LV supply
 - Digital and analog: 1.8 V
 - Gecco board: 2.7 V
 - Amplifier: 1.2 V
- Threshold: 100 mV
- With respect to injection voltages at (COL #, ROW #)
 - Column (0, 1, 2, 3, 4, 5, 6) & Row 0
 - 150 mV – 800 mV in steps of 50 mV
 - Time windows: ~ 90 secs for 1k events

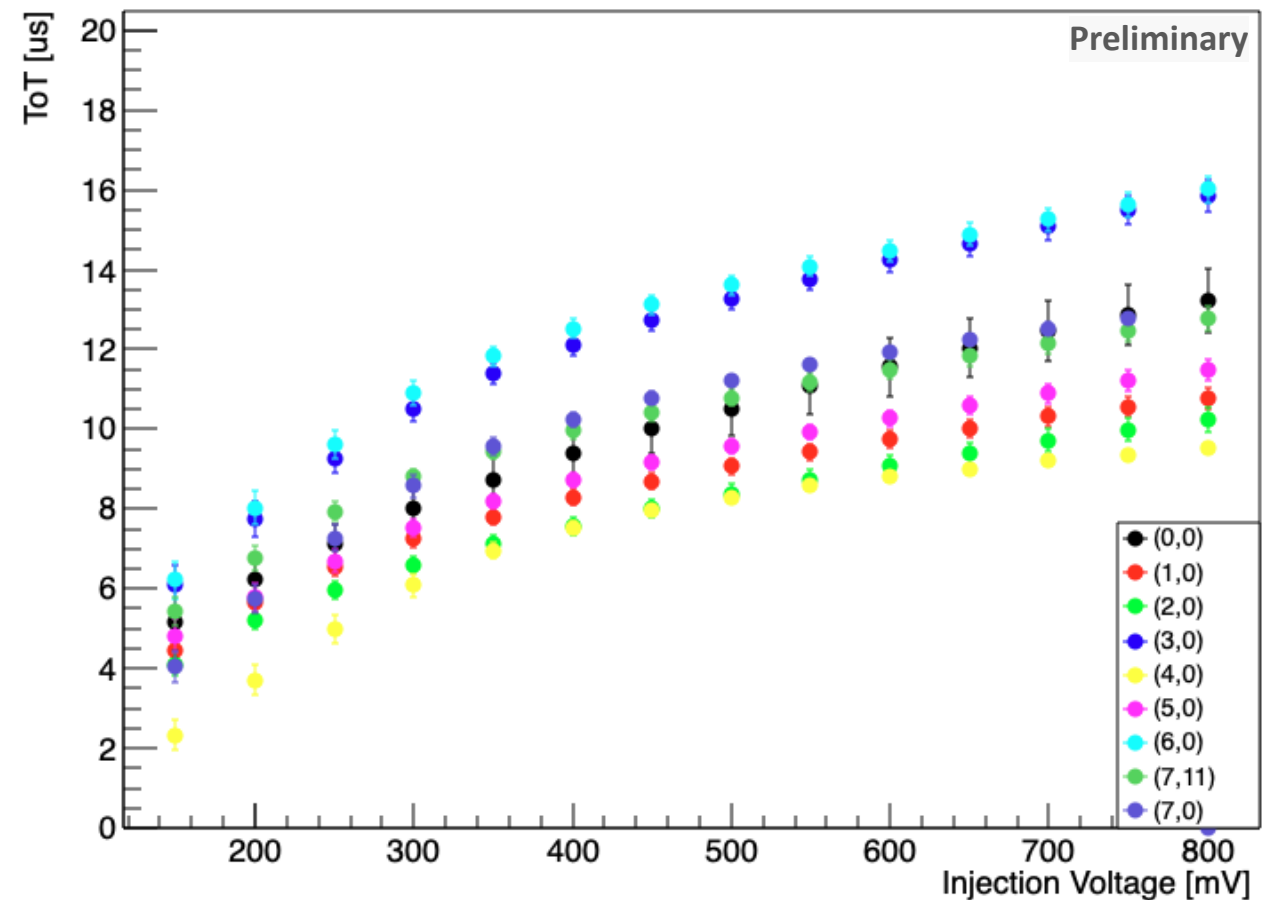
Injection Scan: Injection Voltage vs ToT [us]

- Column (0, 1, 2, 3, 4, 5, 6, 7) & Row 0 and [7, 11]
- ToT [us] as a function of injection voltages from 150 mV to 800 mV in steps of 50 mV.
- Fitting using gaussian function then mean and sigma values from fitting results are plotted.

- ToT spectra w.r.t various injection voltages at pixel (0, 0)



- ToT mean values as a function of various injection voltages

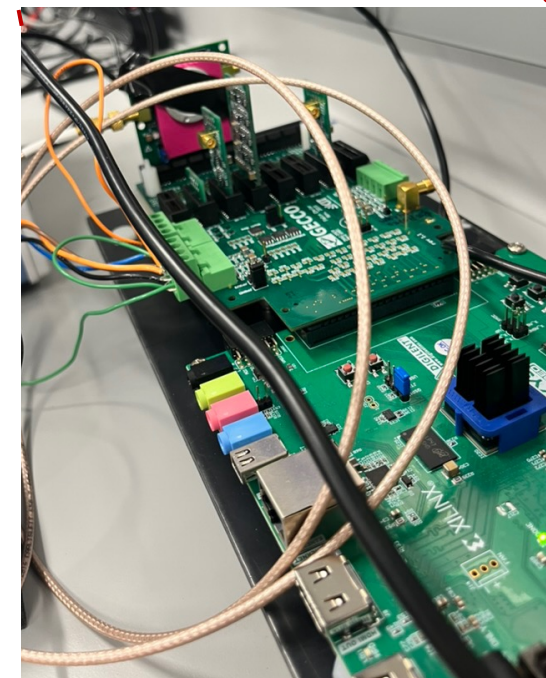
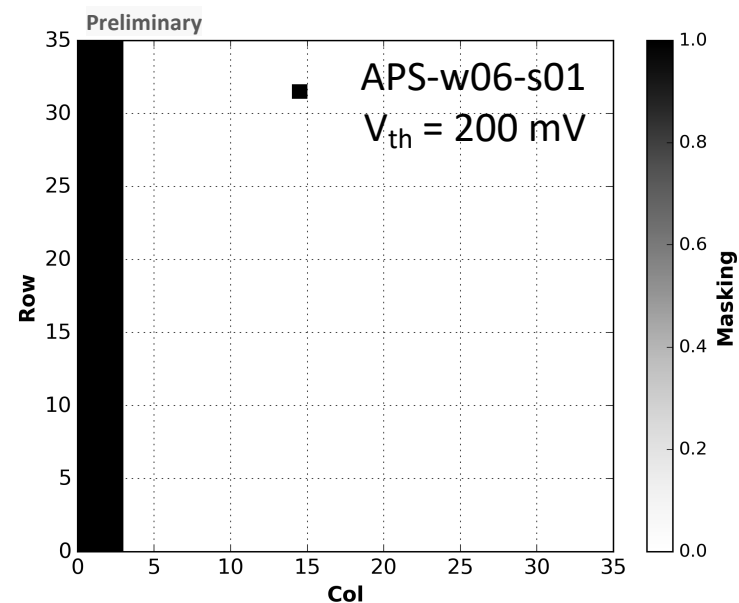
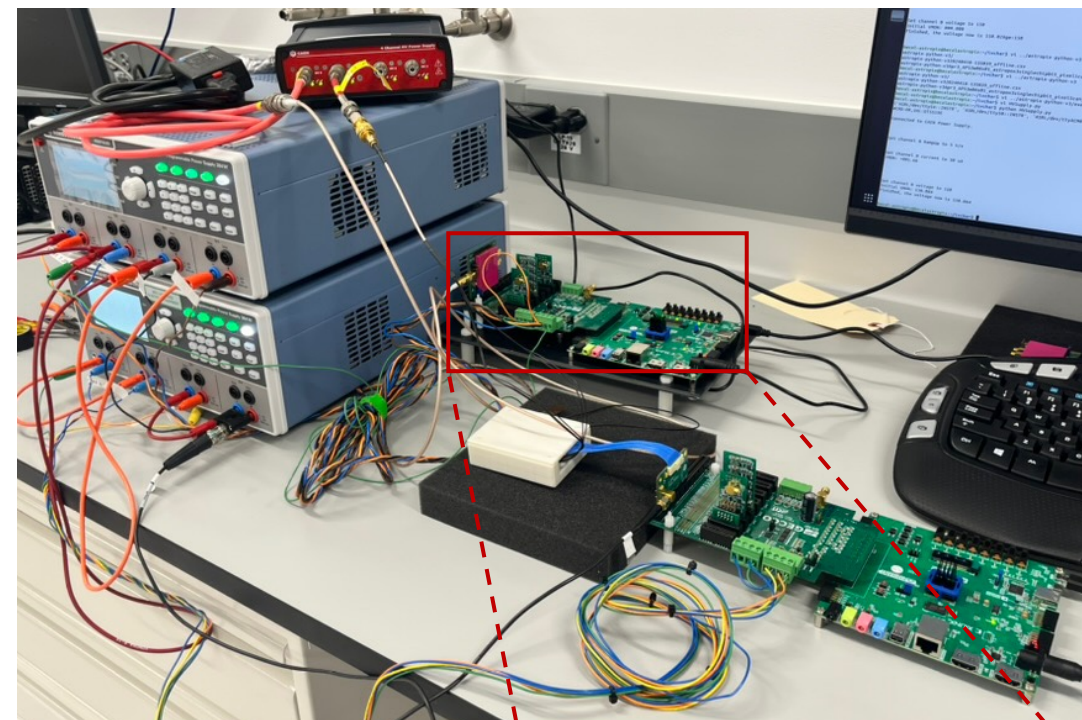


R&D Goal

- M2: AstroPix chips prepared at the bench

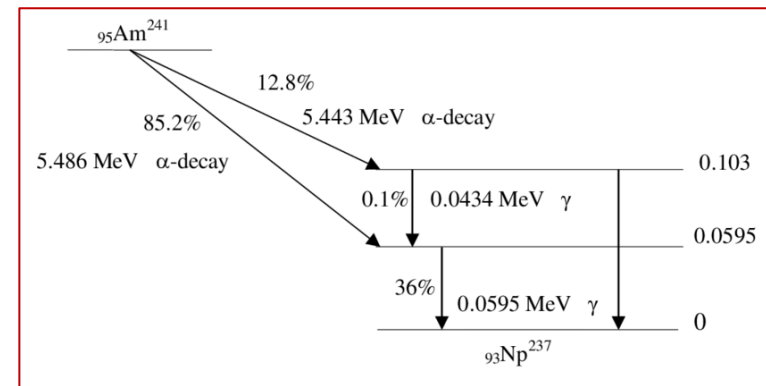
1-3. Source Test using Am-241

- Used Astropix v3 single chips: APS3w06s01
- Bias voltage: - 150 V
- LV supply
 - Digital and analog: 1.8 V
 - Gecco board: 2.7 V
 - Amplifier: 1.2 V
- Used threshold = 200 mV
- Masked first three column pixels



Am-241 Result: 32 × 35 pixels

- Data-taking for 12 hours with 32 × 35 pixels (1,120 pixels)
- ToT distribution for all pixels after applying to scale factor (from proton result)
- Full peak, fitted with gaussian function

