# Update from CLAS12 High-Luminosity TF

Florian Hauenstein Nov 10, 2023 CLAS Collaboration Meeting





## Why Higher Luminosity



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- Catchup on statistics (existing data about factor two lower statistics than expected)
- Gain time for long remaining physics program
- Opportunities for new, low-rate reactions

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- Catchup on statistics (existing data about factor two lower statistics than expected)
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- Opportunities for new, low-rate reactions
- Phase-1 Upgrade:

~2\*10<sup>35</sup>cm<sup>-2</sup>s<sup>-1</sup> with charged particle reconstruction efficiency > 85%

### How to Achieve Phase-1 Goals?

- Improve FD Tracking with AI and Denoising
- Assessment of detector performances at higher luminosities
- New, fast µWELL tracker in front of R1 DC



### Improved FD Tracking with AI and Denoising



- Close to required slope of 0.001/nA
- FD tracking without µRWELL ready for higher luminosities

### **Detector Performance at Higher Luminosities**

- PMT Detectors (P/ECAL, C/FTOF, L/HTCC, CND, RICH)
  - Operational at x2 lumi
  - Main concern: Aging of PMTs (C/FTOF, ECAL) —> replacement plan in development
  - Non-linearity studies at higher luminosities planned for RGK (CTOF)
- DC
  - RGC and RGD study: Operational at x2 lumi
  - Concern: long term stability at x2 lumi -> monitoring in development
  - In progress: Study effect of higher luminosity on resolutions
- CVT
  - Improved stability of BMT at high lumi with new gas
  - SVT okay for x2 lumi
  - Main concern: Spare tiles for BMT —> plan for new tiles in development
  - Need: Improved tracking efficiency with AI based background rejection
- FT:
  - Operational at x2 lumi
  - Concerns:
    - FTC: Light reduction from radiation damage —> more frequent energy calibrations
    - FTH: Gain dependence of SiPm current —> Redesign of FTH FEE

### New µRWELL Tracker



- Micro-Pattern Gaseous Detector (MPGD)
  - Amplification in wells
  - Spark protection due to resistive layer
  - Low-mass, good spatial and timing resolution
- 2D capacitive sharing readout



K. Gnanvo, NIM A1047, 167782 (2023)

### Our Prototype

- Largest µRWELL build so far
- 2D-U/V strip readout with 10 deg stereo angle
  - pitch 1mm
  - various strip width (to find optimal combination)
- Electronics APV25 and SRS







### Full Design (in progress)





Segment 1

Segment 2

Segment 3

#### **During Installation**



**Readout** 

**Figures from Chris Guthrie** 

## **µRWELL in GEMC**

Stors U-V Beam direction



- Implemented upstream of R1-DC (M. Bondi)
  - 6 sectors with either 1 or 2 double-sided detectors
  - each detector: 3 layers capacitive sharing with U-V readout
- µRWELL simulations
  - Coatjava geometry service:
    - GEANT4 volumes for GEMC
    - Definition of strip lines in local and global frame
  - GEMC:
    - **Geometry** full detector completed and merged. Prototype geometry implemented
    - **Digitization** effective description of detector response. Implementation completed and included. More tuning to be done based on prototype and readout electronics tests.

Slide by Raffaella

### Combined µRWELL-DC Track Finding



#### Slide by Raffaella

### **µRWELL+DC** Tracking

- Forward Tracking modified to use URWell as an additional region (R0)
- First studies of impact on resolution in different tracking scenarios completed
- Waiting on GEMC background to continue track finding and fitting optimization



#### Slide by Raffaella

T. Cao

### **Prototype Testing- Status April**

- HV current leaks from left to right —> only half operational
- Low efficiency in cosmic setup -> 1%
- No clear signal during beam tests in HallB and high noise
- Improvements April May:
  - seal of gas inlays —> bubbles observed on exit
  - improved trigger latency —> no signal clipping
  - improved cosmic setup —> GEM and trigger scintillators closer
  - HV filter box
  - signal noise suppression with faraday cage with wrapped Al foil



### Setup at EEL



- Particles need to pass 10x10 GEM and scintillators —> defines area of µRWELL
- Efficiency is relative to GEM which is assumed to be 100% efficient

### Cosmic Signal with uRWELL









### Cosmic Signal with uRWELL and GEM





Some events have signal in both detectors

### Analysis Procedure (guidance from Kondo)

- Select hits  $3\sigma$  away from noise/pedestal
- Clusters with at least 2 hits

Note: clusters with 1 hits are ignored in the analysis.



#### **ADC U-clusters**



**GEM XY** distribution is

uniform as expected!

#### Event has at least two hits in GEM X and Y



#### Plots by Rafo

### **2D Distribution**



### **Efficiency Results**



Drift Voltage = 400 V



- Increase of efficiency with voltage
- Difference between U and V efficiency
- BUT: low overall efficiency

### **Efficiency Results**

Ar:CO<sub>2</sub> 80:20





- Larger efficiency for 80:20 due to more gain
- Still difference between U and V efficiency
- Expected efficiency >90%!

