

Winter Hall A Collaboration

Jan 16, 2024, 8:00 AM → Jan 17, 2024, 7:25 PM US/Eastern

ARC 231/233 & Cebaf Center Auditorium (Jefferson Lab)

Arun Tadepalli (Jefferson Lab), Bishnu Pandey (Virginia Military Institute), Carlos Ayerbe Gayoso (Old Dominion University),
Mark Jones (Jefferson Lab), Michael Nycz (University of Virginia)

BDX

Dark Matter search in a beam-dump experiment at Jefferson Lab

*M. Battaglieri (INFN)
on behalf of the BDX Collaboration*

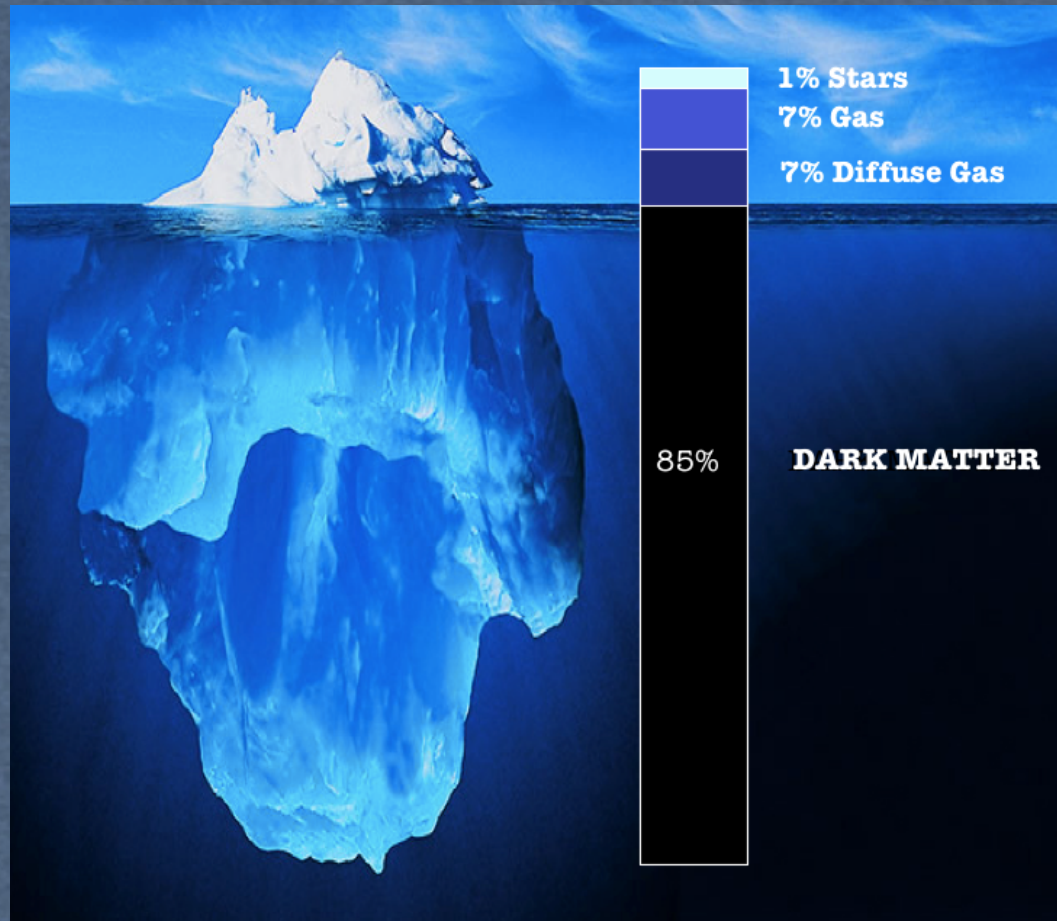


Jefferson Lab's accelerator site

Dark Matter (DM) vs Baryonic Matter (BM)

Compelling astrophysical indications about DM existence

★ How much DM w.r.t. BM?



★ Does DM participate to non-gravitational interactions?

★ Is DM a new particle?

★ Constraint on DM mass and interactions

- should be 'dark' (no em interaction)
- should weakly interact with SM particles
- should provide the correct relic abundance
- should be compatible with CMB power spectrum

... assuming that the gravity is not modified and DM undergoes to other interactions

★ We can use what we know about standard model particles to build a DM theory

Use the SM as an example: $SM = U(1)_{EM} \times SU(2)_{Weak} \times SU(3)_{Strong}$

Particles, interactions and symmetries

Known particles
& new force-carriers

Particles:
quarks, leptons

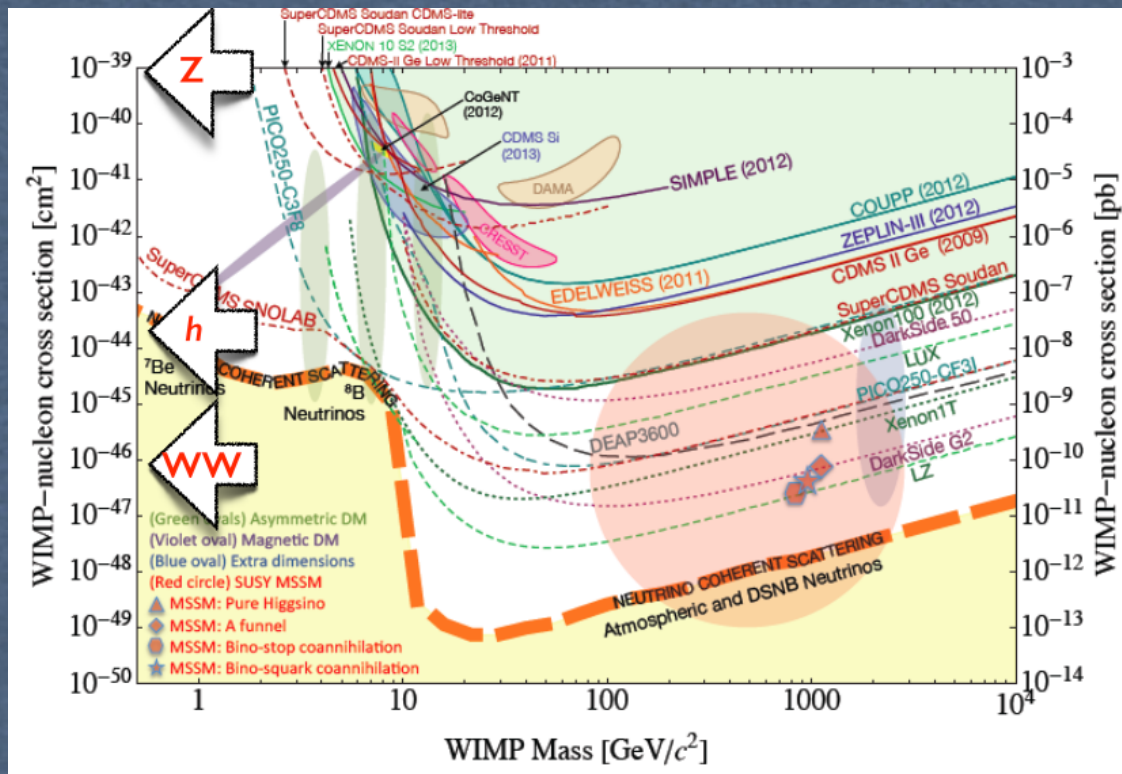
Force-carriers:
gluons, γ , W, Z, graviton (?), Higgs, ...

Two options:

- ★ **New matter** interacting through the **same forces**
- ★ **New matter** interacting through **new forces**

Exploring the WIMP's option

★ Experimental limits



Slow-moving cosmological weakly interacting massive particles

- DM detection by measuring the (heavy) nucleus recoil
- Constraints on the interaction strength from the DM Direct Detection limits
 - Scattering through Z boson ($\sigma \sim 10^{-39} \text{cm}^2$): ruled out
 - Approaching limits for scattering through the Higgs ($\sigma \sim 10^{-45} \text{cm}^2$)
 - Close to irreducible neutrino background
- * No signal observed in Direct Detection
- * Experiments have (almost) no sensitivity to (light) DM ($< 1 \text{ GeV}$)

Direct Detection

1 MeV

1 GeV

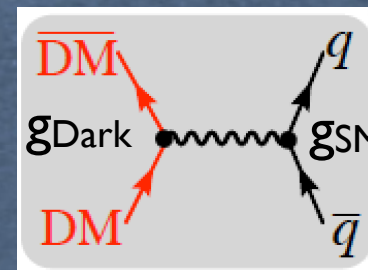
Mz

10 TeV

WIMPs

WIMPs paradigm is not the only option
(keeping the DM thermal origin)

$$\langle \sigma v \rangle \sim g_{\text{Dark}}^2 g_{\text{SM}}^2 M_{\text{DM}}^2 / M_{\text{mediator}}^4$$



