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Deuteron Disintegration at Large Missing Momenta

C. Yero

Science at the Luminosity Frontier: Jefferson Lab 22 GeV Workshop

January 23 - 25, 2023

FIU

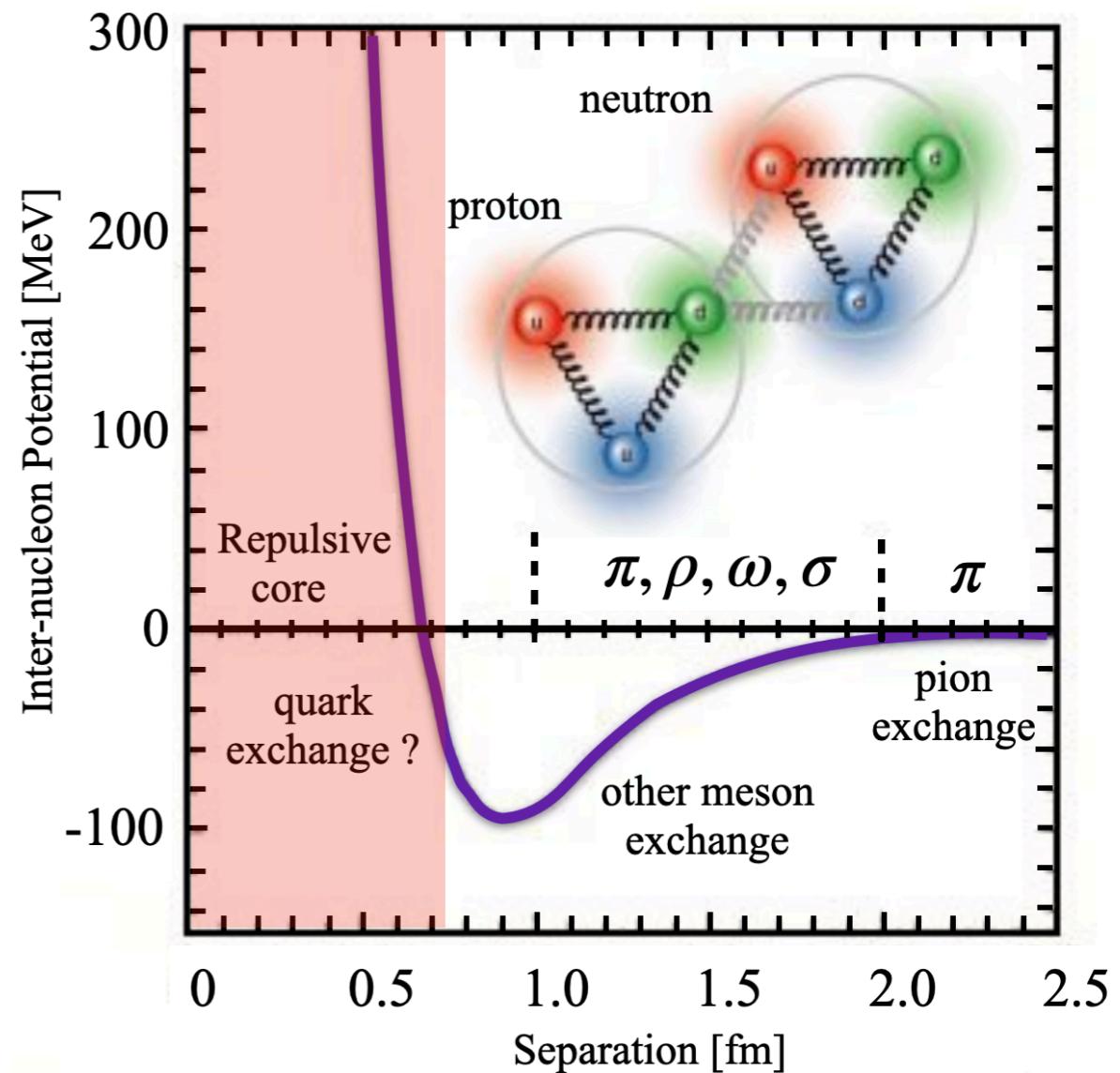


Jefferson Lab
Exploring the Nature of Matter


OLD DOMINION
UNIVERSITY

Motivation

- How does the nuclear force work?
- What is the structure of nuclear matter?
- Short-range part of NN interaction is NOT well understood !!!
 - difficult to generate NN potentials at high nucleon momentum / short distance
 - difficult to calculate heavy nuclei from NN potentials
 - lack of experimental data on nuclei at high momentum or short distances



$d(e, e' p)n$ break-up at \sim PWIA

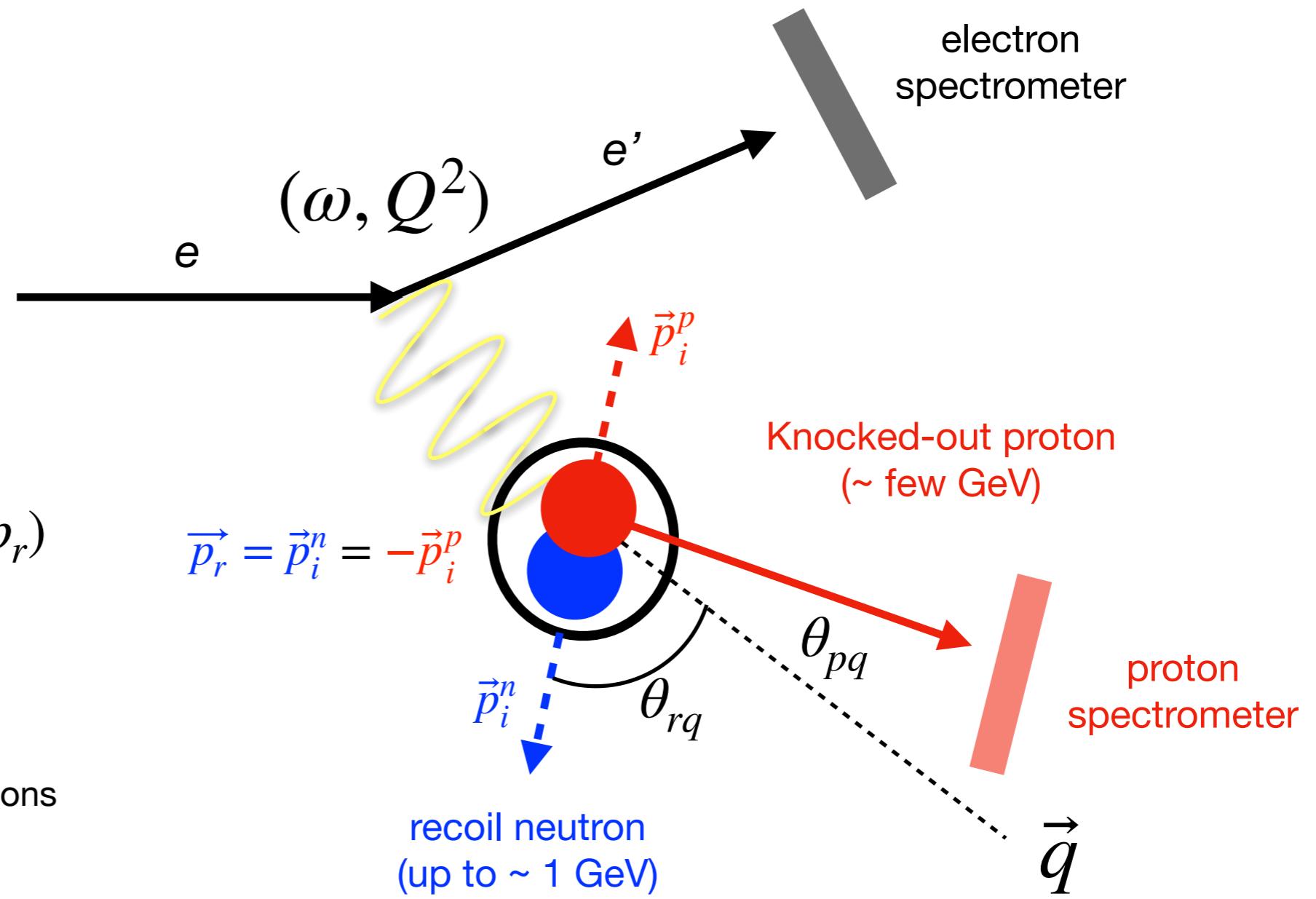
PWIA:

- no re-interaction between ejected nucleons

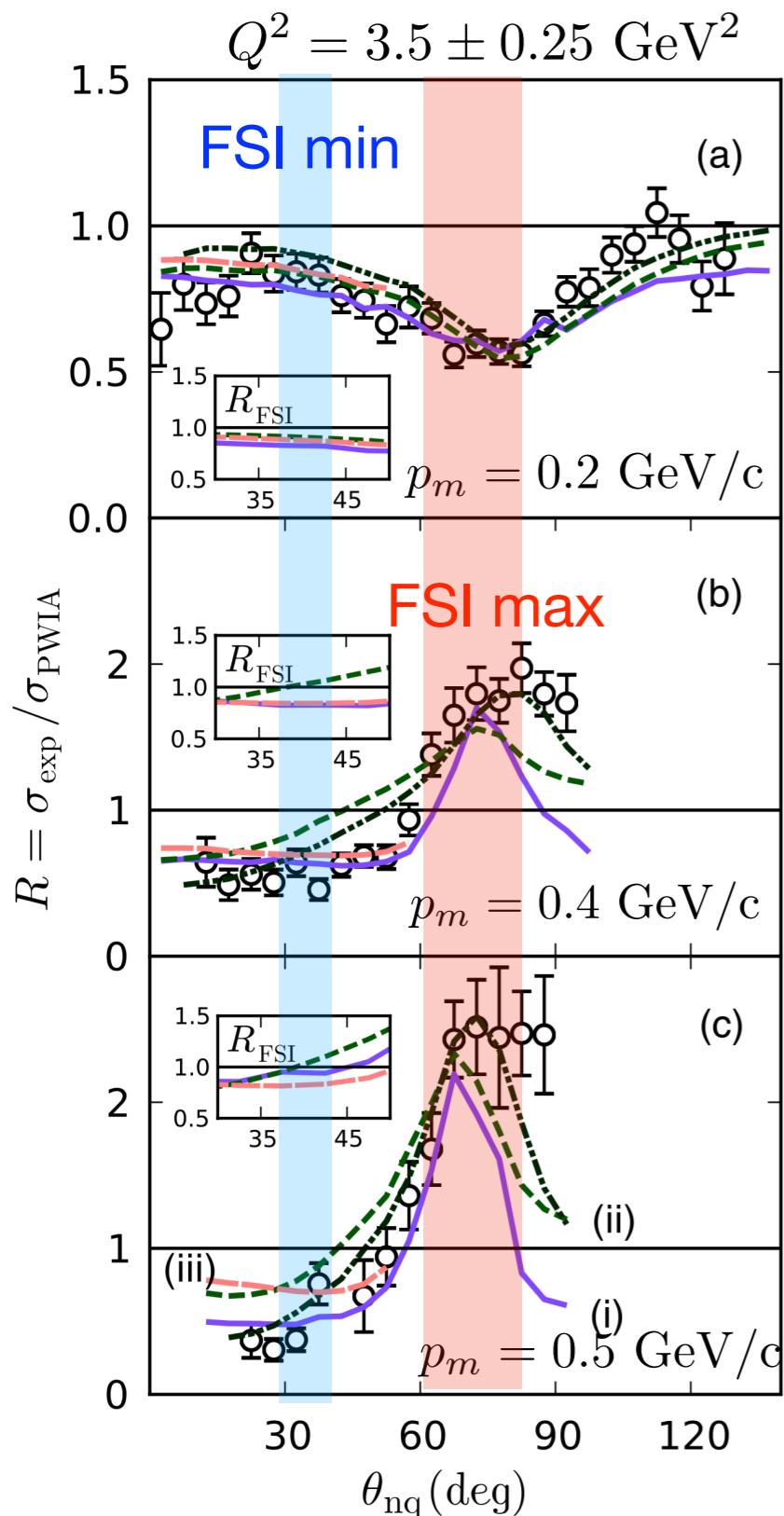
$$\frac{d^5\sigma}{d\omega d\Omega_e d\Omega_p} = k\sigma_{ep}\rho_m(p_r)$$

Experimental momentum distributions

$$\sigma_{red} \equiv \frac{\sigma_{exp}}{k\sigma_{ep}} \approx \rho_m(p_r)$$

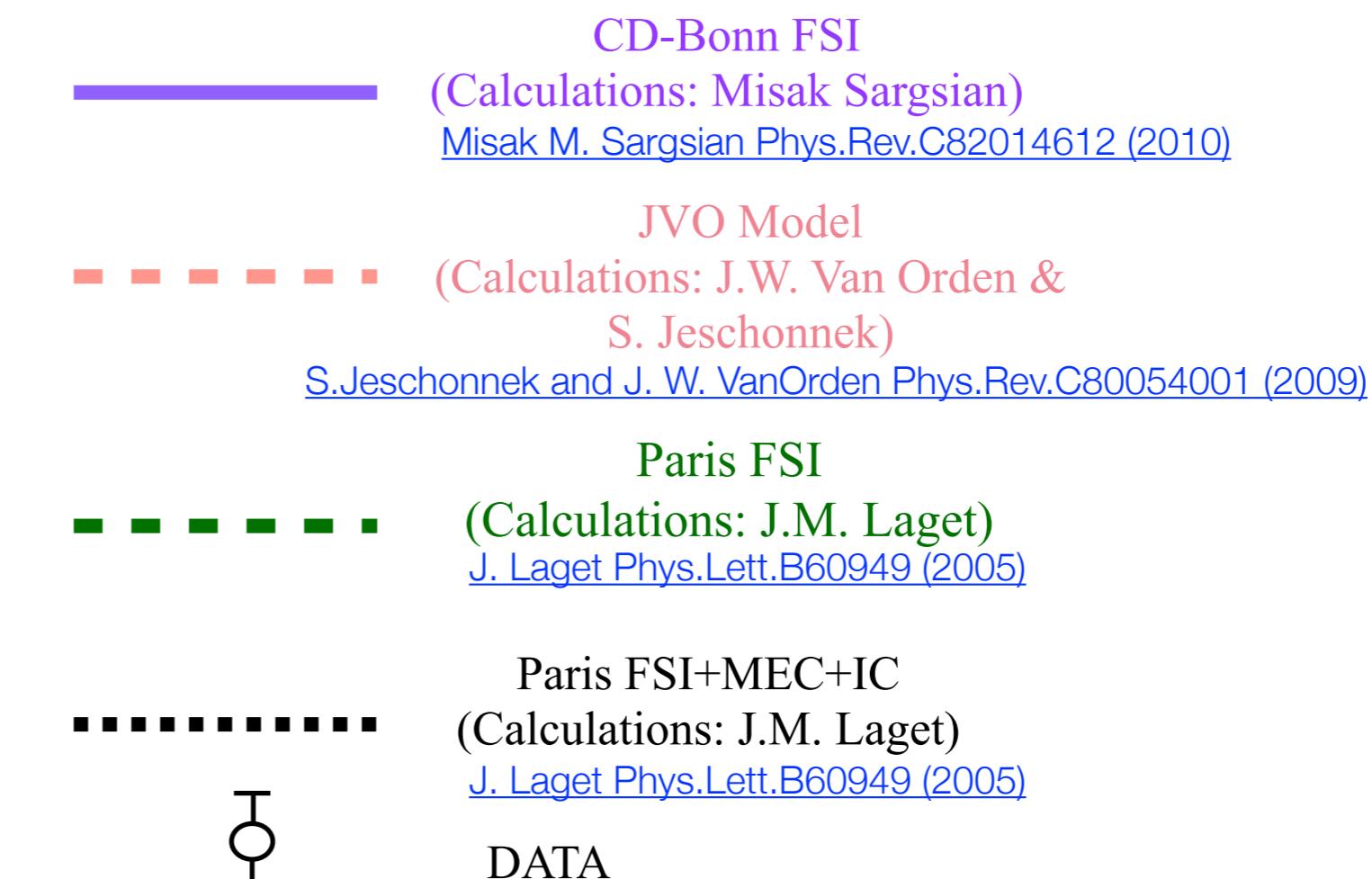


first d(e, e'p)n at high Q^2 ($> 1 \text{ GeV}^2$)