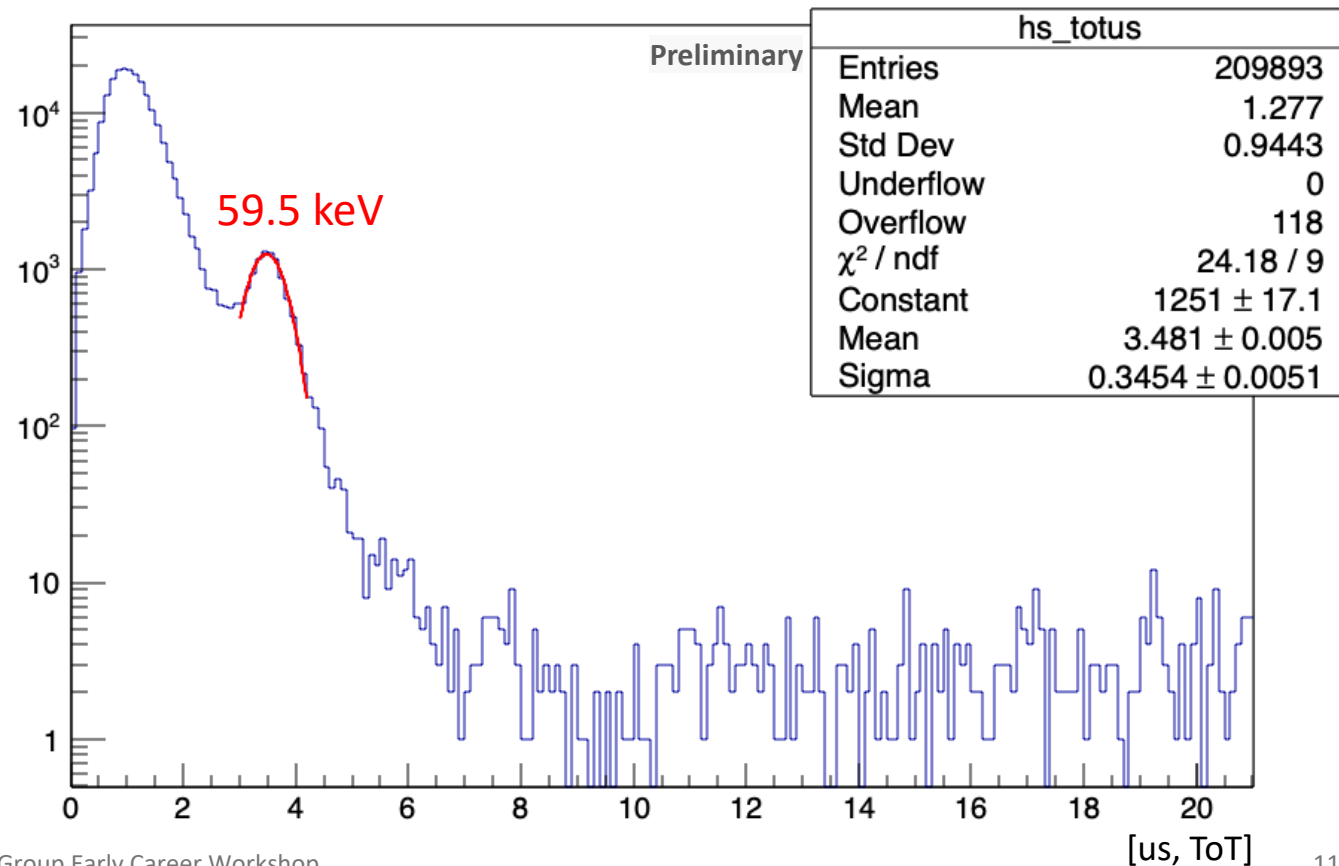
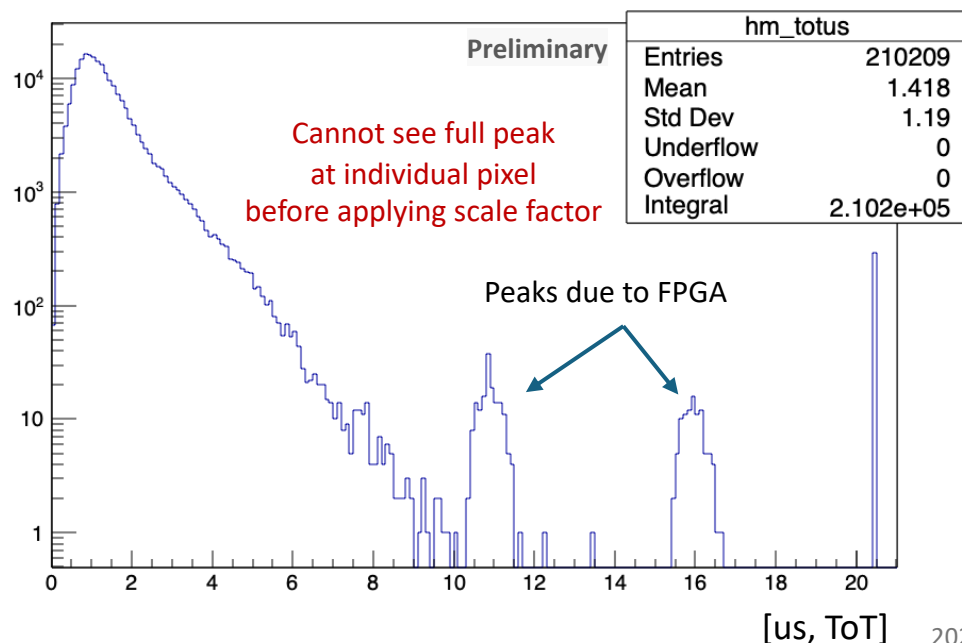


• A **calibration constant** is calculated:
 $59.5 \text{ [keV]} / 3.481 \text{ [us, TOT]} = 17.09 \text{ [keV]/[us, ToT]}$

• $E_{\text{res}} \text{ (FWHM)} = 2.35 \times \frac{\text{Sigma}}{\text{Mean}} = 23.3 \% \text{ @ } 59.5 \text{ keV}$

ToT distribution of all pixels after applying scale factor

ToT distribution of all pixels **before** applying scale factor



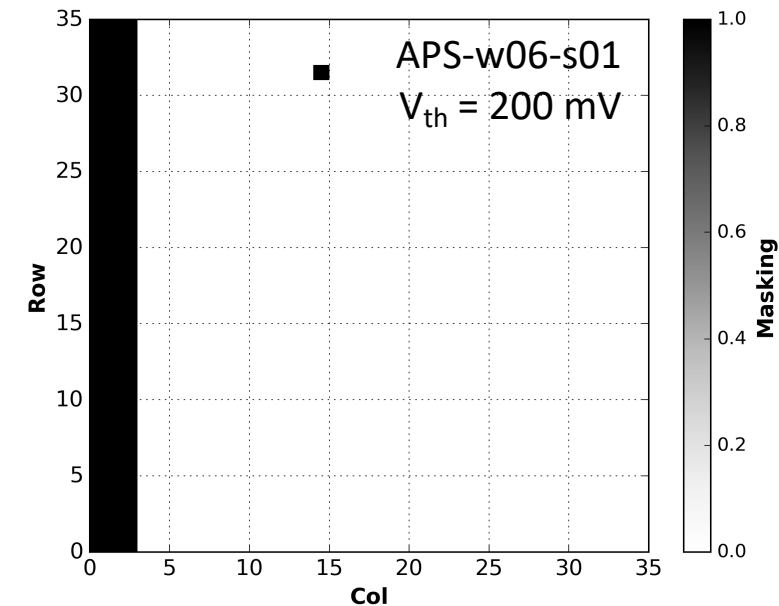
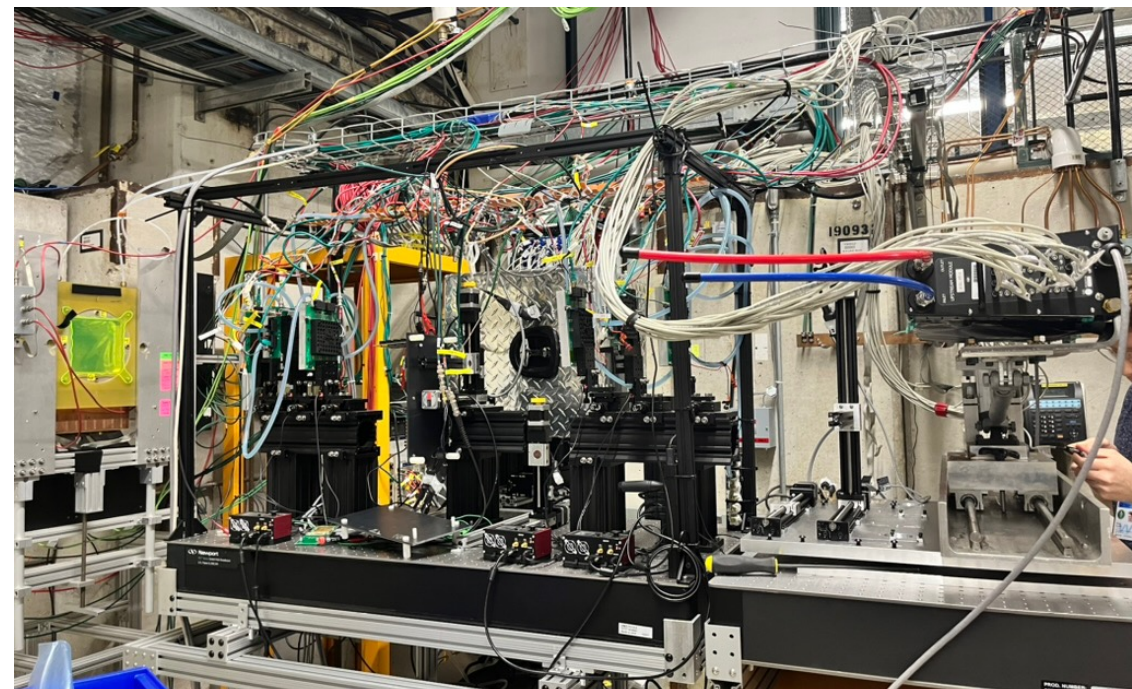
Beam Test R&D Goal

- M6: Integrated system commissioned with proton beam
 - AstroPix v3 single chip with Baby BCAL
 - Double layer with two AstroPix v3 single chips

2. Beam Test at FNAL

On June 13-18

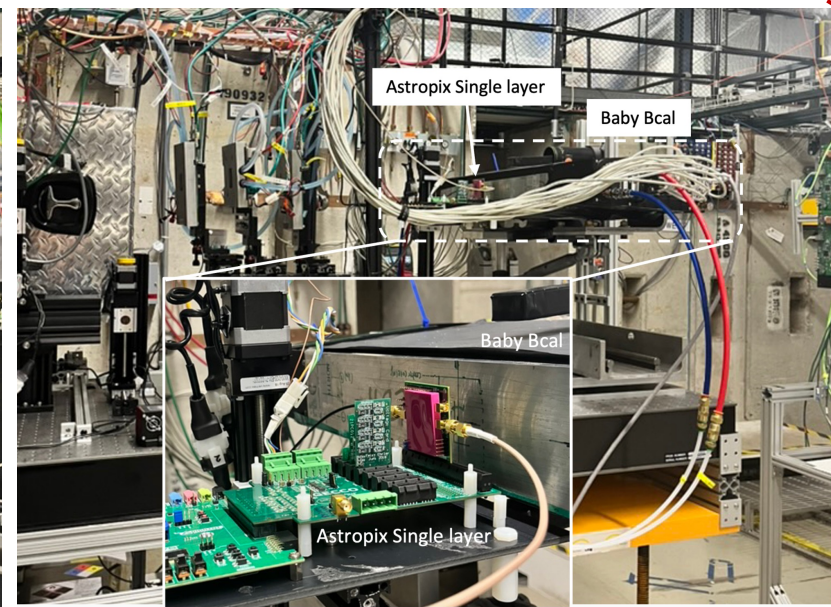
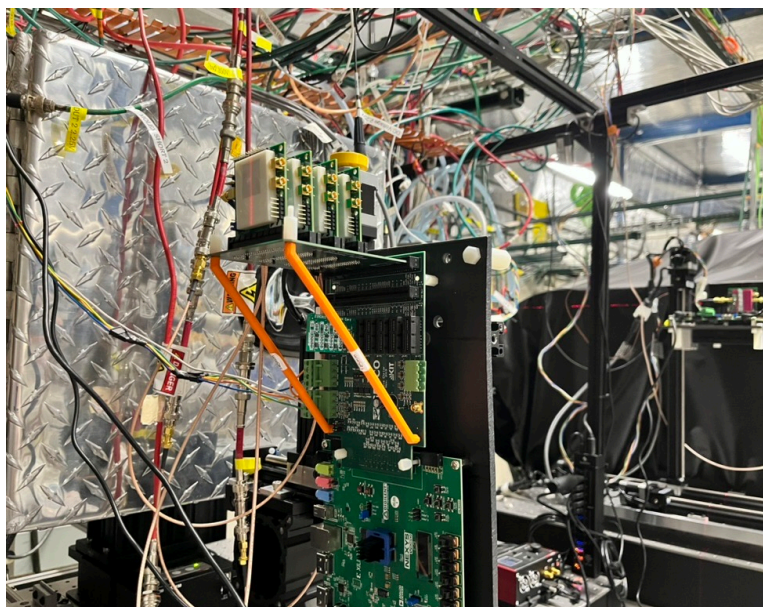
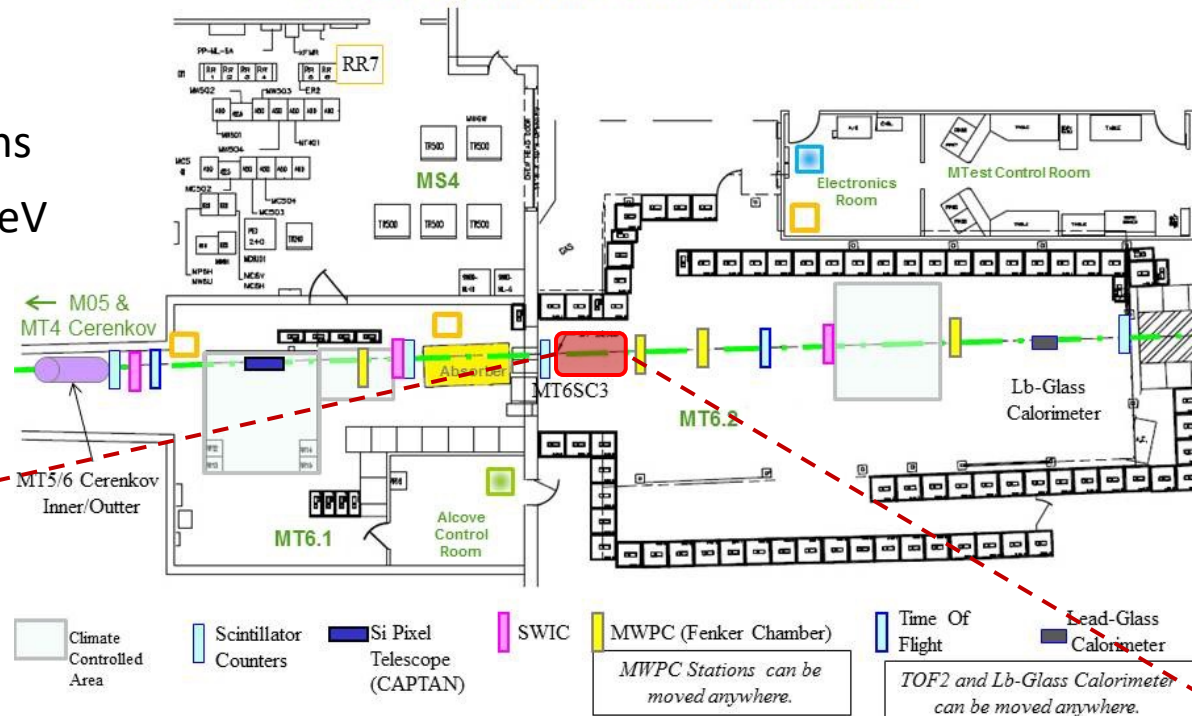
1. Single layer in front of Baby Bcal
 1. MIP response
 2. Astropix+Baby Bcal: coincident event
2. Double layer using two v3 single chips
 1. Coincident hits