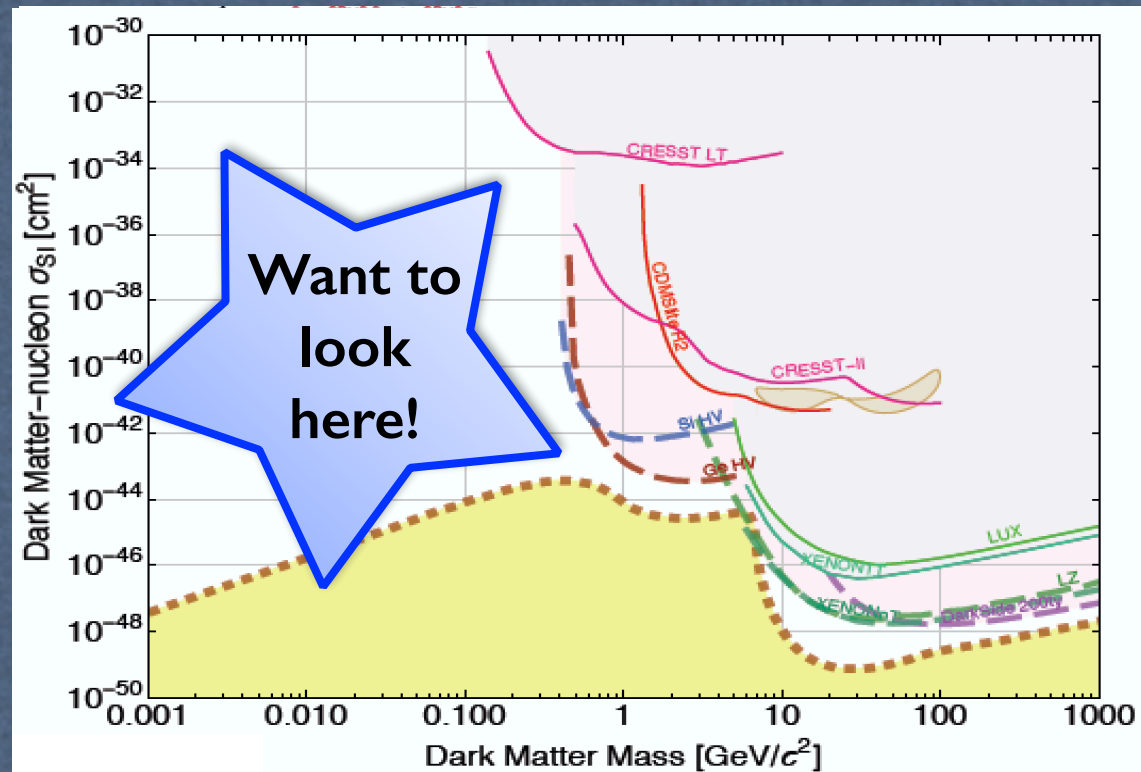
Light Dark Matter

Light Dark Matter ($< \text{TeV}$) naturally introduces light mediators

New interaction

Light Dark Matter

★ Experimental limits



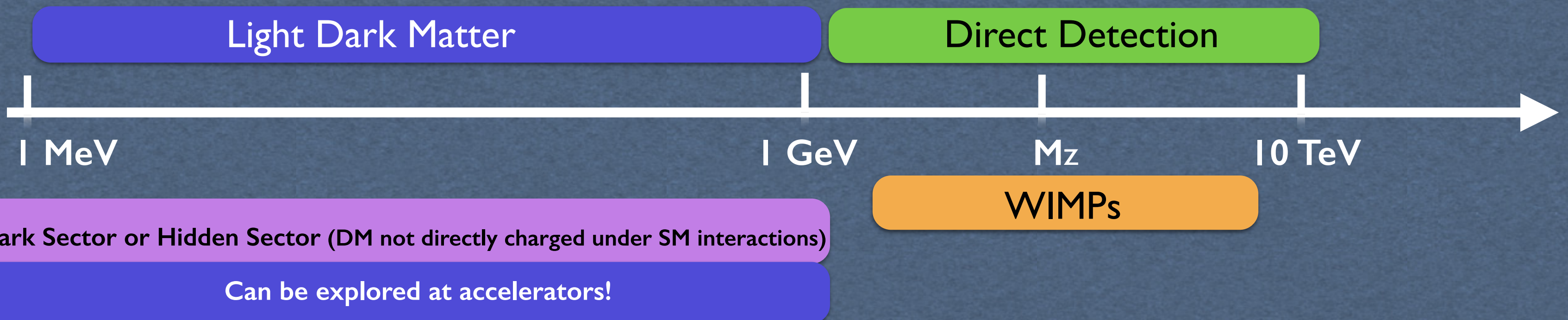
Light Dark Matter with a (almost) weak interaction (new force!)

- Direct Detection is difficult
 - Low mass elastic scattering on heavy nuclei produces small recoil
 - eV-range recoil requires a different detection technology
 - Directionality may help to go behind existing limits at large masses

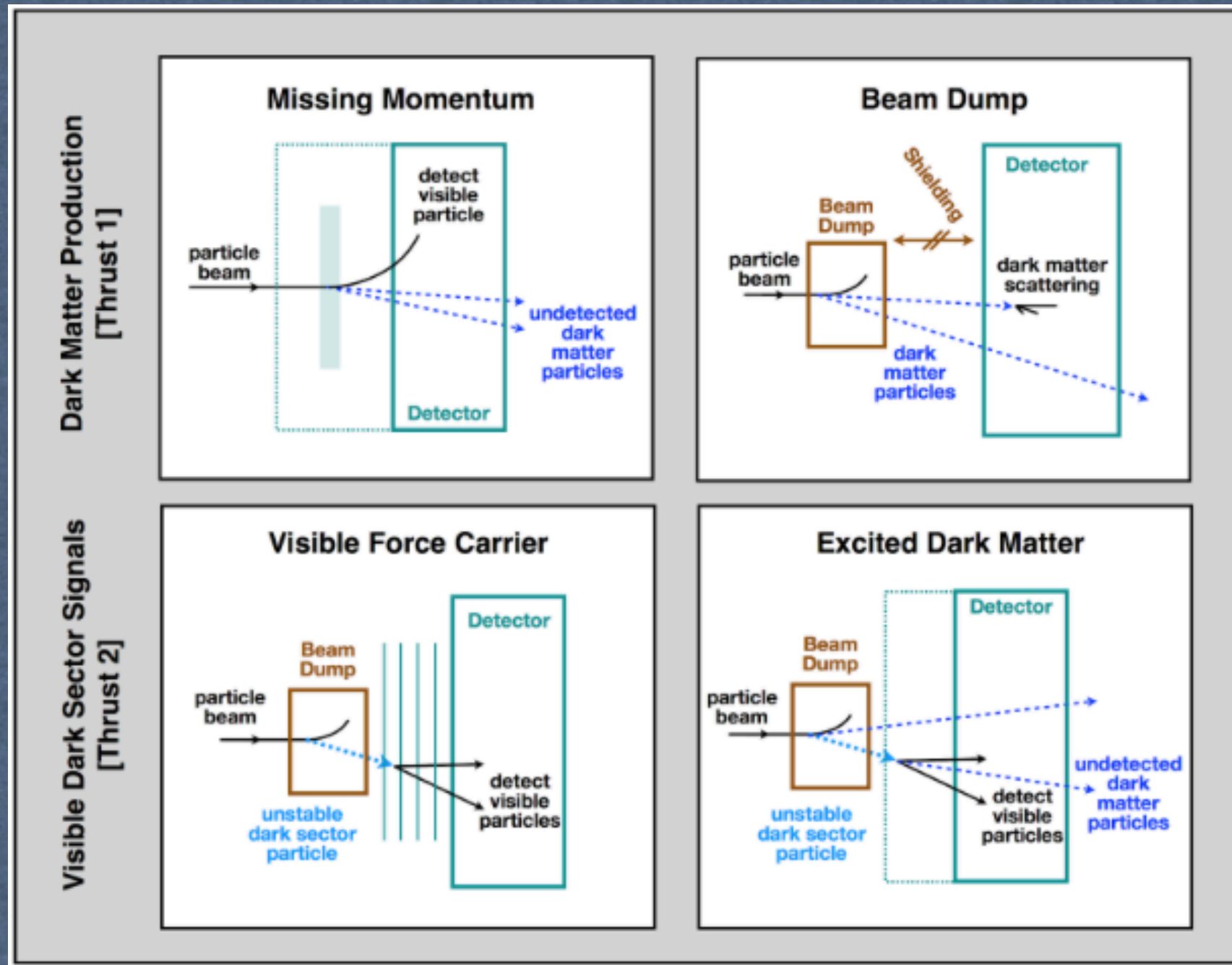
Accelerators-based DM search

covers an unexplored mass region extending the reach outside the classical DM hunting territory

- **High intensity**
- **Moderate energy**



Experimental techniques



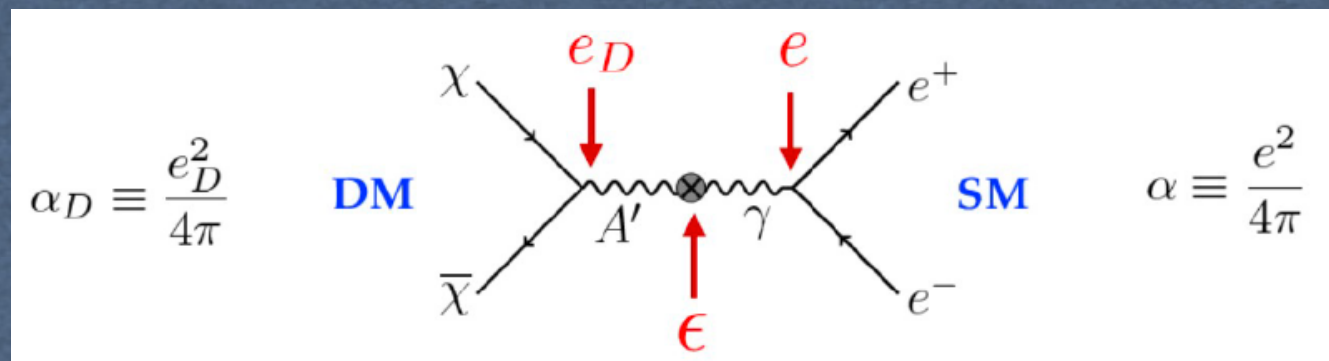
Fixed target vs. collider

	Fixed Target	e^+e^- colliders
Process		
Luminosity	$10^{11} e^-$	$10^{11} e^-$ $10^{11} e^+$
Cross-Section	$\sigma \sim \frac{\alpha^3 Z^2 \epsilon^2}{m^2} \sim O(10 \text{ pb})$	$\sigma \sim \frac{\alpha^2 \epsilon^2}{E^2} \sim O(10 \text{ fb})$
	<ul style="list-style-type: none"> • high backgrounds • limited A' mass 	<ul style="list-style-type: none"> • low backgrounds • higher A' mass
	<ul style="list-style-type: none"> * $1/M_{A'}$.vs. $1/E_{\text{beam}}$ * Coherent scattering from Nucleus ($\sim Z^2$) 	