[Boeglin et al. \(Hall A\) Phys.Rev.Lett. 107, 262501 \(2011\)](#)

[K. S. Egiyan et al. \(CLAS\) Phys. Rev. Lett. 98, 262502 \(2007\)](#)



FSI strongly anisotropic (angular-dependent):

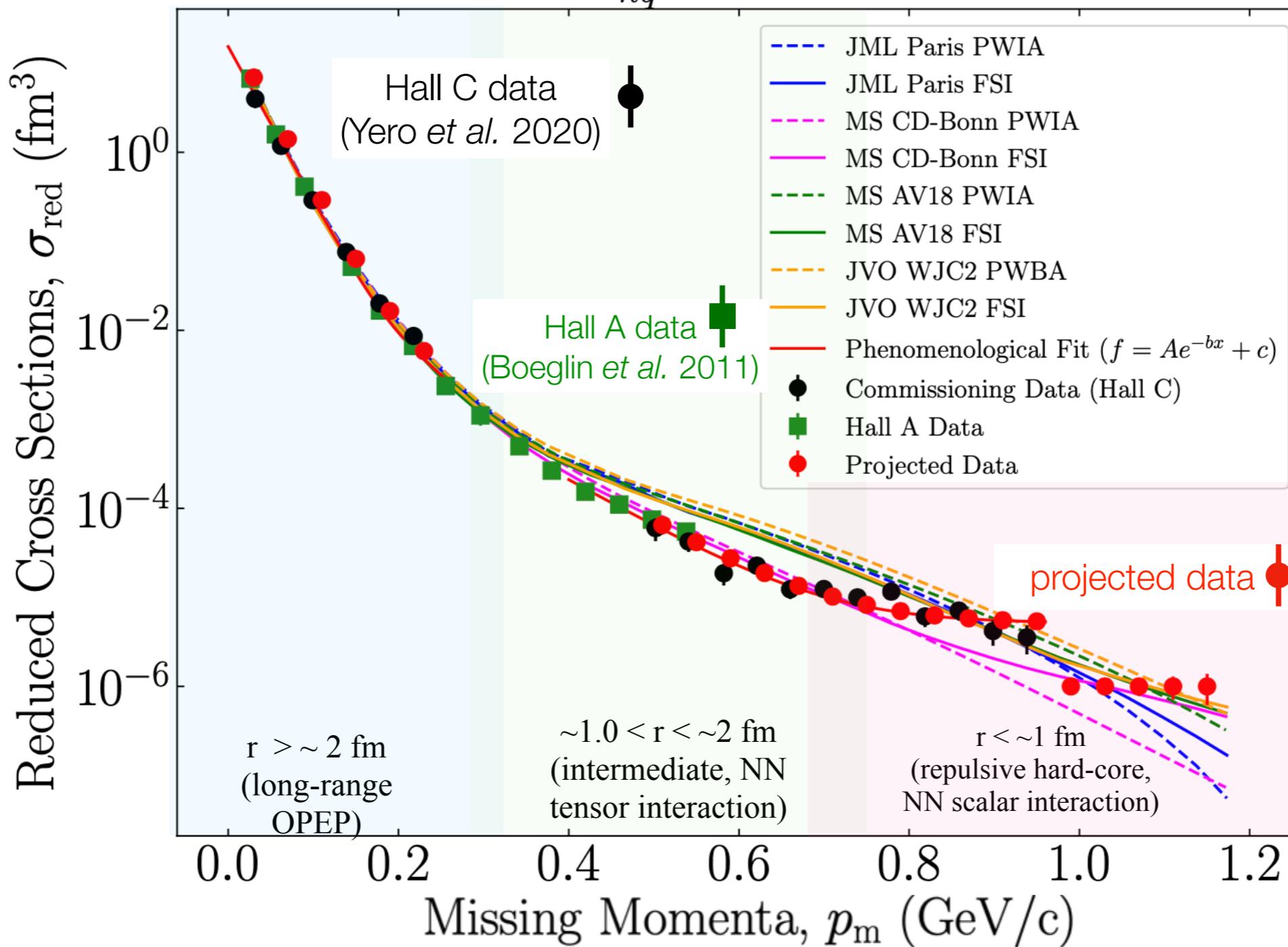
- Sargsian uses GEA, Laget uses fully relativistic
- **FSI peak at $\theta_{nq} \sim 70^\circ$**
- minimal FSI at $\theta_{nq} \sim 35 - 45^\circ$

GEA theory:

[L. L. Frankfurt, M. M. Sargsian, and M. I. Strikman Phys.Rev.C561124 \(1997\)](#)

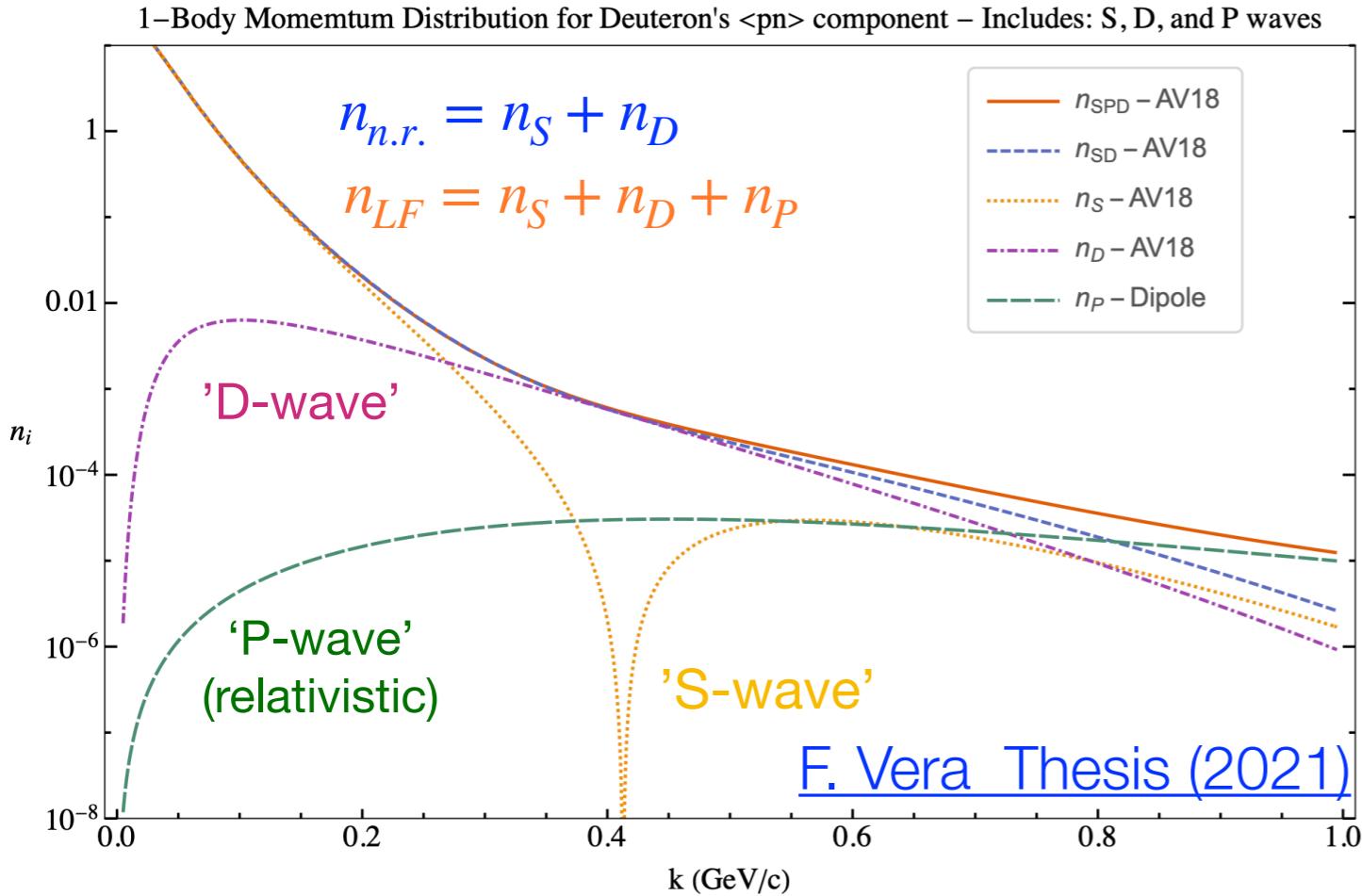
experiment: probing the deuteron NN core

$$\theta_{nq} = 35 \pm 5^\circ$$



- non-relativistic theory calc. using **CD-Bonn** (M. Sargsian) reproduce data up to $p_m \sim 0.7$ GeV/c
- no model reproduces data $p_m > 0.7$ GeV/c (non-nucleonic degrees of freedom?, quarks?)

theory: probing the deuteron NN core

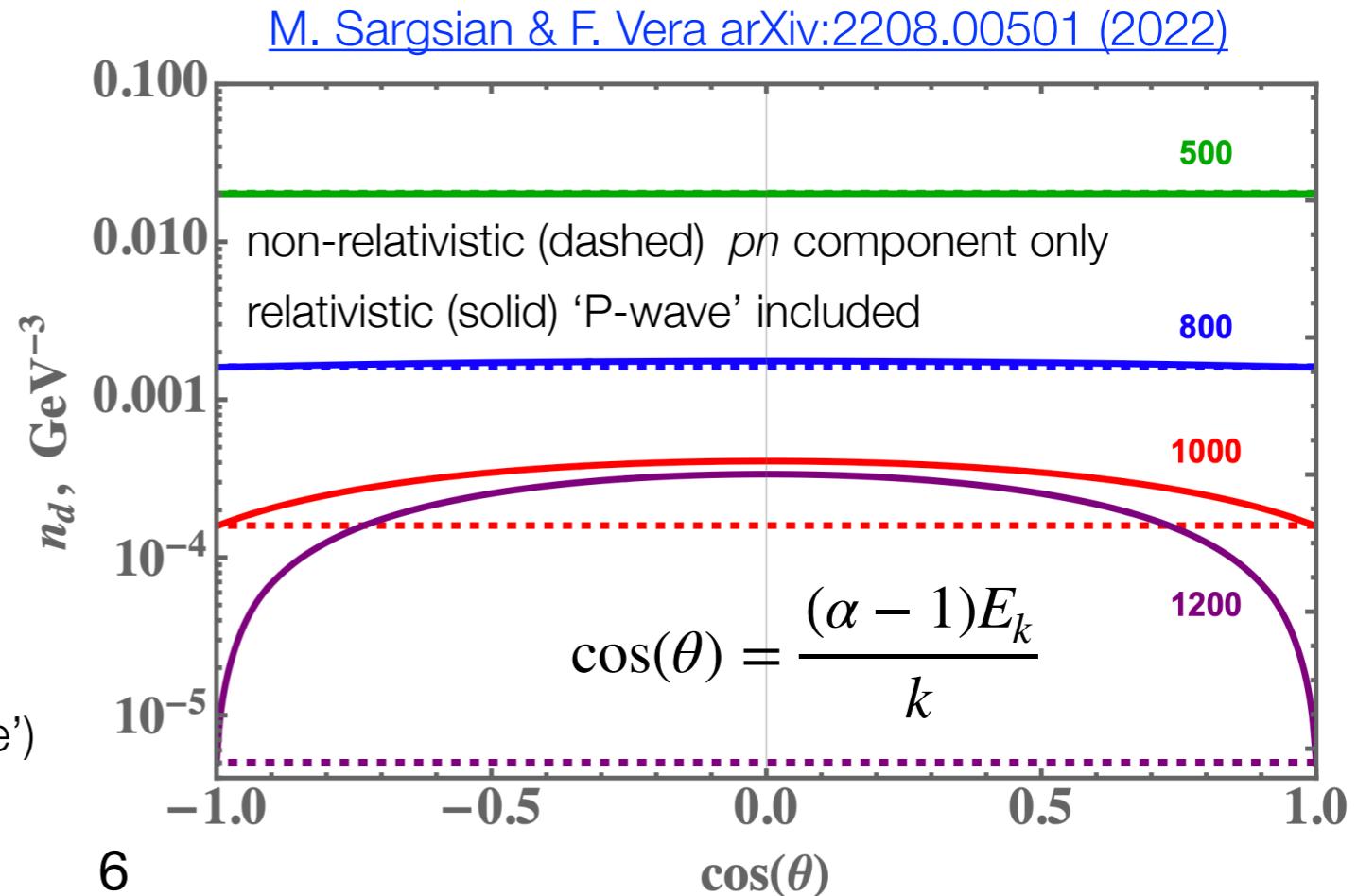


- non-relativistic deuteron w.f. (only S + D wave)
- fully relativistic calculation of deuteron w.f. in the LF give rise to a ‘P-wave’ -like component
- P-wave starts to dominate at $\sim k \sim 800$ MeV/c
 - ▶ angular dependence of momentum distributions (n_d)
 - ▶ n_d modified by non-nucleonic components

