

July 24—28, 2023, Jefferson Lab PAC51, Newport News, VA, US

# JLab E12-15-008

An isospin dependence study of the  $\Lambda$ -N interaction through the high precision spectroscopy of  $\Lambda$  hypernuclei with electron beam

*Contact Person:*

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*Presented by:*

**T. Gogami** (Kyoto University, Japan)

on behalf of **the JLab Hypernuclear Collaboration**

**July 27, 2023**

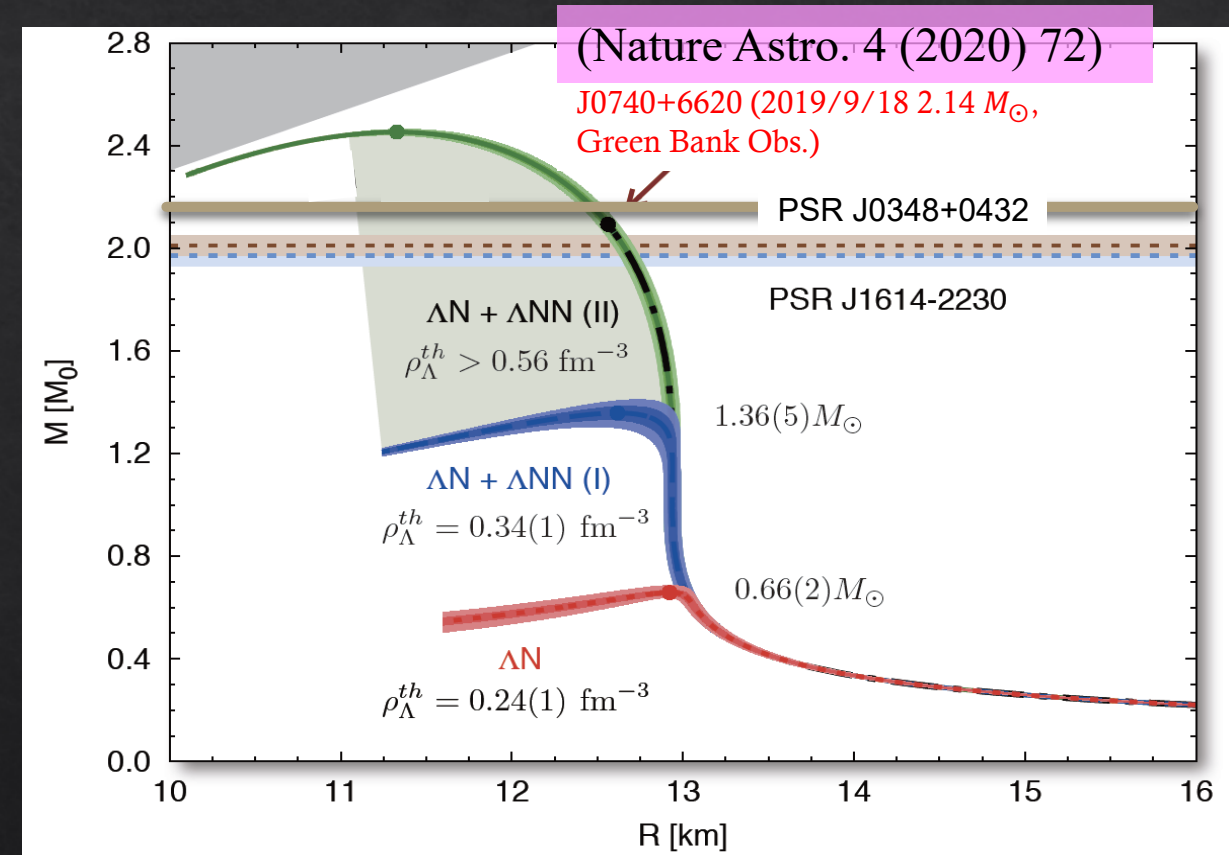
# One page summary of physics motivation

## ◆ EOSs with hyperons are too soft to sustain neutron stars with $\geq 2M_{\odot}$ (HYPERON PUZZLE)

- Repulsive forces by  $\Lambda$  NN?
- E12-15-008 is the first attempt to investigate the  $\Lambda$  NN force ( $|\Delta B_{\Lambda}| \leq 100$  keV)

## ◆ Isospin dependent $\Lambda$ N interaction

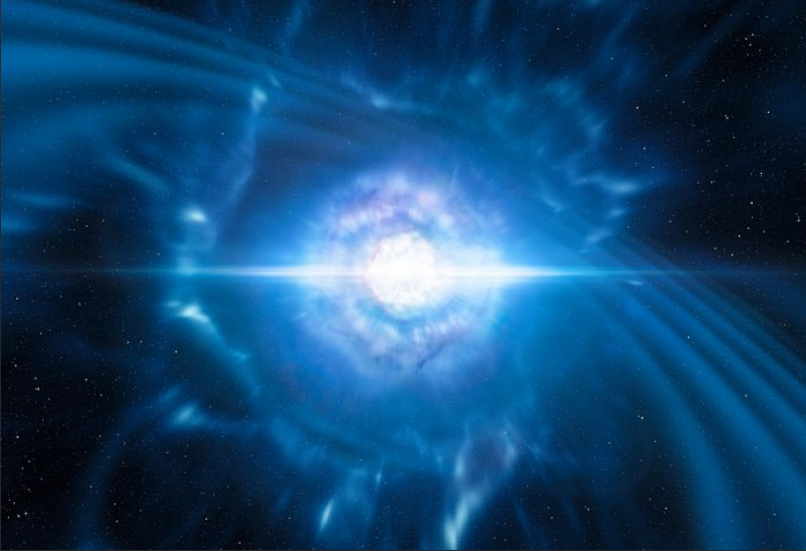
- ◆  $\Lambda$  N- $\Sigma$  N coupling
- ◆ Related to the  $\Lambda$  N charge symmetry breaking
- There is no data at medium heavy mass region
- E12-15-008 will provide the first data ( $Z = 19: A = 40$  vs. 48)



D. Lonardoni et al., *Phys. Rev. Lett.* 114, 092301 (2015)

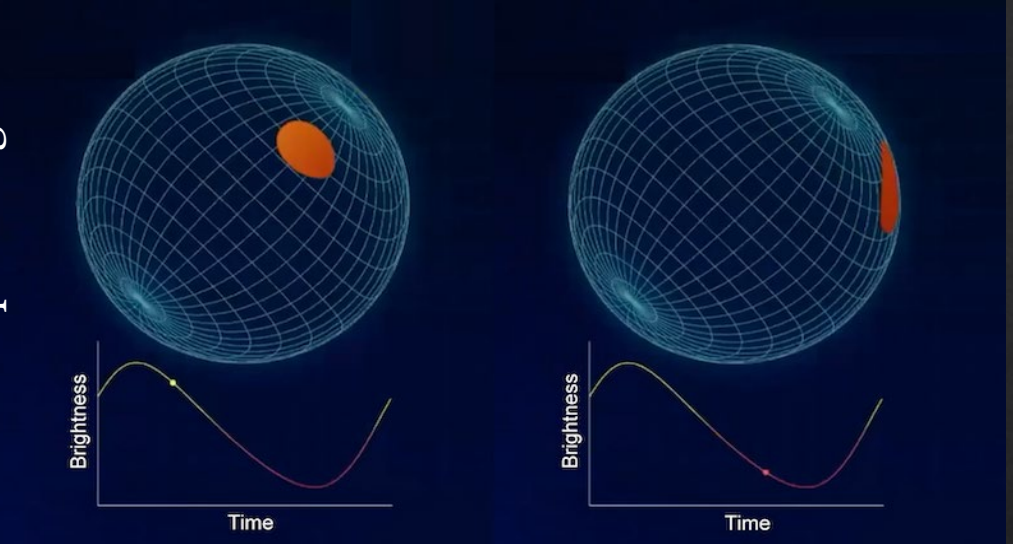
# New astronomical observations

CC4.0 ESO/L.  
Calçada/M. Kornmesser



Gravitation Wave from neutron star mergers  
LIGO/Virgo PRL 119, 161101 (2017)

Goddard Space Flight Center



NICER : NS x-ray hot spot measurement  
Physics 14, 64 (Apr. 29, 2021)

**Macroscopic** features of NS : Tidal deformabilities, masses and radii

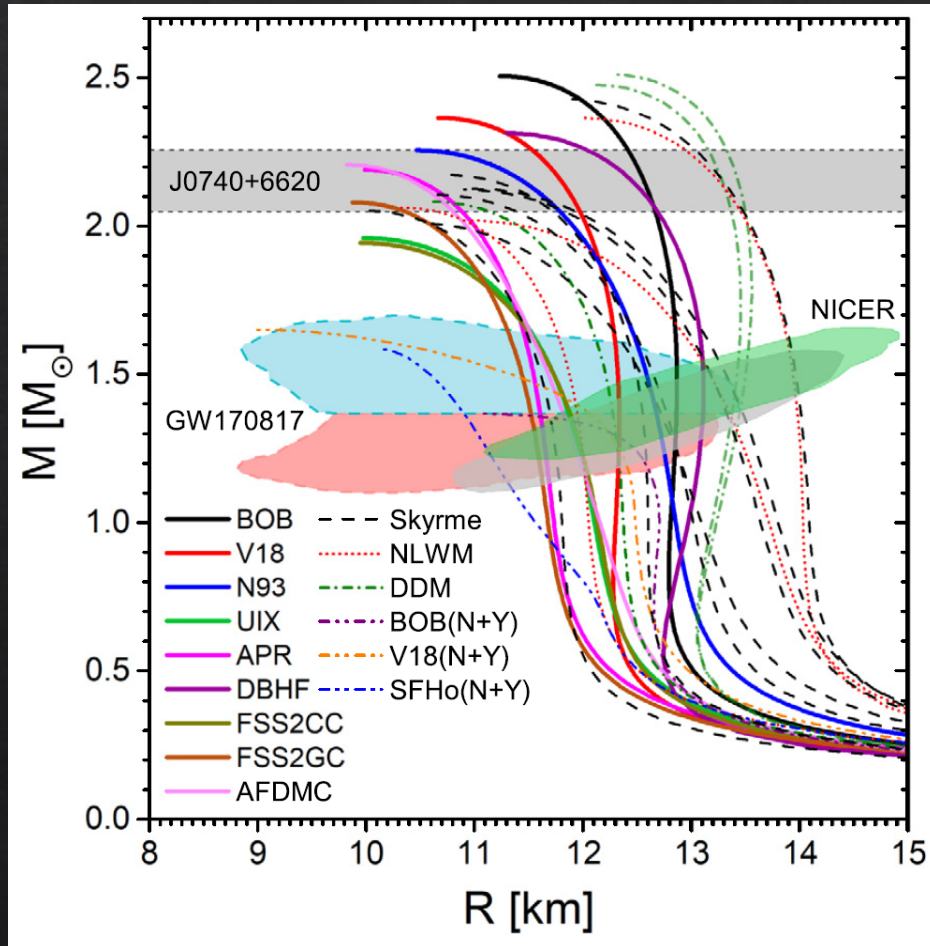
vs.

**Microscopic** investigation of NS: Inner composition →

**HYPERNUCLEAR  
SPECTROSCOPY**

# New constrains from astronomical observations

C.F. Burgio et al. Prog. Part. Nucl. Phys 120 (2021) 103879.



Macroscopic understanding of NS made great progress.