Dark Photon Signatures

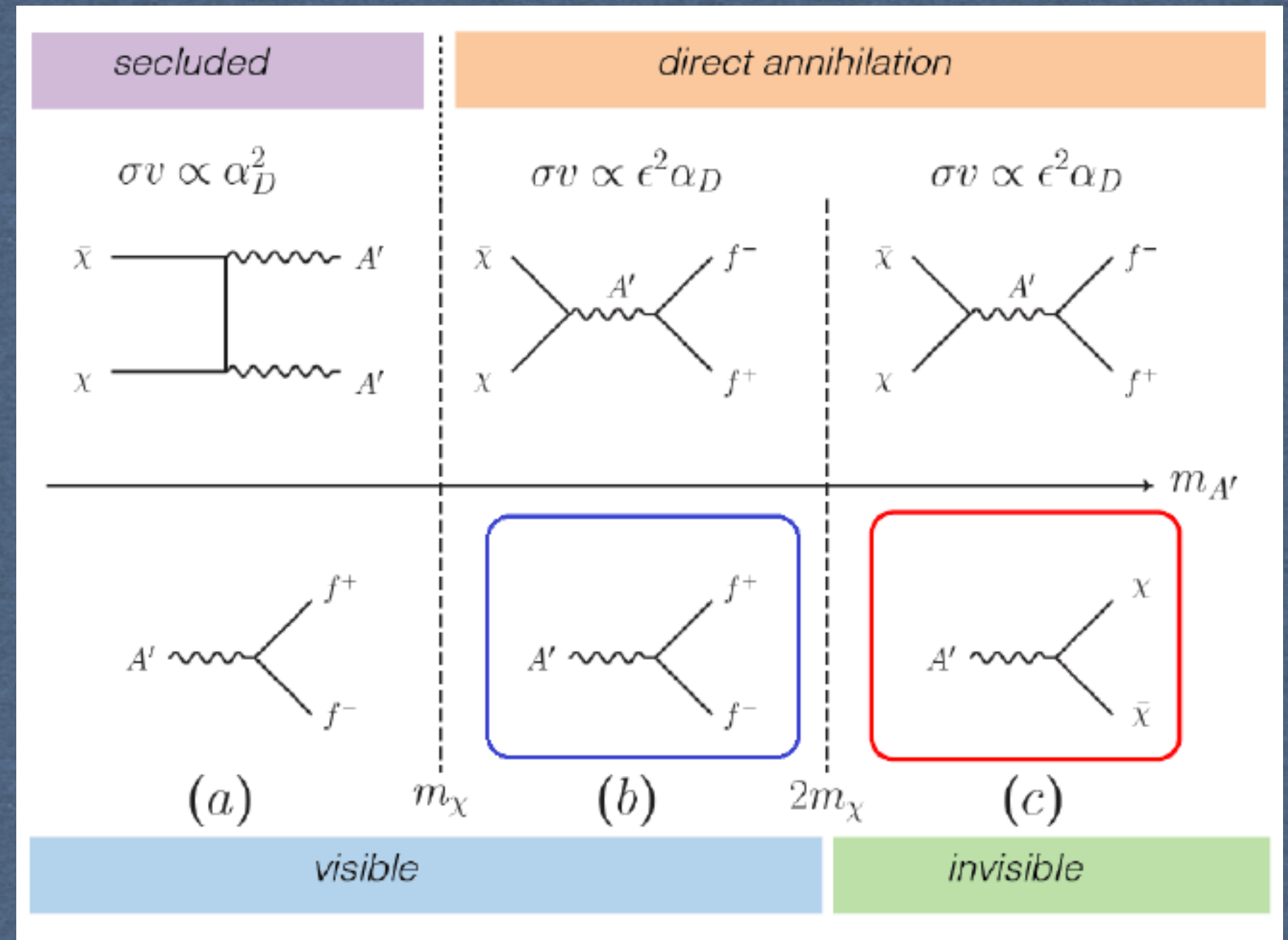
Vector mediated Light Dark Matter

- Vector-Portal: DM-SM interaction mediated by U(1) gauge-boson (dark photon or A') couples to electric charge



A' interaction scenarios

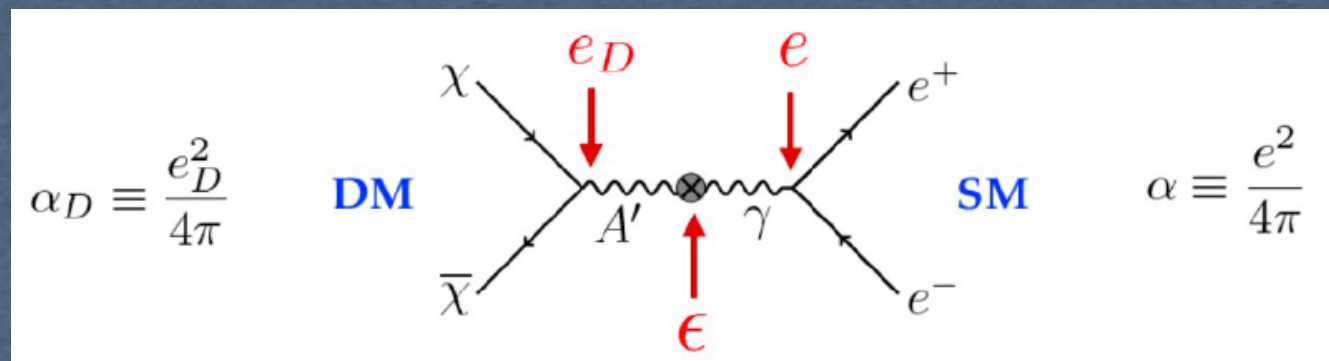
- **Secluded**: no constraints by cosmology for accelerator based experiments. Any ϵ allowed
- **Visible decay**: final state contains SM particles
- **Invisible decay**: A' decays to Dark Sector invisible particles



Dark Photon Signatures

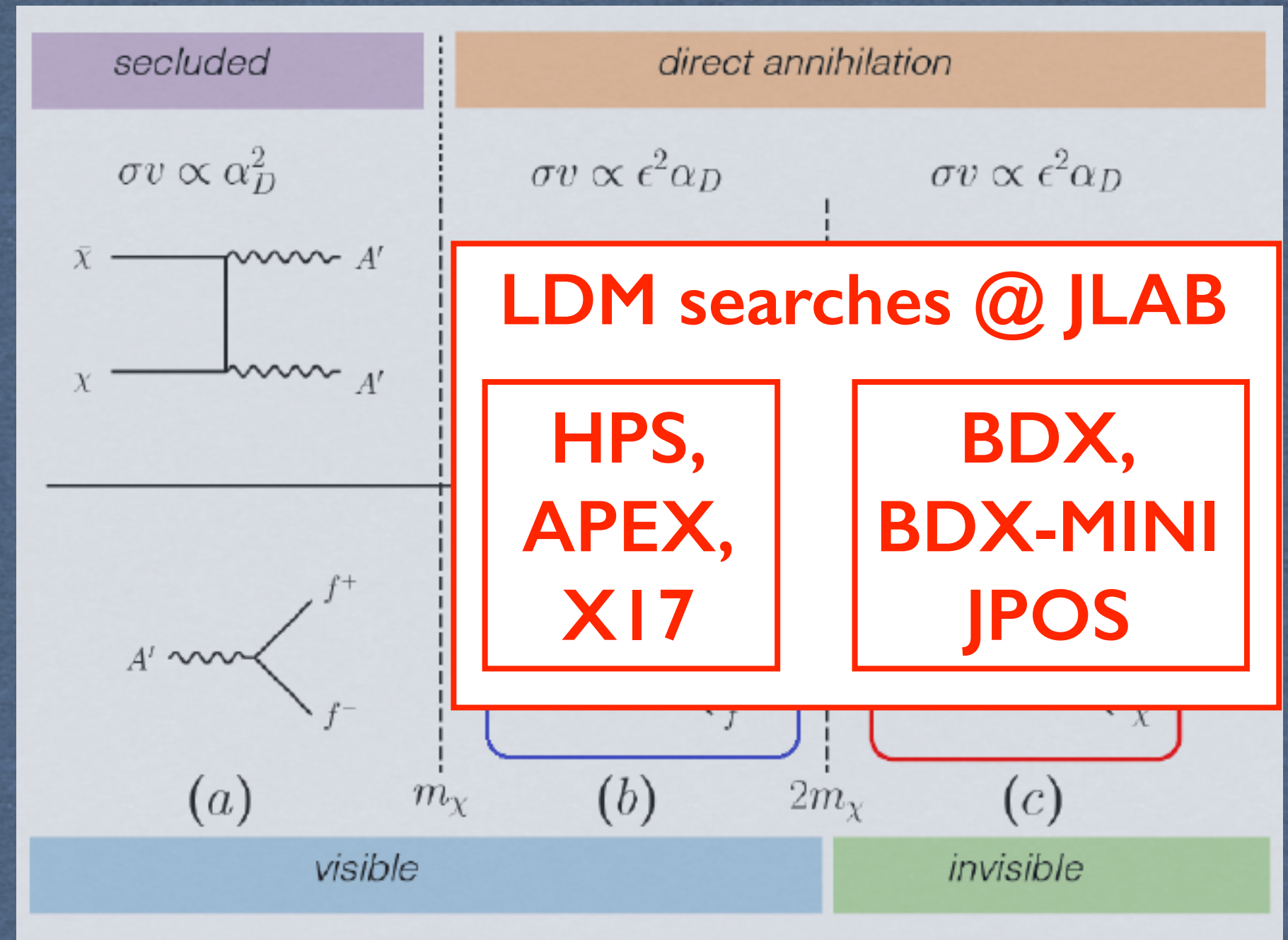
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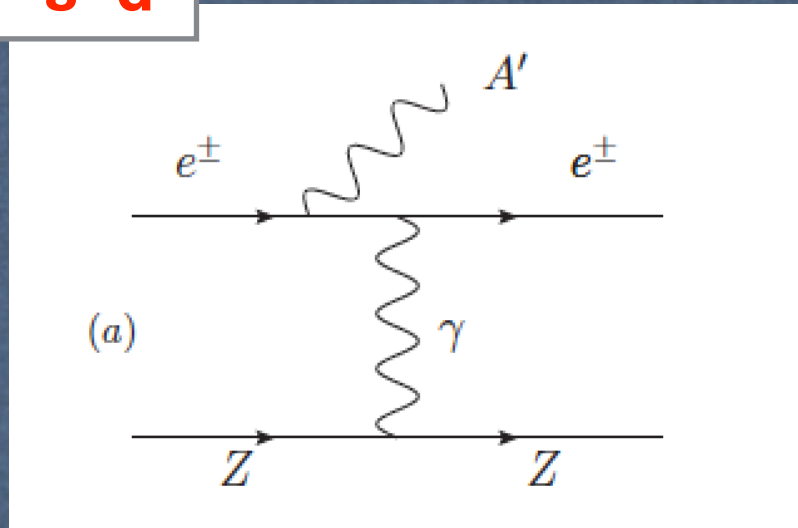
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A' Production mechanisms - e[±]