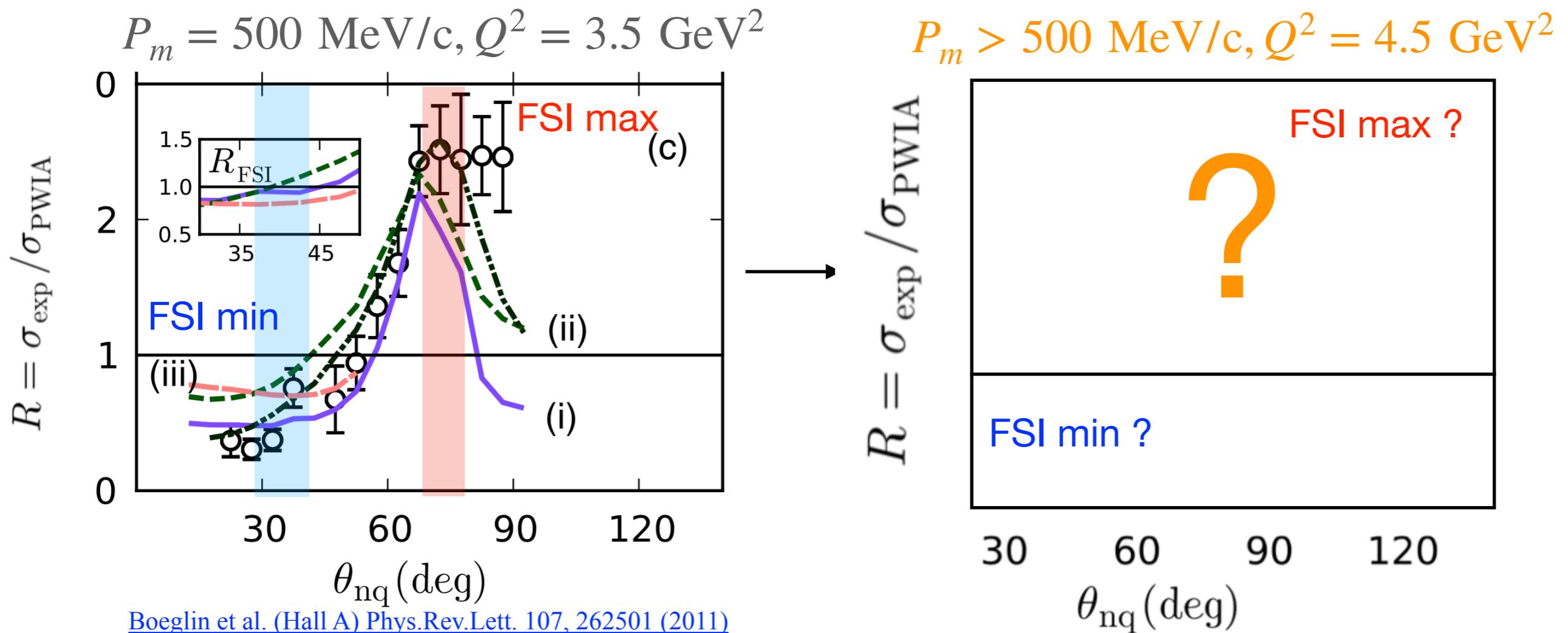
$$\Psi_d = \boxed{\Psi_{pn}} + \boxed{\Psi_{\Delta\Delta} + \Psi_{NN^*}} \\ + \boxed{\Psi_{hc} + \Psi_{NN\pi} \dots}$$

non-relativistic relativistic (non-nucleonic)

- relativistic deuteron more complex than simple $pn \rightarrow pn$ transition
- possible non-nucleonic transitions (relativistic ‘P-wave’)
 $\Delta\Delta \rightarrow np, NN^* \rightarrow np, hc \rightarrow np, NN\pi \rightarrow np$



need angular distributions at high Q^2 and $P_m > 500 \text{ MeV}/c$



- d(e, e'p) measurements are approaching NN core ($\sim 1 \text{ GeV}/c$) \rightarrow Yero et. al 2020
 - ▶ need for extending angular distribution $> 500 \text{ MeV}/c$ to demonstrate FSI are still small at $\theta_{rq} \sim 35 - 45 \text{ deg}$,
- we explore the possibilities of taking advantage of a beam energy upgrade (22 GeV) for these measurements (2nd part of this talk, remaining slides)

Optimizing d(e, e'p) Kinematics (in a future 22 GeV upgrade)

- 1) How high in beam energy can we go while keeping Q^2 fixed?
—> Does it really work out ?
- 2) Consider e- arm spectrometer upgrade (up to ~ 20.5 GeV/c)
—> with upgrade, kinematics can reach $E_b = 22$ GeV while keeping $Q^2 = 4.5$ GeV 2
- 3) Consider luminosity upgrade (double the beam current 80 -> 160 uA)
—> used baseline kinematics ($E_b = 11$ GeV, $Q^2 = 4.5$ GeV 2) to estimate increase in rates

Simulation Reaction: d(e,e'p), electrons: SHMS, protons: HMS

Fixed Parameters: 80 uA, 1 hour, 20 cm-long deuterium target,
Model: AV18 FSI (SIMC)

Angular Settings: (θ_{rq} ~27, ~49, ~65) deg at fixed $Q^2 = 4.5$, $P_m = 1 \pm 0.02$ GeV/c
(3 settings provide wide angular coverage for FSI studies)

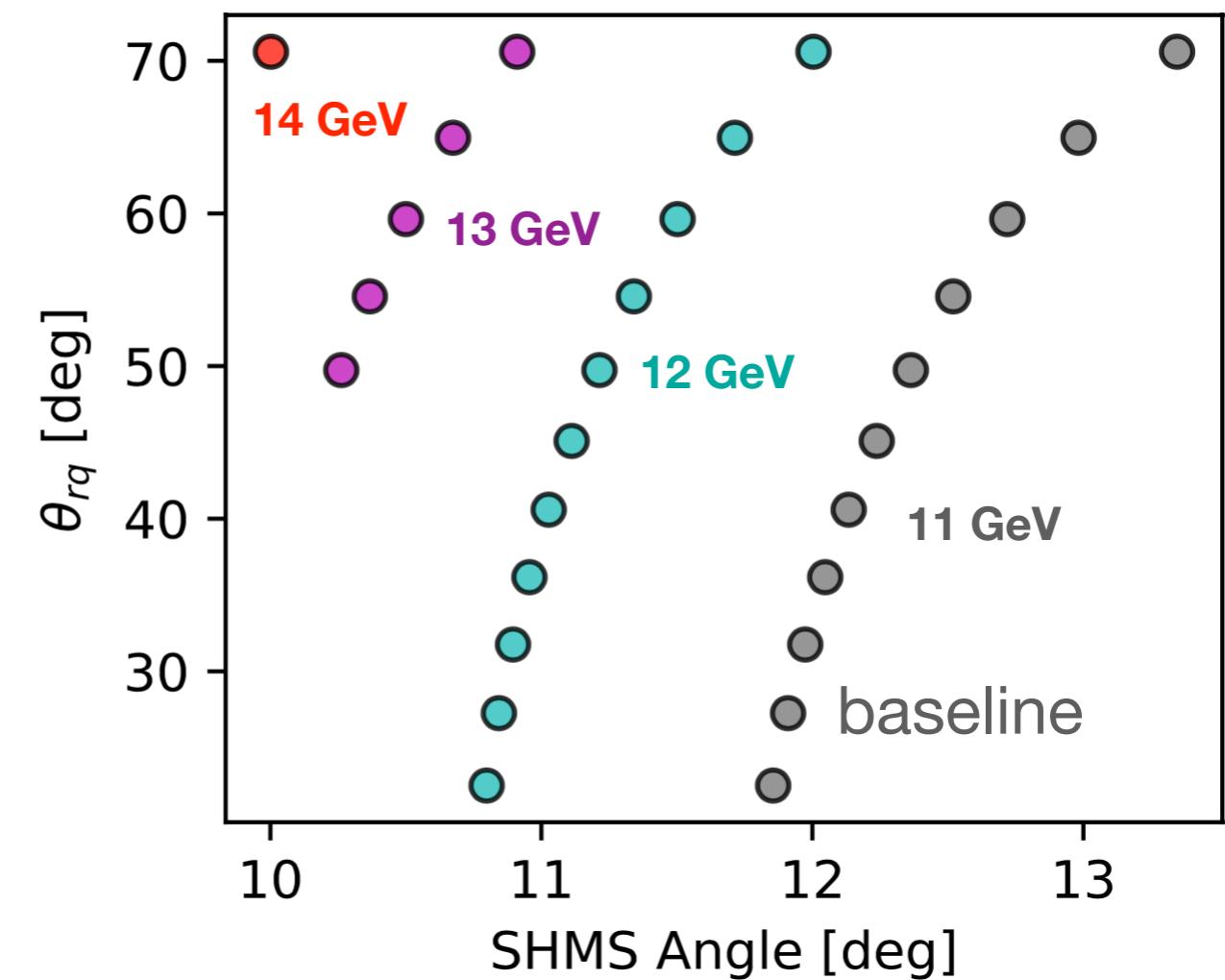
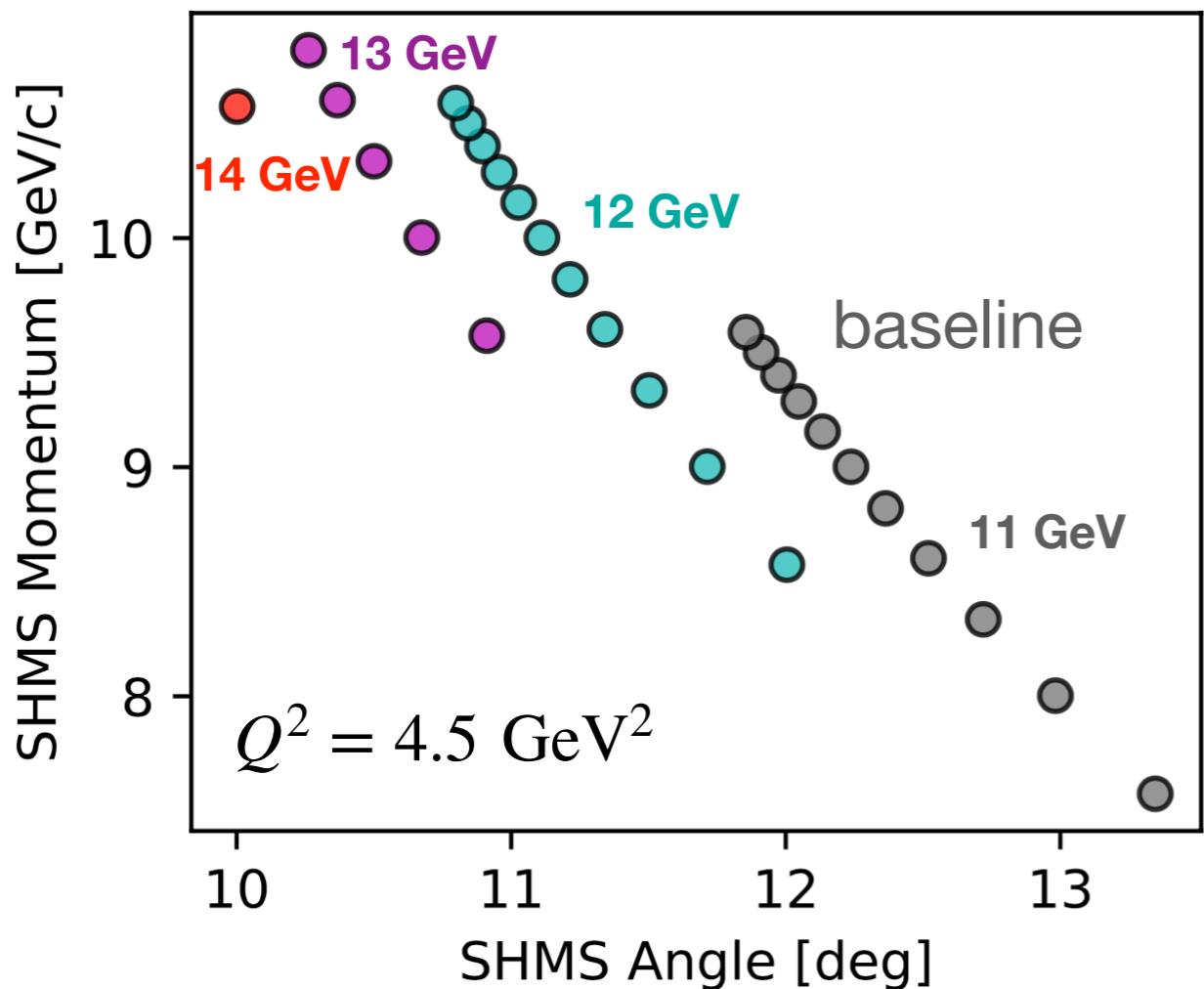
Baseline Kinematics

Baseline

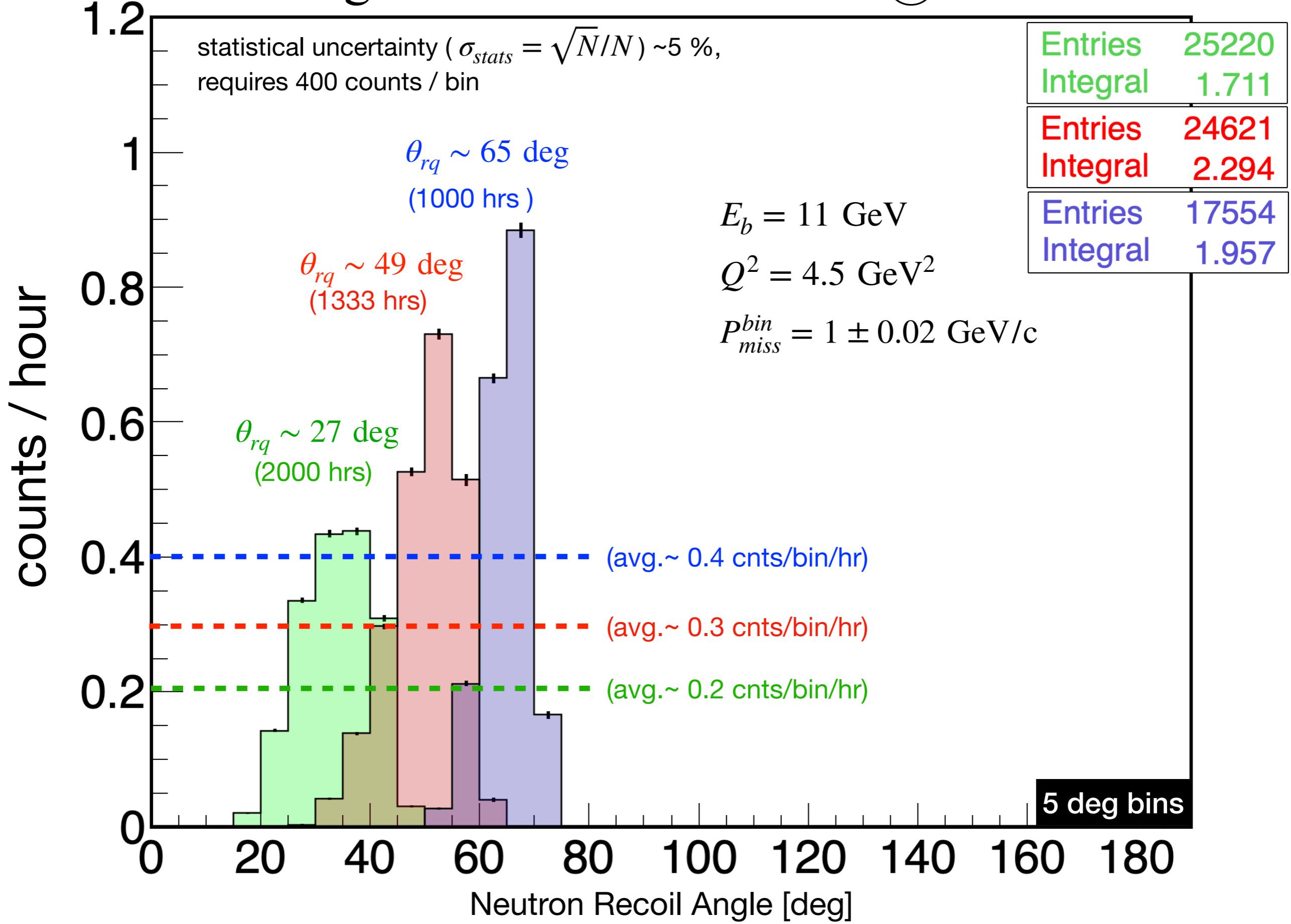
- use existing spectrometer + no beam energy upgrade