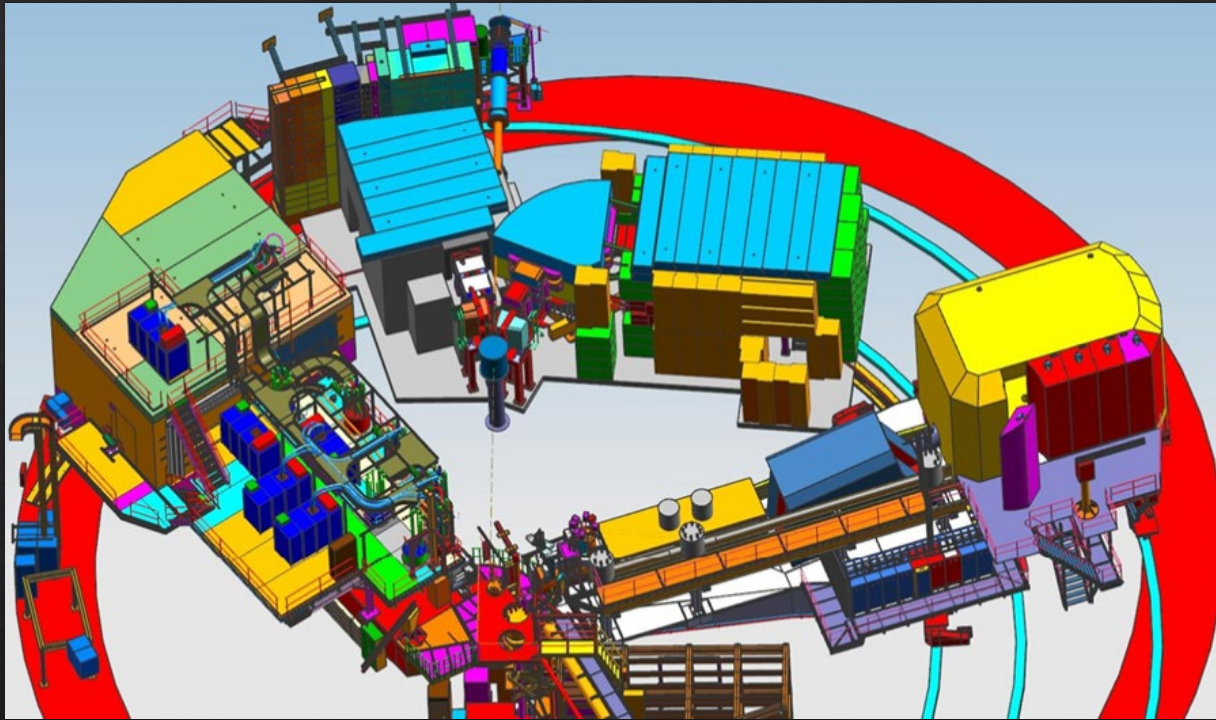
But, we need to know why NSs can be so heavy and large.

→ Microscopic study (nuclear physics exp) has become more important than ever!

# Next experiment at JLab Hall C (FY2026~)

5/10

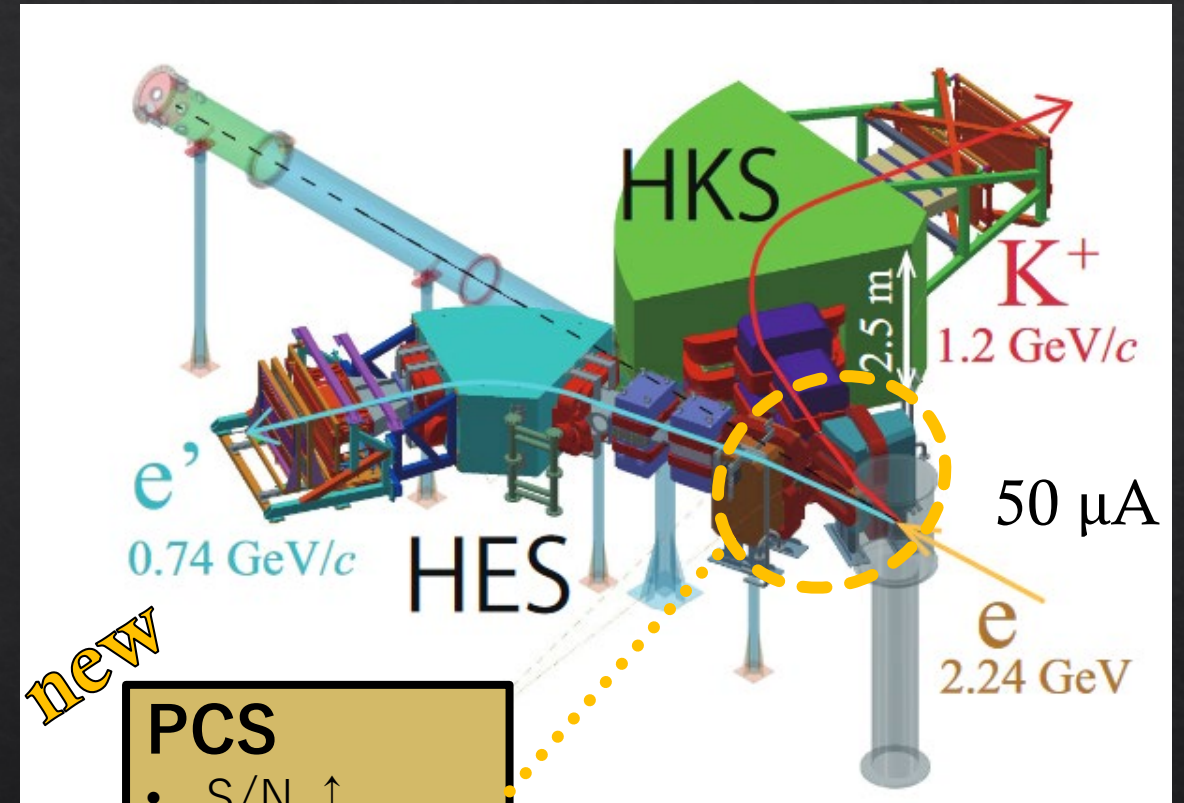
Hall A (original proposal) → Hall C



Steven Lassiter & Bert Metzger, JLab Hypernuclear Collaboration Meeting 2022, online, Dec 2022, [https://wiki.jlab.org/tegwiki/index.php/Hypernuclear\\_CollaborationMeeting\\_2022Dec](https://wiki.jlab.org/tegwiki/index.php/Hypernuclear_CollaborationMeeting_2022Dec)

Schematic from LOI12-23-013 (TG et al., LoI to PAC51):

[https://researchmap.jp/gogami/published\\_papers/42361620/attachment\\_file.pdf](https://researchmap.jp/gogami/published_papers/42361620/attachment_file.pdf)



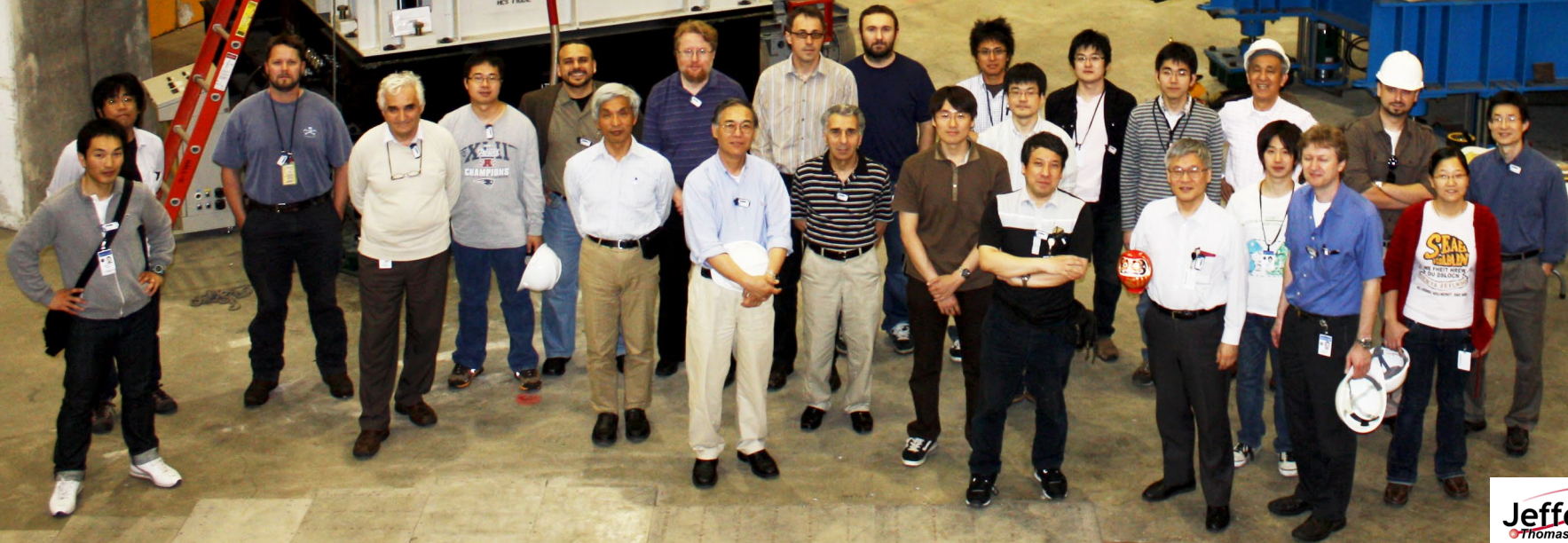
**PCS**

- S/N  $\uparrow$
- Flexibility  $\uparrow$

Existing spectrometers HES and HKS are going to be used

HES

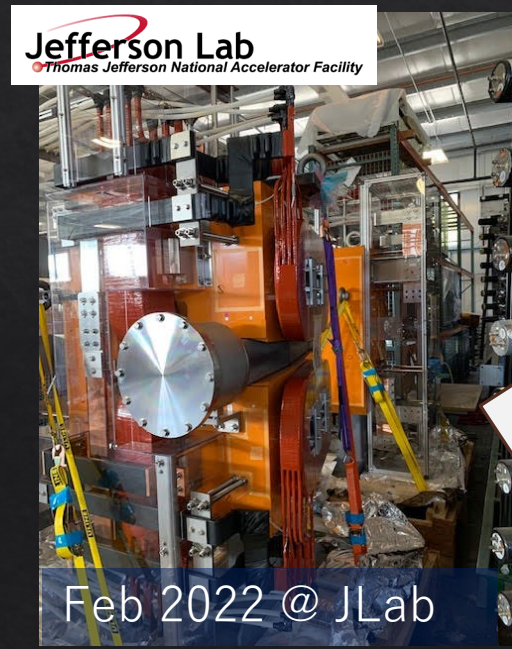
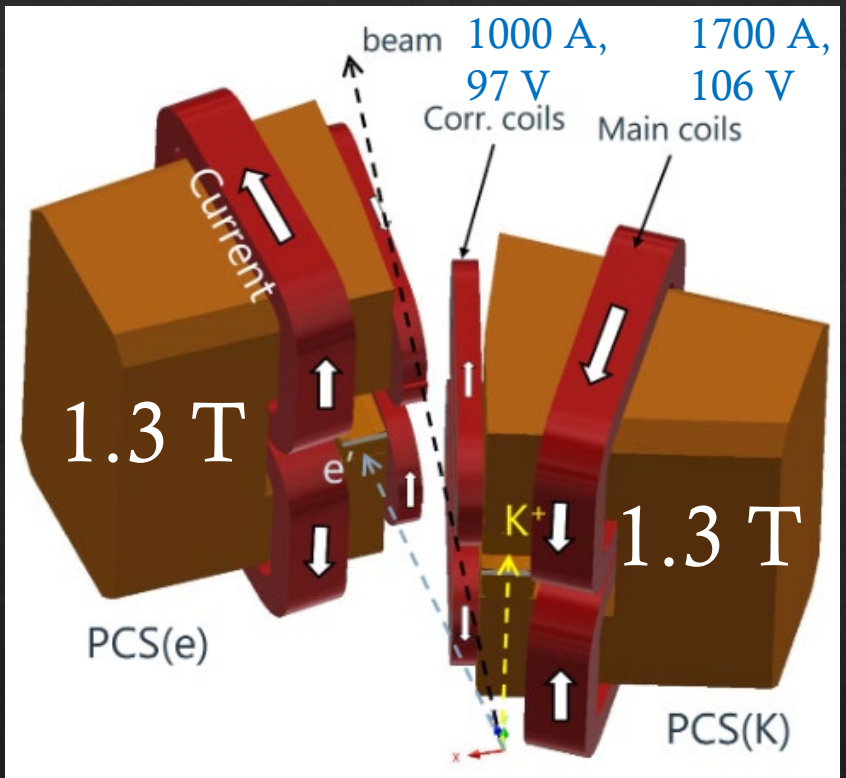
HKS