~ ε² α³



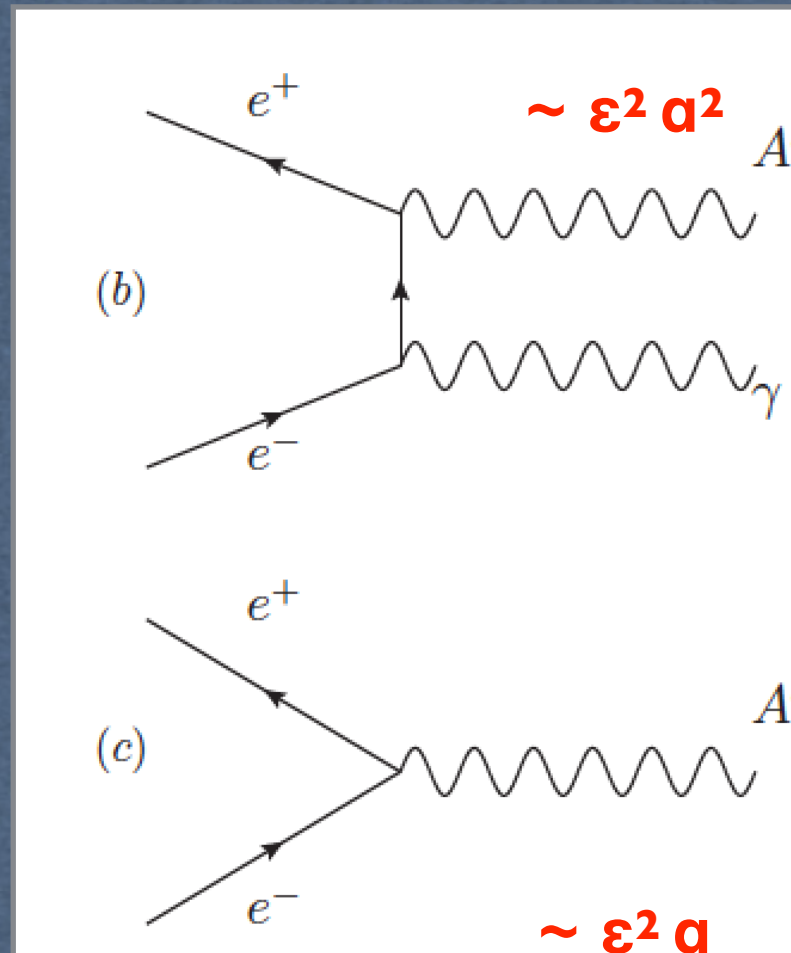
The Weizsacker-Williams approximation (A'-strahlung)

- The first tree-level mechanism proposed

A' Production - resonant/non-resonant production

- Specific for positron annihilation
- A beam dump is a copious source of positrons
- Positrons in the EM shower may have any energy in the range of 0 - E_{beam}

L. Marsicano et al. Phys. Rev. Lett., 121(4) 041802, 2018
L. Marsicano et al. Phys. Rev. D, 98 (1) 015031, 2018



- **NON-RESONANT** annihilation

- **RESONANT** annihilation

$$\sigma_r = \sigma_{\text{peak}} \frac{\Gamma_{A'}^2/4}{(\sqrt{s} - m_{A'})^2 + \Gamma_{A'}^2/4}$$

The BDX experiment

Two step process

I) An electron radiates an A' and the A' promptly decays to a χ (DM) pair

II) The χ (in-)elastically scatters on a e-/nucleon in the detector producing a visible recoil (GeV)

PhysRevD.88.114015 E.Izaguirre,G.Krnjaic, P.Schuster, N.Toro

X production

X detection

elastic on electrons

Inelastic on nuclei

BDX @ JLab

A' yield: $N_{A'} \propto \frac{\epsilon^2}{m_{A'}^2}$

χ cross-section: $\sigma_{\chi e} \propto \frac{\alpha_D \epsilon^2}{m_{A'}^2}$

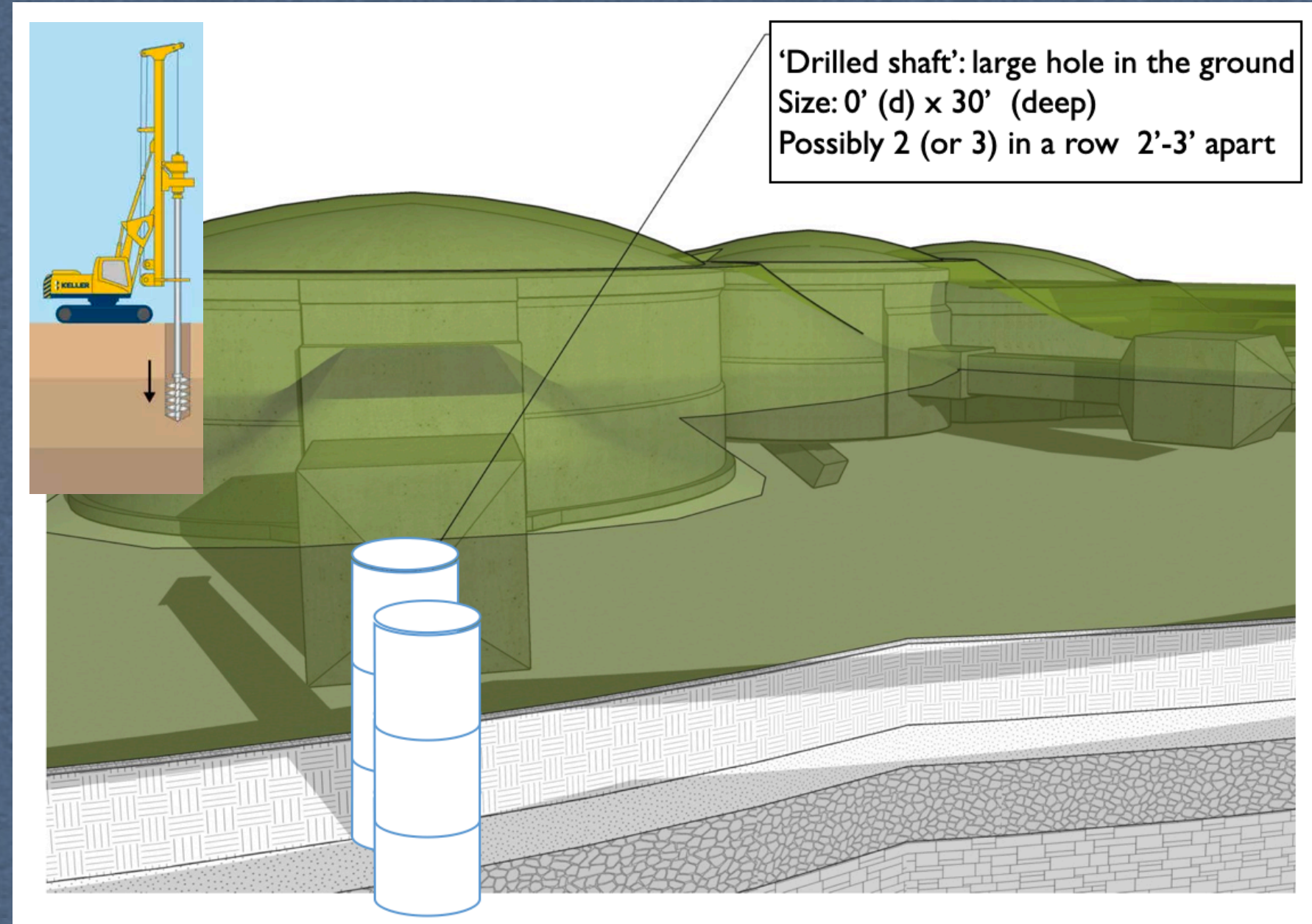
Number of events: $N_\chi \propto \frac{\alpha_D \epsilon^4}{m_{A'}^4}$

- Intense electron beam
- ~ few GeV range energy

Experimental signature in the detector: **X-electron** \rightarrow **EM shower** ~GeV energy