Energy Upgrade

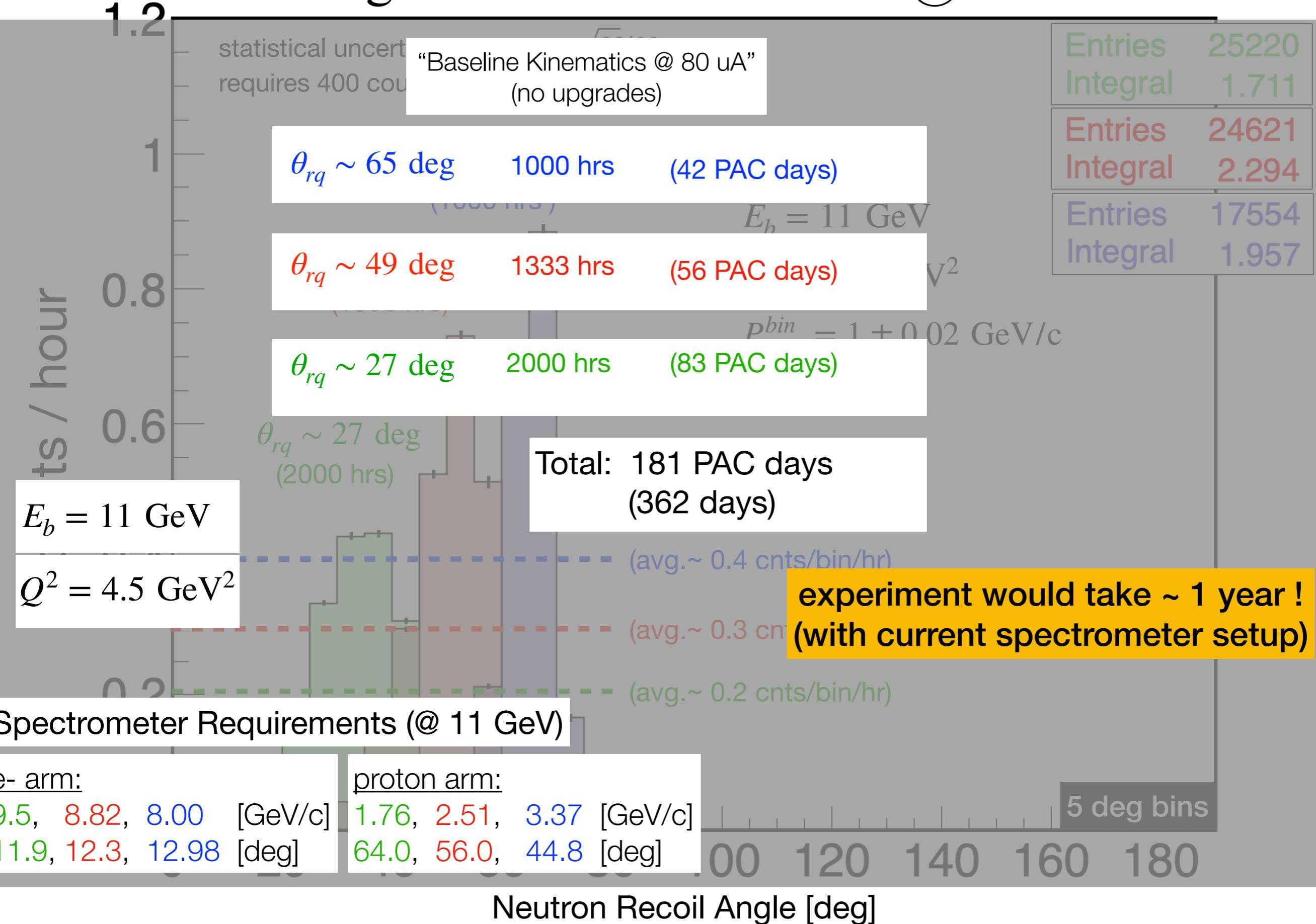
- full angular coverage up to 12 GeV
(limited by spectrometer momentum)



Angular Distribution Rates @ 80 uA



Angular Distribution Rates @ 80 uA



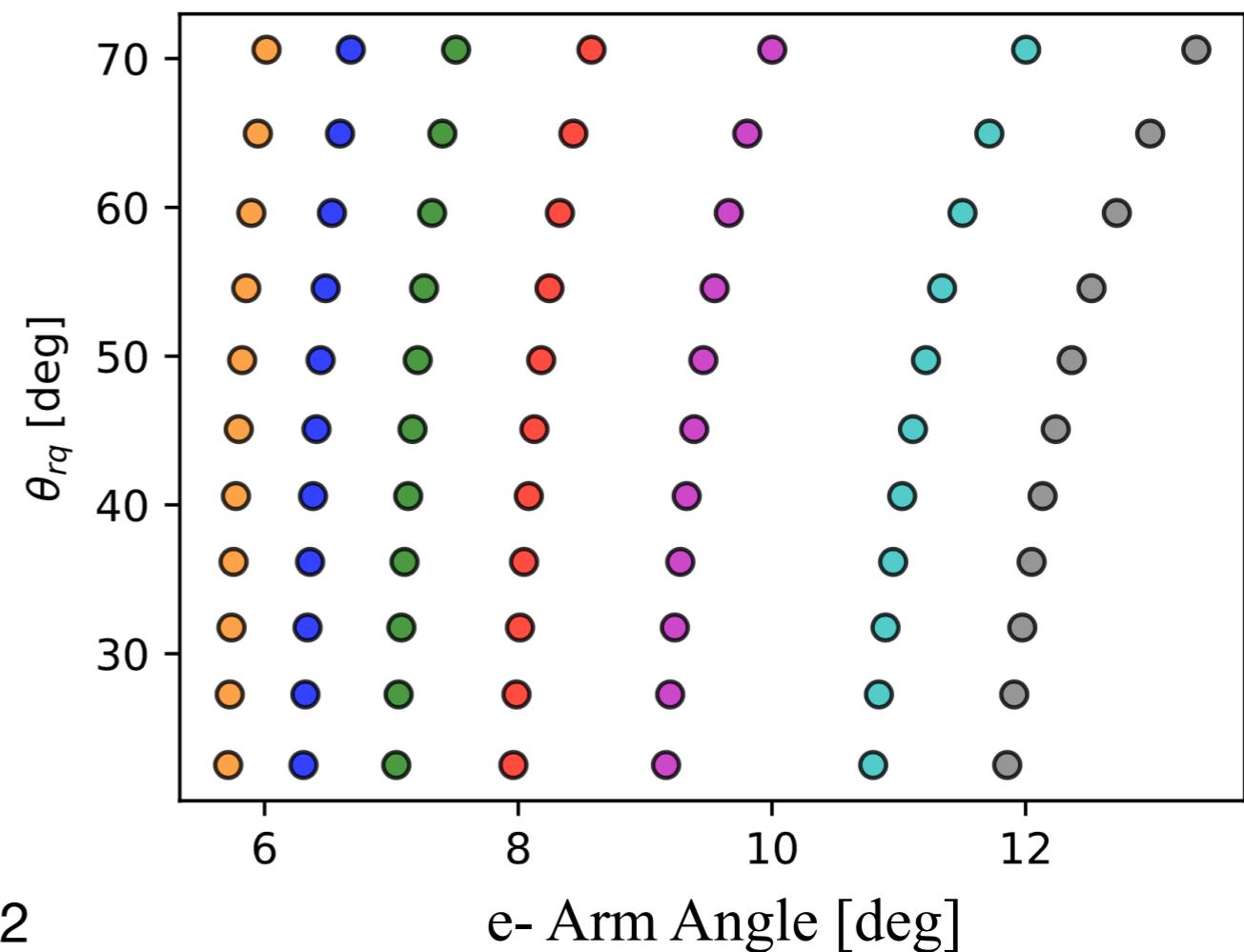
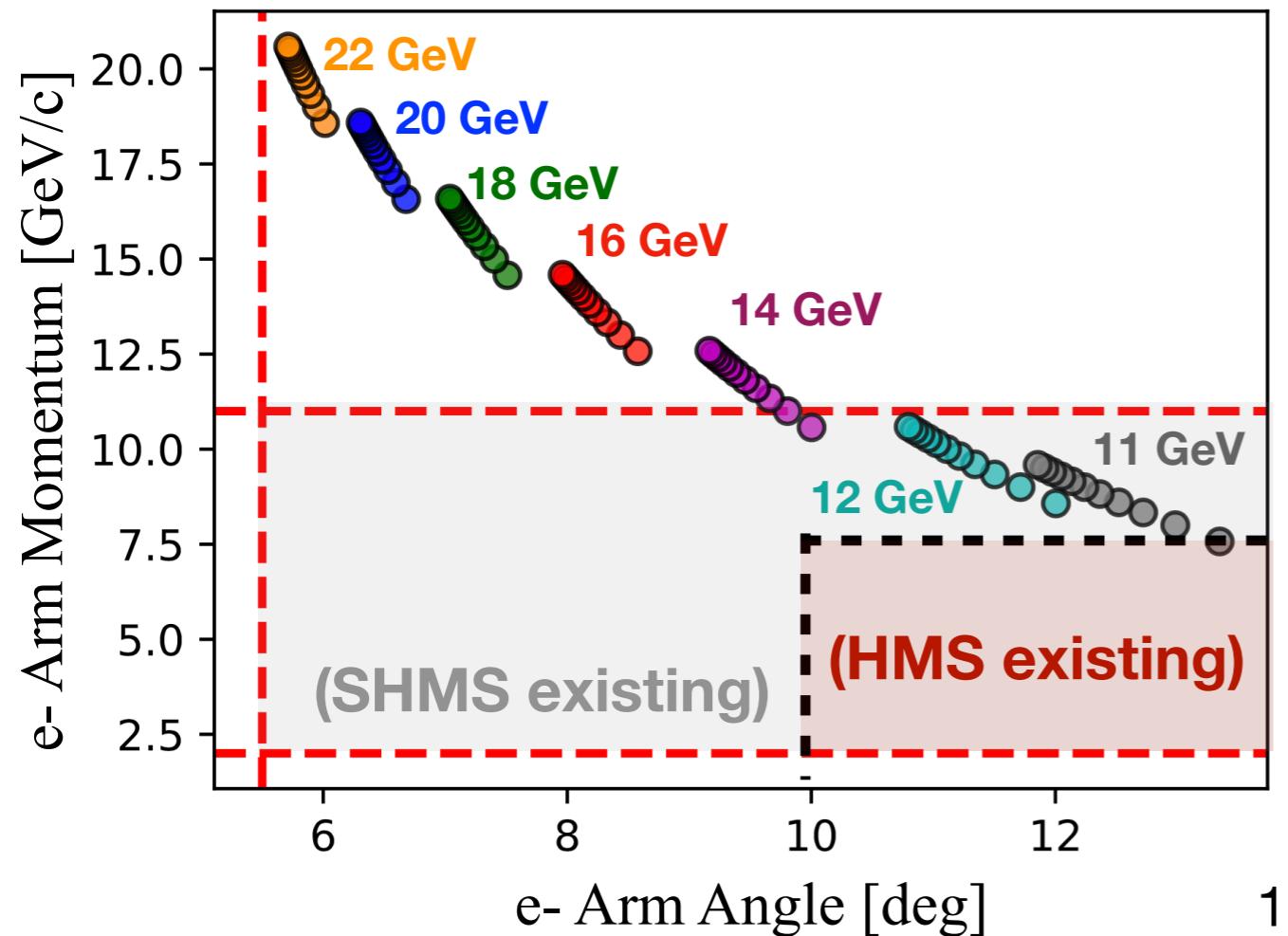
consider electron arm spectrometer upgrade

- If electron arm (HMS?) momentum upgraded up to ~ 20 GeV/c (*wishful thinking . . .*)

- ▶ full angular coverage becomes available at all energies (11-22 GeV) while keeping fixed at $Q^2 = 4.5$ GeV 2
- ▶ **problem:** small e- angles (< 10 deg) required for $E_b \geq 14$ GeV (need to modify HMS/beamline for lower angles)
- ▶ **problem:** SHMS cannot be further upgraded !