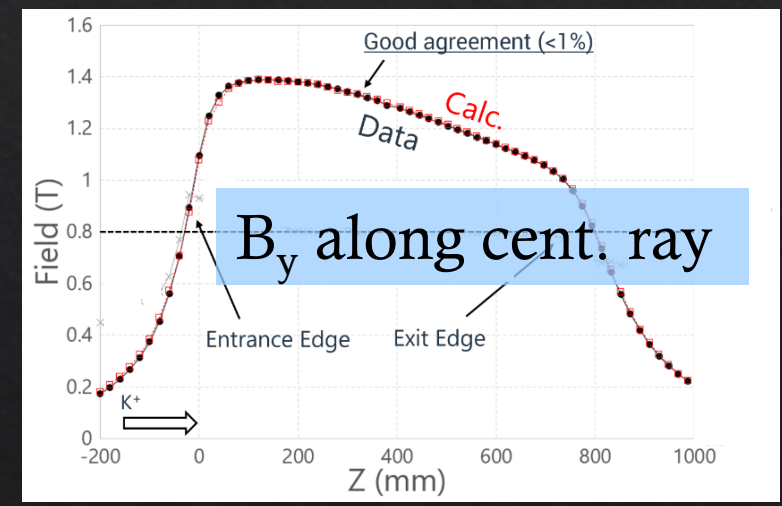
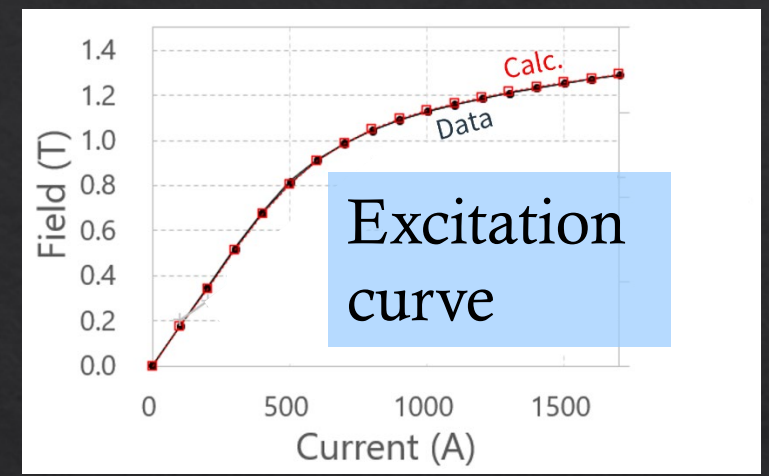
normal conduction

# New magnet PCS

(Pair of charge separation dipole magnets)



Measurement vs. TOSCA calc.



## It worked as expected !!

科研費  
KAKENHI

Supported by KAKENHI:  
No. 17H01121 (33,500,000 JPY)  
No. 18H05459 (144,100,000 JPY)

### > \$ 1 M

# Beam time request (E12-15-008)

Table 6: Updated request of beamtime.

Target (Hyper Nucleus)	Beam current ( $\mu\text{A}$ )	Target thickness ( $\text{mg}/\text{cm}^2$ )	Assumed cross section ( $\text{nb}/\text{sr}$ )	Expected yield (/h)	Num. of events	Req. beamtime (hours)	B.G. rate (/MeV/h)	S/N	Comments
$\text{CH}_2$ ( $\Lambda, \Sigma^0$ )	2	500	1000	8.62	1000	120	0.03	290	Calibration
${}^6\text{Li}$ ( ${}^6_\Lambda\text{He}$ )	50	100	10	—	—	—	—	—	Separate LoI
${}^9\text{Be}$ ( ${}^9_\Lambda\text{Li}$ )	50	100	10	—	—	—	—	—	Separate LoI
${}^{11}\text{B}$ ( ${}^{11}_\Lambda\text{Be}$ )	50	100	30	—	—	—	—	—	Separate LoI
${}^{12}\text{C}$ ( ${}^{12}_\Lambda\text{B}$ )	50	150	90	6.79	1100	168	1.20	5.67	Calibration
${}^{27}\text{Al}$ ( ${}^{27}_\Lambda\text{Mg}$ )	50	150	60 *	1.98	330	168	1.77	1.87	Calibration
Subtotal						456			Calibration
${}^{40}\text{Ca}$ ( ${}^{40}_\Lambda\text{K}$ )	50	150	50	1.13	520	456	2.41	0.47	Physics
${}^{48}\text{Ca}$ ( ${}^{48}_\Lambda\text{K}$ )	50	150	50	0.94	520	552	1.89	0.50	Physics
Subtotal						1008			Physics
Total						1464			

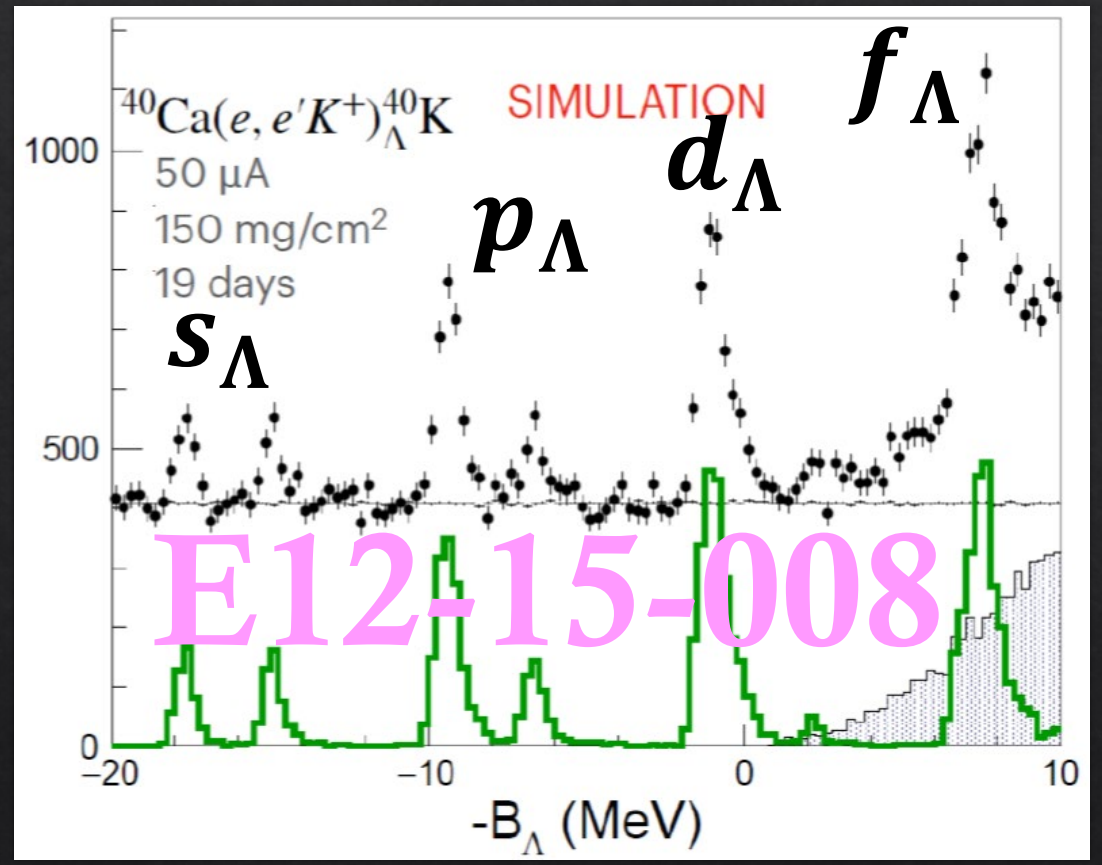
\* for  $0s^\Lambda$   $9/2^+$ ,  $7/2^+$  doublet.

(=61 PAC days)



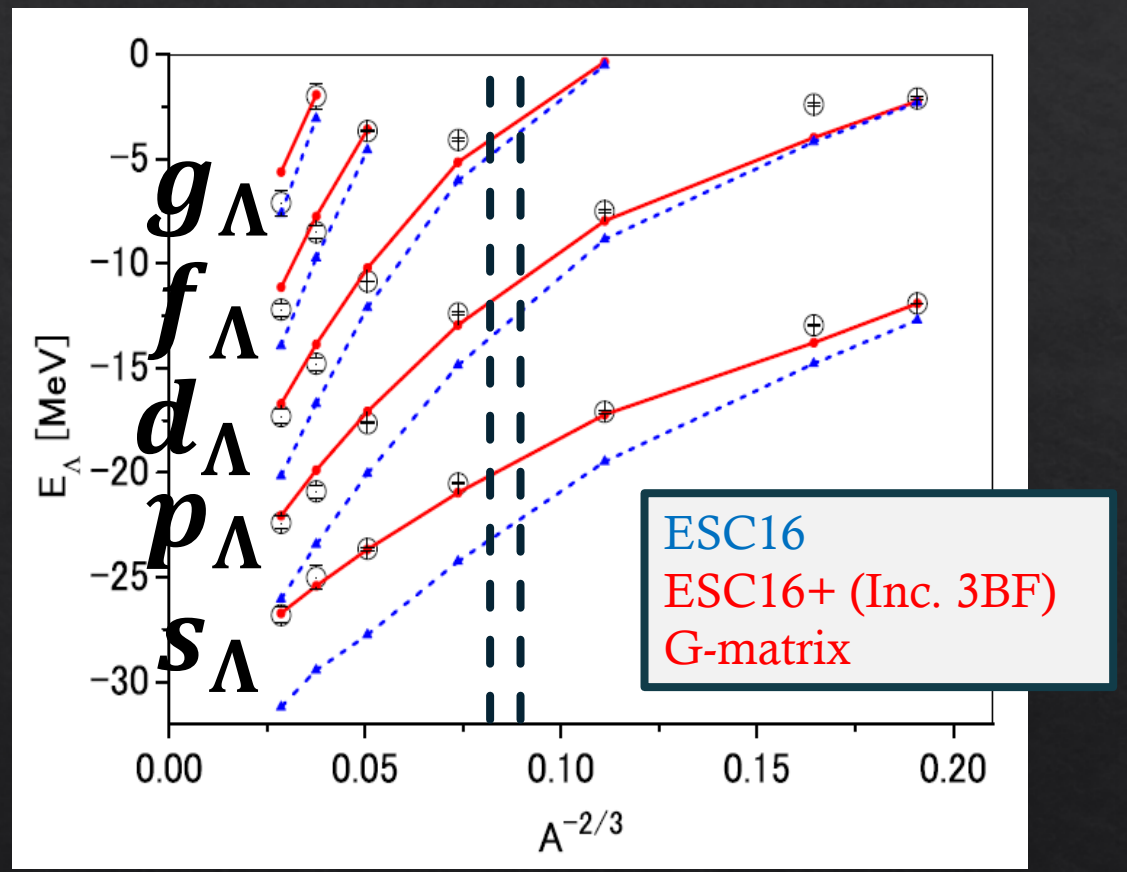
# High accuracy experiment → 3-body force study

Expected spectrum based on Geant4 simulation



Missing mass spectroscopy with the world best accuracy  $|\Delta B_\Lambda| \leq 100$  keV

M.M. Nagels et al., PRC 99 (2019) 044003.



Clearer understanding the 3BF

# Summary (JLab E12-15-008)

10/10

Physics motivation → **E12-15-008 is unique, and the importance has been more significant**

◆ **EOSs with hyperons are too soft to sustain neutron stars with  $\geq 2M_{\odot}$  (HYPERON PUZZLE)**

→ Repulsive forces by  $\Lambda$  NN?

