



Stuart Fegan University of York June 17th, 2024





## Outline

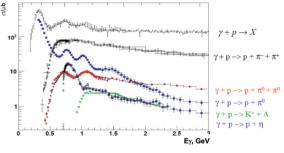
- 1 Introduction
  - A World of Polarisation (Observables)
  - JLab, CLAS and FROST
- 2 Analysis
  - Event Selection
  - Observable Extraction
  - Results
- 3 Outlook
  - Unfinished Business

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#### Motivation

- Baryon Spectroscopy is the study of excited nucleon states
- Finding some states can be difficult in a simple "bump hunt"; many are wide and overlap
- Use alternative means; coupling strength to a reaction channel, manifestation in experimental observables, etc. to aid searches



R. Beck and U. Thoma, EPJ Web Conf 134, 04003 (2017)

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# A World of Polarisation (Observables)

- Looking for polarisation observables on strangeness photoproduction
- Many possible channels, but this talk will focus on  $K^+\Lambda$

$$\gamma p \to K^+ \Lambda \to K^+ p \pi^-$$

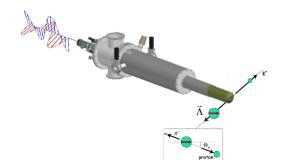
■ 16 observables for single meson photoproduction, arising from the scattering amplitudes of the interaction and the particles which carry polarisation

• "Single":  $\sigma$ ,  $\Sigma$ , P, T

■ Beam-Target: E, F, G, H

■ Beam-Recoil:  $O_X$ ,  $O_Z$ ,  $C_X$ ,  $C_Z$ 

■ Target-Recoil:  $T_X$ ,  $T_Z$ ,  $L_X$ ,  $L_Z$ 



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# PARA/PERP, Target too

- With a polarised beam and target, can access the single and beam-target double observables
  - Single:  $\sigma$ ,  $\Sigma$ , P, T
  - Beam-Target: E, F, G, H
- And more with recoil (i.e. with a self-analysing hyperon)



This work has two goals:

- Verify the beam asymmetry,  $\Sigma$
- Perform first measurements in this channel of the beam-target observable, G

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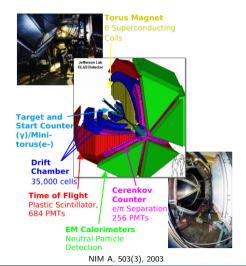
#### Jefferson Lab

- US DoE facility, Newport News, VA
- Superconducting RF accelerator electron beams up to 6 GeV, 12 GeV since a 2017 upgrade
- Three Experimental Halls, a Fourth added in the upgrade
- Tagged real photon beam facility available in the 6 GeV era in Hall B. using a secondary bremsstrahlung photon beam

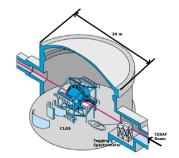


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#### Hall B and CLAS



- CEBAF Large Acceptance Spectrometer (1995-2012)
- Multi layered and segmented
- Toroidal magnetic field



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## g9a/FROST

- Data from g9a run period: November 2007 to February 2008
- Linearly and circularly polarised photon beams on a longitudinally polarised target
- Polarisation direction regularly flipped during run
- Nine coherent peak settings in linear polarisation, spanning energy range 0.7 to 2.3 GeV



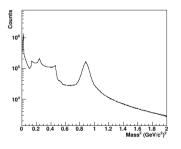
■ In this case, the reduced cross section can be expressed as:

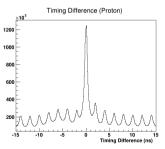
$$\frac{d\sigma}{d\Omega} = \sigma_0 \{1 - P_{lin} \Sigma cos(2\phi) + P_z(P_{lin} Gsin(2\phi))\}$$

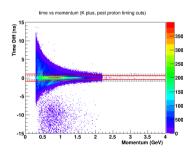
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# Analysis - Particle ID

- Initial particle ID via combination of charge and time-of-flight mass
- Select candidate Protons and Kaons
- Apply photon-to-particle timing difference cuts eliminate misidentification







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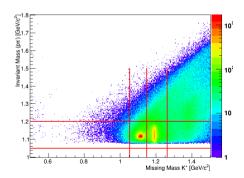