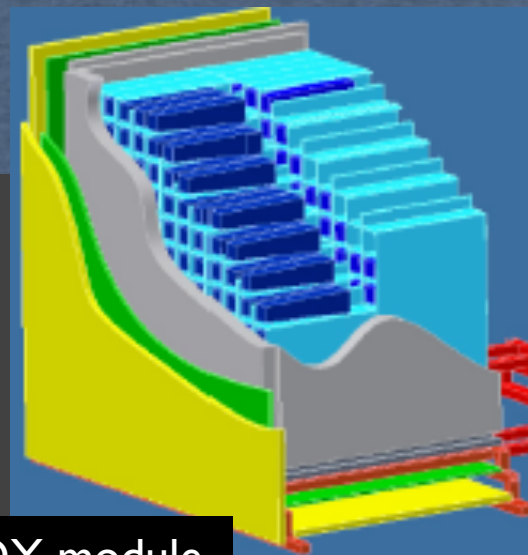
BDX @ JLab

- ★ JLab Hall-A offers the best condition for the BDX experiment
 - A high energy beam: 11 GeV
 - The highest available electron beam current: $\sim 65 \mu\text{A}$
 - The highest integrated charge: 10^{22} EOT (41 weeks)
- ★ Fully parasitic wrt Hall-A physics program
- ★ Drilled shaft downstream of Hall-A BD
- ★ Approved by JLab PAC-46 in July 2018 with maximum scientific rate (A) and waiting for scheduling
- ★ Expected to run in parallel to the Moeller experiment (2026-2029)
- ★ Presented, discussed, and included in SNOWMASS-21 report (RF6-RF0)
- ★ BDX would take advantage of the future 11 GeV positron beam and 20+ GeV upgrade
- ★ BDX Collaboration: more than 100 researchers from 18 institutions (US, Italy, Germany, UK, Korea) signed the BDX proposal



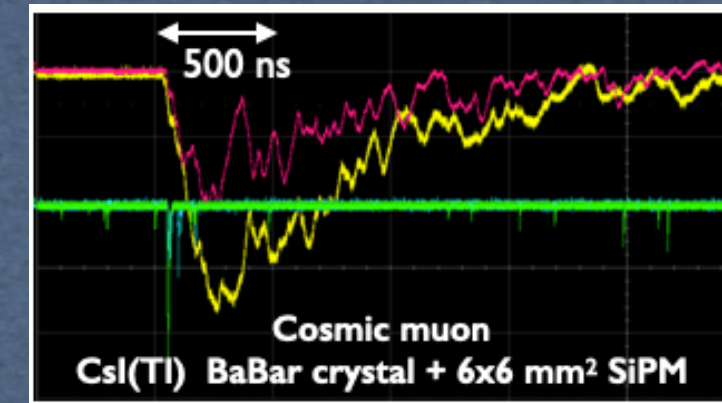
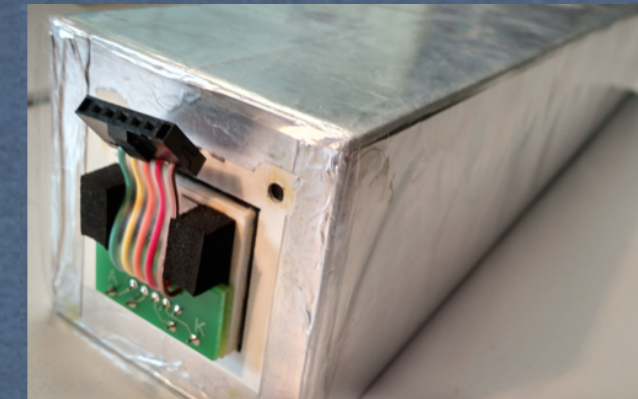
The BDX detector

- Modular design
- 8 modules each having 10x10 crystals
- 800 CsI(Tl) crystals
- 6x6 mm² Hamamatsu SiPM readout + fADC electronics
- Inner and Outer veto: plastic scintillator + WLS fibres, SiPMs



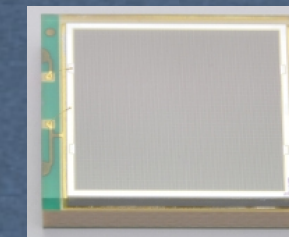
A BDX module

E.M. Calorimeter



Crystals available from BABAR em calorimeter

- Size: (5x5)cm² front face, (6x6)cm² back face, 30cm length
- 820 crystals available from end cap
- Decay time: fast 900ns, slow 4000ns
- LY= 50k γ /MeV



SiPM readout

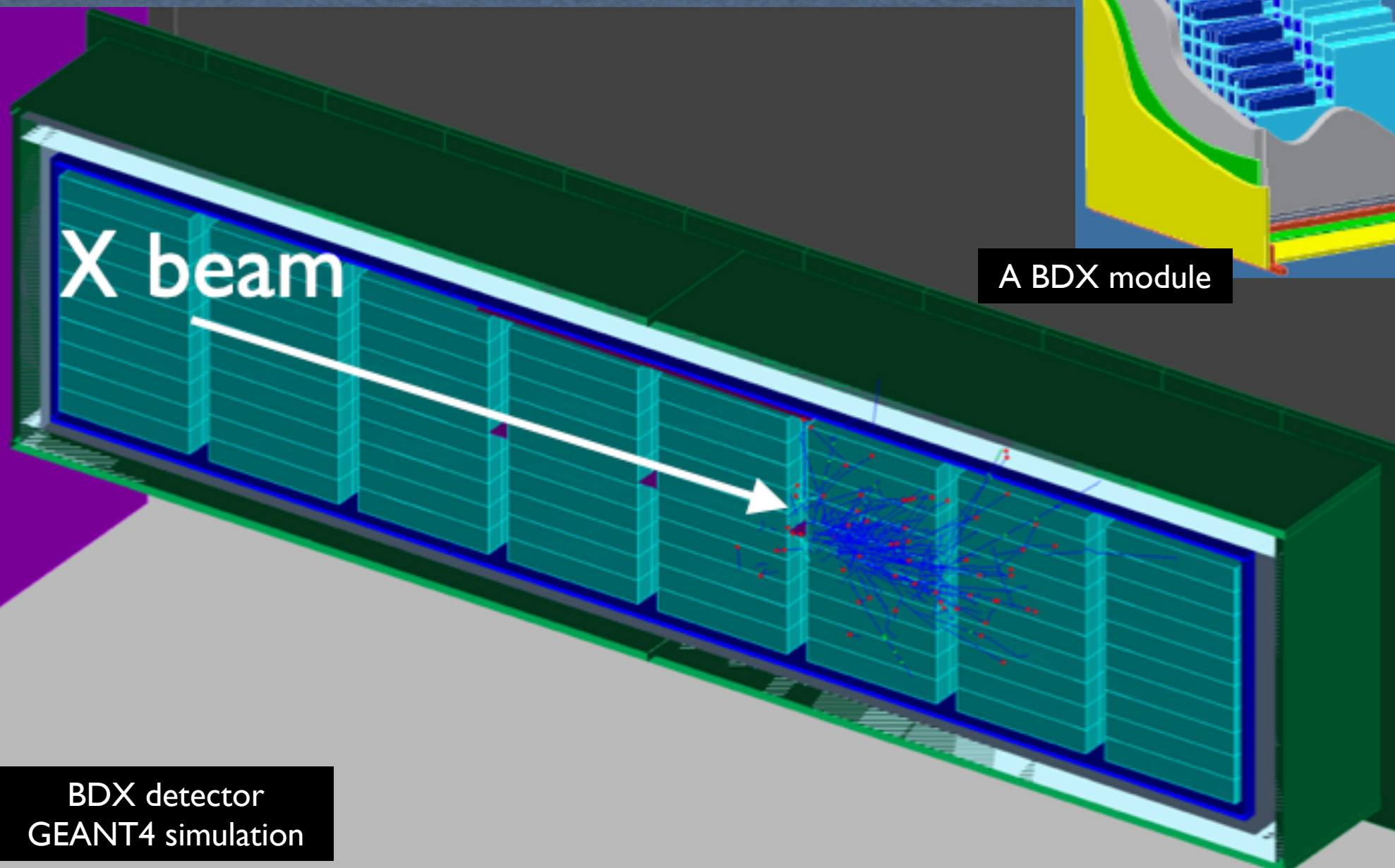
- Size: 6x6 mm², 25 μ m, 57.6k cells, trenched, pde=25%
- SPE capability
- CsI(Tl): 40 pe/MeV
- Time resolution: ~6ns (MIPs)

M.Bondi et al. NIM.A, 867:148-153, 2017

Veto

Inner/Outer veto:

- 1cm (clear) plastic scintillator + WLS fibers
- 3x3 SiPM readout
- LY= 15-50 pe/MIP

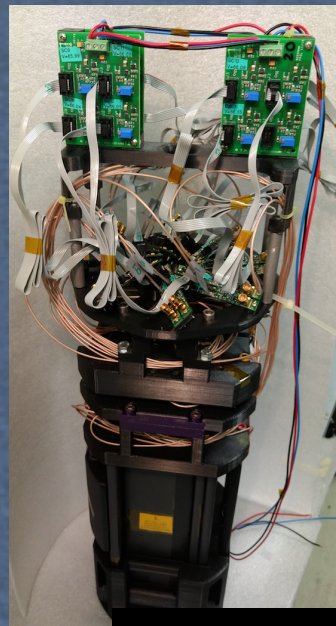


BDX detector
GEANT4 simulation

Background

- Beam unrelated: cosmic rays (BDX-PROTO)
- Beam related: n and muons (BDX-HODO)
- High stat simulations (FLUKA and GEANT4) validated with measurements