→ **E12-15-008 is the first attempt to investigate the  $\Lambda$  NN force**

(Importance of the microscopic study of NS is more significant due to the recent progress of astronomical observations)

◆ **Isospin dependent  $\Lambda$  N interaction**

◆  $\Lambda$  N- $\Sigma$  N coupling

◆ Related to the  $\Lambda$  N charge symmetry breaking

→ There are no data at medium heavy mass region

→ **E12-15-008 will provide the first data**

Preparation status and outlook

◆ All spectrometer systems (HKS+HES+PCS) are ready

◆ Detectors are being tested w/ cosmic rays

◆ We plan ERR in the beginning of 2024

◆ Long stay of Japanese members is planned to be from sometime in 2024

◆ We aim to start the experiment in 2026



**We thank the JLab staff for the great support**

# Q&A

# Requested 4 points to be addressed by jeopardy presentation

1. Is there any new information that would affect the scientific importance or impact of the experiment since it was originally proposed?
  - **New Gravitational Waves from NS mergers and NICER : Macroscopic features of NS**
  - **Microscopic understanding of NS has become more important!**
2. If the experiment has already received a portion of its allocated beam time, the spokespersons should present the status of the analysis of the existing data and the projected result for the final complete data set. The goal is to show the physics impact of the beam time requested in the jeopardy update.
  - **Not applicable**

# Requested 4 points to be addressed by jeopardy presentation

3. What is the status of the collaboration in terms of institutes, committed staff, and prospective students?

Currently, core members of JLab hypernuclear collaboration working on optimization of experiment design, commission/development of magnets, spectrometer.

Ten faculties of U Tokyo, Kyoto U., Tohoku U., Hampton U., Florida Int. U., INFN/Rome, Virginia M. Inst. Four Ph-D students and four master's students are working on the projects. One – two years before the beamtime, Japanese group members start to stay at JLab. More members of Hall A/C collaboration will participate in preparation works and taking shifts. (Hypernuclear workshop was organized at JLab in March, 2023 to share interests with potential collaborators at JLab. )

4. Should the remaining beam time allocation and experiment grade be reconsidered?

- **Scientific significance is more than the original E12-15-008 (Grade A).**
- **Relocation of the experimental from Hall-A to Hall-C makes longer beamtime.**

# Answer to questions from Prof. Sawada (1)

Q1) Because several years have passed since the original proposal was approved and please show theoretical and experimental progress on the subject during the period.

A1) Hyperon puzzle was widely recognized and it is general consensus that some mechanism needs to make the EoS of massive neutron stars stiffer. In addition to AFDMC, various theoretical calculations based on different potential models with the three-body repulsive LNN force were performed for hypernuclei.

**Theoretical studies made significant progresses.**

- Meson exch. Potential (Nijmegen) + G-matrix M.M. Nagels et al., PRC 99 (2019) 044003.
- Chiral EFT + G-matrix J.Haidenbauer, I.Vidana, EPJA (2020) 56:55.
- Hyper AMD M.Isaka et al., PRC94, 044310 (2016), PRC 95, 044308 (2017)

Experimentally, **Charge Symmetry Breaking study made great progress for light ( $A=4$ ) hypernuclei**, and large CSB was observed for  $^4_{\Lambda}\text{H}$  and  $^4_{\Lambda}\text{He}$  at MAMI and J-PARC. But there is still **no CSB study or iso-spin dependence study for medium-heavy hypernuclei**. At J-PARC,  $\Sigma N$  scattering experiment was successfully performed and it becomes possible to obtain direct information about the hyperon-nucleon two-body force. Femtoscopy study also give important Information about two-body baryonic force. **A better understanding of the two-body baryon force is important to a more reliable analysis of the  $\Lambda NN$  three-body force. Therefore, these progresses will help our analysis.**

# Answer to questions from Prof. Sawada (2)

Q2) You claim that from the viewpoint of international competition this experiment should be implemented early. Please explain the international situation.

A2) At **J-PARC**, hadron experimental facility extension project (HEF-ex) has been discussed seriously and KEK Project Implementation Plan (PIP 2022) awarded the first priority to the project. Hopefully the project will be budgeted soon. Precise spectroscopy of hypernuclei with the  $(\pi^+, K^+)$  reaction is one of flagship experiments at J-PARC HEF-ex.  $(\pi^+, K^+)$  and  $(e, e'K^+)$  reactions produce iso-spin partner hypernuclei and thus these studies are **complementary** and **both of them are quite important to study Charge Symmetry Breaking** of hypernuclei with a sub-MeV resolution. It is important to establish **high-precision spectroscopy techniques ( $<0.5$  MeV, FWHM resolution) for medium-heavy hypernuclei under high counting rate** at JLab before J-PARC HEF-ex. Experiences at JLab will significantly contribute to the hypernuclear study at J-PARC HEF-ex project.

At **MAMI**, high resolution spectroscopic study with decay p spectroscopy is on going. It is currently limited to  $A=3,4$  hypernuclei, but it would be possible to extend the study at JLab as Lol12-23-011 proposes.

At **GSI**, WASA-FRS Hyp-HI is exploring study of light hypernuclei with heavy-ion beams and **FAIR** PANDA is aiming to study multi-L hypernuclei with anti-proton beams.

# Answer to questions from Prof. Sawada (3)

Q3) With the move from Hall-A to Hall-C of the experiment, we got a new TAC report on the experiment. If you have comments on the TAC report, please let us know.

A3) We would like to add comments to the TAC report.

As TAC recognizes, a cryogenic gas target need a vertical bending spectrometer and modification of HES to be a vertical one is not simple task.

Considering heavy load of design of new cryogenic gas target to the JLab target group and potential difficulties of the HES modification, we are concentration only on the solid targets including lead as the first campaign of the next JLab hypernuclear program. Solid targets with a movable ladder system were already used in the previous Hall-C hypernuclear experiments and it is not very complicated. We have been working on design of cryogenic target with the JLab target expert and continue to ask his help for the solid target design.

Possibility of  $B_4C$  target instead of a thin foil of Boron will be considered carefully. We need isotopically enriched boron target and availability of  $^{11}B_4C$ . We will evaluate which is better option to use  $^{11}B_4C$  or a  $^{11}B$  foil, considering effect of carbon data contamination the boron calibration data.



# Backup

## *Requested 4 points to be addressed by jeopardy presentation*

1) Is there any new information that would affect the scientific importance or impact of the experiment since it was originally proposed?

### **New astronomical observations**

PSR J1614 – 2230 ( $M = 1.928 \pm 0.017 M_{\odot}$ ) (Nature 467 (2010) 1081, APJ 832 (2016) 167)

PSR J0348 + 0432 ( $M = 2.01 \pm 0.04 M_{\odot}$ ) (Science 340 (2013) 1233232)

**PSR J0740 + 6620 ( $M = 2.14^{+0.10}_{-0.09} M_{\odot}$ )** (Nature Astro. 4 (2020) 72)

→ New observation of heaviest neutron star ever observed.

Existence of massive neutron stars ( $\geq 2 M_{\odot}$ ) is more certain!

# HYPERON Puzzle

Based on our knowledge of baryonic force,  
**Hyperon naturally appear at high density ( $\rho \sim 2,3\rho_0$ )**