e- arm requirements



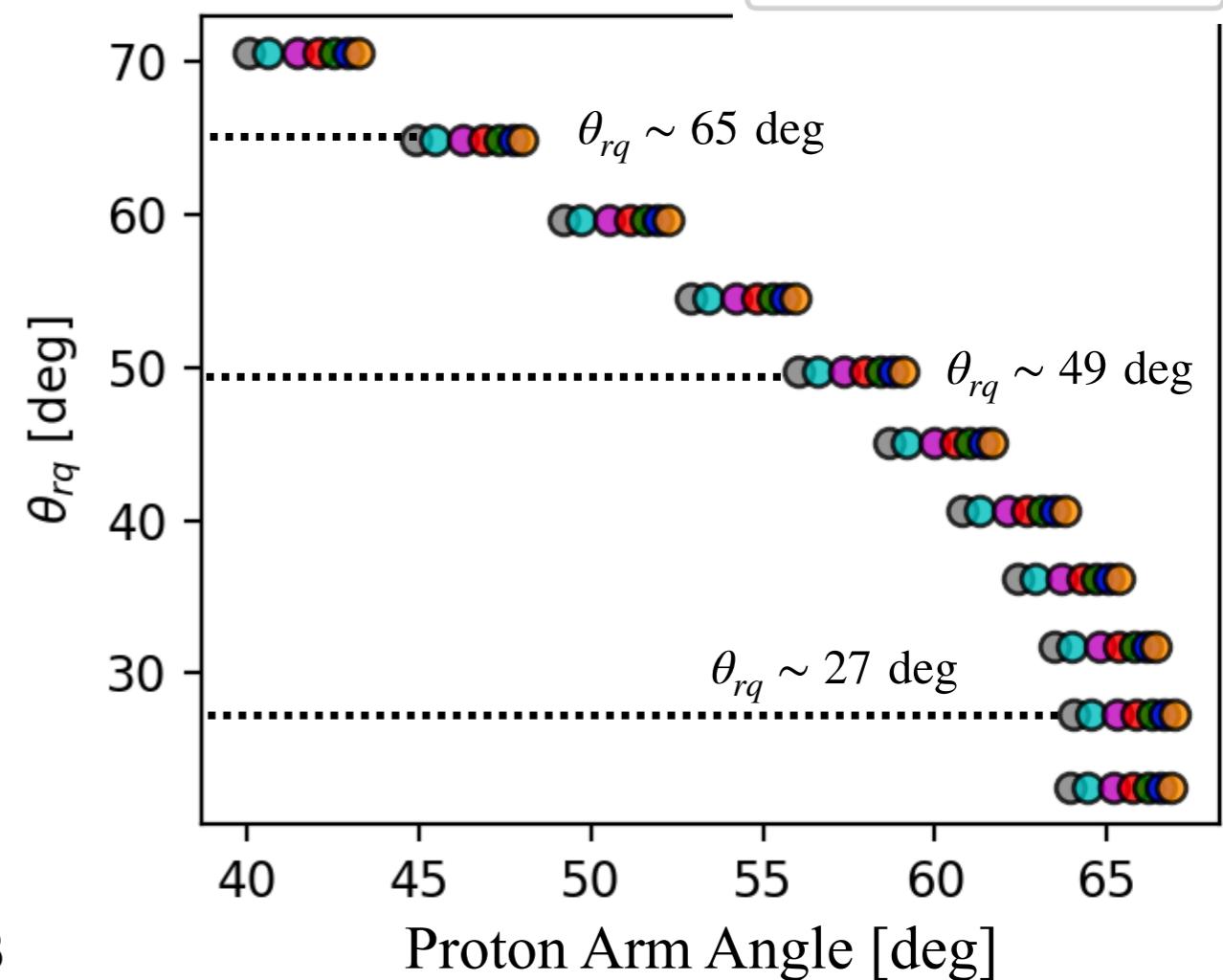
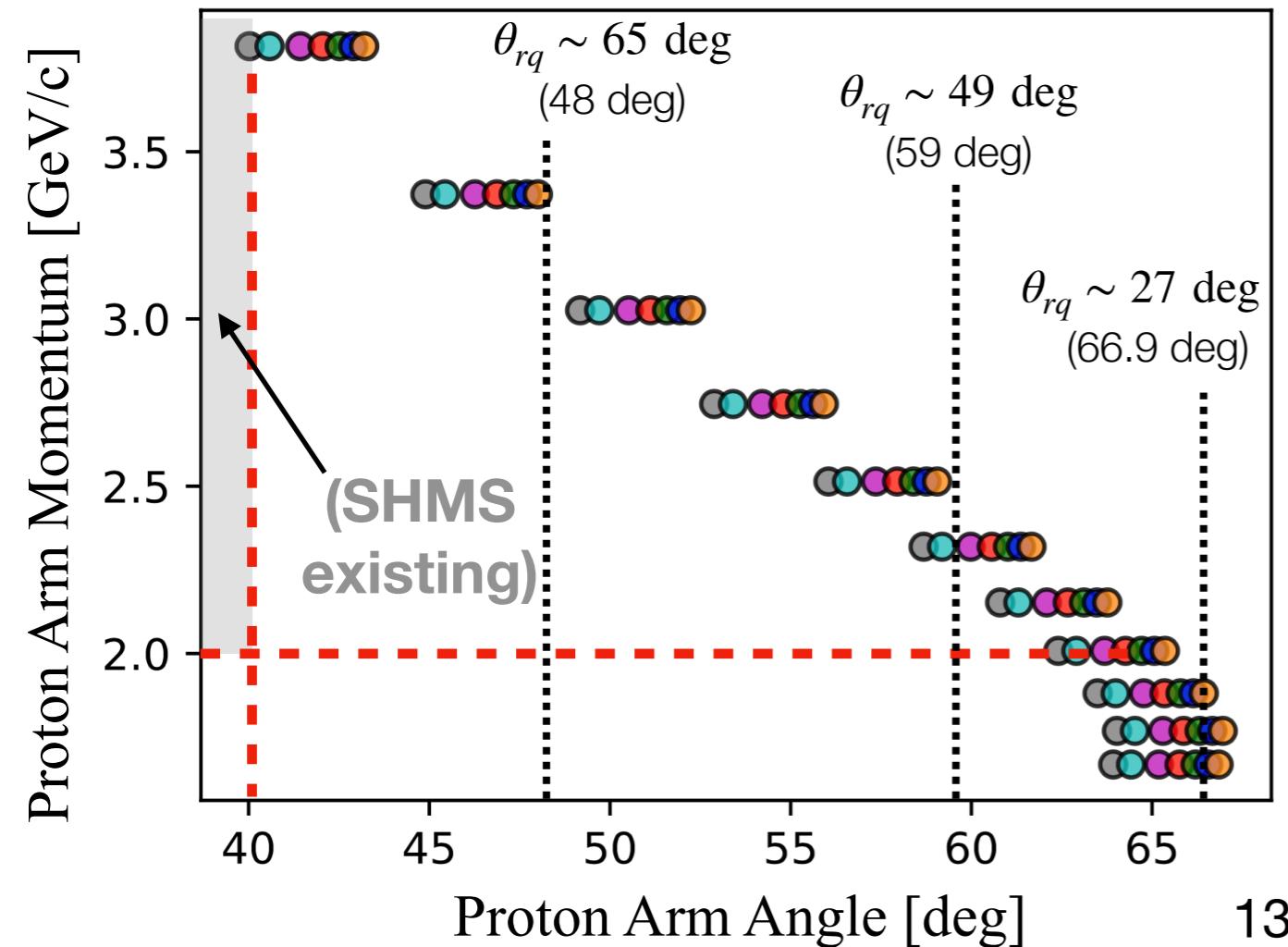
consider proton arm spectrometer upgrade

- SHMS **CANNOT** be used as proton arm

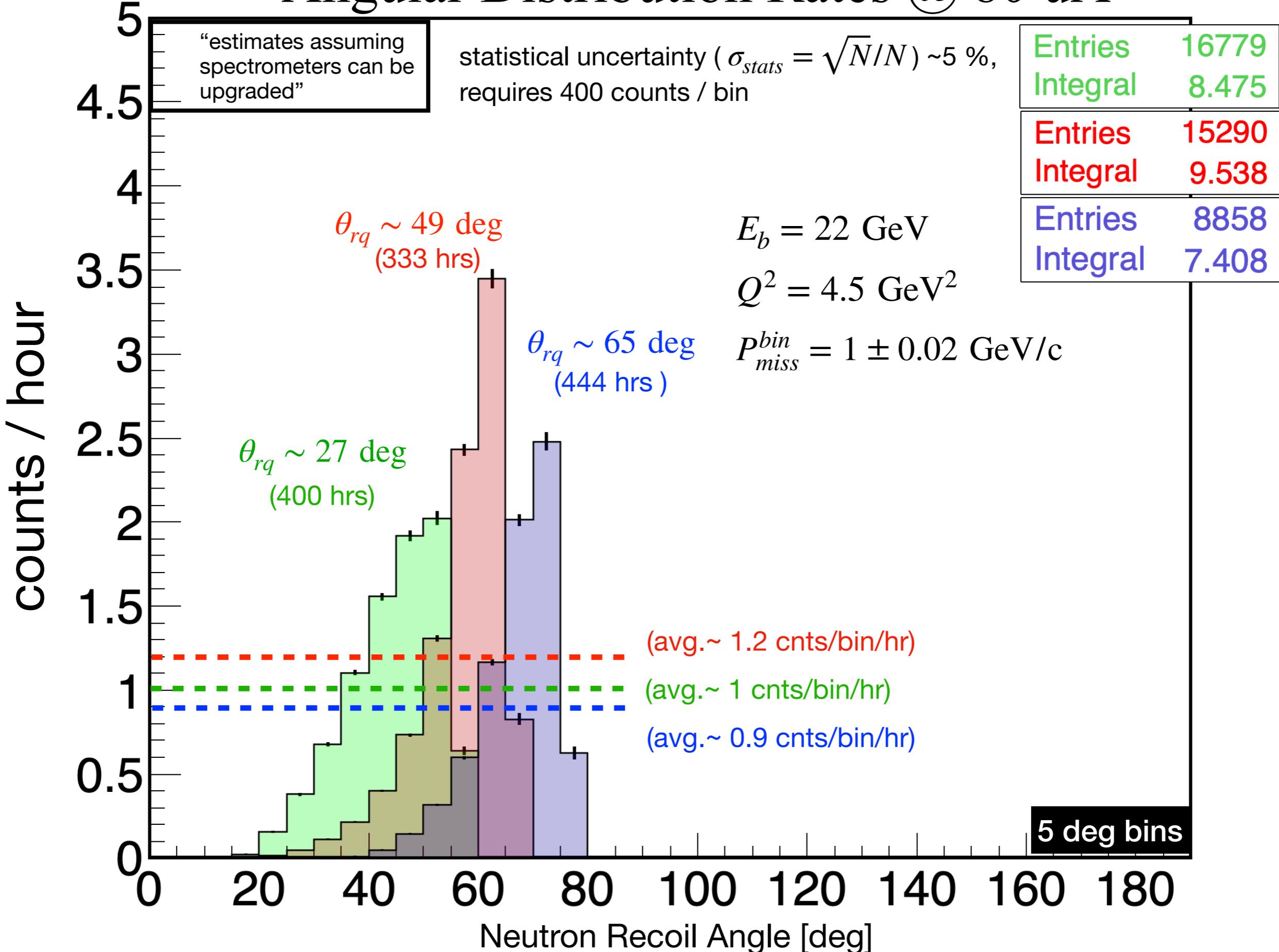
- ▶ limited mainly by the maximum angle (~40 deg) while keeping fixed at $Q^2 = 4.5 \text{ GeV}^2$
- ▶ Limited partially by momentum (cannot go below 2 GeV/c)
- ▶ **problem:** large proton angles (>40 deg) required for all kinematics
- ▶ **key question:** What can be used as proton arm? (BigBite, ... ?)