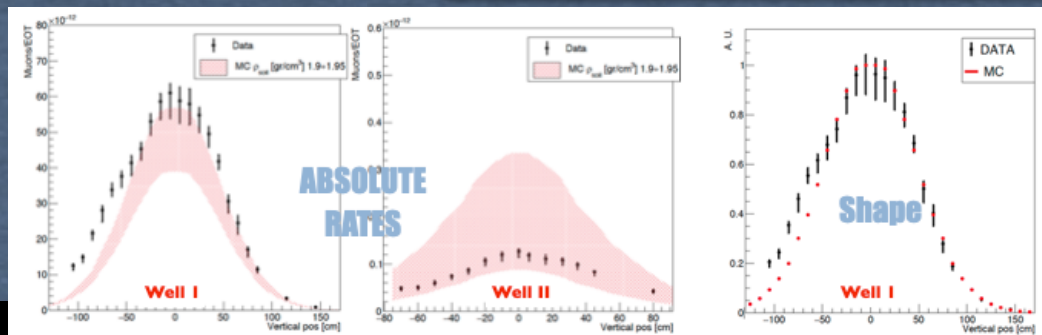
BD muons bg measurement



BDX-HODO



M. Battaglieri et al. NIM.A, 925:116-122, 2019



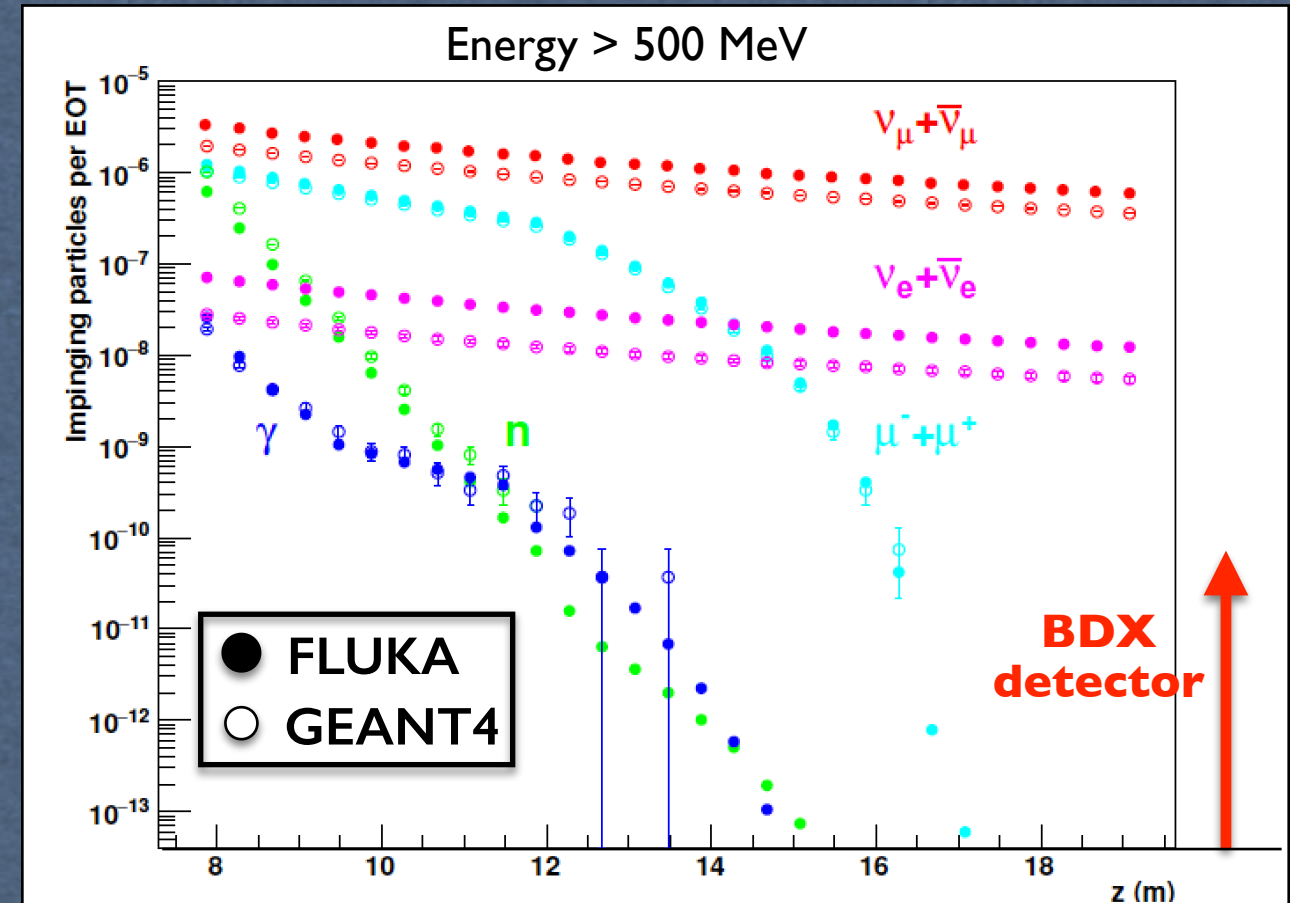
Cosmic muons

- Rate ($E_{Thr} \sim 300 \text{ MeV}$) < 2 counts
- Conservatively extrapolated from the (lower E) non-0 counts region
- Measured rate scaled to the JLab set-up (x800 crystals)



BDX prototype

- ★ No μ , n and γ with $E > 100 \text{ MeV}$ are found at the detector location
- ★ Validated by an in-situ measurement using BDX-HODO



★ Neutrino

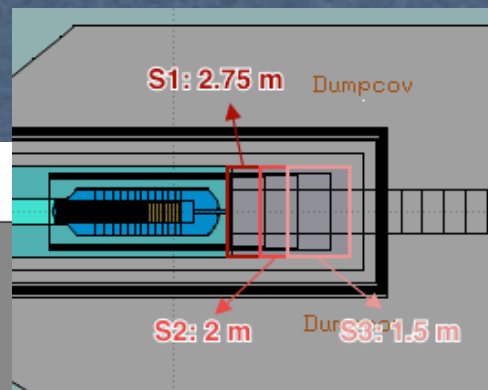
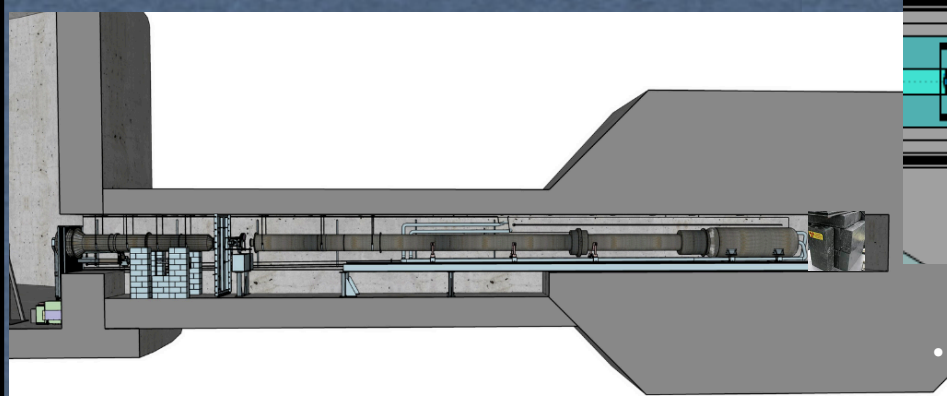
- $\pi \rightarrow \mu \nu_\mu$ $\mu \rightarrow e \nu_\mu \nu_e$
- Mainly low energy ($< 60 \text{ MeV}$) from decay at rest
- Some ν produced in HadShower and boosted to BDX detector

Non-negligible contribution by ν CC interaction

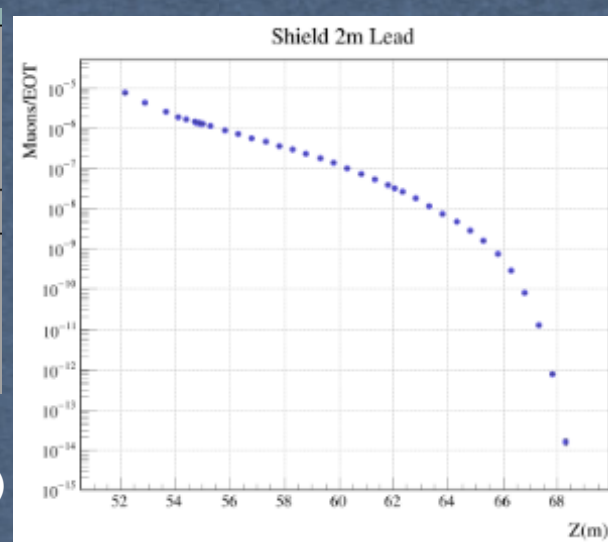
$\nu + N \rightarrow X + e^-$
 $\nu + e^-$ suppressed

Neutrino irreducible bg represents the ultimate limitation for BDX (~10 counts in BDX life time)

BD muons bg simulations



• lead wall (1.5m - 2.0m)



BDX sensitivity

Beam time request (parasitic to Hall-A ops)

- 10^{22} EOT (65 uA for 285 days)
- BDX can run parasitically with any Hall-A $E_{\text{beam}} > 10$ GeV experiments (e.g. Moeller)