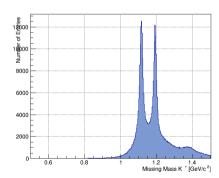
## Channel Identification

■ Select channel of interest:

$$\gamma p o K^+ \Lambda o K^+ p \pi^-$$

- Non exclusive selection reconstruct pion from detected proton and kaon
- Hyperons identified via kaon missing mass and proton pion invariant mass





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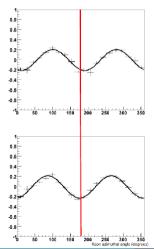


# Extracting Observables

- Binned fitting on asymmetries of two states of beam polarisation (PARA and PERP)
- Technique has been extensively employed in the JLab N\* program and similar experiments worldwide
- Recall that on a linpol beam and a longitudinally polarised target:

$$rac{d\sigma}{d\Omega} = \sigma_0 \{1 - P_{lin} \Sigma cos(2\phi) + P_z(P_{lin} Gsin(2\phi))\}$$

■ A  $cos(2\phi) + sin(2\phi)$  fit to a PARA/PERP asymmetry can be used to extract  $\Sigma$  and G for each state of target polarisation



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# Extracting Observables II

- Binned asymmetry fitting relies on a number of assumptions
- In lower statistics channels, or datasets where PARA and PERP have large variations between them in flux, polarisation, etc., reliable observable extraction is more challenging
- Using a maximum likelihood approach, we can extract observables event-by-event, and better account for variation between polarisation states
- We can define the likelihood function for each event as:

$$L_i = c_i [1 - P_{lin,i} \Sigma cos(2\phi_i) + P_{z,i} (P_{lin,i} Gsin(2\phi_i))] A$$

■ And extract observables by maximising the log-likelihood function:

$$\log L = b + \sum_{i} \log [1 - P_{lin,i} \Sigma cos(2\phi_i) + P_{z,i}(P_{lin,i}Gsin(2\phi_i))]$$

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# Dilution of Observables

- Parameters extracted from  $cos(2\phi) + sin(2\phi)$  fits are the free proton value, diluted with a carbon contribution (and beam and target polarisations)
- lacktriangle i.e. for the  $\Sigma$  observable, we actually measure  $\Sigma_{Butanol}$ , from which we can infer the free proton value

$$\Sigma_{Proton} = rac{1}{N_{Proton}} (N_{Butanol} \Sigma_{Butanol} - N_{Carbon} \Sigma_{Carbon})$$

■ For G, carbon in the target is unpolarised and we measure  $G_{Butanol}$ , estimating the free proton value via;

$$G_{Proton} = \frac{N_{Butanol}}{N_{Proton}} (N_{Butanol} G_{Butanol})$$

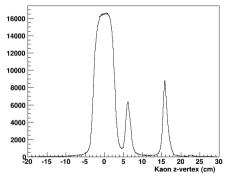
- The 'N' terms represent event yields per bin corresponding to the relevant material
- These must be estimated for Carbon and Proton...

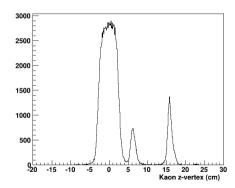
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# Target Selection

- FROST target contains Butanol (left), Carbon (centre) and Polythene (right)
- Resolvable from Kaon z-vertex after particle and channel identification



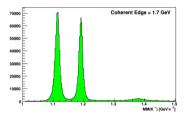


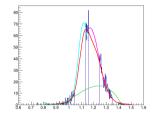
 Only Butanol is polarised, other targets used to account for nuclear backround and dilution effects of unpolarised nuclei in butanol

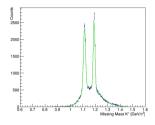
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# Estimating Carbon







- Controlling systematic uncertainties, particularly on a measurement of G, requires a robust method of accounting for Carbon
- We know from data on proton targets (left) that shape under the hyperon peaks on butanol (right) is almost entirely from bound nucleon effects
- Parameterise, use to estimate amount of Carbon events in each bin on Butanol