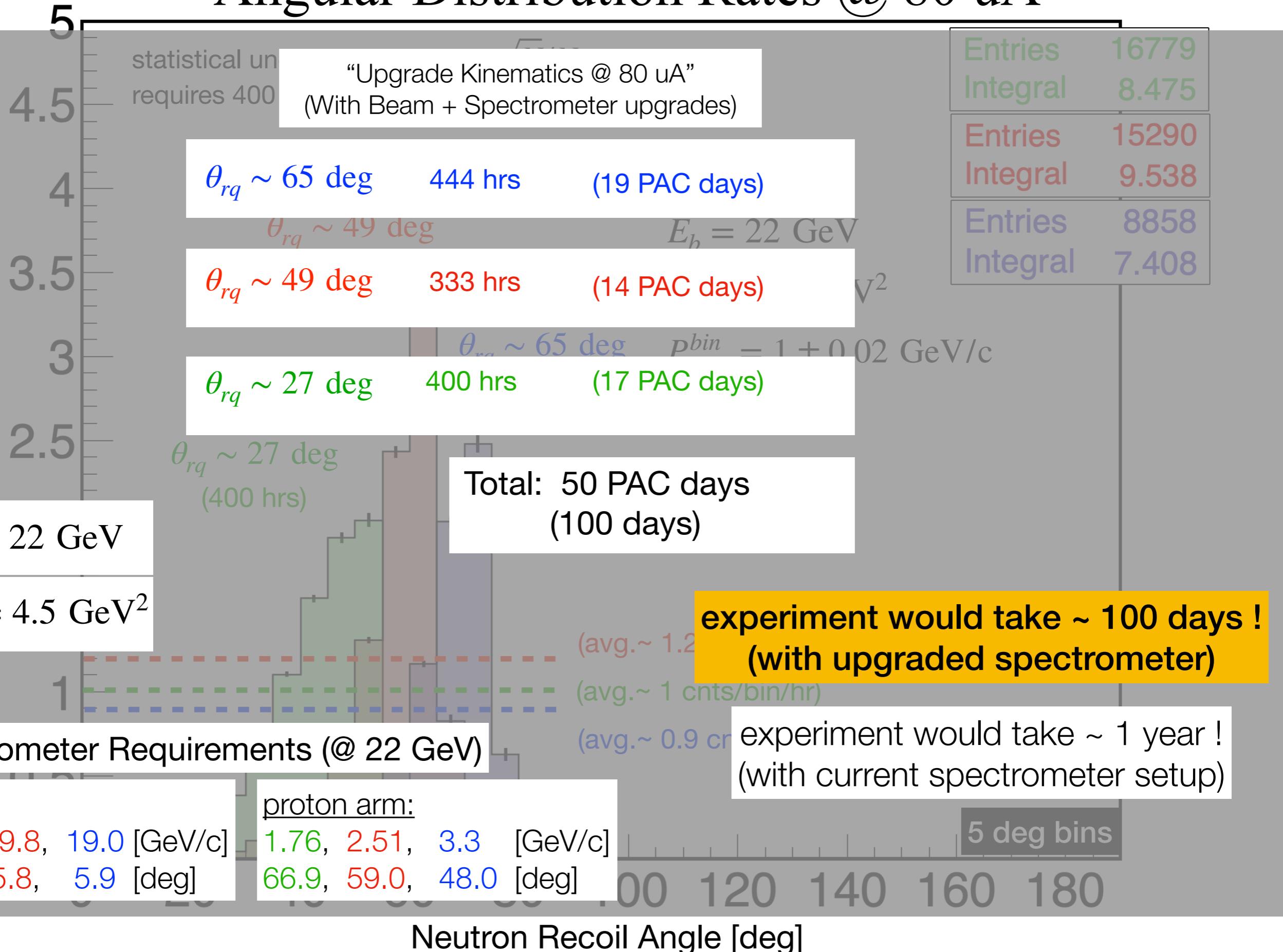
proton arm requirements



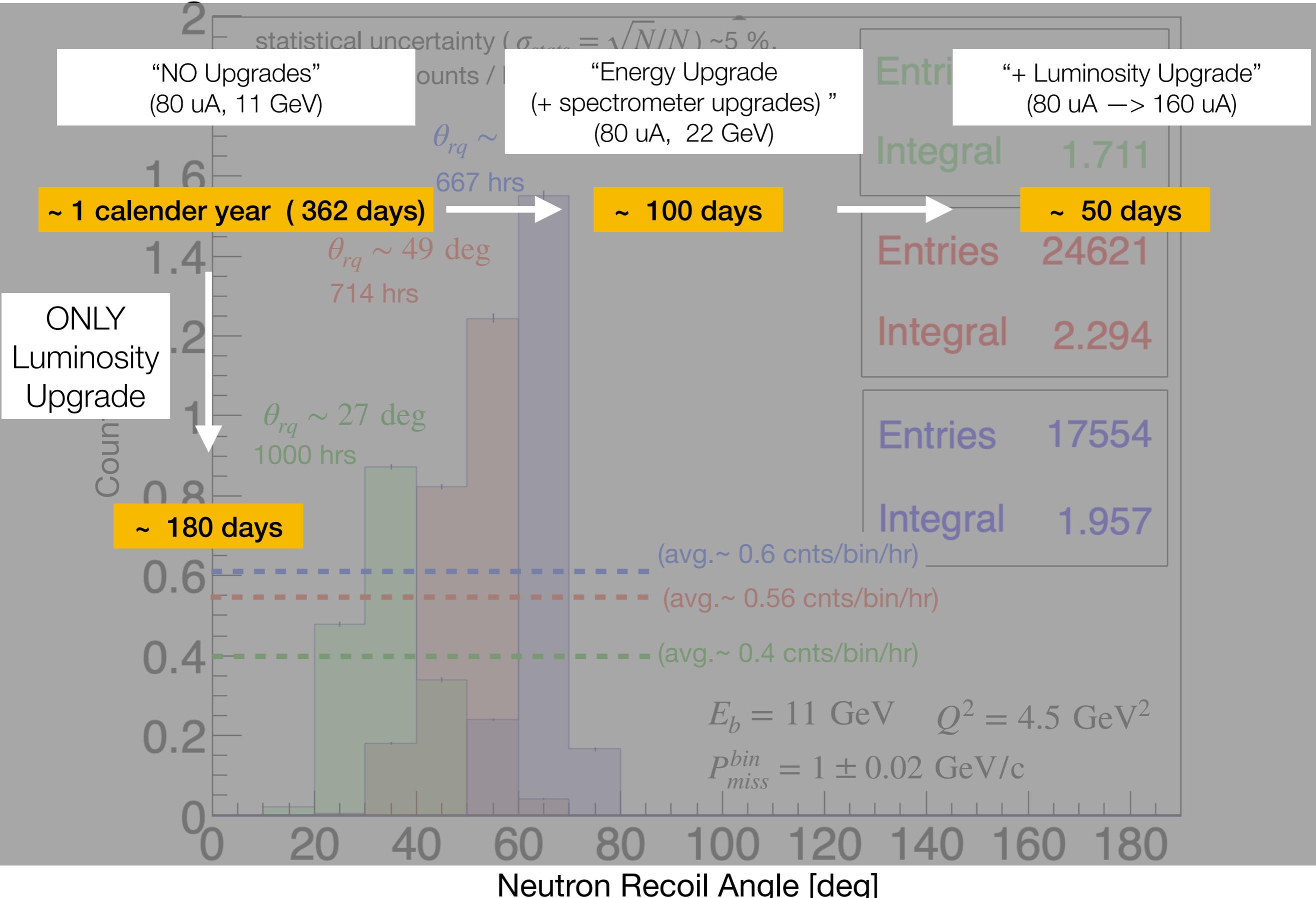
Angular Distribution Rates @ 80 uA



Angular Distribution Rates @ 80 uA



Consider Luminosity Upgrade (80 uA \rightarrow 160 uA)



Summary

- need to measure angular distributions (θ_{rq}) beyond $P_r \sim 500 \text{ MeV/c}$
- with existing spectrometer, can only take advantage up to 12 GeV (while fixing $Q^2 \sim 4.5 \text{ GeV}^2$)
 - ▶ can take ~ 1 calendar year for a $\sim 5\%$ measurement per angular bin $\theta_{rq} \pm 5 \text{ deg}$
(unrealistic beam time requirement)
- to take advantage of full 22 GeV upgrade requires spectrometer upgrades
 - ▶ Our measurements require:
 - electron: 18.5 - 20.5 GeV/c @ 5.7 - 6 deg
 - proton: 1.6 - 3.8 GeV/c @ 43 - 66.8 deg
 - ▶ Possible / impossible spectrometer upgrades:
 - HMS can be used as e- arm, if momentum can be upgraded (by how much? 15, 21 GeV/c ?)
(HMS is limited to 10 deg, can it be modified for lower angles?)
 - SHMS cannot be further upgraded (can't be used as e- arm),
and can't be used as proton arm (due to its limitation in angle and momentum)
 - Can Big Bite spectrometer be used as proton arm? Any other ideas?
- Could it be more cost-effective to propose a luminosity upgrade (80 uA \rightarrow 160 uA ?)

Acknowledgements

I would like to thank Drs. Werner Boeglin, Mark Jones and Misak Sargsian for their insight and useful discussions on this topic



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Back-Up Slides

“Baseline” Kinematics

(Without spectrometer upgrade)

| Eb | Pr | xbj | kf | th_e | Pf | th_p | q | th_q | th_rq | th_pq | Q2 |
|-------|--------|--------|--------|---------|--------|---------|--------|---------|---------|---------|--------|
| 11.00 | 1.0000 | 0.7000 | 7.5742 | 13.3458 | 3.8155 | 40.0272 | 4.0294 | 25.7154 | 70.5957 | 14.3118 | 4.5000 |
| 11.00 | 1.0000 | 0.8000 | 8.0025 | 12.9823 | 3.3729 | 44.8930 | 3.6722 | 29.3112 | 64.9598 | 15.5817 | 4.5000 |
| 11.00 | 1.0000 | 0.9000 | 8.3355 | 12.7192 | 3.0258 | 49.1733 | 3.4058 | 32.6062 | 59.6289 | 16.5670 | 4.5000 |
| 11.00 | 1.0000 | 1.0000 | 8.6020 | 12.5198 | 2.7456 | 52.8847 | 3.2016 | 35.6212 | 54.5680 | 17.2635 | 4.5000 |
| 11.00 | 1.0000 | 1.1000 | 8.8200 | 12.3635 | 2.5141 | 56.0471 | 3.0418 | 38.3777 | 49.7378 | 17.6694 | 4.5000 |
| 11.00 | 1.0000 | 1.2000 | 9.0016 | 12.2376 | 2.3192 | 58.6796 | 2.9143 | 40.8976 | 45.0951 | 17.7820 | 4.5000 |
| 11.00 | 1.0000 | 1.3000 | 9.1554 | 12.1341 | 2.1525 | 60.7968 | 2.8112 | 43.2022 | 40.5911 | 17.5946 | 4.5000 |
| 11.00 | 1.0000 | 1.4000 | 9.2871 | 12.0474 | 2.0079 | 62.4047 | 2.7265 | 45.3119 | 36.1692 | 17.0928 | 4.5000 |
| 11.00 | 1.0000 | 1.5000 | 9.4013 | 11.9737 | 1.8811 | 63.4940 | 2.6563 | 47.2456 | 31.7588 | 16.2483 | 4.5000 |
| 11.00 | 1.0000 | 1.6000 | 9.5012 | 11.9104 | 1.7688 | 64.0296 | 2.5974 | 49.0208 | 27.2617 | 15.0088 | 4.5000 |
| 11.00 | 1.0000 | 1.7000 | 9.5894 | 11.8553 | 1.6683 | 63.9256 | 2.5475 | 50.6533 | 22.5203 | 13.2722 | 4.5000 |

Not so “Optimum” Kinematics (Without spectrometer upgrade)

| Eb | Pr | xbj | kf | th_e | Pf | th_p | q | th_q | th_rq | th_pq | Q2 |
|-------|--------|--------|---------|---------|--------|---------|--------|---------|---------|--------|--------|
| 14.00 | 1.0000 | 0.8000 | 7.6719 | 17.1034 | 6.7668 | 26.6915 | 7.0388 | 18.6960 | 70.2595 | 7.9955 | 9.5000 |
| 14.00 | 1.0000 | 0.9000 | 8.3750 | 16.3646 | 6.0562 | 30.1773 | 6.4141 | 21.5852 | 64.7956 | 8.5921 | 9.5000 |
| 14.00 | 1.0000 | 1.0000 | 8.9375 | 15.8378 | 5.4863 | 33.3306 | 5.9270 | 24.3016 | 59.4273 | 9.0289 | 9.5000 |
| 14.00 | 1.0000 | 1.1000 | 9.3977 | 15.4427 | 5.0188 | 36.1481 | 5.5390 | 26.8574 | 54.1203 | 9.2907 | 9.5000 |
| 14.00 | 1.0000 | 1.2000 | 9.7813 | 15.1351 | 4.6281 | 38.6225 | 5.2247 | 29.2618 | 48.8297 | 9.3608 | 9.5000 |
| 14.00 | 1.0000 | 1.3000 | 10.1058 | 14.8888 | 4.2964 | 40.7411 | 4.9664 | 31.5227 | 43.4944 | 9.2184 | 9.5000 |
| 14.00 | 1.0000 | 1.4000 | 10.3839 | 14.6869 | 4.0113 | 42.4818 | 4.7514 | 33.6479 | 38.0255 | 8.8339 | 9.5000 |
| 14.00 | 1.0000 | 1.5000 | 10.6250 | 14.5184 | 3.7633 | 43.8036 | 4.5706 | 35.6451 | 32.2798 | 8.1585 | 9.5000 |
| 14.00 | 1.0000 | 1.6000 | 10.8359 | 14.3756 | 3.5455 | 44.6216 | 4.4172 | 37.5217 | 25.9904 | 7.0999 | 9.5000 |

*NOTE: baseline kinematics (Eb=11 GeV) are actually better than optimal (Eb=14 GeV) since, optimal requires Q2 → 9.5 GeV^2, which significantly reduces count rates

Kinematics

(With spectrometer upgrade)

electron proton

| Eb | Pr | xbj | kf | th_e | Pf | th_p | q | th_q | th_rq | th_pq | Q2 |
|--------------|--------|--------|---------|--------|--------|---------|--------|---------|---------|---------|--------|
| 22.00 | 1.0000 | 0.7000 | 18.5742 | 6.0154 | 3.8155 | 43.1983 | 4.0294 | 28.8865 | 70.5957 | 14.3118 | 4.5000 |
| 22.00 | 1.0000 | 0.8000 | 19.0025 | 5.9471 | 3.3729 | 48.0037 | 3.6722 | 32.4219 | 64.9598 | 15.5817 | 4.5000 |
| 22.00 | 1.0000 | 0.9000 | 19.3355 | 5.8956 | 3.0258 | 52.2394 | 3.4058 | 35.6724 | 59.6289 | 16.5670 | 4.5000 |
| 22.00 | 1.0000 | 1.0000 | 19.6020 | 5.8554 | 2.7456 | 55.9167 | 3.2016 | 38.6531 | 54.5680 | 17.2635 | 4.5000 |
| 22.00 | 1.0000 | 1.1000 | 19.8200 | 5.8231 | 2.5141 | 59.0519 | 3.0418 | 41.3825 | 49.7378 | 17.6694 | 4.5000 |
| 22.00 | 1.0000 | 1.2000 | 20.0016 | 5.7966 | 2.3192 | 61.6624 | 2.9143 | 43.8804 | 45.0951 | 17.7820 | 4.5000 |
| 22.00 | 1.0000 | 1.3000 | 20.1554 | 5.7744 | 2.1525 | 63.7614 | 2.8112 | 46.1668 | 40.5911 | 17.5946 | 4.5000 |
| 22.00 | 1.0000 | 1.4000 | 20.2871 | 5.7556 | 2.0079 | 65.3539 | 2.7265 | 48.2611 | 36.1692 | 17.0928 | 4.5000 |
| 22.00 | 1.0000 | 1.5000 | 20.4013 | 5.7394 | 1.8811 | 66.4301 | 2.6563 | 50.1818 | 31.7588 | 16.2483 | 4.5000 |
| 22.00 | 1.0000 | 1.6000 | 20.5012 | 5.7254 | 1.7688 | 66.9544 | 2.5974 | 51.9456 | 27.2617 | 15.0088 | 4.5000 |
| 22.00 | 1.0000 | 1.7000 | 20.5894 | 5.7132 | 1.6683 | 66.8405 | 2.5475 | 53.5683 | 22.5203 | 13.2722 | 4.5000 |