**Too soft EOS. NS cannot support mass of  $2 M_\odot$**



**Contradict to astronomical observations.  
 (One more massive NS after PAC44)**



**Need additional repulsive force  
 ( $\Lambda NN$  3-body repulsive force)  
 Make stiffer EOS**

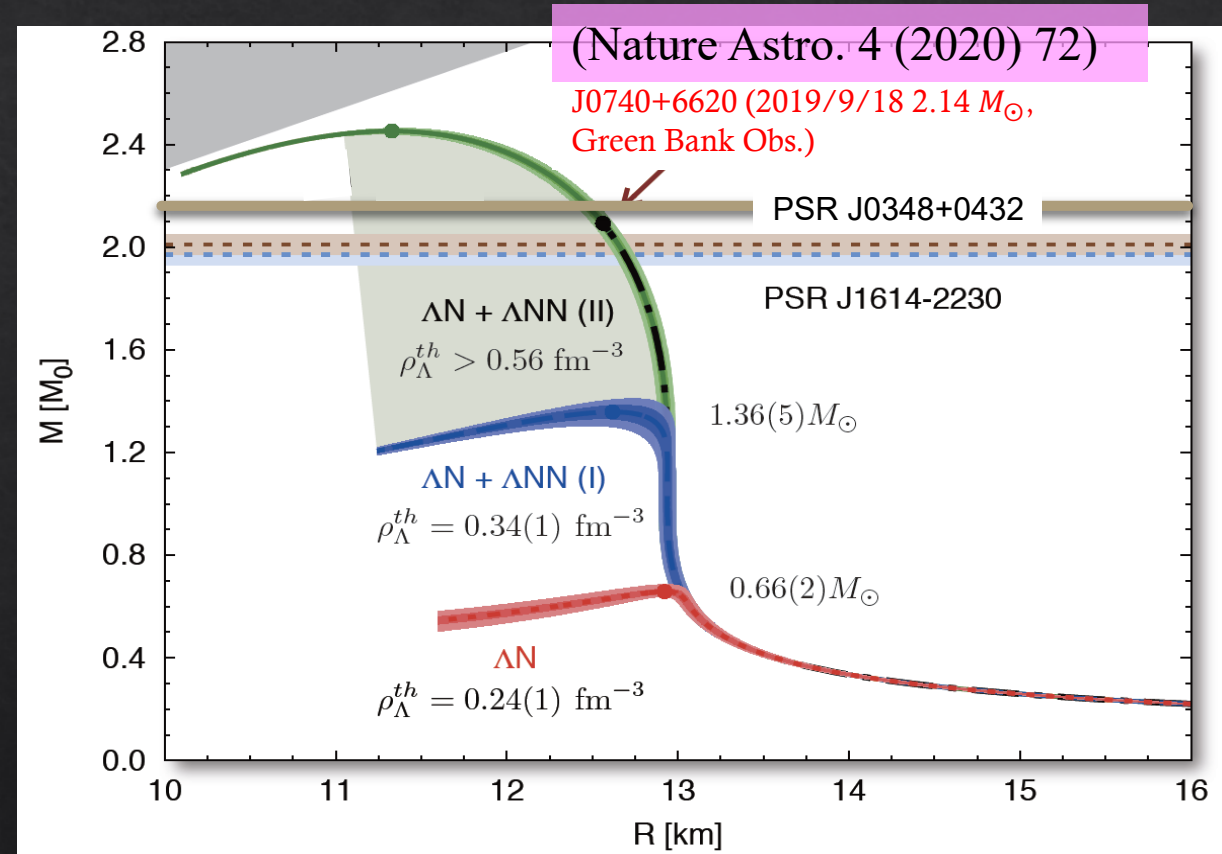


**Neutron star : Large  $(N - Z)/A \geq 0.9$  and Large A**

**Iso-spin dependence  
 E12-15-008**

A dependence  
 E12-20-013

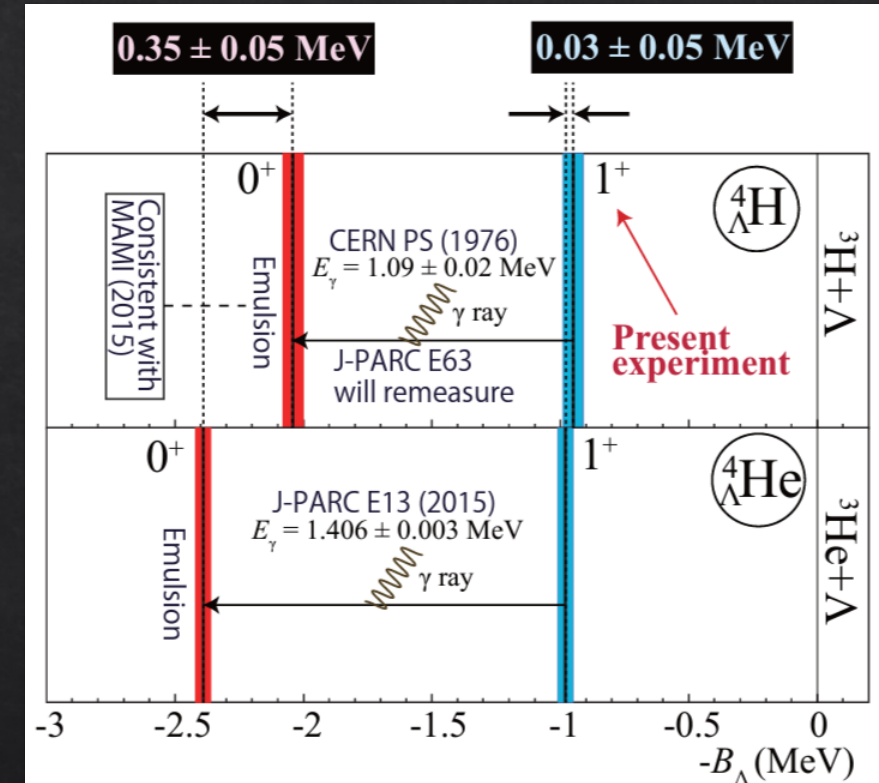
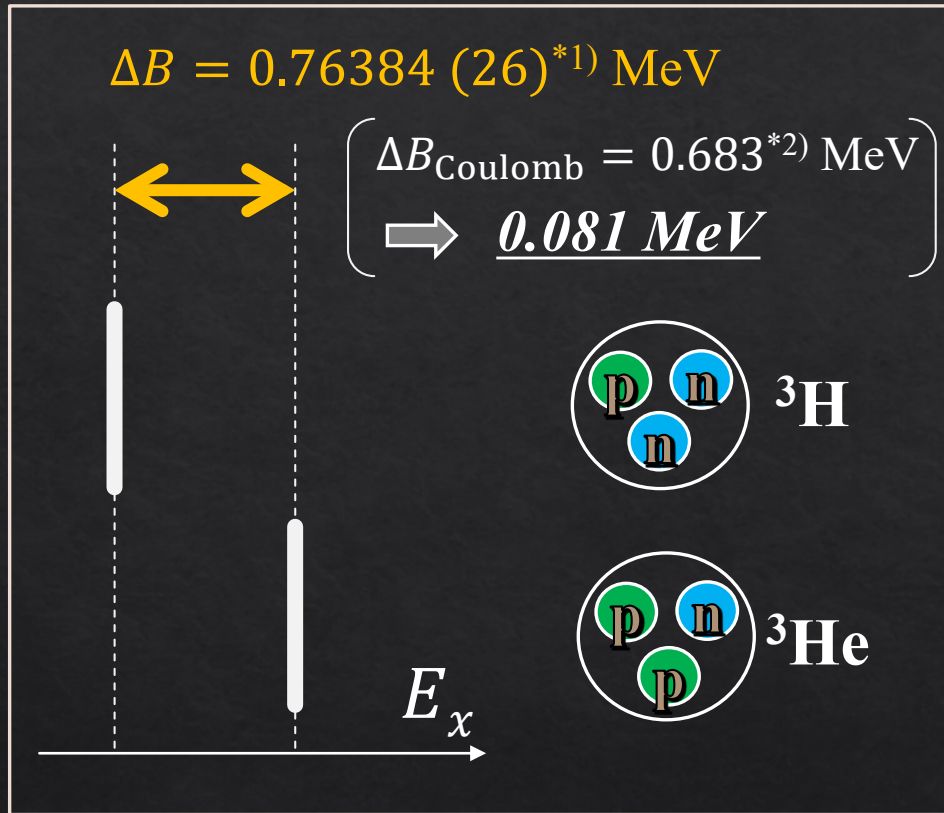
D. Lonardon et al., Phys. Rev. Lett. 114, 092301 (2015)



PAC44 approval of E12-15-008 with grade-A

# Charge Symmetry Breaking (CSB), the mystery

Figure from proposal of [JLab E12-19-002](#)



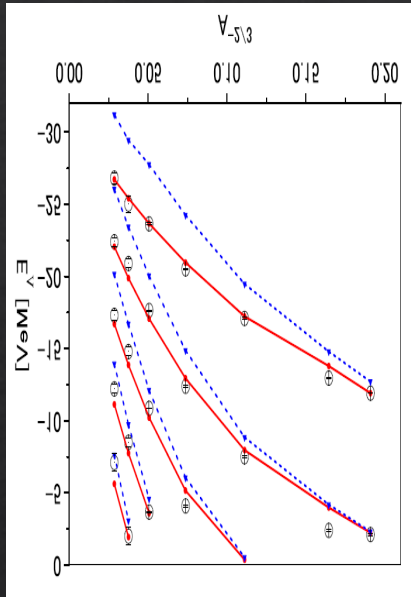
\*1) J.H.E.Mattauch *et al.*, *Nucl. Pys.* **67**, 1 (1965).

\*2) R.A.Brandenburg, S.A.Coon *et al.*, *NPA294*, 305 (1978).

Data of other systems are necessary to pin down the origin of CSB

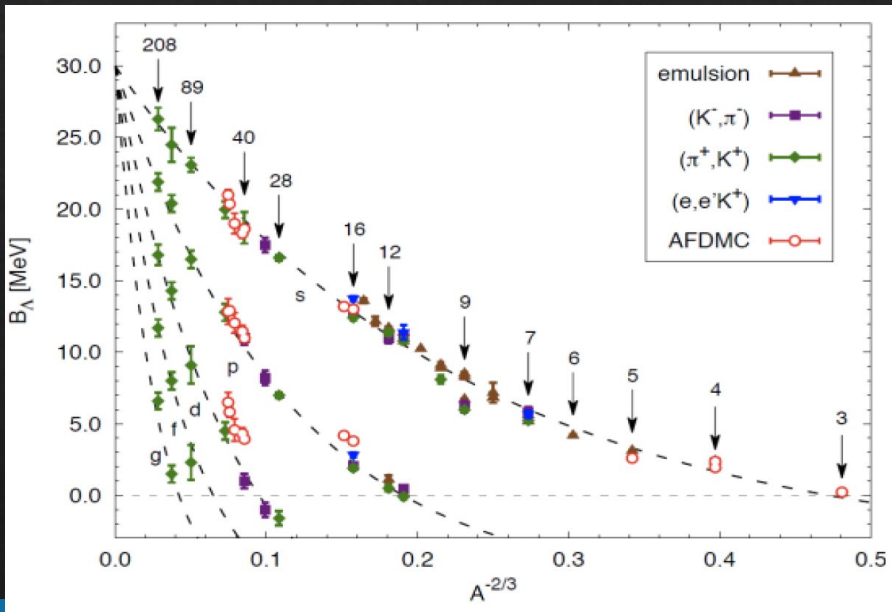
# $\Lambda$ Single Particle Energies of $\Lambda$ Hypernuclei by Various Calculations

M.M. Nagels et al., PRC 99 (2019) 044003.



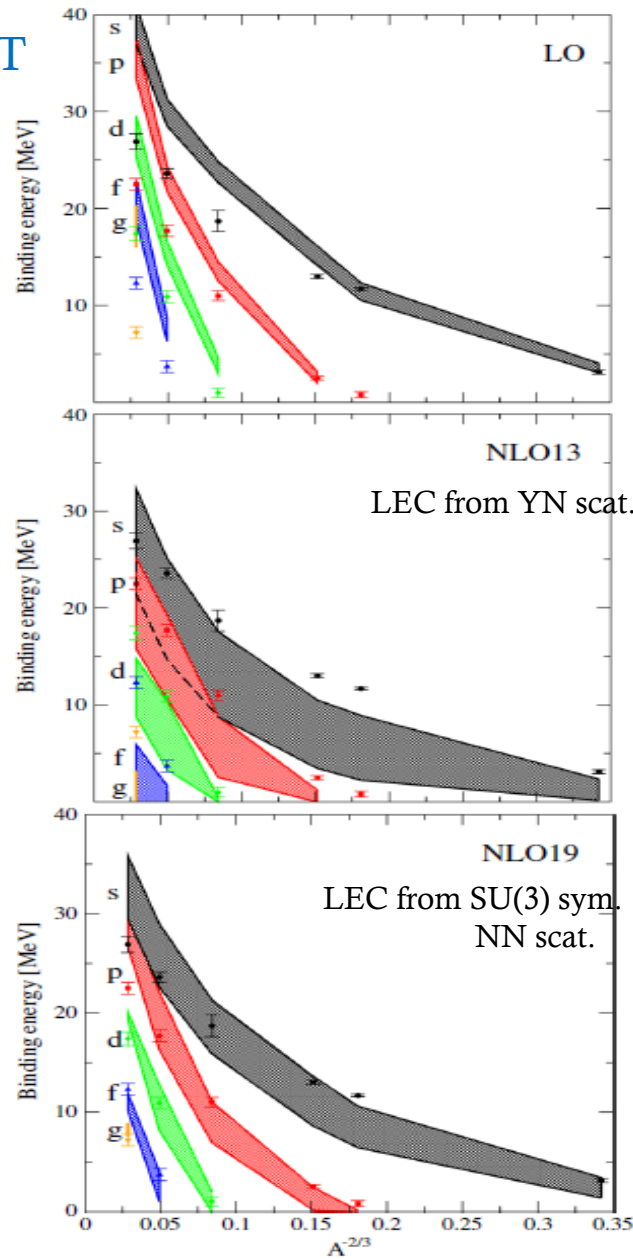
ESC16  
ESC16+ (Inc. 3BF)  
G-matrix

D.Lonardonni and F. Pederiva, arXiv:1711.07521.

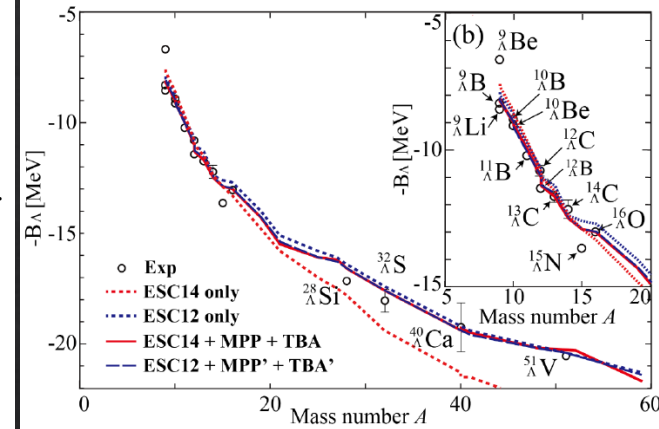


AFDMC

ChEFT

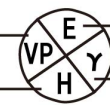


Hyper-AMD



M. Isaka et al.,  
PRC94, 044310 (2016),  
PRC 95, 044308 (2017)

J.Haidenbauer, I.Vidana, EPJA (2020) 56:55.