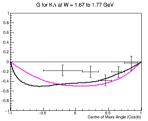
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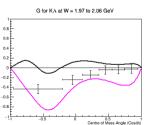


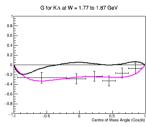
- CLAS-approved results from the maximum likelihood technique for the G observable on  $K^+\Lambda$
- Comparison to Bonn-Gatchina (pink line) and Jülich Bonn (black line) model predictions (fron 2017)
- Actively seeking new fits

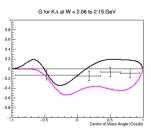
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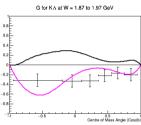
### Results for G

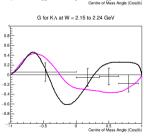












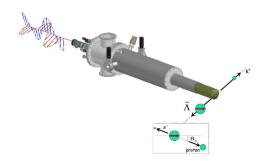


# The Future: Target-Recoil Observables

The FROST polarised target also enables access to target-recoil double observables:

$$\frac{d\sigma}{d\Omega} = \sigma_0 \{ 1 + P_{y'}P + P_x(P_{x'}T_{x'} + P_{z'}T_{z'}) + P_y(T + P_{y'}\Sigma) - P_z(P_{x'}L_{x'} + P_{z'}L_{z'}) \}$$

- Polarisation of recoiling hyperon can be inferred from decay angle distribution
- Can exploit this to measure beam-target observables  $T_X$ ,  $T_Z$ ,  $L_X$ ,  $L_Z$
- Requires evaluation of detector acceptance