Beam Test Setup @FNAL

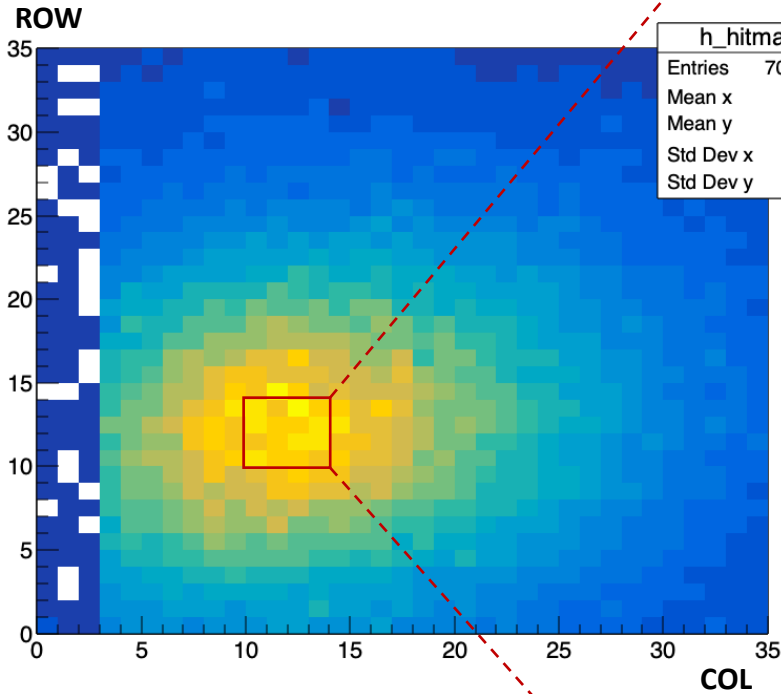
- 'Mtest' Beam line provides particles from 120 GeV protons to secondaries (electrons, muons, and pions) of ~ 200 MeV
 - Single 4.2 s long spill per every 60 seconds
- FNAL instrumentation system
 - Segmented Wire Ionization Chambers (SWICs);
2 mm wire spacing for tracking
 - Scintillator counter for triggering
 - Two Cherenkov detectors for PID

MTest Beam line Instrumentation

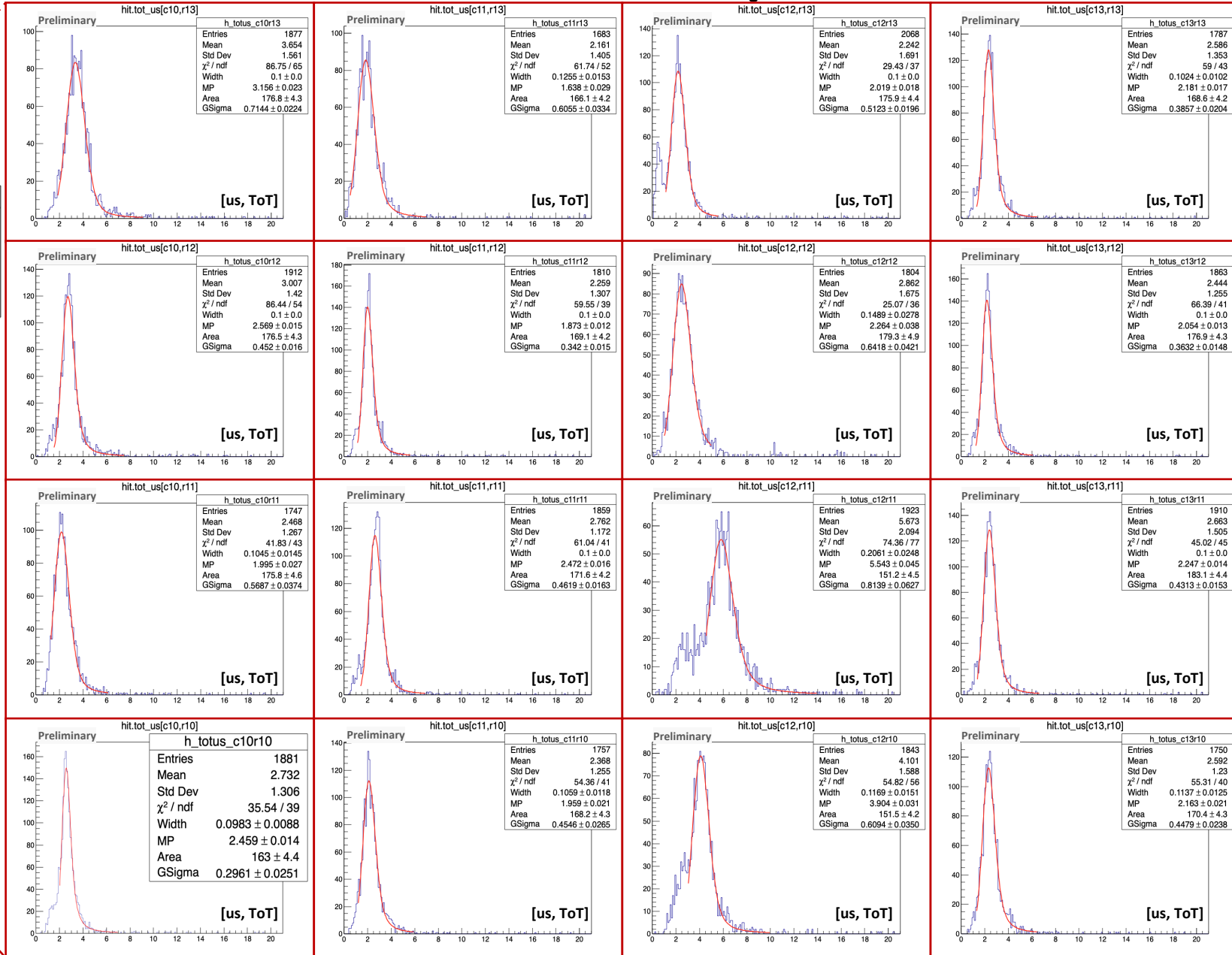


120 GeV Proton: Astropix, ToT Distributions of 4x4 pixels

- 4x4
- Col 10-13
- Row 10-13



- Fit with landau convoluted with gaussian function on ToT distribution

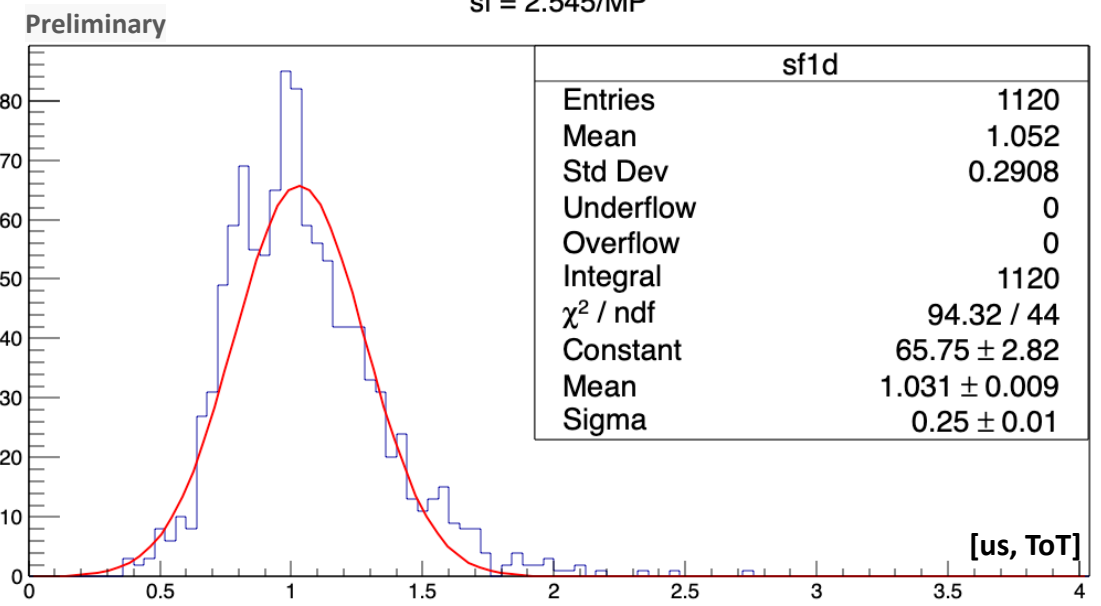
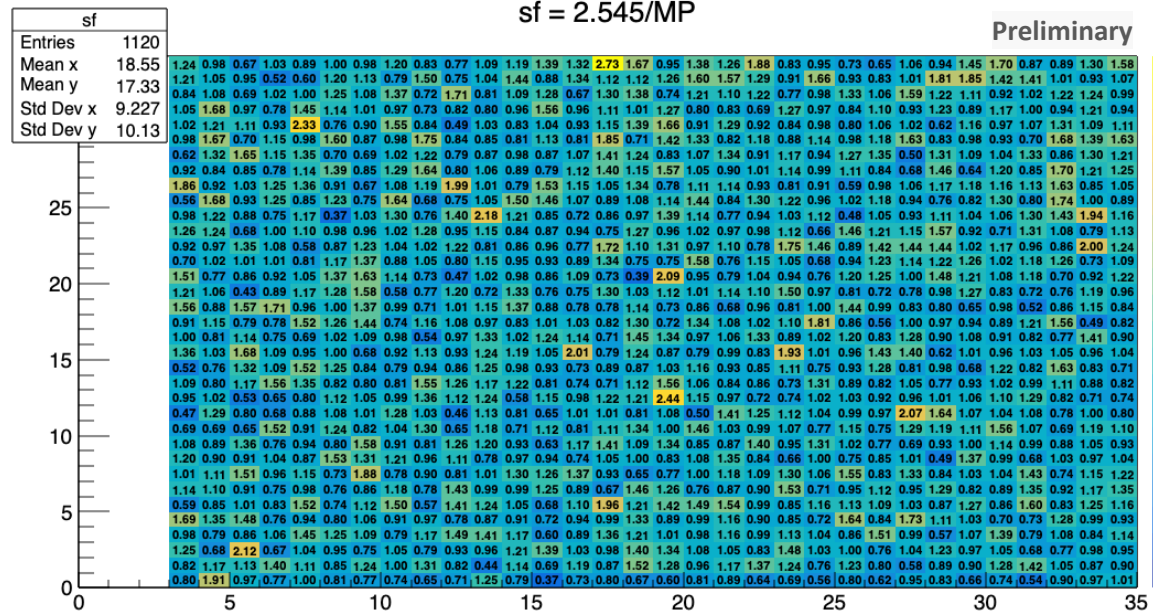
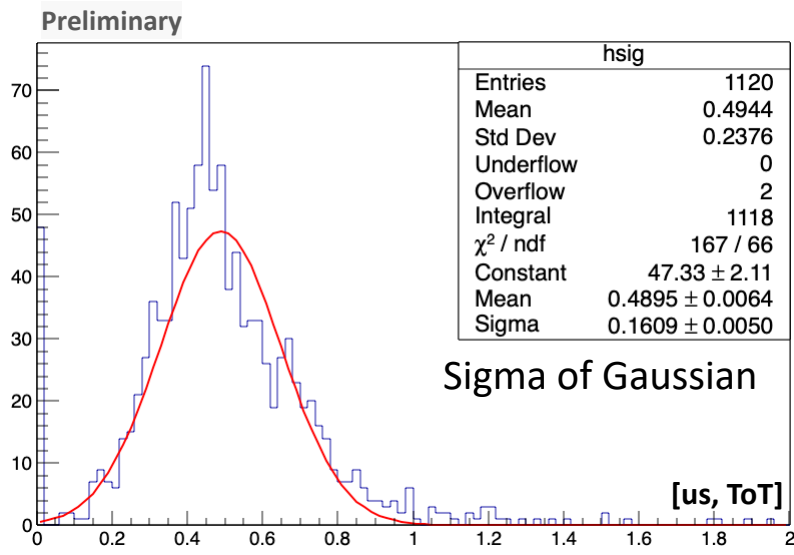
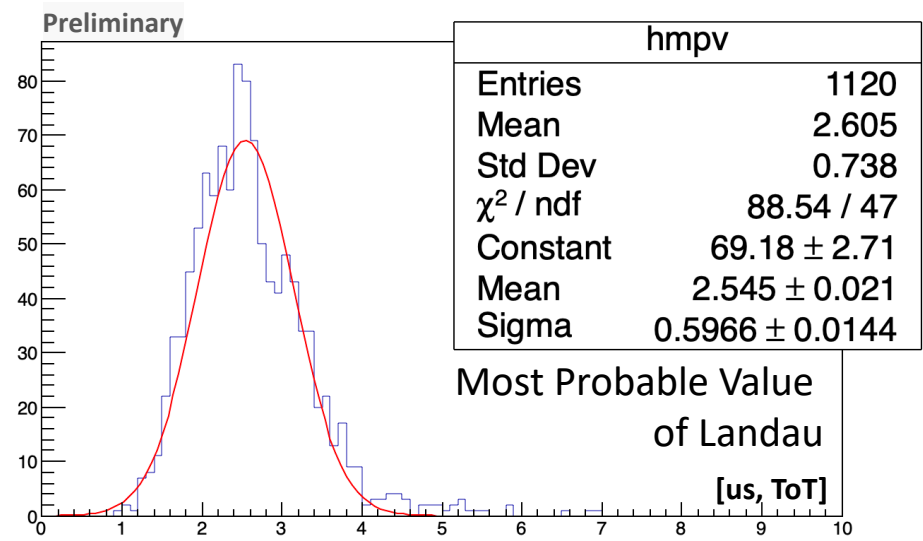