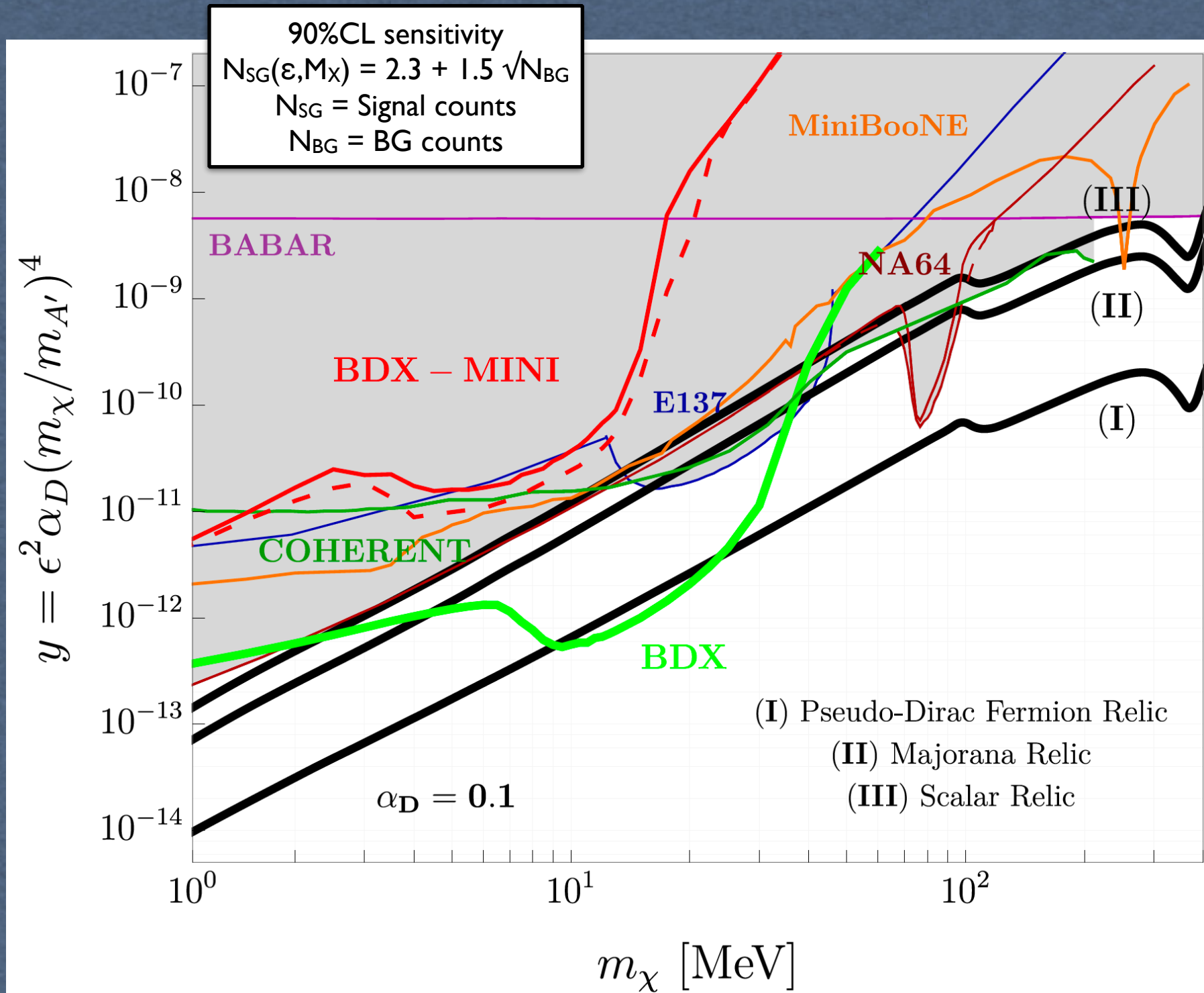
Beam-related background

Energy threshold	N_v (285 days)
300 MeV	~10 counts

Cosmic background

Energy threshold	\sqrt{Bg} (285 days)
300 MeV	<2 counts

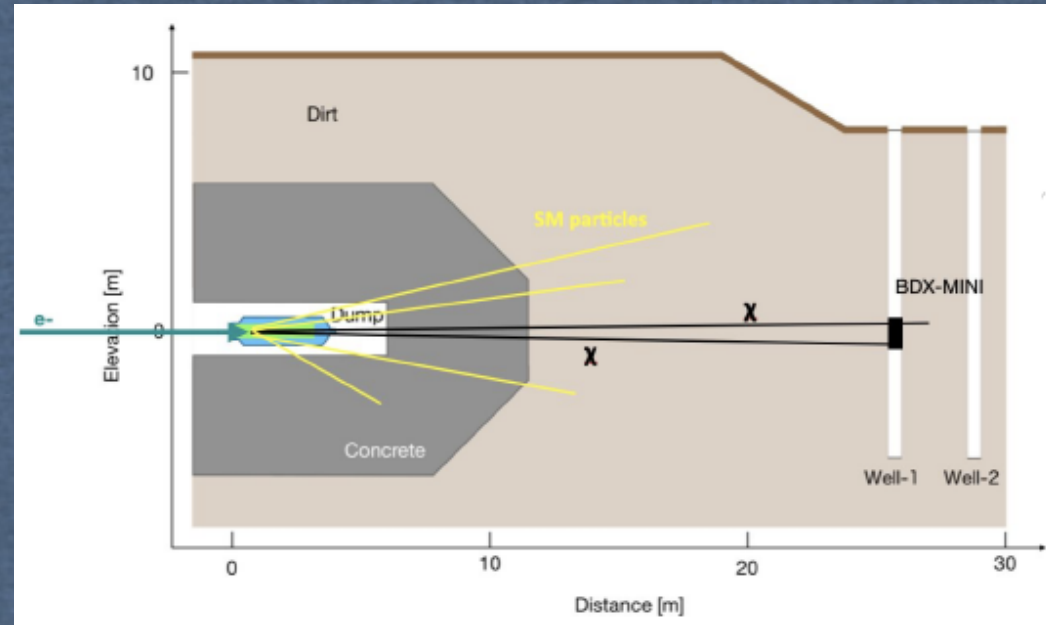
- Calculation includes resonant positron annihilation
- Sensitivity to inelastic LDM is not shown



**The sensitivity of BDX exceeds more than 10x the existing limits on LDM production
 Such tight exclusions will set limits on LDM mechanisms or render an important null result**

BDX-MINI @JLab

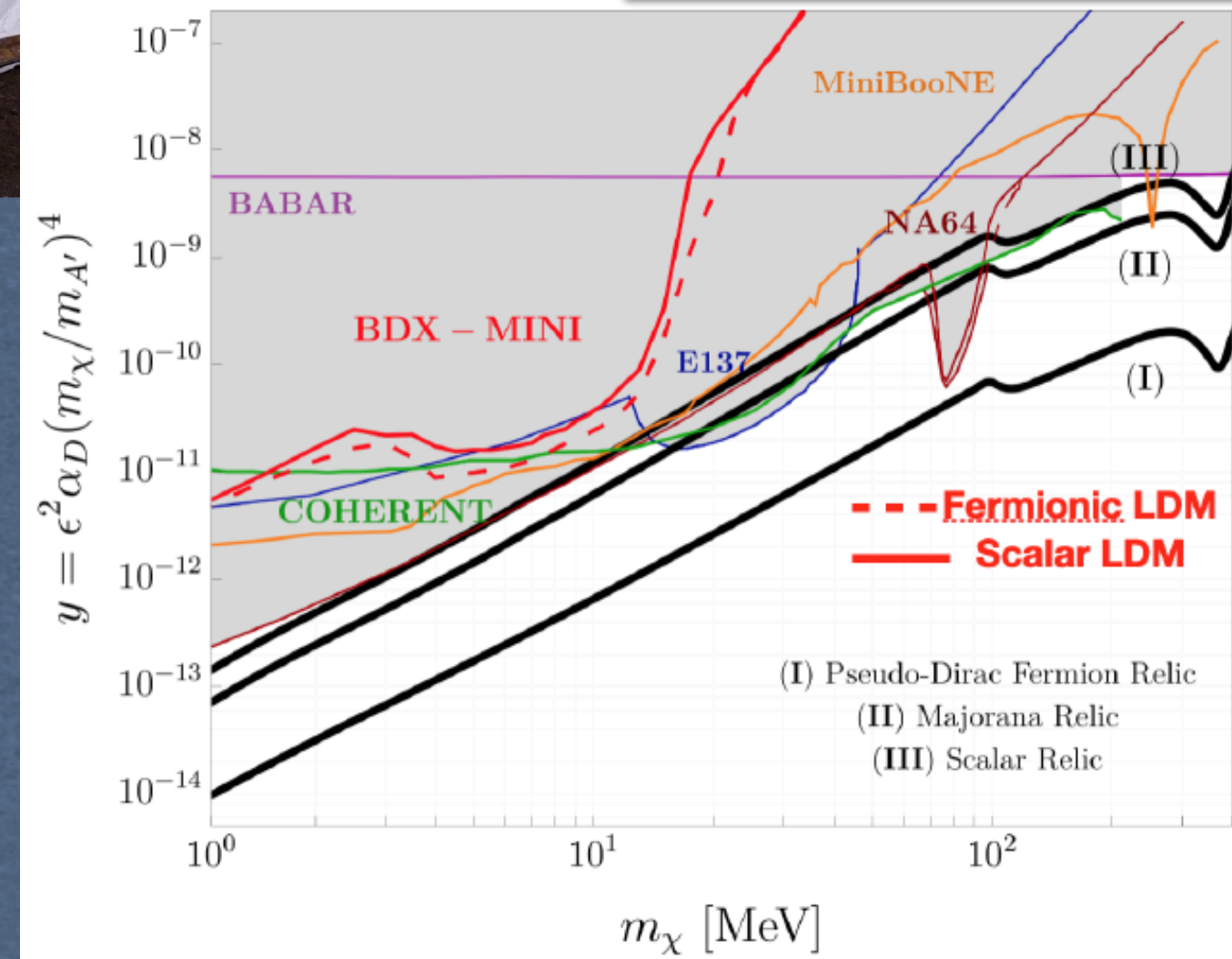
BDX-MINI: pilot experiment to prove the validity and feasibility of the BDX experiment