Hypernucleus		T<0	T=0	T>0	CSB study		
					Now	New JLab	J-PARC
s-shell	$d N \Lambda$ ( $0^+$ )	${}^4_{\Lambda}H$		${}^4_{\Lambda}He$	○	○	○
	$d N \Lambda$ ( $1^+$ )	${}^4_{\Lambda}H$ <i>new</i>		${}^4_{\Lambda}He$	△	●	●
p-shell	$\alpha N \Lambda$	${}^6_{\Lambda}He$ <i>new</i>		${}^6_{\Lambda}Li$ <i>new</i>			●
	$\alpha NN \Lambda$	${}^7_{\Lambda}He$	${}^7_{\Lambda}Li^*$	${}^7_{\Lambda}Be$	○	○	○
	$\alpha d N \Lambda$	${}^8_{\Lambda}Li$		${}^8_{\Lambda}Be$	○	○	○
	$\alpha d NN \Lambda$	${}^9_{\Lambda}Li$ <i>new</i>	${}^9_{\Lambda}Be$	${}^9_{\Lambda}B$		●	●
	$\alpha \alpha N \Lambda$	${}^{10}_{\Lambda}Be$		${}^{10}_{\Lambda}B$ <i>new</i>			●
	$\alpha \alpha NN \Lambda$	${}^{11}_{\Lambda}Be$ <i>new</i>	${}^{11}_{\Lambda}B$ <i>new</i>	${}^{11}_{\Lambda}C$		●	●
	$\alpha \alpha d N \Lambda$	${}^{12}_{\Lambda}B$		${}^{12}_{\Lambda}C$ <i>new</i>			●

Isospin multiplet for CSB study

		T < 0		T = 0		CSB study		
		E12-19-002				Now	New JLab	J-PARC
s-shell	(0 <sup>+</sup> )	${}^4\text{He}$	${}^4_{\Lambda}\text{He}$	${}^4\text{He}$	${}^4_{\Lambda}\text{He}$			
	(1 <sup>+</sup> )	${}^6\text{He}$	${}^6_{\Lambda}\text{He}$ <i>new</i>	${}^6\text{Li}$	${}^6_{\Lambda}\text{Li}$ <i>new</i>			
p-shell	$\alpha$ N $\Lambda$	${}^7\text{He}$	${}^7_{\Lambda}\text{He}$	${}^7\text{Li}^*$	${}^7_{\Lambda}\text{Li}$	${}^7\text{Be}$	${}^7_{\Lambda}\text{Be}$	
	$\alpha$ d N $\Lambda$	${}^8\text{Li}$	${}^8_{\Lambda}\text{Li}$			${}^8\text{Be}$	${}^8_{\Lambda}\text{Be}$	
	$\alpha$ d N N $\Lambda$	${}^9\text{Li}$	${}^9_{\Lambda}\text{Li}$ <i>new</i>	${}^9\text{Be}$	${}^9_{\Lambda}\text{Be}$	${}^9\text{B}$	${}^9_{\Lambda}\text{B}$	
	$\alpha$ $\alpha$ N $\Lambda$	${}^{10}\text{Be}$	${}^{10}_{\Lambda}\text{Be}$	${}^{10}\text{B}$	${}^{10}_{\Lambda}\text{B}$ <i>new</i>	${}^{10}\text{C}$	${}^{10}_{\Lambda}\text{C}$	
	$\alpha$ $\alpha$ N N $\Lambda$	${}^{11}\text{Be}$	${}^{11}_{\Lambda}\text{Be}$ <i>new</i>	${}^{11}\text{B}$	${}^{11}_{\Lambda}\text{B}$ <i>new</i>	${}^{11}\text{C}$	${}^{11}_{\Lambda}\text{C}$	
	$\alpha$ $\alpha$ d N $\Lambda$	${}^{12}\text{B}$	${}^{12}_{\Lambda}\text{B}$			${}^{12}\text{C}$	${}^{12}_{\Lambda}\text{C}$ <i>new</i>	

LOI12-23-013

E12-19-002

Proposing

Proposing

E94

E94

Isospin multiplet for CSB study

To "complete"

J-PARC E63

J-PARC E94

JLab LOI12-23-013

- YN scat. exp.
- Femotoscropy

CSB

${}^3_\Lambda\text{H}$  lifetime puzzle

JLab E12-19-002

JLab LOI12-23-011

$nn\Lambda$  bound puzzle

Invariant mass spectroscopy by HI beam @LHC, RHIC, GSI

JLab C12-20-013 (C2)

JLab LOI12-23-016

Many Body effect (Cluster, deformation)

- Space observation
- Graviton wave meas.

Neutron star puzzle

Strangeness	2B	Coupled channel	3B
-1		$\Lambda N - \Sigma N$	
-2		$\Sigma N - \Lambda\Lambda$	

J-PARC E70

J-PARC E75

J-PARC E96

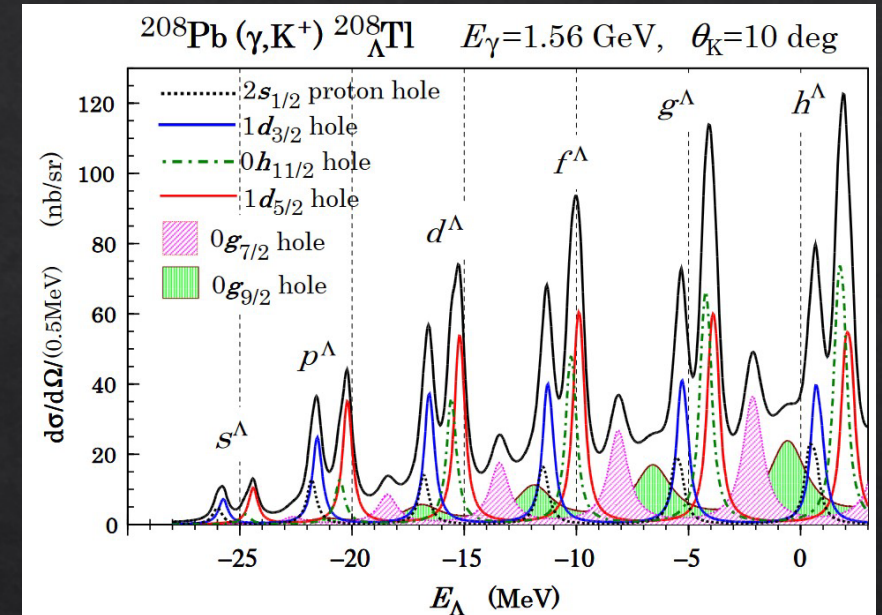
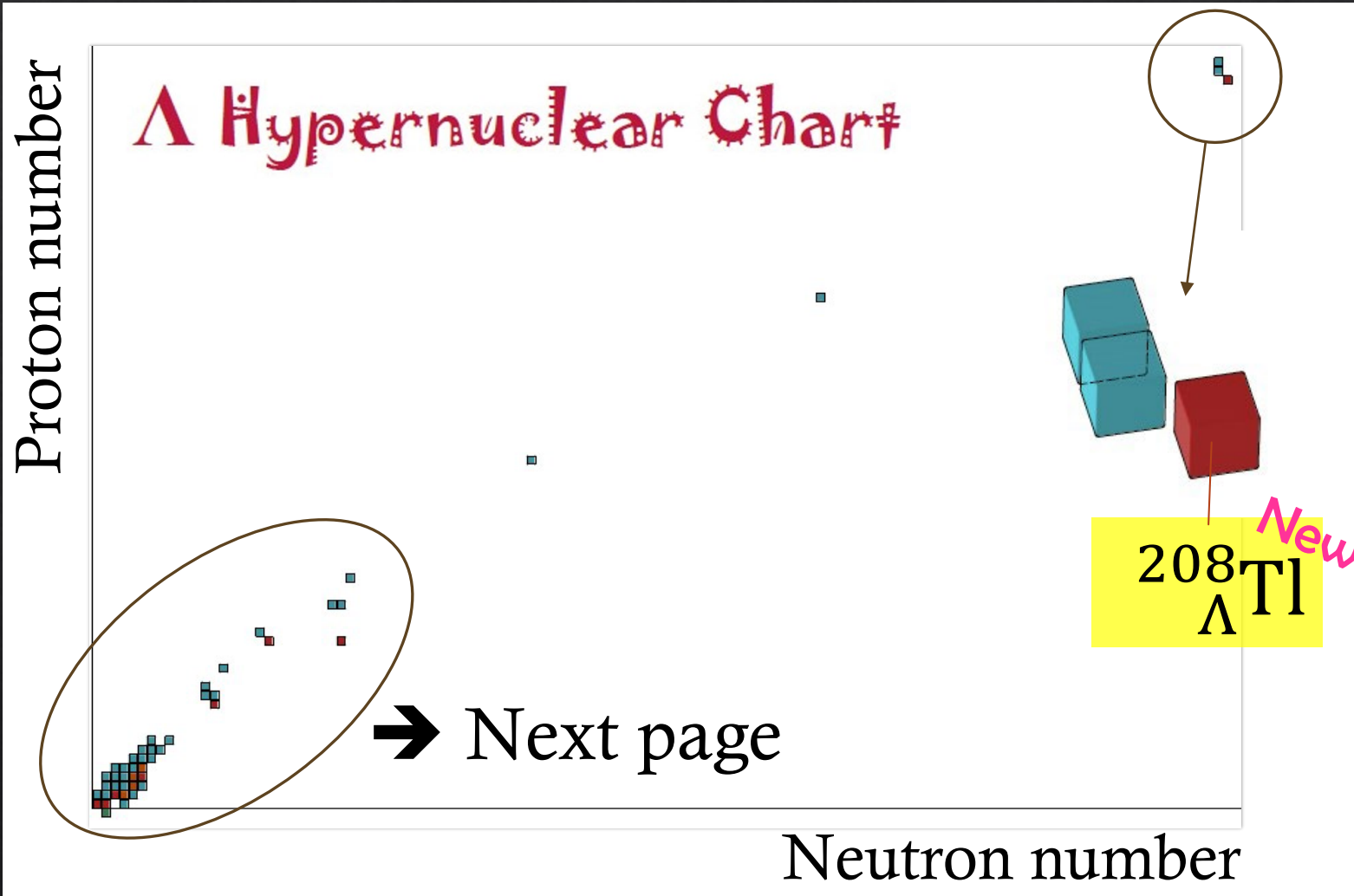
JLab E12-15-008

JLab E12-20-013



# Next JLab experiments; light to heavy hypernuclei

T. Motoba, [JPS Conf. Proc. 17, 011003 \(2017\)](#)



**JLab E12-20-013**

→  **$\Lambda$ NN force**

F. Garibaldi et al.,  
[EPJ Web Conf. 271, 01007 \(2022\)](#).



# Next JLab experiments; light to heavy hypernuclei

New

LOI12-23-013

→ [Link](#)



Contact person  
Prof. T. Gogami  
(Kyoto Univ.)