120 GeV Proton: Astropix, 32x35 pixels

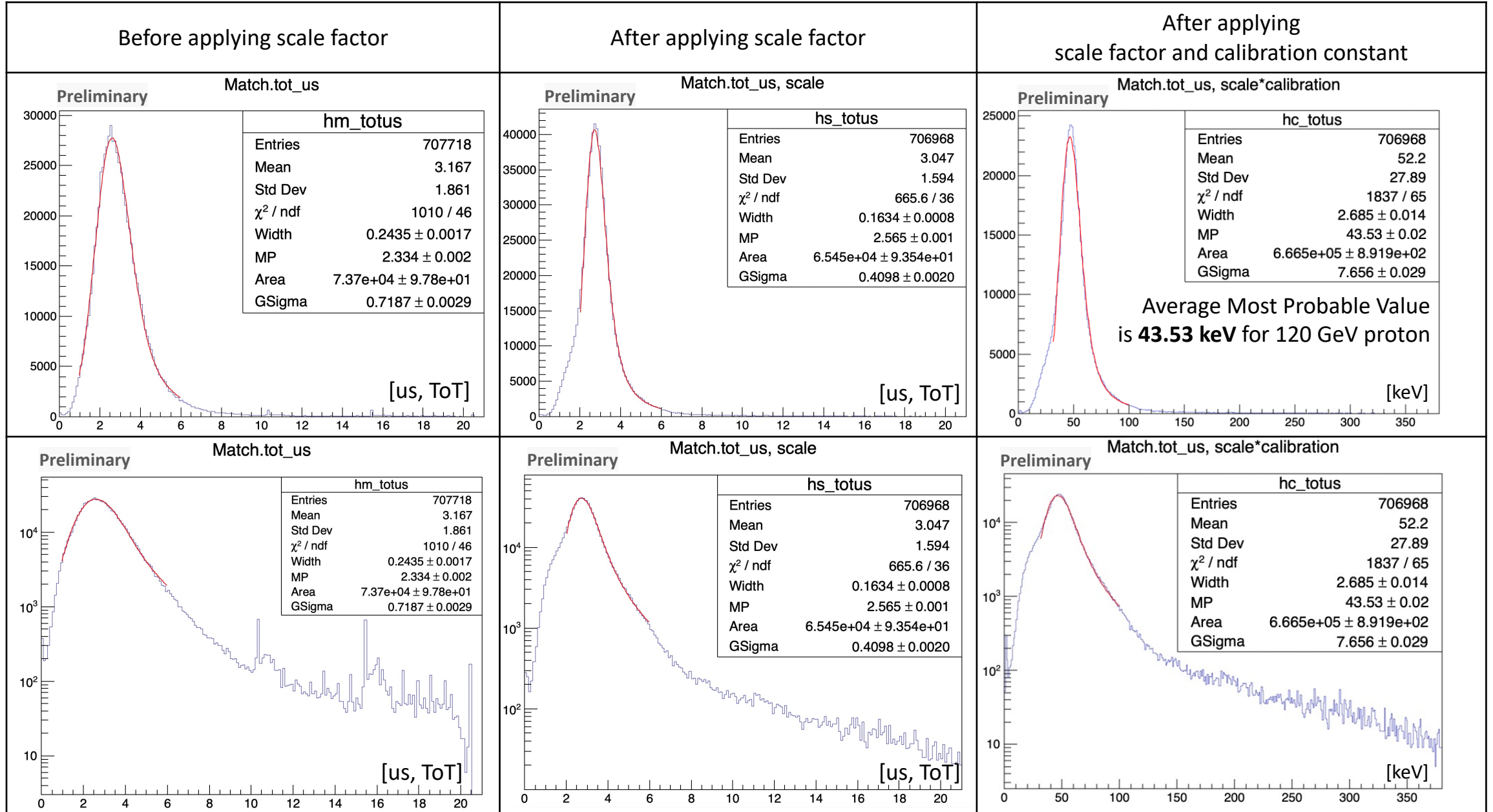
- 32x35 (1,120 pixels)
 - Col 3-24
 - Row 0-35

- Fit with landau convoluted with gaussian function on ToT distribution for all pixels

- Scale factor (sf), defined as mean of Most probable value for all pixels / Most probable value of each pixel

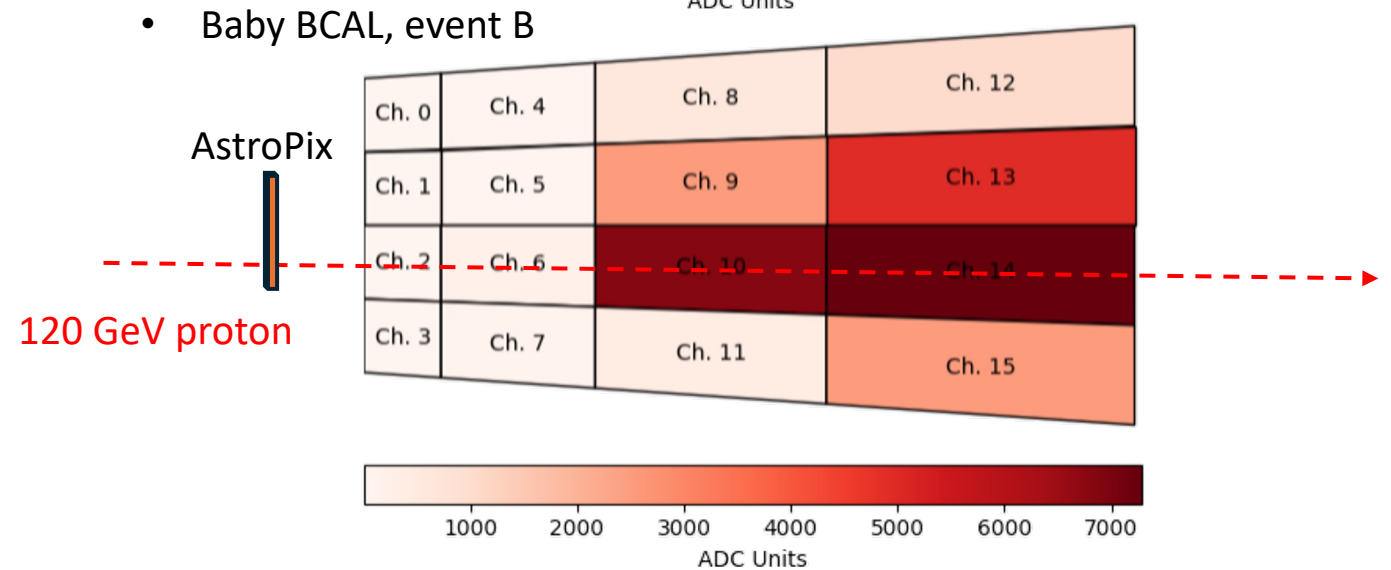
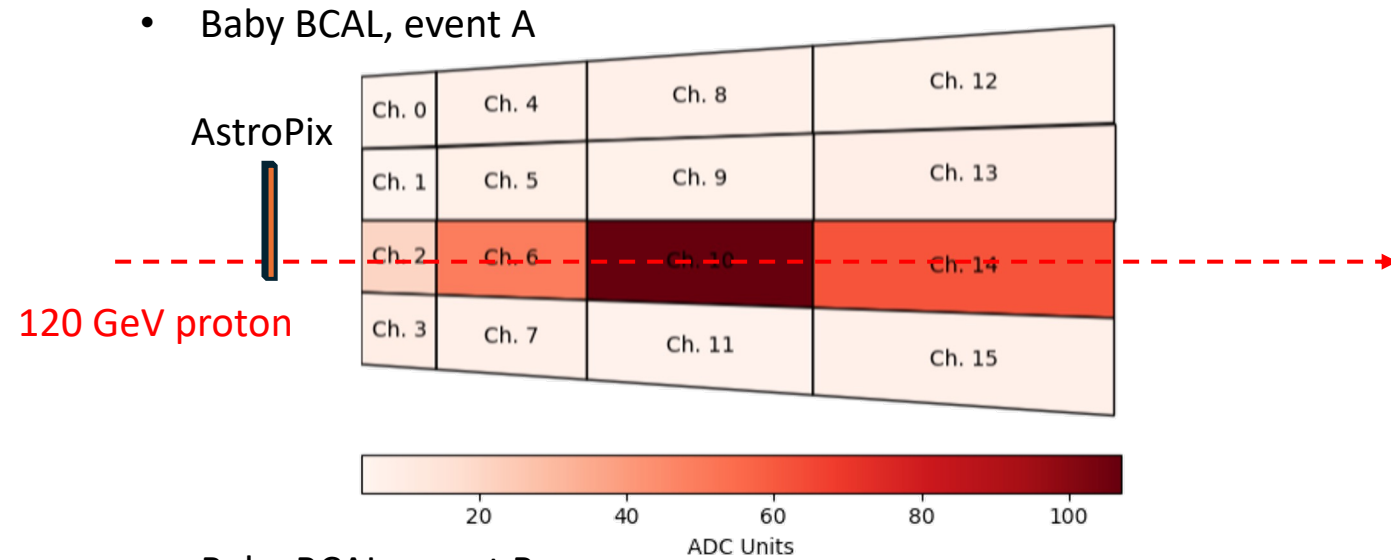
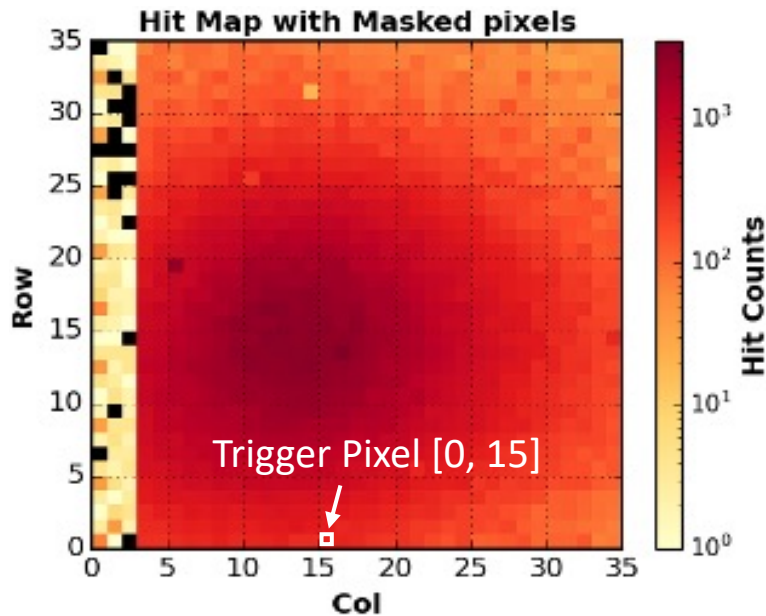
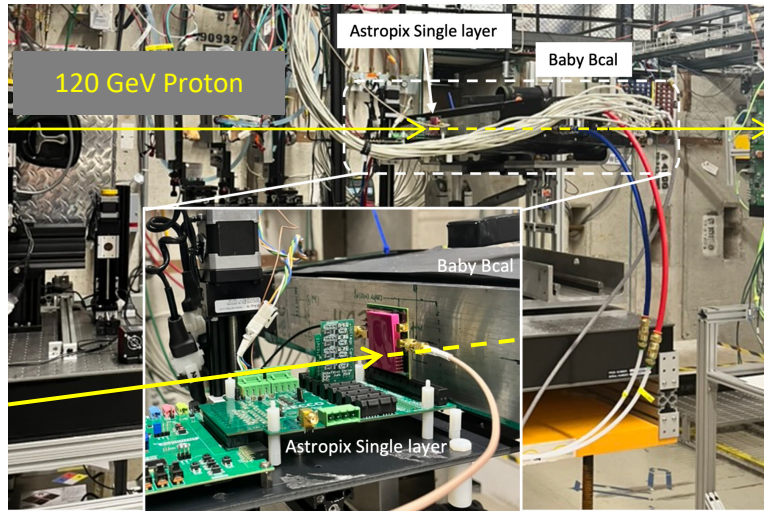