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Data Lab

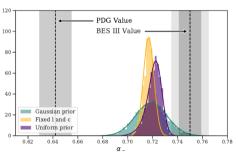


## Identities and Interpretations

$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	Observable	W (MeV)	$\theta$ (deg)	Data	Lab
$d\sigma/d\Omega$ 1617-2290 32-148 920 CEBAF d $\sigma/d\Omega$ 1617-2198 26-154 90 ELSA 1617-2198 26-154 90 ELSA 1625-2395 27-154 1674 CEBAF 1625-2333 26-143 1377 CEBAF 1649-1906 31-144 66 GRAAL 721-2190 37-134 314 CEBAF 1649-230 13-49 45 Spring-8 1646-2280 13-49 30 Spring-8 1646-2280 13-49 30 Spring-8 1646-2280 13-49 30 Spring-8 1649-2906 31-144 66 GRAAL 1721-2180 37-134 314 CEBAF 1649-2906 31-144 66 GRAAL 1721-2180 37-134 314 CEBAF 1649-2906 31-144 66 GRAAL 1721-2180 37-134 314 CEBAF 1649-2906 31-144 CEBAF 1649-2906 31-144 66 GRAAL 1721-2180 37-134 314 CEBAF 1649-2906 31-144 66 GRAAL 1649-2906 31-144		1610-2390	18-162	701	ELSA
do   dΩ		1612-1896	66-143	1306	MAMI
1625-2295   27-154   1674   CEBAF     1628-2533   26-143   1377   CEBAF     1934-2310   13-41   78   Spring-8     1649-1906   31-144   66   GRAAL     1721-2180   37-134   314   CEBAF     1946-2300   13-19   45   Spring-8     1946-2300   13-19   30   Spring-8     1946-2300   13-19   30   Spring-8     1941-2230   13-20   30   Spring-8     1647-2290   26-154   233   CEBAF     1649-1906   31-144   66   GRAAL     1669-22017   41-139   12   EISA     1669-22017   41-139   12   EISA     1669-22017   41-139   12   EISA     1721-2180   37-134   314   CEBAF     1649-1906   31-144   66   GRAAL     1721-2180   37-134   314   CEBAF     1649-1906   31-144   66   CEBAF     1649-1906   31-144   66   GRAAL     1649-1906		1617-2290	32-148	920	CEBAF
1628-2533   26-143   1377   CEBAF	$d\sigma/d\Omega$	1617-2108	26-154	90	ELSA
1934-2310   13-41   78   Spring-8		1625 - 2395	27-154	1674	CEBAF
1649-1906   31-144   66 GRAAL   1721-2180   37-134   314 CEBAF   5 prings   8 prings   1946-2300   13-49   45 prings   8 prings   1946-2300   13-49   45 prings   8 prings   167-2290   26-154   233 CEBAF   1672-290   26-154   233 CEBAF   1699-1906   31-144   66 GRAAL   1660-2280   34-146   30 EISA   1721-2180   37-134   314 CEBAF   1699-1906   31-144   66 GRAAL   1721-2180   37-134   314 CEBAF   1691-1906   31-144   66 GRAAL   1		1628-2533	26-143	1377	CEBAF
1721-2180   37-134   314   CEBAF		1934-2310	13-41	78	Spring-8
Σ         1946-2300         13-19         45         Spring-8           1946-2220         13-19         30         Spring-8           1946-2220         13-19         30         Spring-8           1647-2200         26-145         233         CEDAF           1652-25245         26-143         1497         CEBAF           1669-2907         31-144         66         GRAAL           1696-2220         37-134         314         CEDAF           1721-2180         37-134         314         CEDAF           1691-1906         31-144         66         GRAAL           1721-2180         37-134         314         CEDAF           C <sub>e</sub> 1678-2454         32-139         144         CEDAF           C <sub>e</sub> 1678-2454         32-139         144         CEDAF           O <sub>e</sub> 1649-1906         31-144         66         GRAAL           O <sub>e</sub> 1619-1906         31-144         66         GRAAL	Σ	1649-1906	31-144	66	GRAAL
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$		1721-2180	37-134	314	CEBAF
2011-2238   18-32   4 Spring-8		1946-2300	13-49	45	Spring-8
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$		2041-2238	18-32	4	Spring-8
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		1617-2290	26-154	233	CEBAF
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		1625-2545	26-143	1497	CEBAF
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$C_{i'}$ 1678-2454 32-139 146 CEBAF 0 <sub>x</sub> 1699-1906 31-144 66 GRAAL 1721-2180 37-134 314 CEBAF 0 <sub>c</sub> 1649-1906 31-144 66 GRAAL		1721-2180	37-134	314	CEBAF
$O_x$ 1649-1906 31-144 66 GRAAL 1721-2180 37-134 314 CEBAF 0. 1649-1906 31-144 66 GRAAL	$C_{x'}$	1678-2454	32-139	144	CEBAF
O <sub>x</sub> 1721-2180 37-134 314 CEBAF 1649-1906 31-144 66 GRAAL	$C_{z'}$	1678-2454	32-139	146	CEBAF
0. 1649-1906 31-144 66 GRAAL	0	1649-1906	31-144	66	GRAAL
0-	$O_x$	1721-2180	37-134	314	CEBAF
1721-2180 37-134 314 CEBAF	0	1649-1906	31-144	66	GRAAL
	O <sub>2</sub>	1721-2180	37-134	314	CEBAF

W (MoV)

- Polarisation observable data is relatively sparse in  $K^+\Lambda$
- Possible to verify the weak decay parameter,  $\alpha_-$ , using observables and Fierz identities



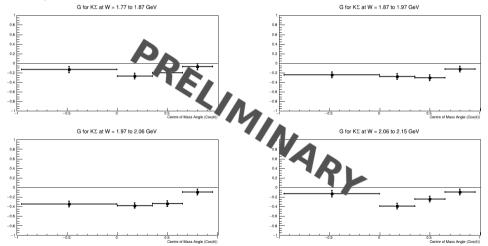
D.G. Ireland et. al., Phys. Rev. Lett. 123, 182301 (2019)

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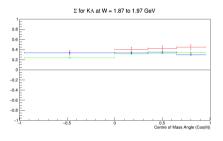


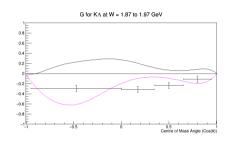
## Other Channels

#### Preliminary Measurements of G on $K^+\Sigma$



### Outlook and Plans





- First measurement of the G observable for  $K^+\Lambda$  photoproduction using maximum likelihood technique on this data
- Results approved by CLAS, seeking new model fits
- Data allows  $K^+\Lambda$  and  $K^+\Sigma$  channels to be studied
- Measuring target-recoil observables may be feasible using this method

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