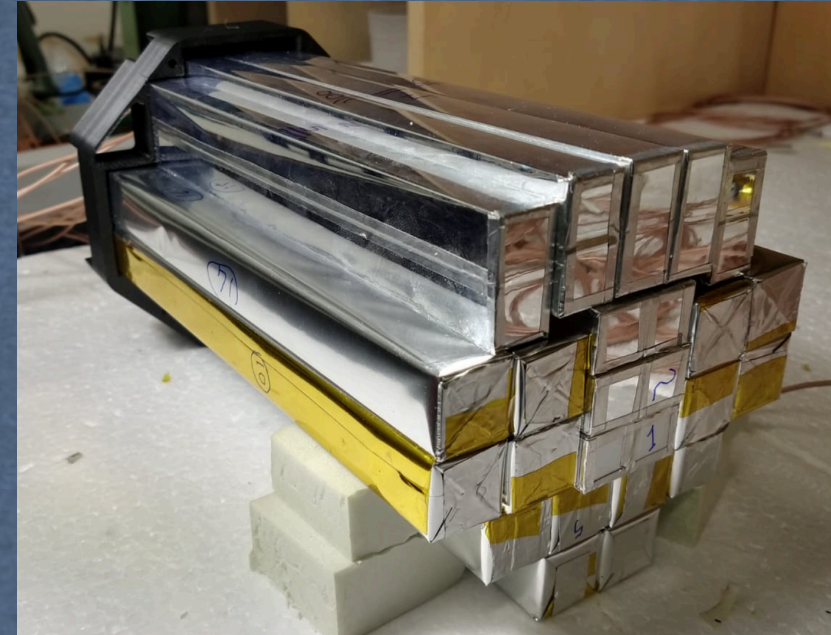
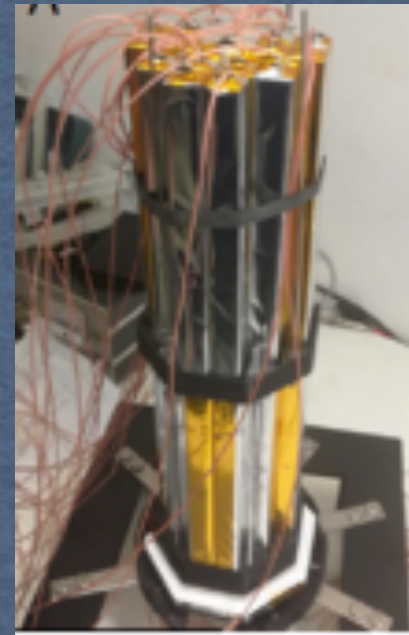
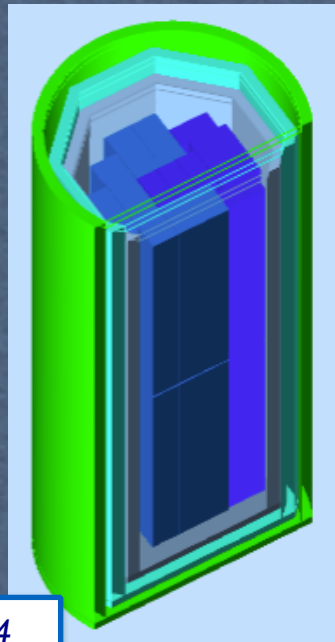
- Two wells dug for bg muon tests
- $E_{\text{beam}}=2.2$ GeV, no muons
- Limited reach but first physics result!

- Installed in March 2019
- Run from Dec 2019 to Aug 2020
- Collected $4e21$ EOT (40% BDX!) in ~ 4 months (+ cosmics)
- Good detector performance with high duty factor

M.Battaglieri et al. *Phys. Rev. D* 106, 072011



- 44 PbWO4 PANDA/FT-Cal crystals ($\sim 1\%$ BDX active volume)
- 6×6 mm² SiPM readout
- 2 active plastic scintillator vetos: cylindrical and octagonal (8 sipm each) + 2x lids + Passive W shielding



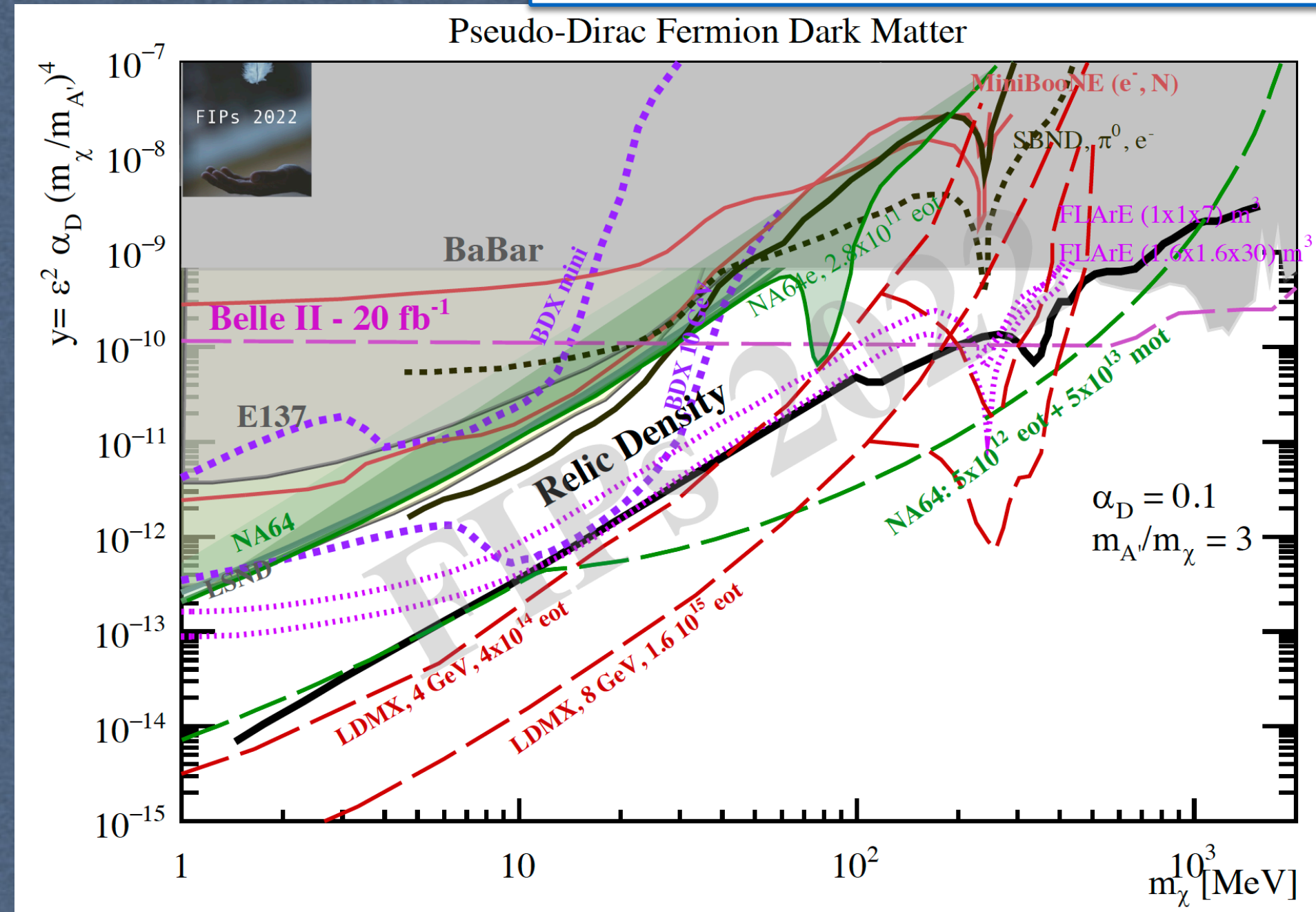
- Data-taking completed, analysis completed
- Results provide exclusion limits similar to the best existing experiments (E137, NA64, BaBar, ...)

M.Battaglieri et al. EPJC (2021) 81:164

LDM searches status

L.G. Krnjaic et al. Snowmass 2021 Rare & Precision Frontier (RF6): 7 2022
 C.Antel et al. Feebly interacting particles: Fips 2022 workshop report, 2023

- CERN-FIPs workshop report, (SNOWMASS22) shows that BDX offers great potential for discovering LDM (or disproving it, in a large area of parameter space)
- Several experiments planned/proposed (LHC, SLAC, Mainz, FNAL, KEK, PSI, LPARC) with a variety of beams (proton, leptons, photons), energies (from 150 MeV to 14 TeV) and experimental techniques (visible, invisible, recoil, ..) with a timeline that reaches ~2042
- BDX has the unique opportunity to be one of the first (significant) experiments to set limits and verify/falsify viable theoretical scenarios (relic LDM, inelastic DM, ...)
- Over the last 5 years, NA64 has pushed the exclusion limit slightly down (it is scheduled to run more!)



Conclusions

- * The existence of Dark Matter is a compelling reason to investigate new forces and matter over a broad range of mass
- * Accelerator-based (Light)DM search provides a unique feature of distinguishing DM signal from any other cosmic anomalies or effects
- * An extensive program to explore the Dark Sector making use of the high-intensity ~ 11 GeV electron beam available at Jefferson Lab is in full swing
- * Discovery or decisive tests of the simplest scenarios will possible in the next 5-10 years!
- * BDX is an optimised beam-dump experiment ready to run at Jefferson lab
- * The BDX concept has been tested with several prototypes and dedicated measurement campaigns
- * The BDX-MINI successful pilot run has demonstrated the potential of the BDX (technique and physics reach)
- * Collecting 10^{22} EOT in 285 days of parasitic running (in parallel to Moeller?) at 11 GeV, the BDX experiment would be >10 times more sensitive than previous experiments