120 GeV Proton: Astropix, Energy Distribution of 32×35 pixels



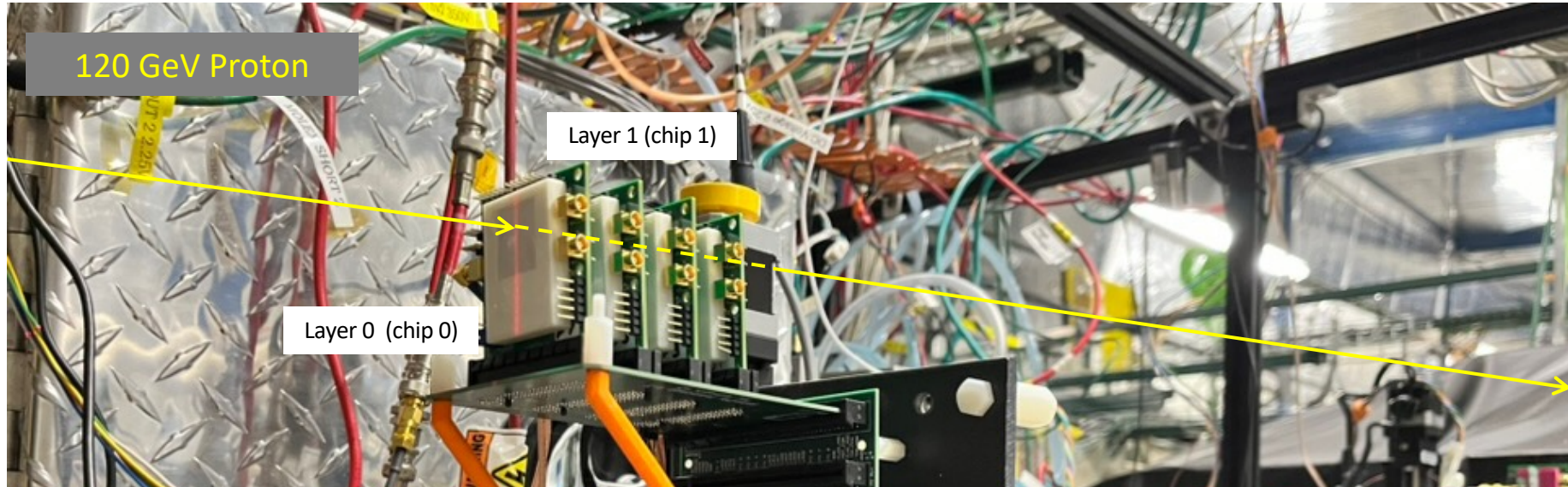
120 GeV Proton: Coincident Event Display of AstroPix and Baby BCAL

- Baby BCAL triggered by analog signal from [0, 15] pixel of Astropix

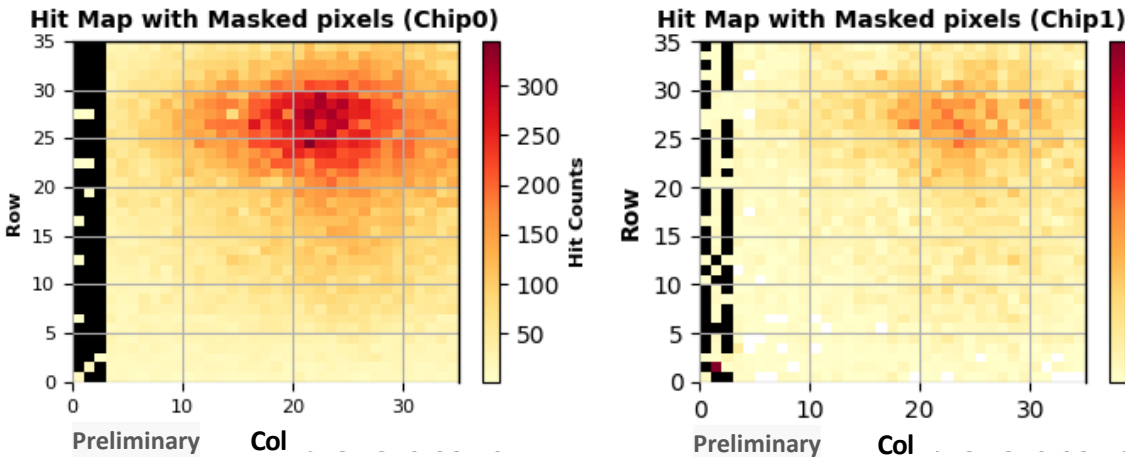


*Example plot of 120 GeV proton event display from an integrated system of Baby BCAL and a single-layer AstroPix v3 chip by Henry Klest

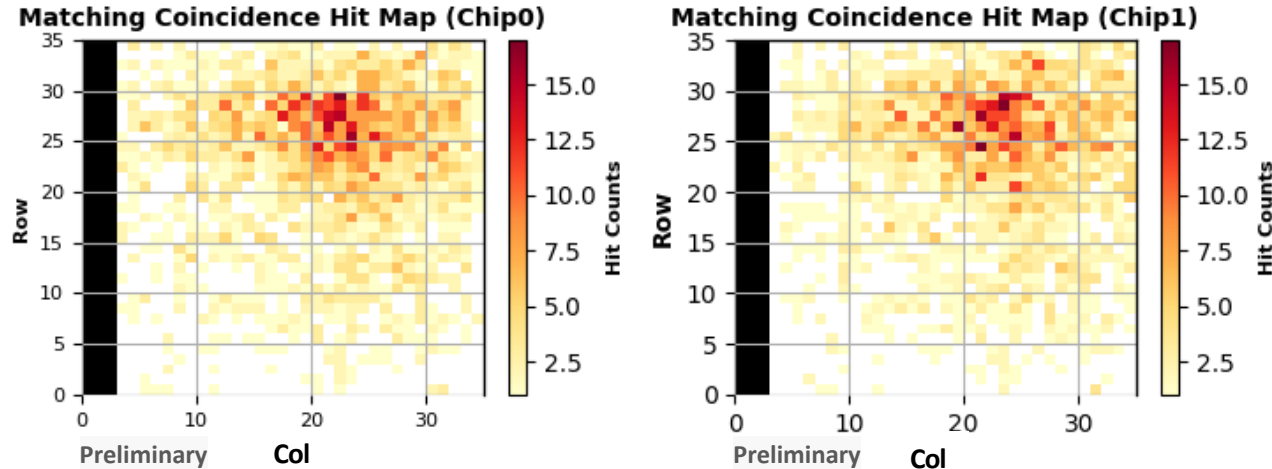
120 GeV Proton: Double Layer, Hit Map of Coincident Hit



- Matching hit using col&row information on same chip



- Matching coincident hit between chip0 and chip1



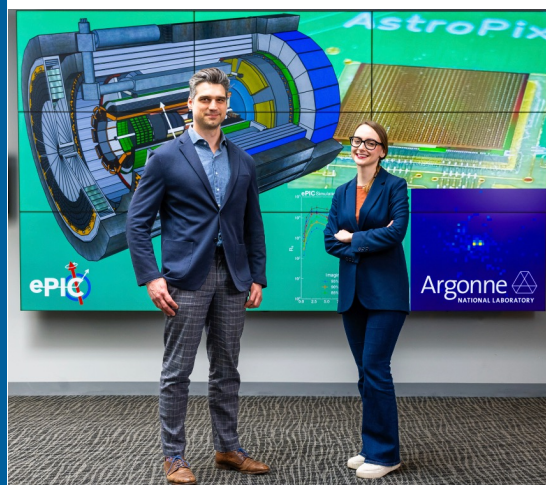
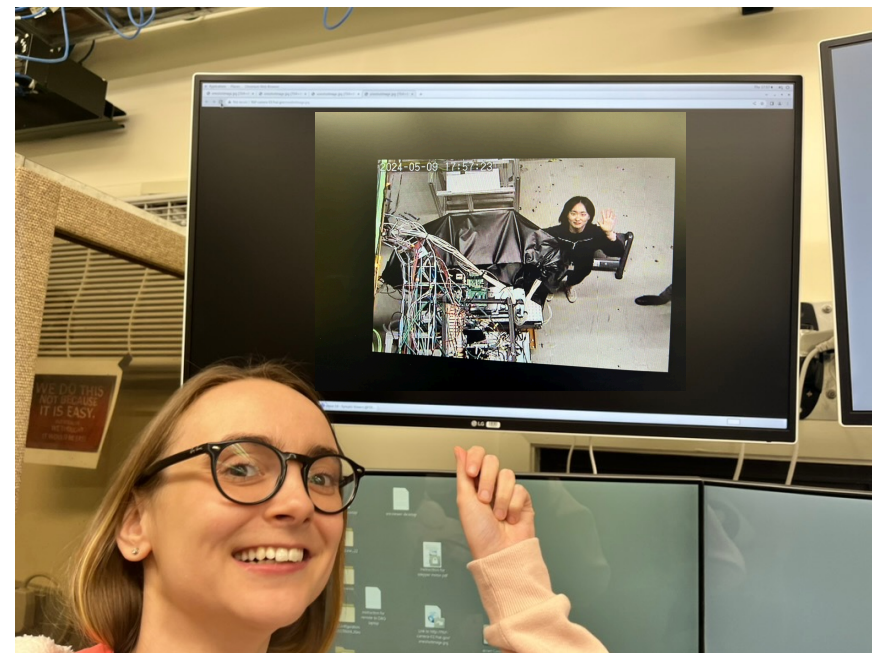
- Matching hits with time stamp < 2 and ToT error $< 10\%$

- Matching hits with location difference < 3 and time stamp difference < 5

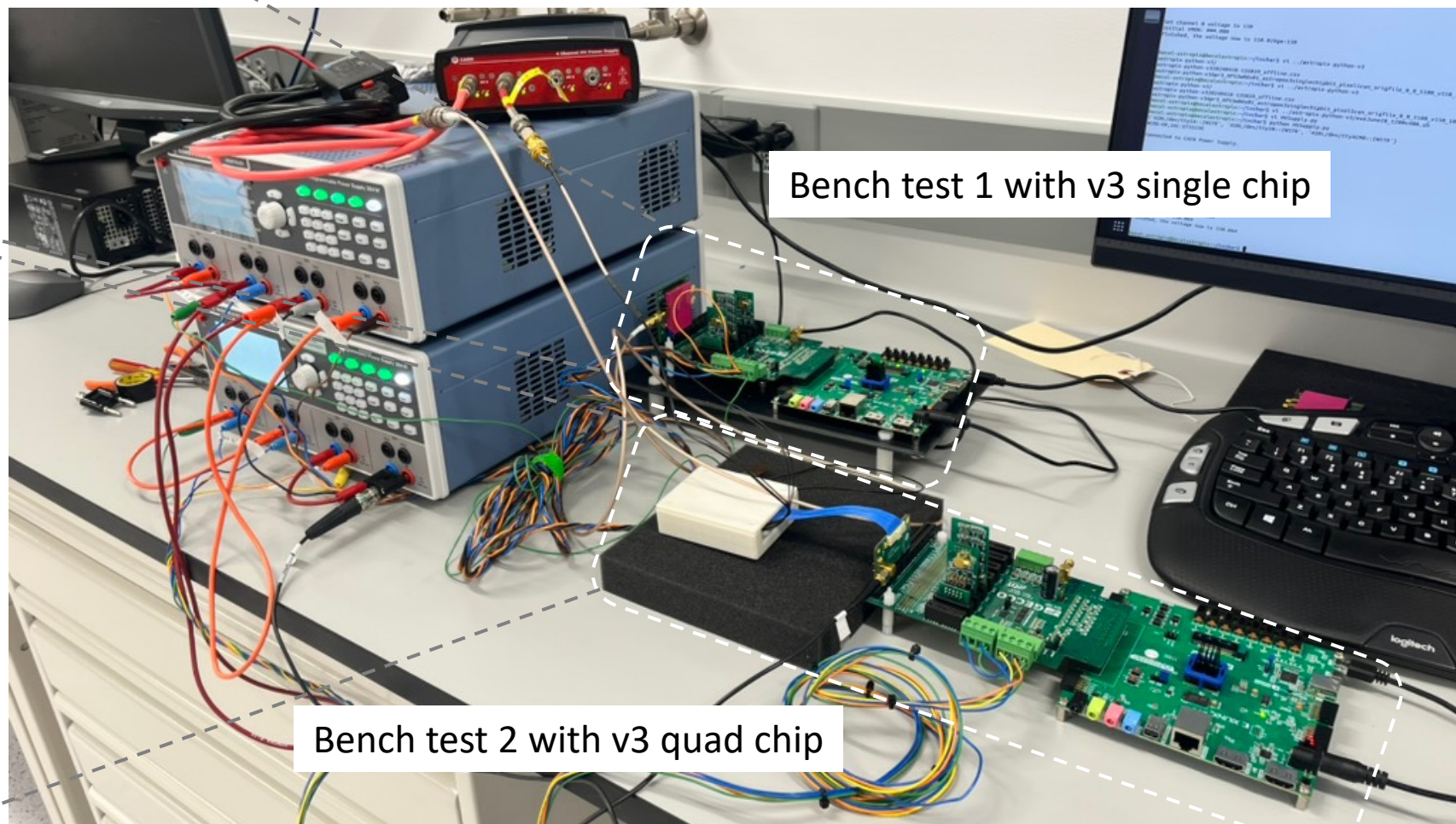
Summary

- The performance of Astropix v3 single chip is measured at ANL:
 - Noise scan: measured noisy pixels
 - Injection scan: measured a relation between injection voltage and individual pixel output
 - Source test: measured a calibration constant of 17.09 [keV]/[us, ToT]
- Using 120 GeV proton beam at FNAL
 - Astropix v3 single chip,
 - Measured beam position by hit map and average MIP deposited energy of 43.53 keV
→ Extract MIP response
 - Astropix and Baby BCAL,
 - Measured coincident hits between Astropix and Baby BCAL
→ Proof-of-concept tested in the beam
 - Double layer of two Astropix v3 single chips
 - Measured coincident hits between double layer
- Collaboration paper is in progress during this summer.