electron arm requirements

P: 18.5-20.5 GeV/c

angle: 5.7 - 6 deg

proton arm requirements

P: 1.6-3.8 GeV/c

angle: 43 - 66.8 deg

Integrated Count Rates Estimates @ 80 uA

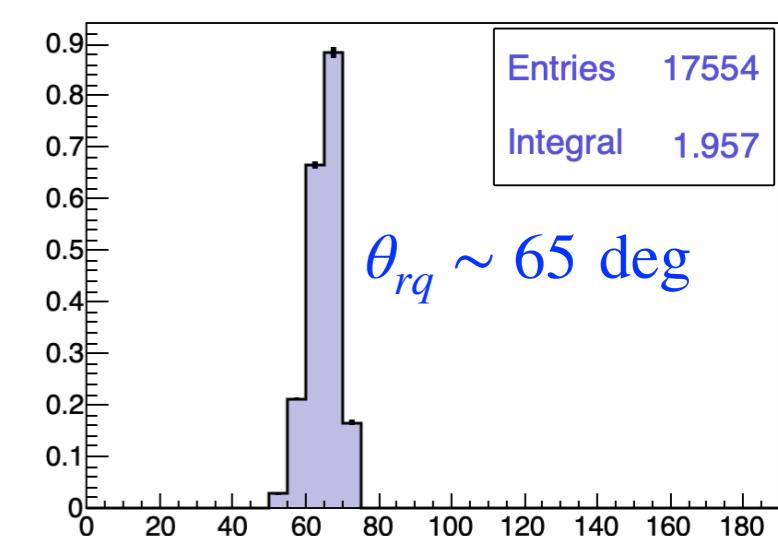
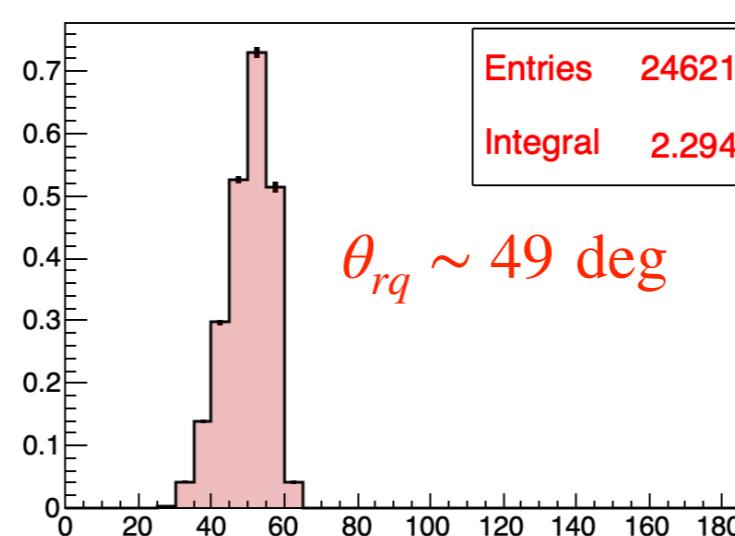
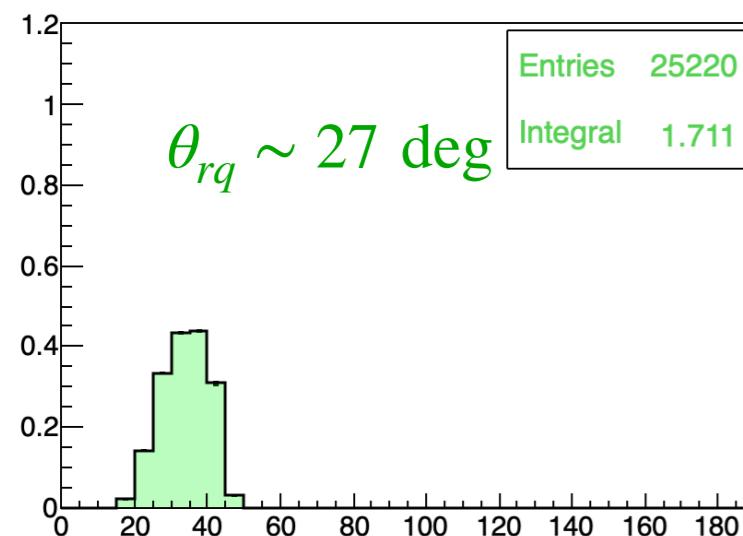
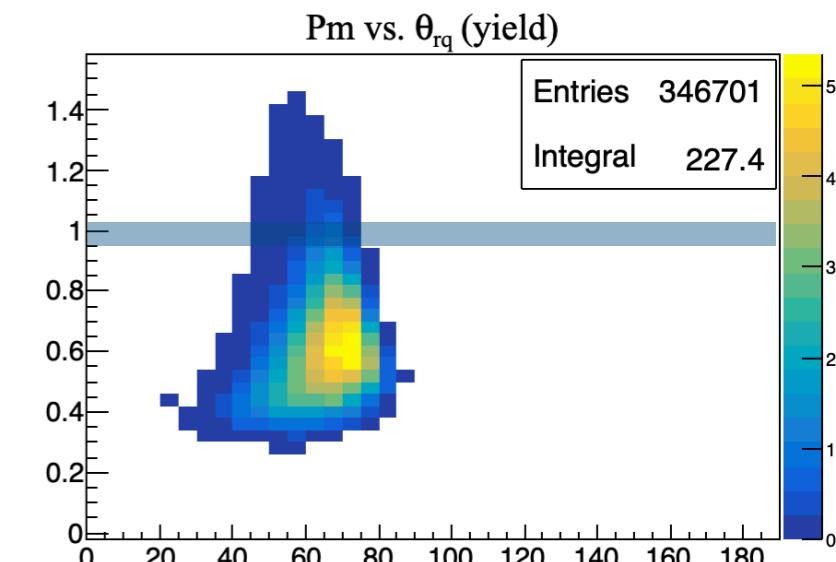
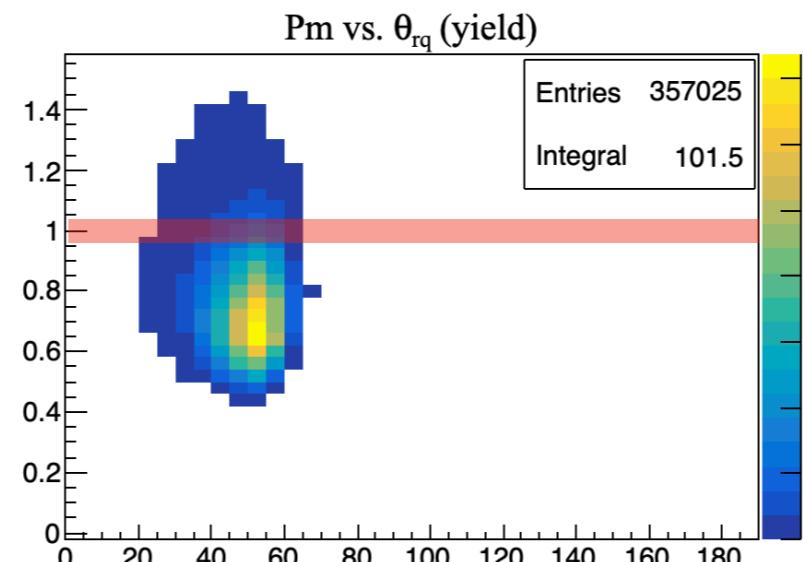
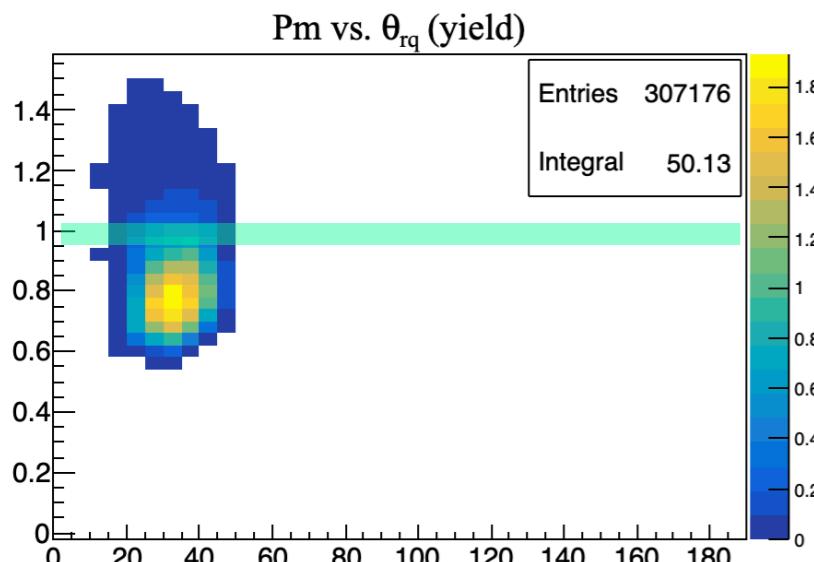
5 deg bins

$E_b = 11 \text{ GeV}$

$Q^2 = 4.5 \text{ GeV}^2$

$P_{miss}^{bin} = 1 \pm 0.02 \text{ GeV/c}$

2D Projections



Integrated Count Rates Estimates @ 80 uA

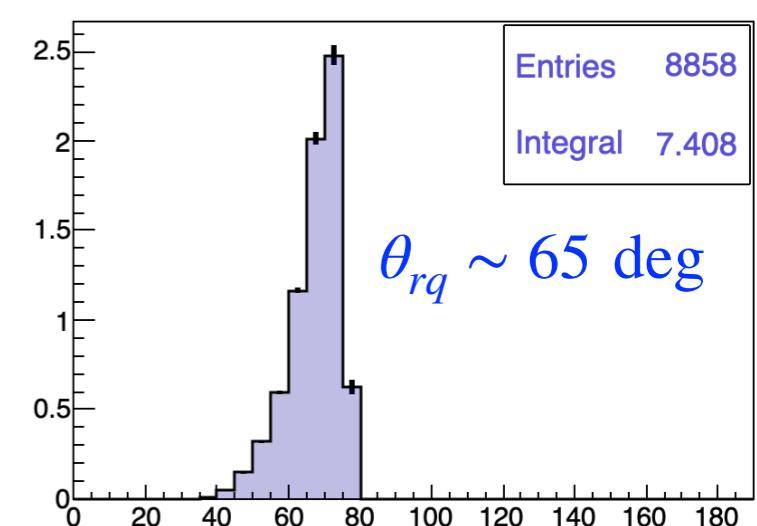
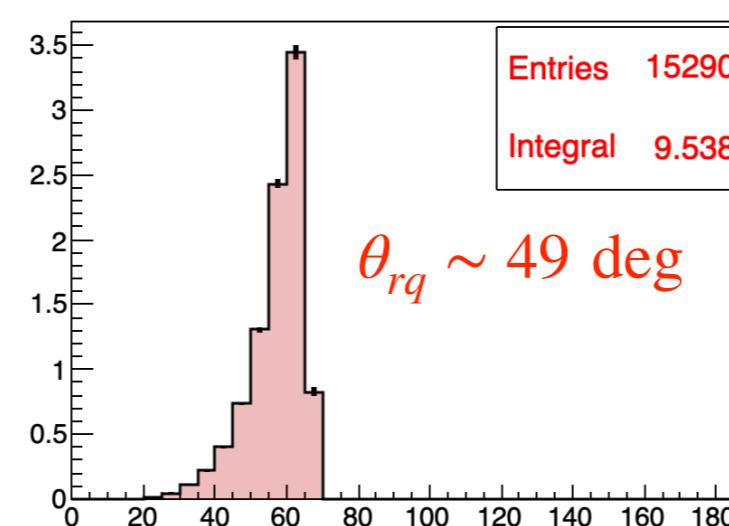
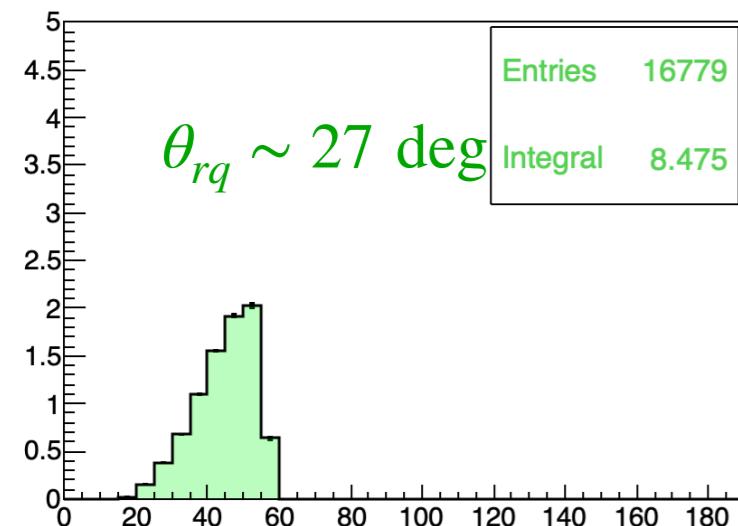
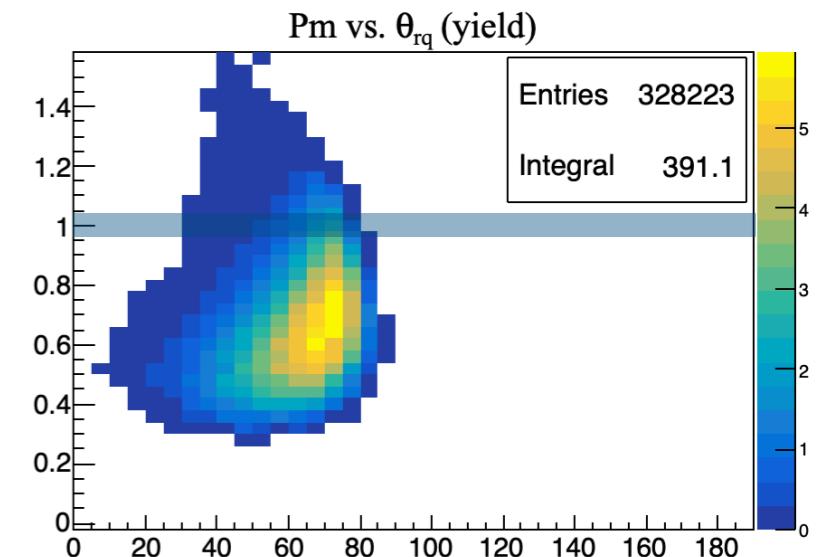
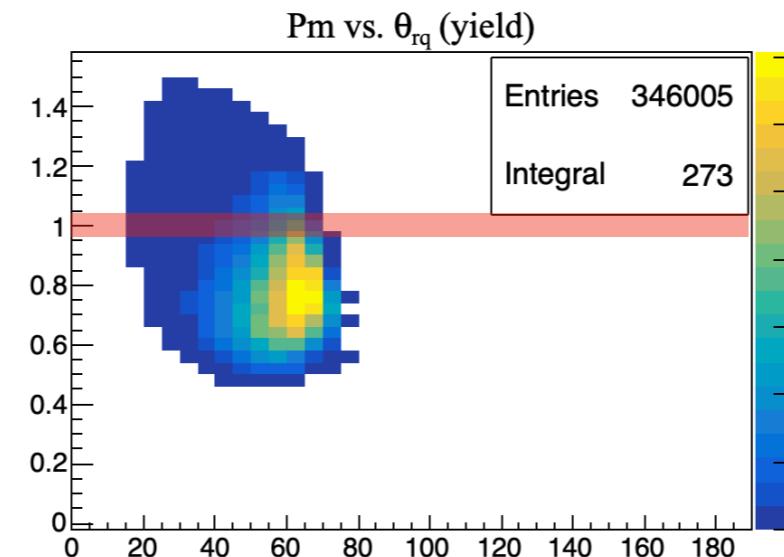
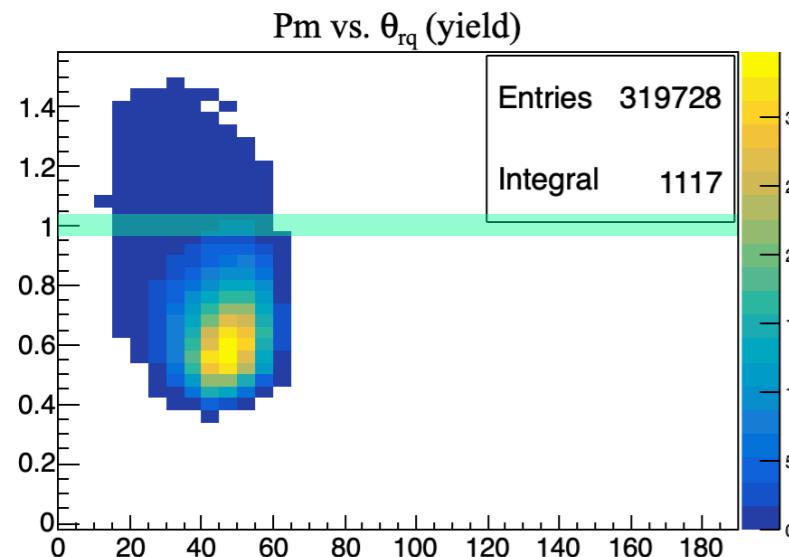
5 deg bins