### Test with Ar:C<sub>4</sub>H<sub>10</sub> (Isobutane) 90:10

80

**Ar:C**<sub>4</sub>**H**<sub>10</sub>

Efficiency [%] - Eff. U cluster Eff. V cluster Eff. Any cluster Mesh HV **Ar:CO<sub>2</sub> 80:20** • 575V for Ar:CO<sub>2</sub> 60 **500V for Ar:C**<sub>4</sub>H<sub>10</sub> 50 **Result with Ar:C4H10**  ~2 times larger efficiency U and V similar 40 **BUT: unstable operation** 30 20 400 450 500 550

 Note: No more measurements with different gas because detector was sent back to CERN for repairs (see later slide)

600

HV Drift [V]

### Efficiency Dependence of Strip Width/Length



#### • 3 GEM positions

- U175µm / V355µm crossing (run 1576)
- U175µm / V550µm crossing (run 1574)
- U550/262µm / V500µm crossing (run 1582)
- Ar:CO<sub>2</sub> 75:25 and same HV on  $\mu$ RWELL

### **Results for the Runs**

run 1582

#### run 1576



- All "V" strips show the same efficiency (~14%), however the strips in 1574 are in average x2 longer than the one for 1582
- U strip efficiency dependence on strip width/length difficult to interpret

run 1574

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#### **Dependence on capacitance of strips??**

run 1574

### U strip capacitance



U strip capacitance vs Area

- Capacitance is high typical values for GEMs with APV readout are 100-200pF —> could explain lower overall efficiency —> loss of signal to noise
- More studies underway



Plot by Rafo

### **Upcoming Plans with Prototype**

- Detector repaired at CERN
  - leakage problem solved (confirmed in EEL)
  - recovered bad HV sectors
- Cosmic measurements on left side with lower capacitance



• Measurements with VMM3



- Study of other gas mixtures and HV
- Next year maybe again in the hall?



### µRWELL Prototype lests at INFN

#### 2D – readout: step by step approach

1. One prototype reads the 2-nd coordinate on the "top" copper layer

Same readout geometry on the ton and the bottom.



RWELL

layer is the object of investigation.



- 1200 μm pitch
- 300  $\mu m$  vs 1000  $\mu m$  strips width
- 10 x 10 cm<sup>2</sup> active surface
- 83 channels







Coppe

Kaptor

Honeycomb/millifoam

Al – Faraday Cage

Glue FR4 glue

glue FR4 glue

### **µRWELL Prototype Tests at INFN**

 $\pi/\mu$ 

#### Test Beam at CERN SPS North Area H8 in June



#### Efficiency

 $10x10 \text{ cm}^2$ 

Top/Bottom

2D readout

 $0 \text{ cm}^2$ 

- Leadout

city Sharing

- CS readout reaches a plateau at higher HV values than standard readout scheme.
- TOP readout is not yet at plateau at 600 V
  (HV was chosen to to be raised to higher values)



e

Reference

Tracker - OUT





#### **NEXT STEP**

• Test the 50x50 cm2 prototype!

#### with X-RAYS and cosmics





November 9, 2023



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### Future Phase-2 Luminosity Upgrade

- CEBAF can deliver 10<sup>37</sup>cm<sup>-2</sup>s<sup>-1</sup> luminosity
- High luminosity needed to study new and low cross section reaction like DDVCS



- Phase-2 update converts CLAS to µCLAS
  - new Moller cone
  - µ-pairs in FD after shielding
  - calorimeter and tracker instead of HTCC
  - tracker inside solenoid for recoils
  - timeframe 7-10y



### Tracker — High-rate capable µRWELL

- LDRD project
- Prototype testing
  - Various resistive layer layouts
  - Capacitive sharing X-Y-U strip readout
  - Thin gap



Cross section view of 3-coordinates X-Y-U strips capacitive-sharing readout

- Software development
  - Implementation in GEMC
  - Hit and track reconstruction algorithm in a high-rate environment
  - Validation with test measurements
- Personnel:
  - 1 Postdocs (80% FTE)
  - Florian Hauenstein, Rafayel Paremuzyan, Kondo Gnanvo

### **Questions?**

## Backup

### Results Ar/CO2 75:25



### Results Ar/CO2 80:20



#### UDMELL Drototupo Tooto of INEN





#### Resolution

- CS readout reaches 100  $\mu m$ resolution at highest HV values (starting from 1200  $\mu m$  pitch)
- TOP readout resolution is fixed at 250-300 μm (pitch is 780 μm)



### **Urwell in HallB**



### Noise Issue fixed



- However, still no clear signal observed :( latency issues?
- Installed GEM behind the trigger bar in FTOF panel 2, cover about 1/90 of FTOF bar

### Signals from GEM in HallB



GEM showed signal but still nothing seen in uRWELL

### **Alternative Design**

strips parallel to sides for segment 1, readout from top, pitch could be larger than 1mm strips parallel for segment 2 with vias to readout on side strips at +/-10deg for Segment 3



### **Generation of Beam Background**

- Using GEMC to generate URWell beam background
- Assessing the accuracy comparing DC GEMC background to data
  - Currently both occupancy and tracking efficiency from luminosity-scan analysis indicate simulated BG is 30% lower than data
  - -Investigation of possible sources in progress





#### Slide by Raffaella