Thank you

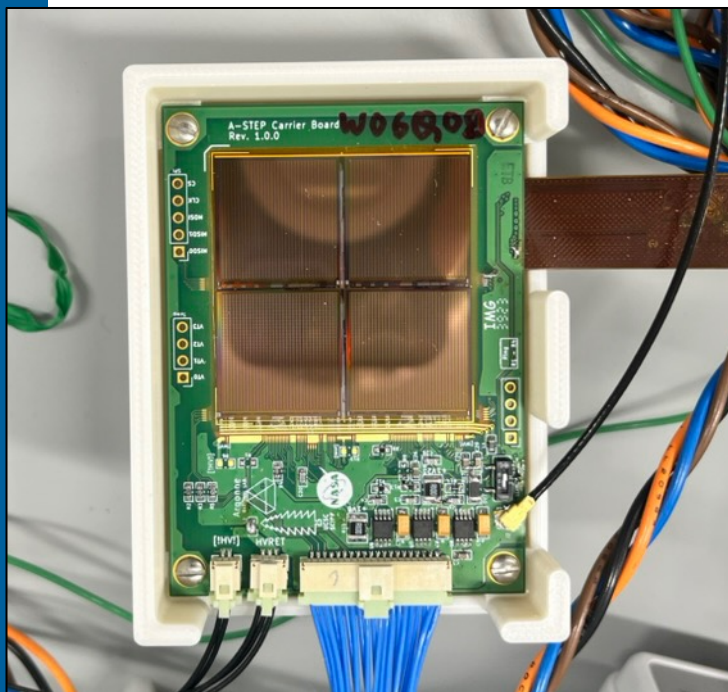


Bench setup at ANL with AstroPix



Bench test 1 with v3 single chip

Bench test 2 with v3 quad chip



EIC Calorimetry Requirements

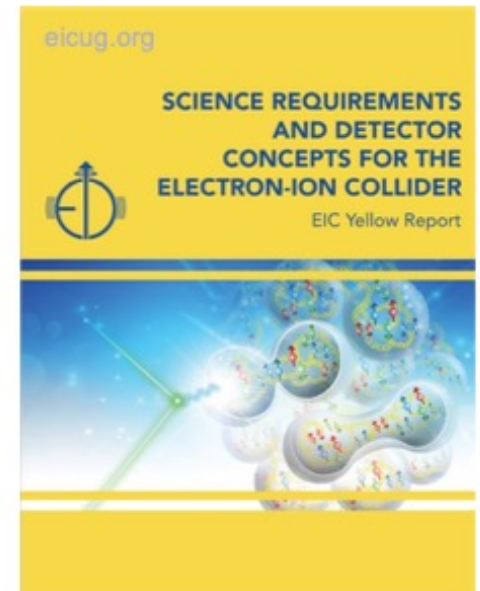
Barrel ECAL in EIC Yellow Report

EIC Community outlined physics, detector requirements, and evolving detector concepts in the [EIC Yellow Report](#).

EIC Yellow Report requirements for Barrel EM Calorimeter

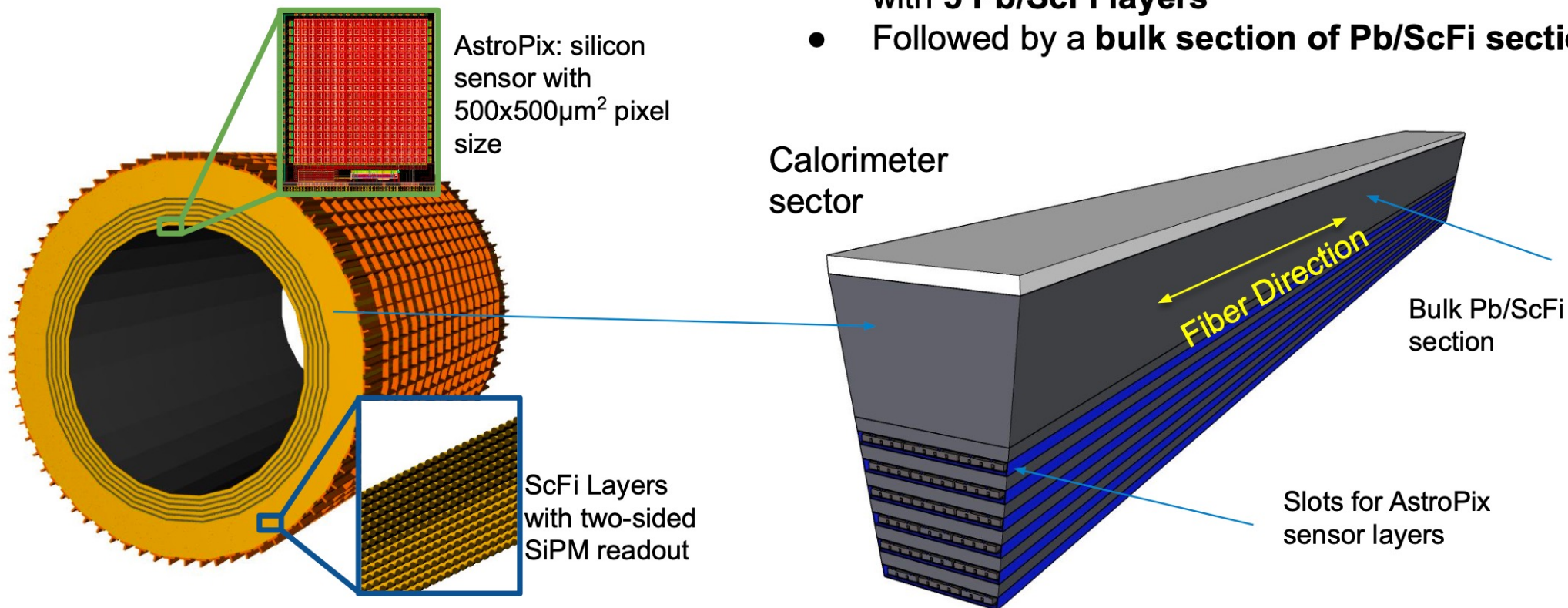
- Detection of electrons/photons to measure **energy and position**
- Require **moderate energy resolution** $(7 - 10)\%/\sqrt{E} \oplus (1 - 3)\%$
- Require **electron-pion separation up to 10^4** at low momenta in combination with other detectors
- Discriminate between **π^0 decays and single γ up to ~ 10 GeV**
- **Low energy photon** reconstruction ~ 100 MeV

Challenges: e/ π PID, γ/π^0 discrimination, available space



Geometry

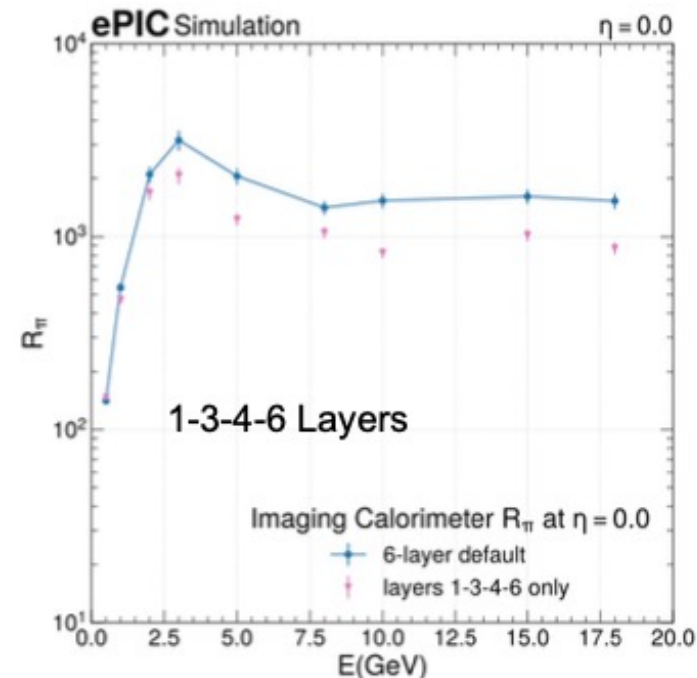
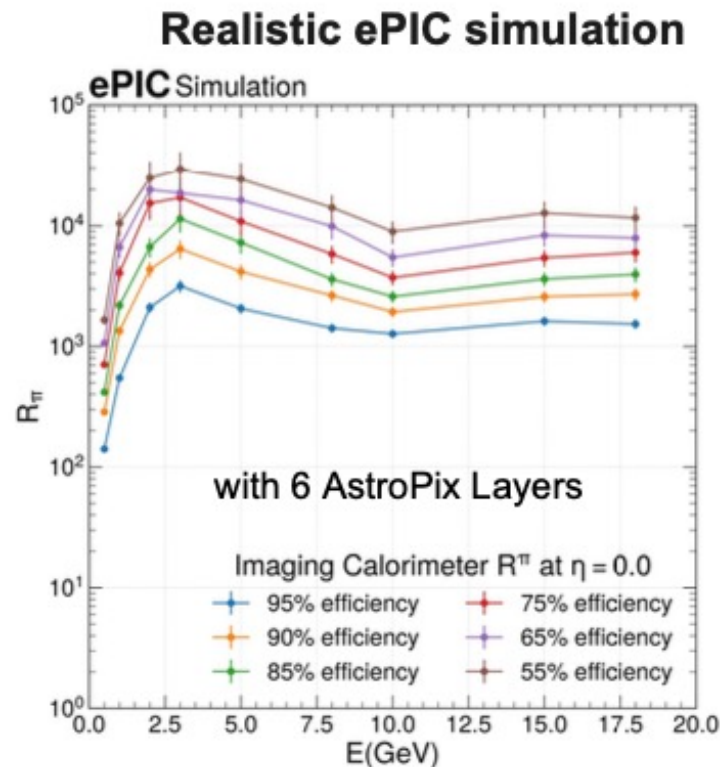
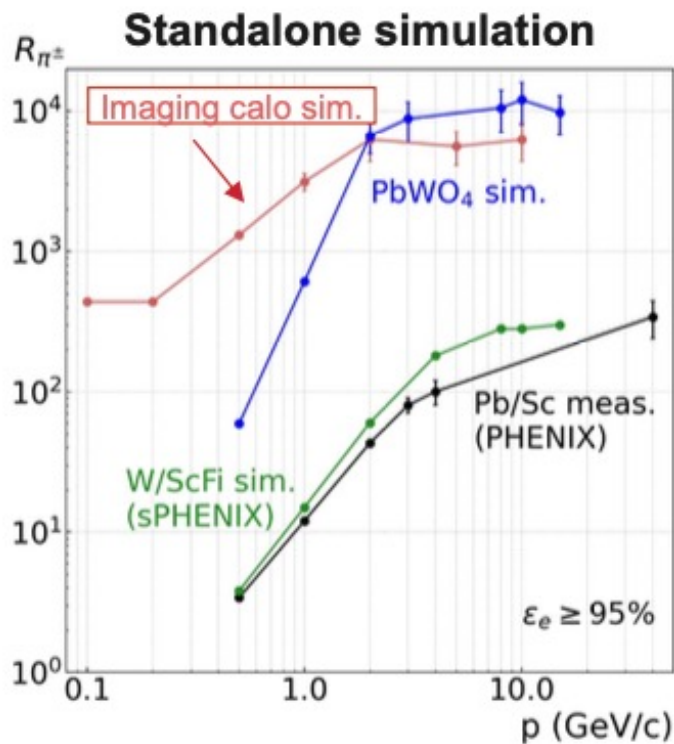
- **4(+2) layers of imaging Si sensors interleaved with 5 Pb/ScFi layers**
- **Followed by a bulk section of Pb/ScFi section**



Energy resolution - Primarily from Pb/ScFi layers (+ Imaging pixels energy information)

Position resolution - Primarily from Imaging Layers (+ 2-side Pb/ScFi readout and radial segmentation)

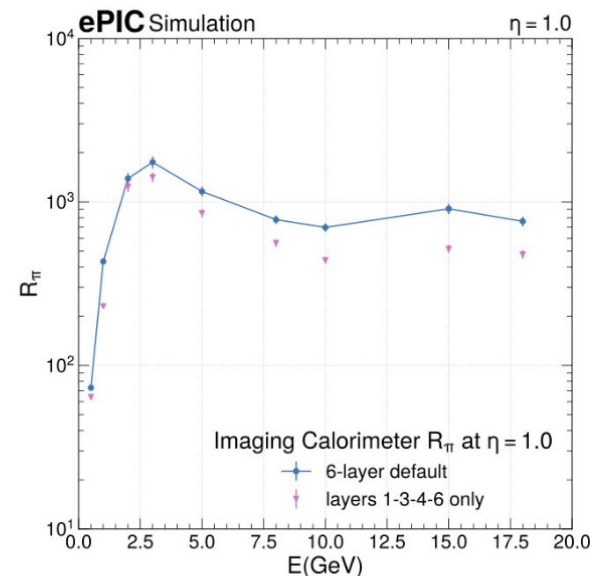
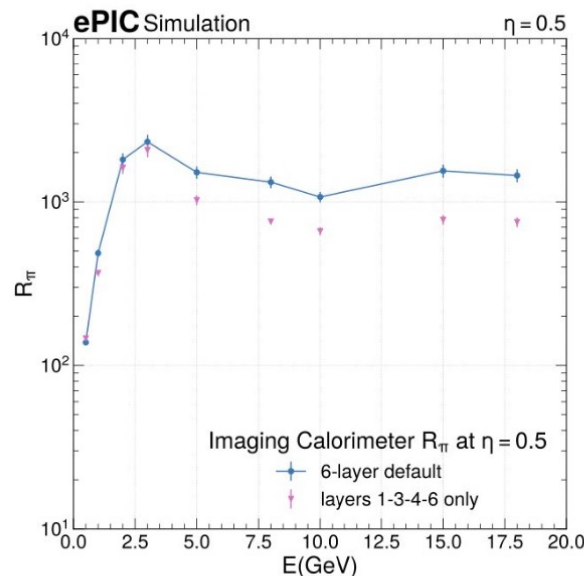
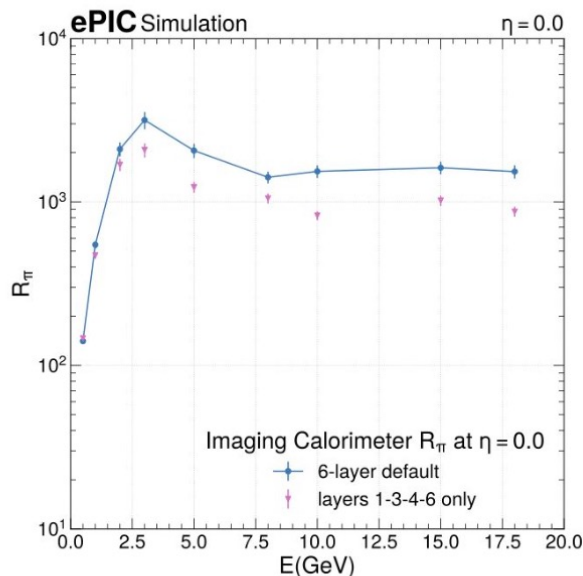
Barrel Imaging Calorimeter: Performance Example



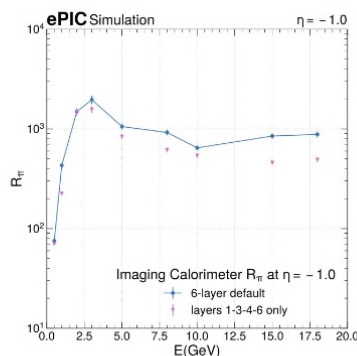
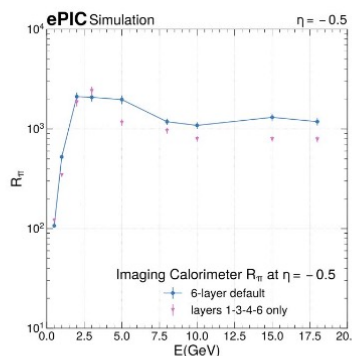
- **Goal:** Separation of electrons from background in Deep Inelastic Scattering (DIS) processes
- **Method:** **E/p cut (Pb/ScFi) + Neural Network** using **3D position and energy info** from imaging layers
- e- π separation exceeds 10^3 in pion suppression at **95% efficiency** above 1 GeV in realistic conditions!