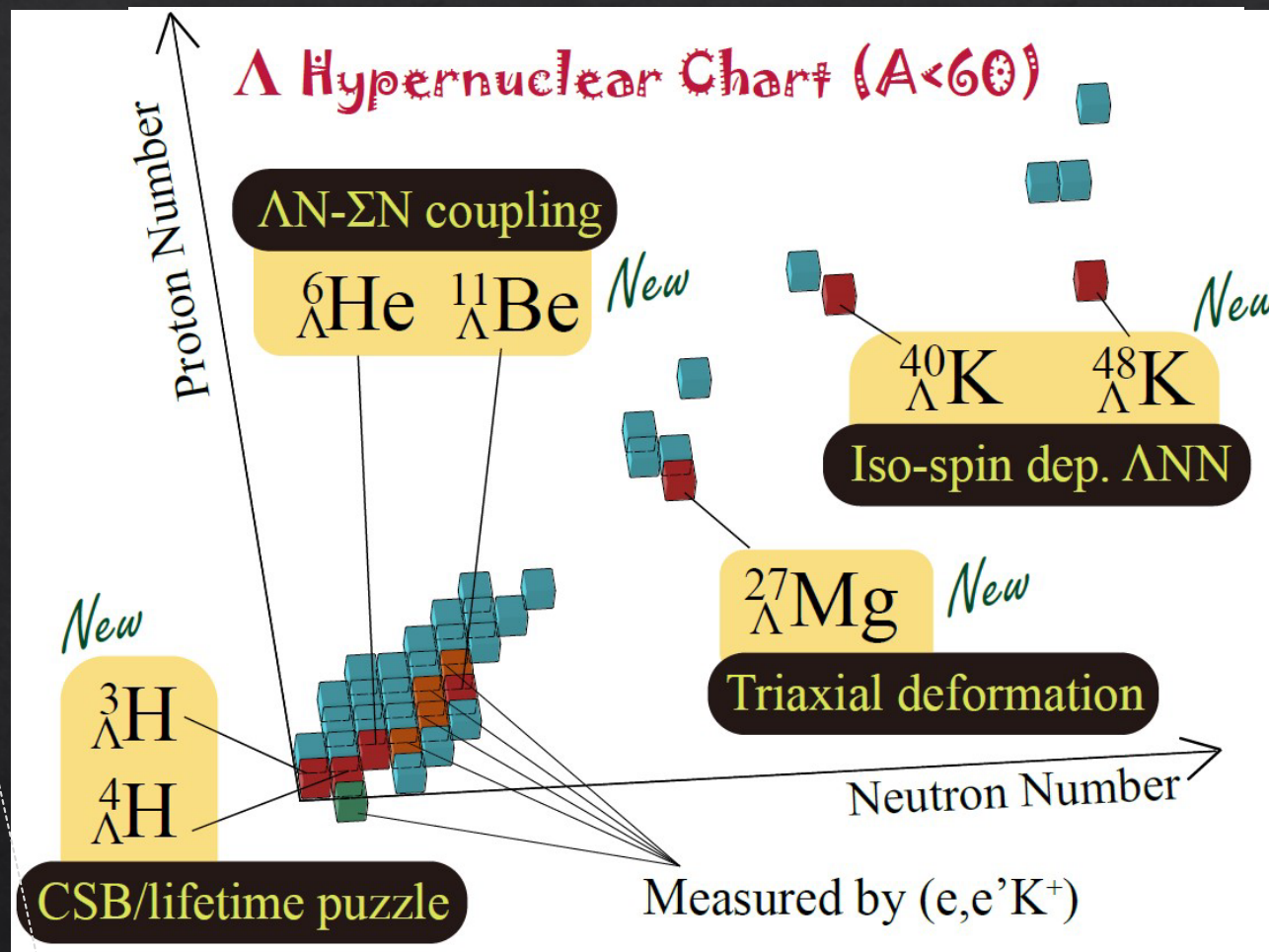
E12-19-002

- Proposal → [Link](#)
- TG et al., [EPJ Web Conf. 271, 01001 \(2022\)](#).

New

LOI12-23-011

Presentation by Prof. S. Nagao, “[Mass Determination of Light Hypernuclei with the Decay Pion Spectroscopy: from MAMI to JLab](#)”, Hall A/C Summer meeting, June 30, 2023



Contact person  
Prof. S. Nagao (U. Tokyo)

E12-15-008

→ [Link](#)



Contact person  
Prof. S. N. Nakamura  
(U. Tokyo)

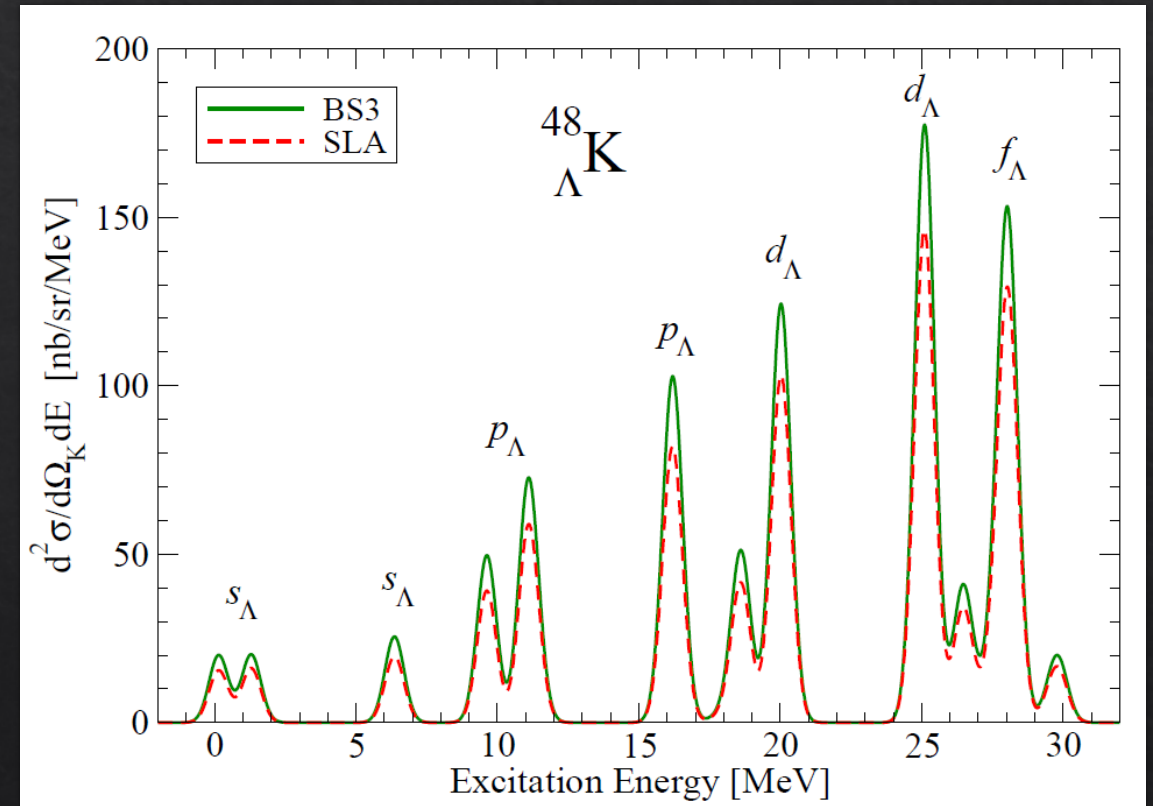
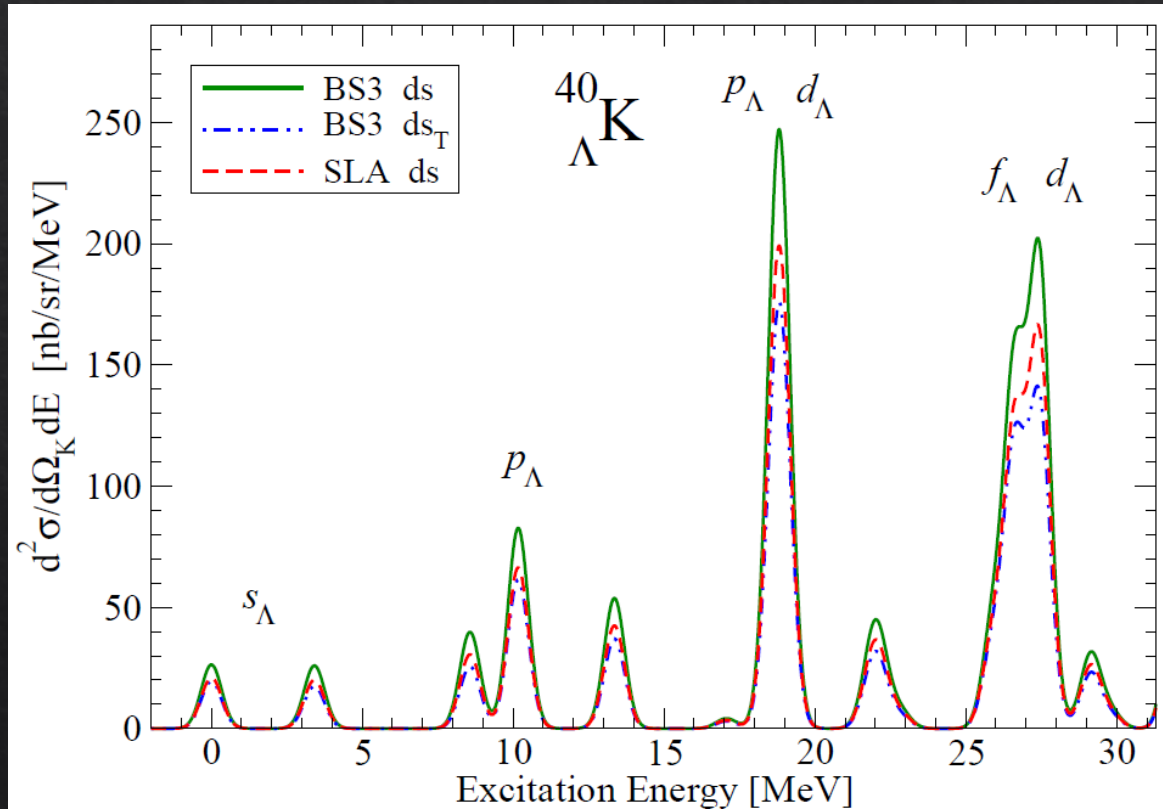
New

LOI12-23-016

Presentation by Prof. M. Isaka, “[Triaxial deformation of nuclei probed by  \$\Lambda\$  particle](#)”, Hall A/C Summer meeting, June 30, 2023

# Theoretical calculation, $A = 40$ and $48$

P. Bydžovský, D. Denisova, D. Petrellis, D. Skoupil, P. Veselý,  
G. De Gregorio, F. Knapp, N. Lo Iudice, [arXiv:2306.01308](https://arxiv.org/abs/2306.01308) [nucl-th] (2023)



# The CSB in p-shell $\Lambda$ hypernuclei

Phenomenological CSB:

E. Hiyama *et al.*, *Phys. Rev. C* **80**, 054321 (2009).

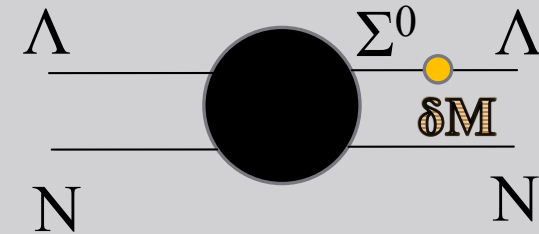
$\Lambda$ N- $\Sigma$ N coupling:

A. Gal, *Phys. Lett. B* 744, 352 (2015).

## Phenomenological $\Lambda$ N CSB interaction

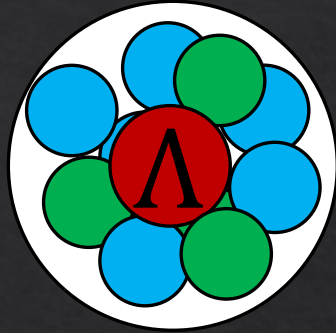
$$V_{\Lambda N}^{\text{CSB}}(r) = -\frac{\tau_z}{2} \left[ \frac{1+P_r}{2} \left( v_0^{\text{even,CSB}} + \sigma_\Lambda \cdot \sigma_N v_{\sigma_\Lambda \cdot \sigma_N}^{\text{even,CSB}} \right) e^{-\beta_{\text{even}} r^2} + \frac{1-P_r}{2} \left( v_0^{\text{odd,CSB}} + \sigma_\Lambda \cdot \sigma_N v_{\sigma_\Lambda \cdot \sigma_N}^{\text{odd,CSB}} \right) e^{-\beta_{\text{odd}} r^2} \right]$$

## Explicit inclusion of the $\Lambda$ N- $\Sigma$ N coupling



- ✓ Little effect on p-shell  $\Lambda$  hypernuclear system
- ✓ Large CSB in the  $A = 12$  system ( ${}^{12}_\Lambda\text{B}$  and  ${}^{12}_\Lambda\text{C}$ ) ??