Performance with reduced number of layers

e/π separation at 95% efficiency



4-layer alternate:
layers 1-3-4-6

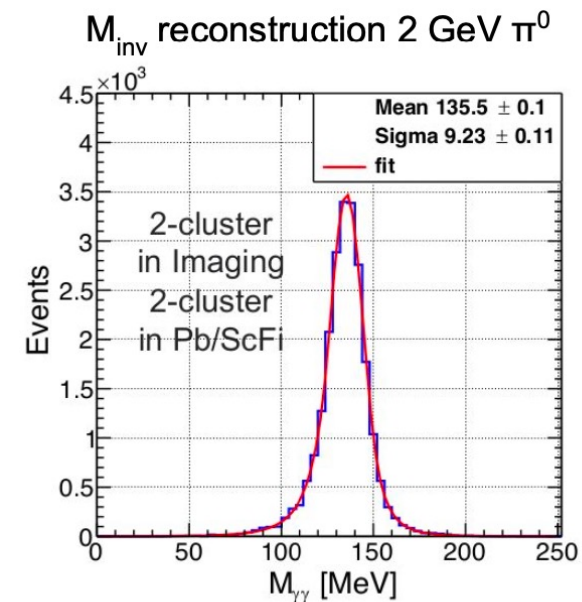
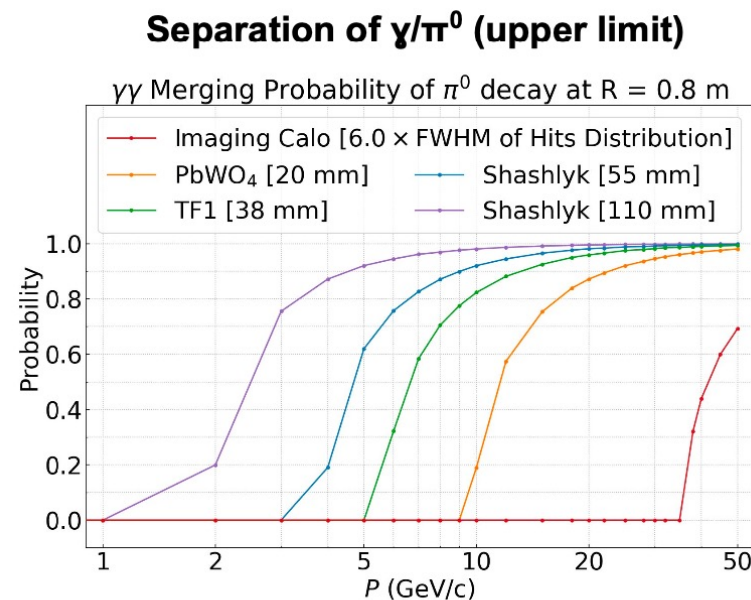
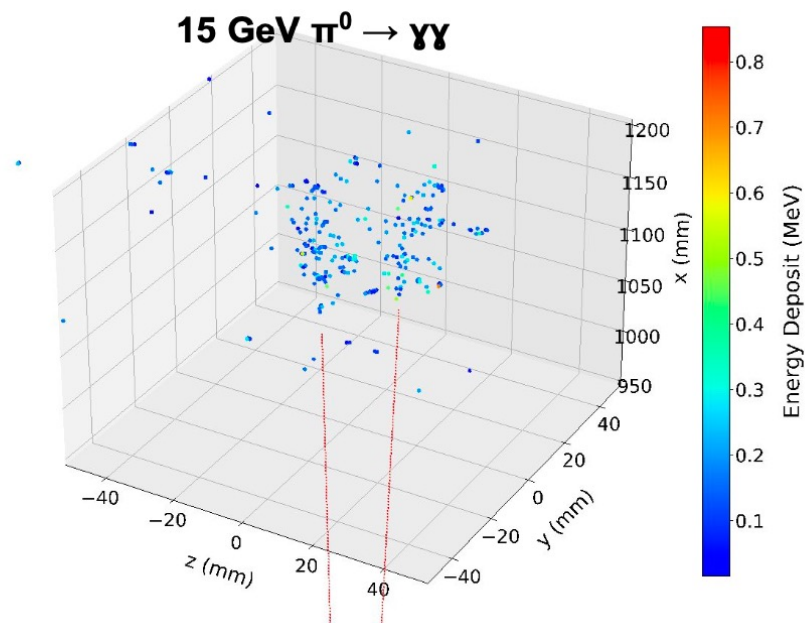


Default configuration exceeds 10^3 pion rejection almost everywhere

4-layer alternate still performs relatively well at lower energies (where most rejection is needed), larger degradation at higher energies

4-layer alternate seems workable compromise.

Neutral Pion Identification



- **Goal:** Discriminate between π^0 decays and single γ from DVCS, neutral pion identification
- Precise position resolution allow for excellent separation of γ/π^0 based on the 3D shower profile
- Reconstruction of 2 GeV π^0 invariant mass as a testing ground for cluster energy splitting

Separation of two gammas from neutral pion well above required 10 GeV

γ/π^0 Separation - Exploratory Studies

Convolutional neural network utilizing energy and spatial information from AstroPix layers

- Started from **10 GeV/c at $\eta = 0$** - the upper limit for γ/π^0 from YR

No proper **topological clustering algorithm** in the ePIC reconstruction yet

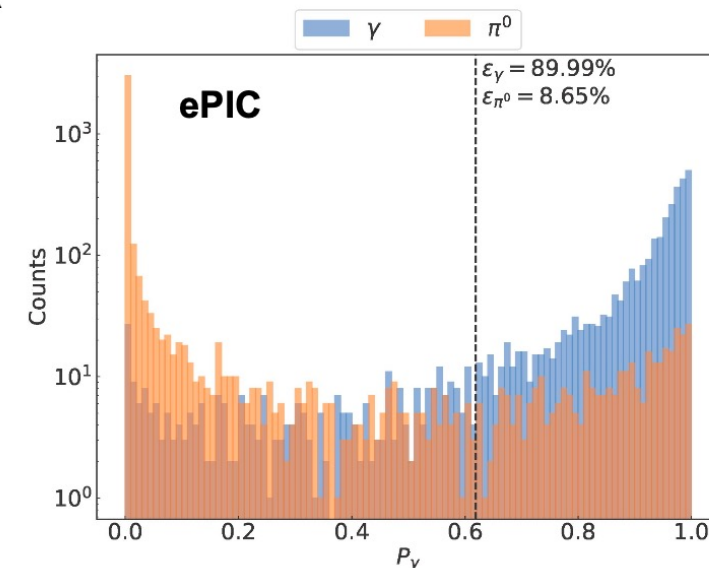
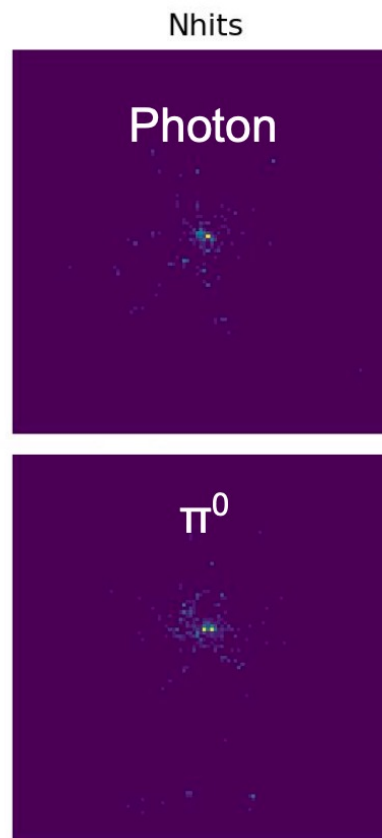
With a quick study we easily achieved

10 GeV/c particles - **91.4%** rejection of π^0 at **90%** efficiency of γ (better than PbWO_4 crystal with 20mm block size)

4-layer alternate is workable (still better than theoretical limit on a crystal calorimeter!), but reduced π^0

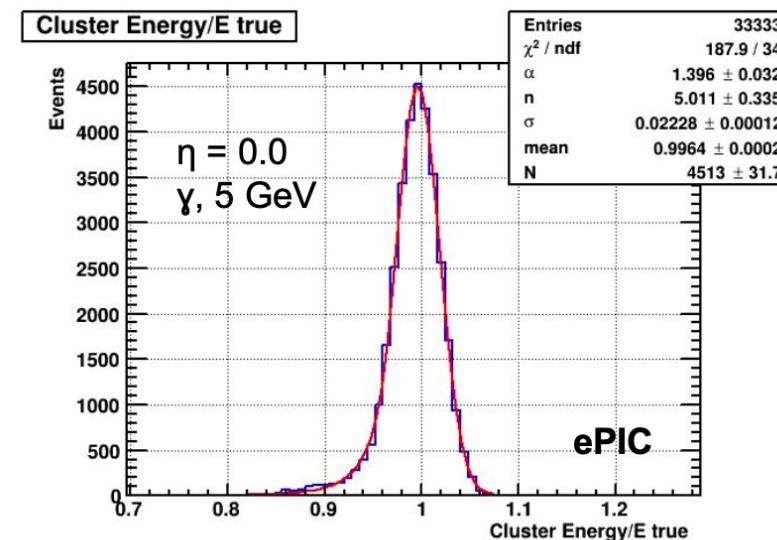
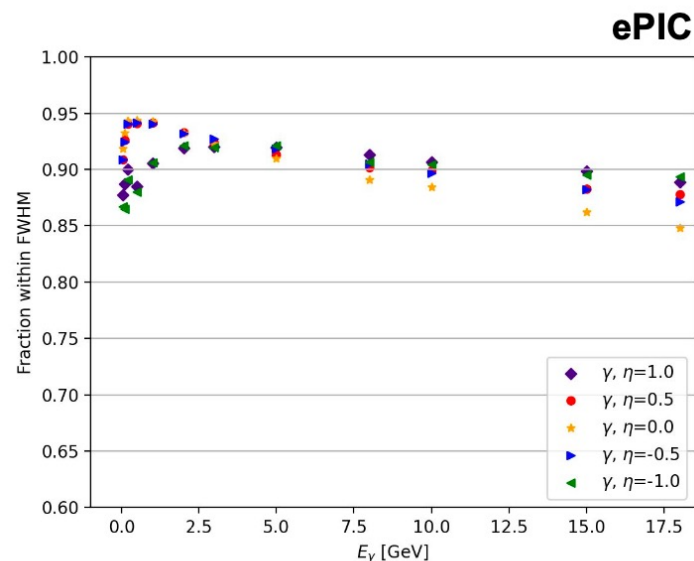
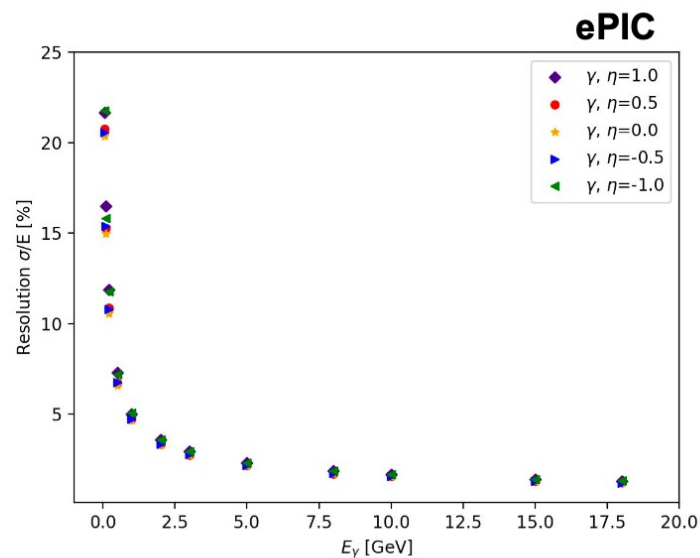
Full study is ongoing:

- Implementing optimized topological clustering for AstroPix layers
- Significant improvements expected



Configuration	γ efficiency	π^0 rejection at 10 GeV/c
6-layer default	90%	11.5
4-layer alternate	90%	5.4

Energy Resolution - Photons



Fit parameters

η	a/\sqrt{E} [%]	b [%]
-1	5.1(0.01)	0.47(0.03)
-0.5	4.77(0.01)	0.38(0.02)
0	4.67(0.01)	0.40(0.02)
0.5	4.75(0.01)	0.39(0.02)
1	5.1(0.01)	0.41(0.02)

- Based on Pb/ScFi part of the calorimeter
- Resolution extracted from a Crystal Ball fit σ

GlueX Pb/ScFi ECal: $\sigma = 5.2\% / \sqrt{E} \oplus 3.6\%$ NIM, A 896 (2018) 24-42

- 15.5 X_0 , extracted for integrated range over the angular distributions for π^0 and η production at GlueX ($E_\gamma = 0.5 - 2.5$ GeV)
- Measured energies not able to fully constrain the constant term

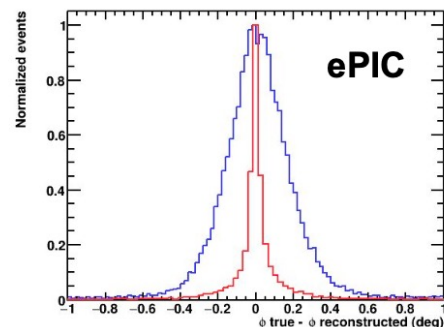
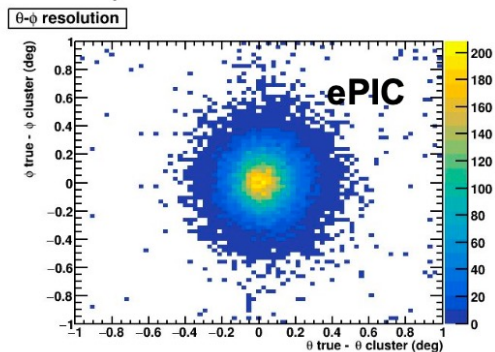
Simulations of **GlueX prototype** in ePIC environment agree with data at $E_\gamma < 0.5$ NIM, 596 (2008) 327-337

Position Resolution

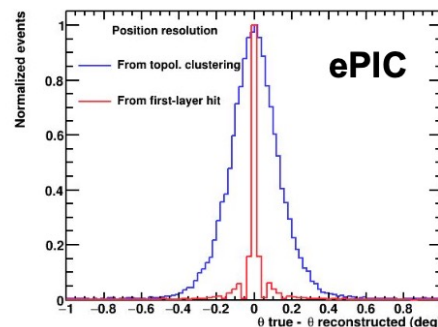
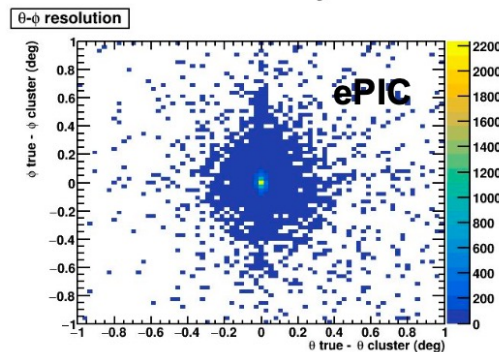
with 6 AstroPix Layers

Example of $\theta - \phi$ resolution for 5 GeV photons

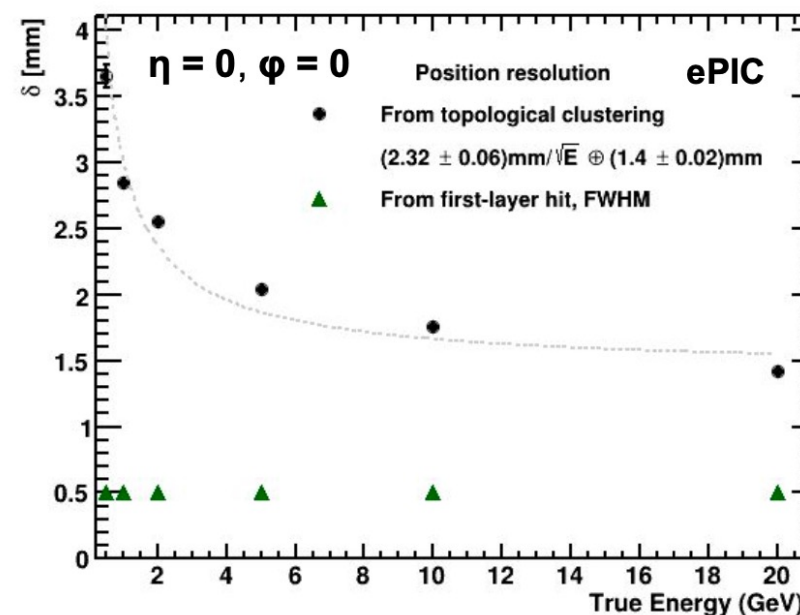
Only information from clusters



Clusters + first-layer hit



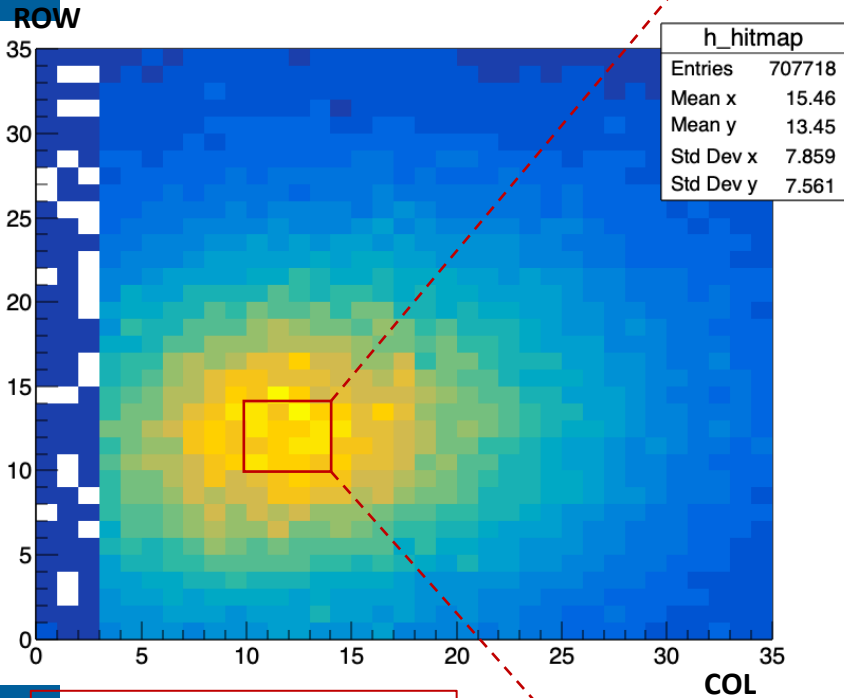
Position resolution for photons Particles thrown perpendicular to the calo surface



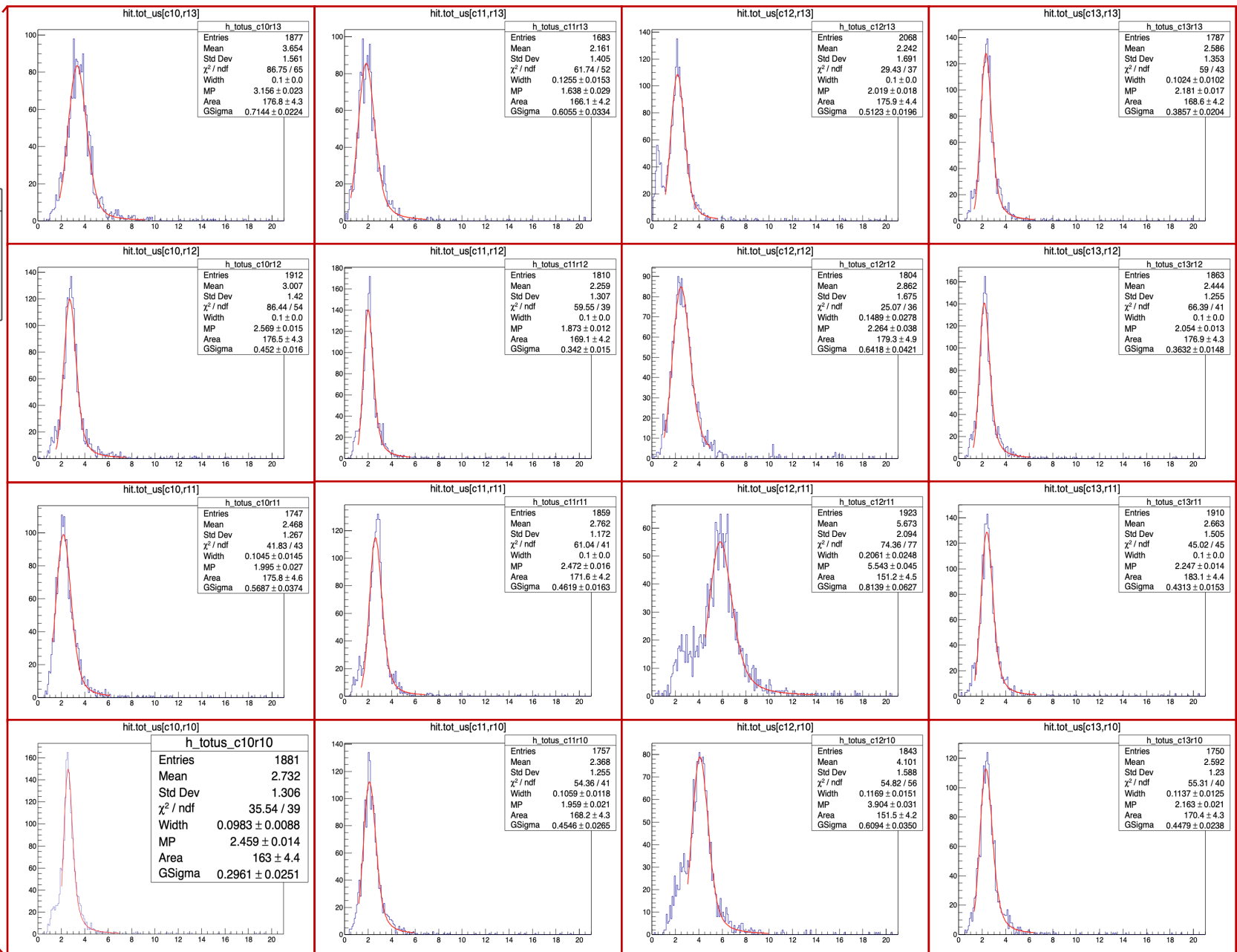
- Clusters from Imaging Si layers reconstructed with 3D topological algorithm
- **Cluster level information:** $\sigma_{\text{position}} = (2.32 \pm 0.06)\text{mm}/\sqrt{E} \oplus (1.4 \pm 0.02)\text{mm}$ at $\eta=0$
- **First-layer hit information added:** $\sigma_{\text{position}} = \sim 0.5\text{mm}$ (pixel size)

120 GeV Proton: ToT distributions (June14)

- Fit with landau convoluted with gaussian function on ToT distribution



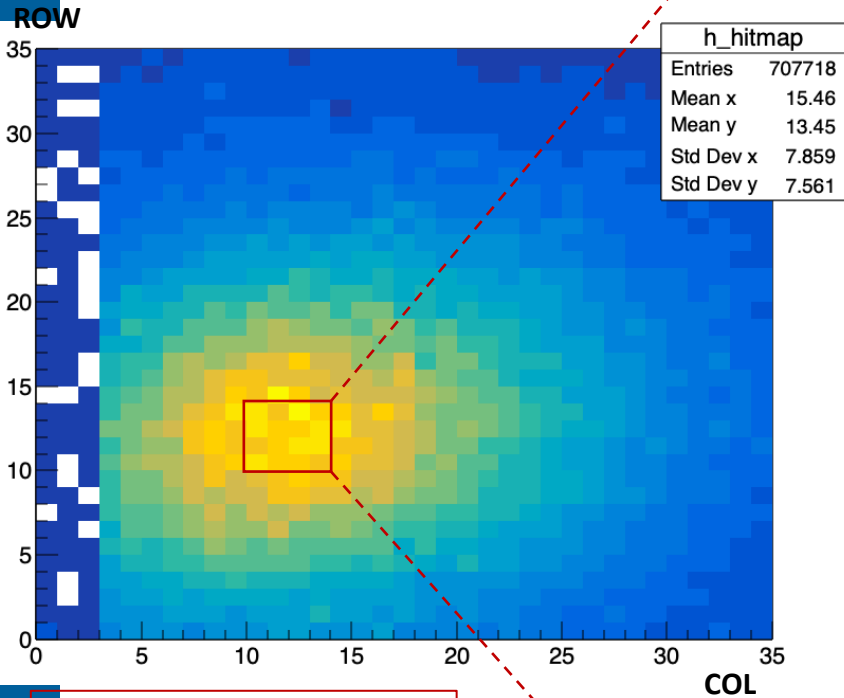
h_hitmap	
Entries	707718
Mean x	15.46
Mean y	13.45
Std Dev x	7.859
Std Dev y	7.561



- 4 x 4
 - Col 10 – 13
 - Row 10 - 13

120 GeV Proton: ToT distributions (June14), **scaled**

- Fit with landau convoluted with gaussian function on ToT distribution



- 4 x 4
 - Col 10 – 13
 - Row 10 - 13