# $^{12}_{\Lambda}\text{C}$ shift problem



$$B_{\Lambda}^{\text{emul.}}(^{12}_{\Lambda}\text{C}) = 10.37 \text{ MeV}$$

(The mean of 6 events)

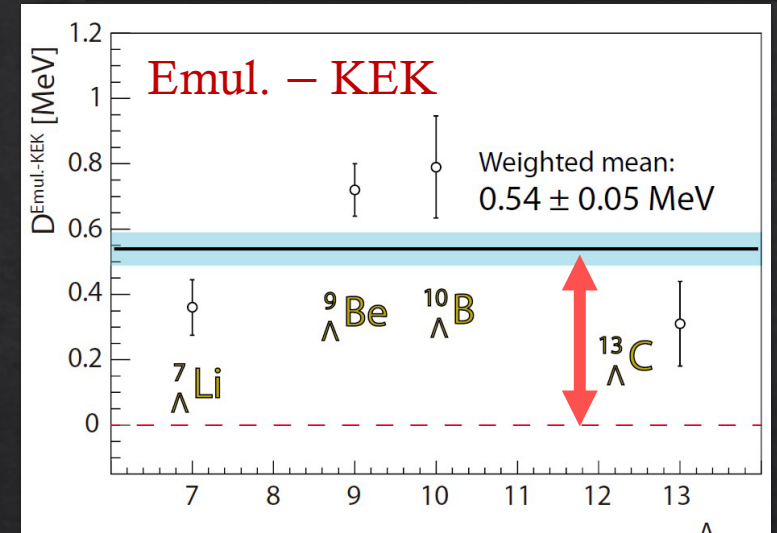
**0.5—0.6 MeV correction is necessary**

**Direct measurement**  
(J-PARC94)

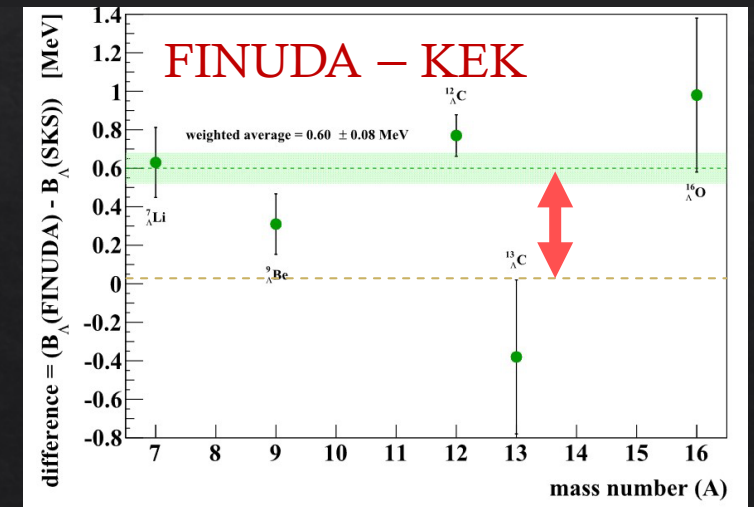
Indirect measurement

KEK: calibrated by  $B_{\Lambda}^{\text{emul.}}(^{12}_{\Lambda}\text{C})$

TG et al., PRC 93, 034314 (2016)

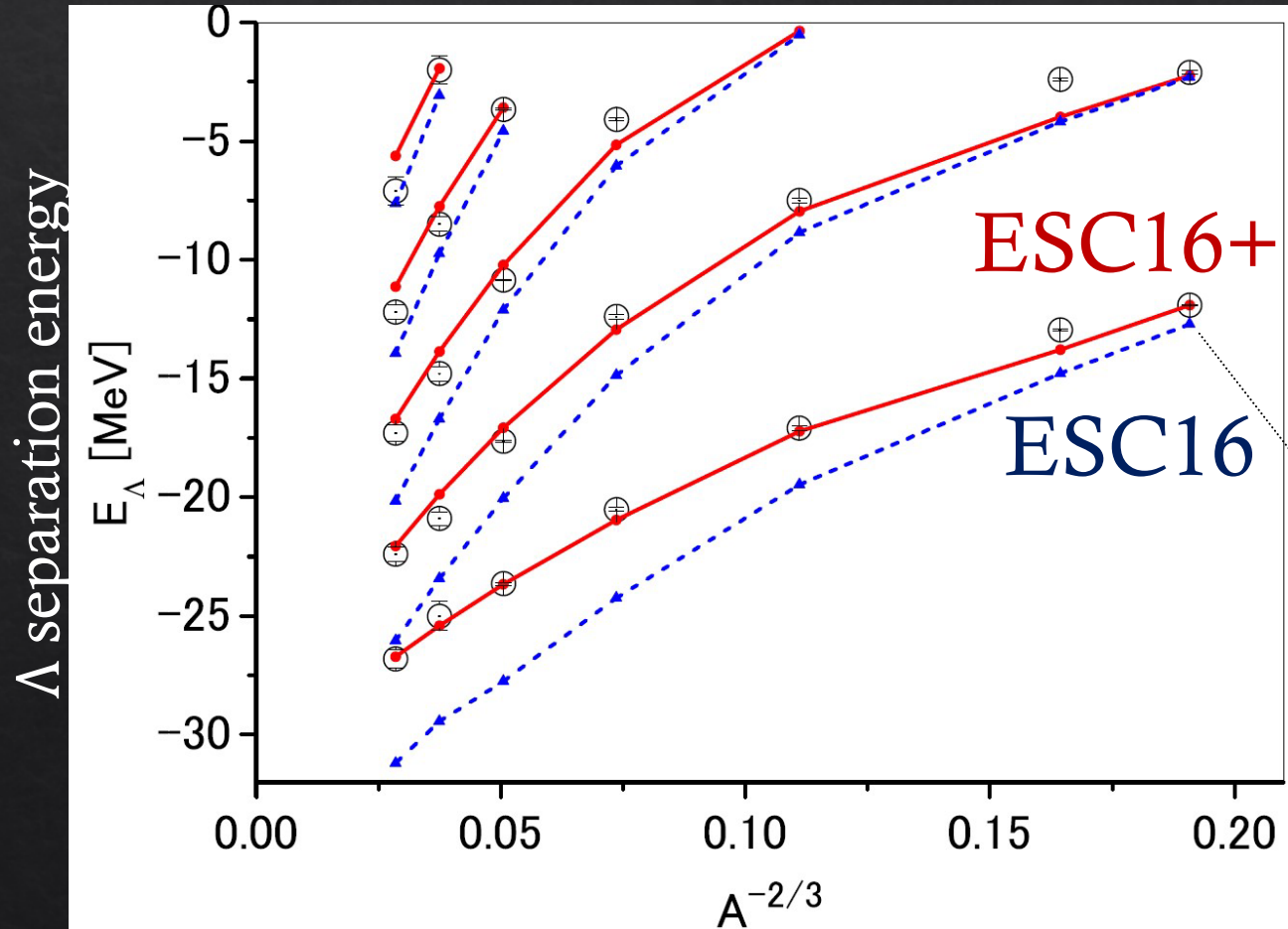


E. Botta et al., NPA 960, 165—179 (2017)



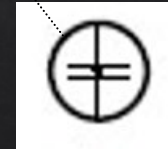
# 0.5 MeV shift became standard

M. M. Nagels et al., Phys. Rev. C **99**, 044003 (2019)



(A: Mass number of core nucleus)

$B_{\Lambda}^{\text{emul.}}(^{12}_{\Lambda}\text{C})$  was used for the energy calibration of the  $(\pi^+, K^+)$  experiments

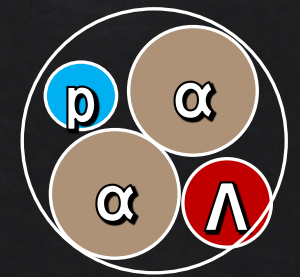
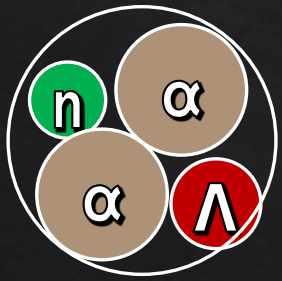
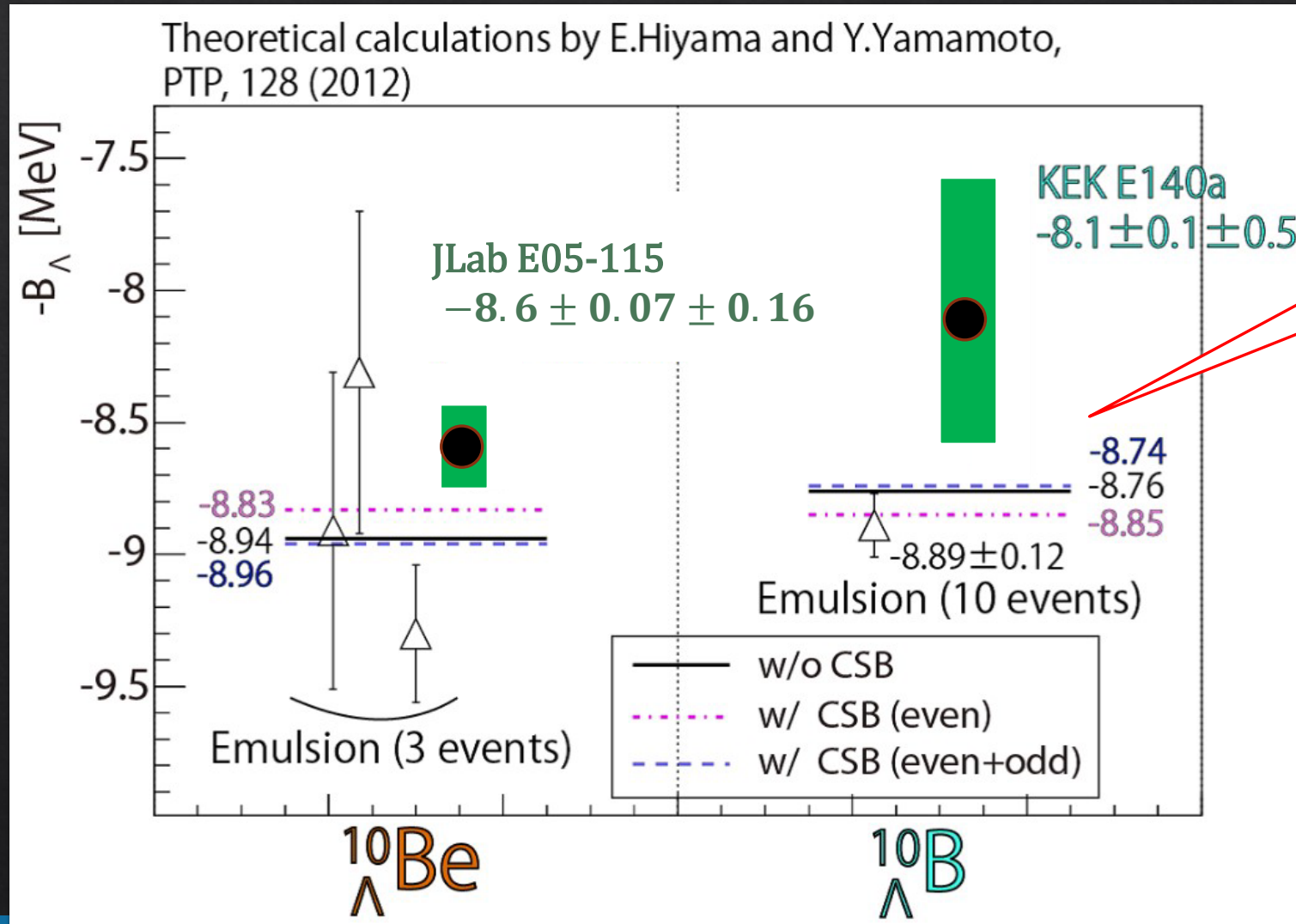


Data with 0.5 MeV correction

to be compared with theoretical calculations

# CSB in the $A = 10$ iso-doublet HN

J-PARC  
E94



Jun 2022

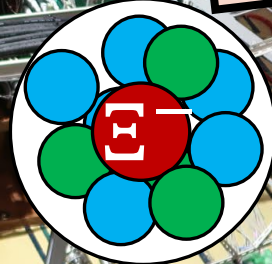


Q2 Q1

1.05 GeV/c

D

$\pi^+$



J-PARC  
E94

$K^+$

2 m

S-2S

0.72 GeV/c