$E_b = 22 \text{ GeV}$

$Q^2 = 4.5 \text{ GeV}^2$

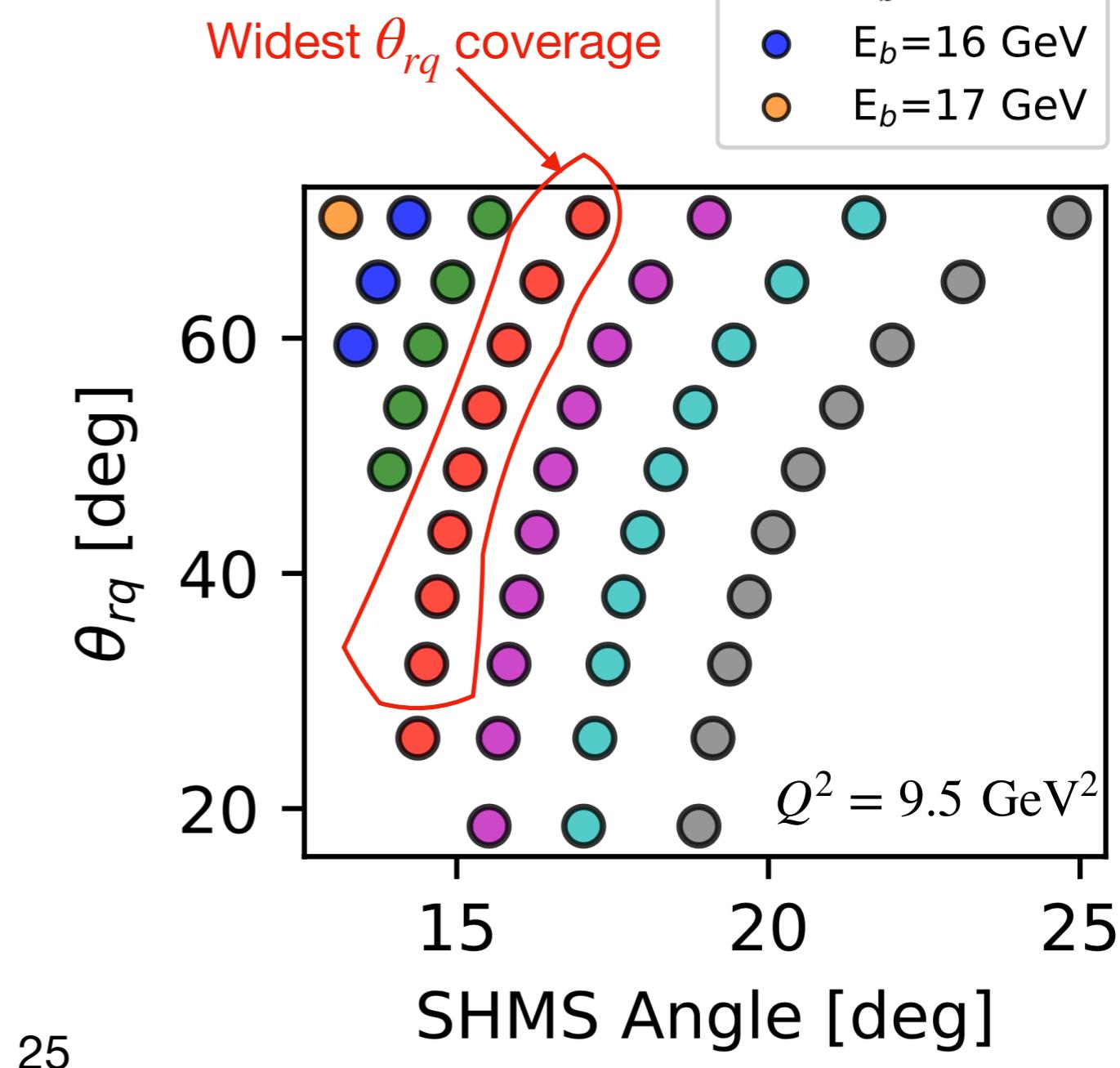
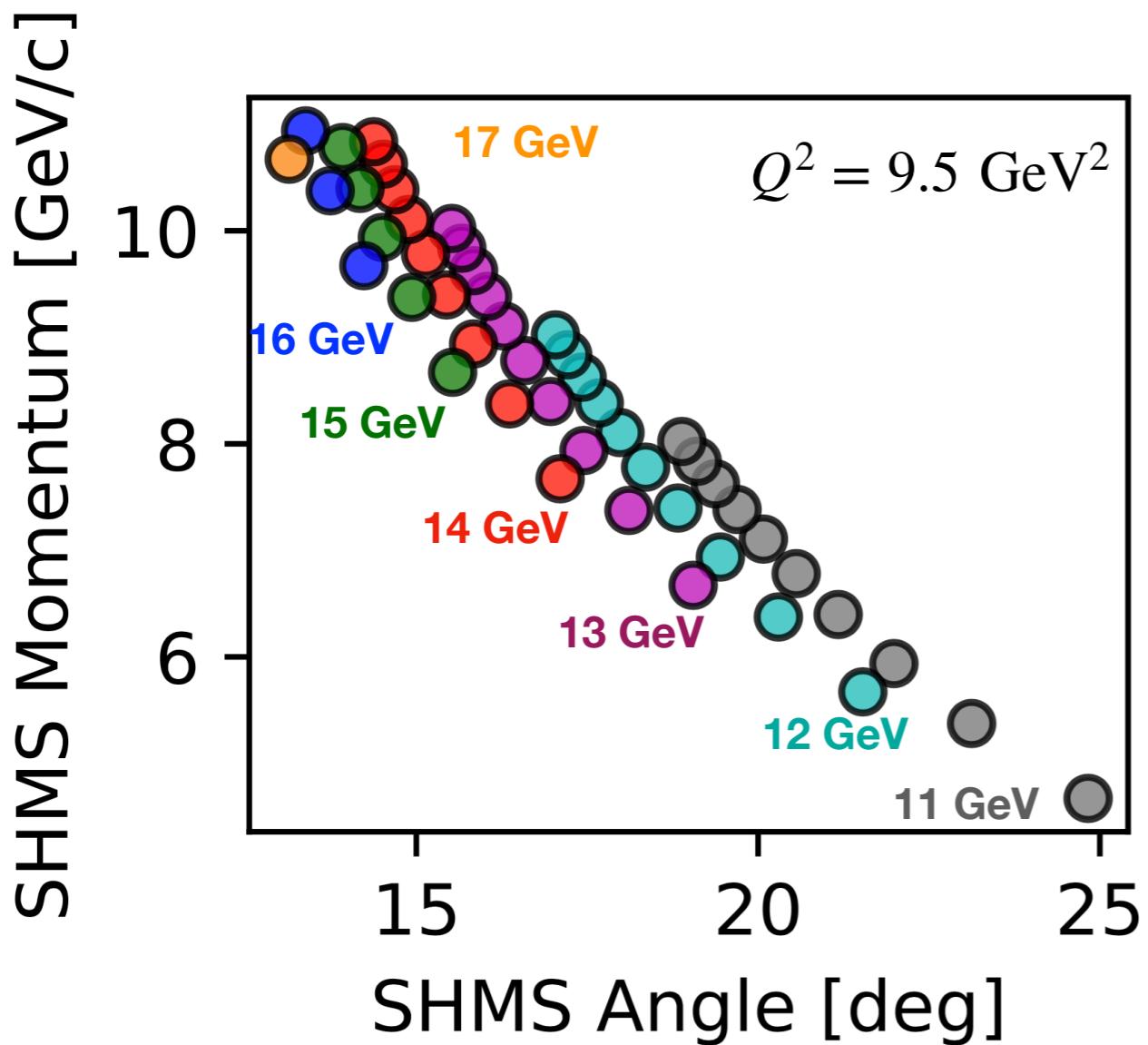
$P_{miss}^{bin} = 1 \pm 0.02 \text{ GeV/c}$

2D Projections

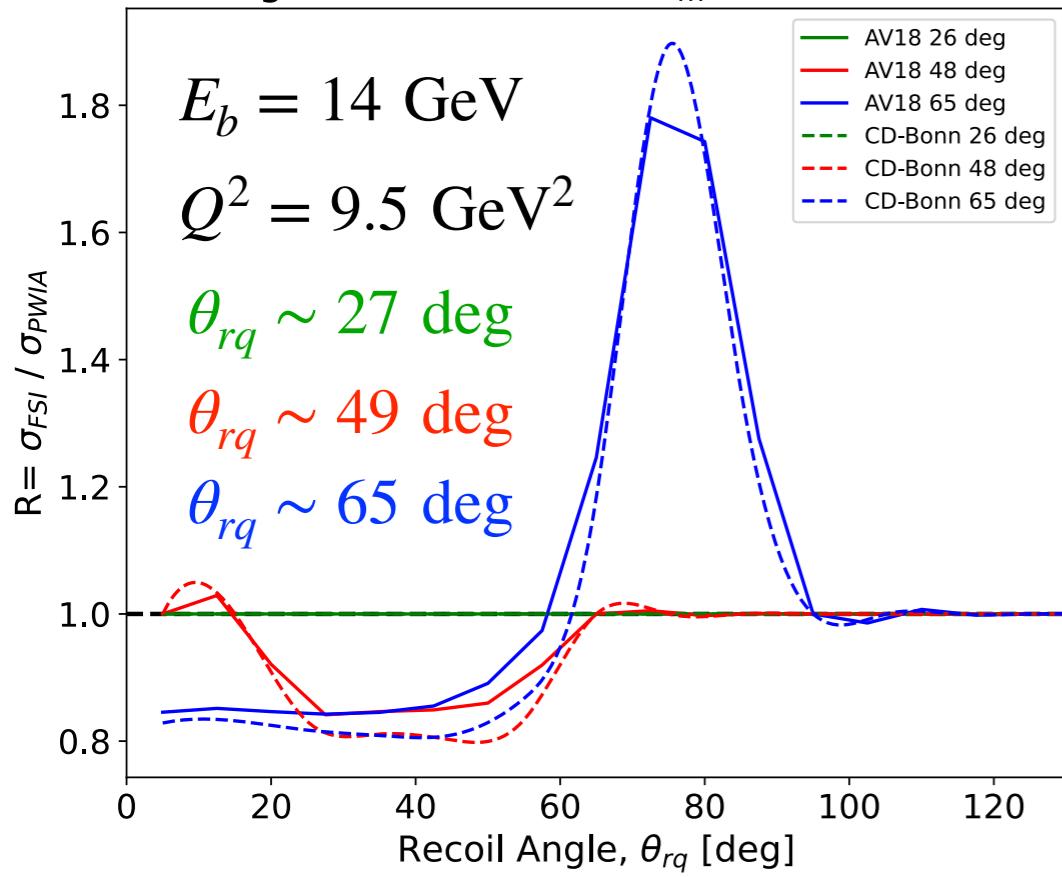
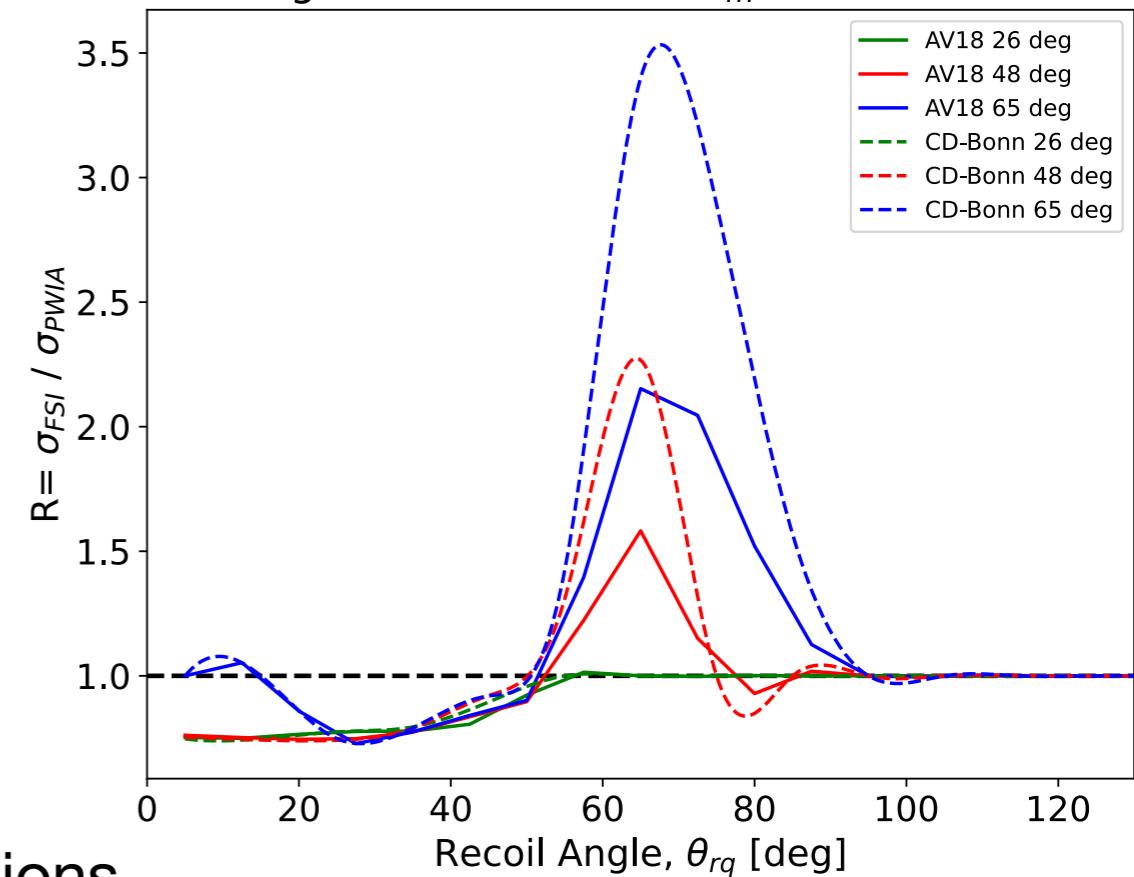


Find Optimum Kinematics (at higher Q²)

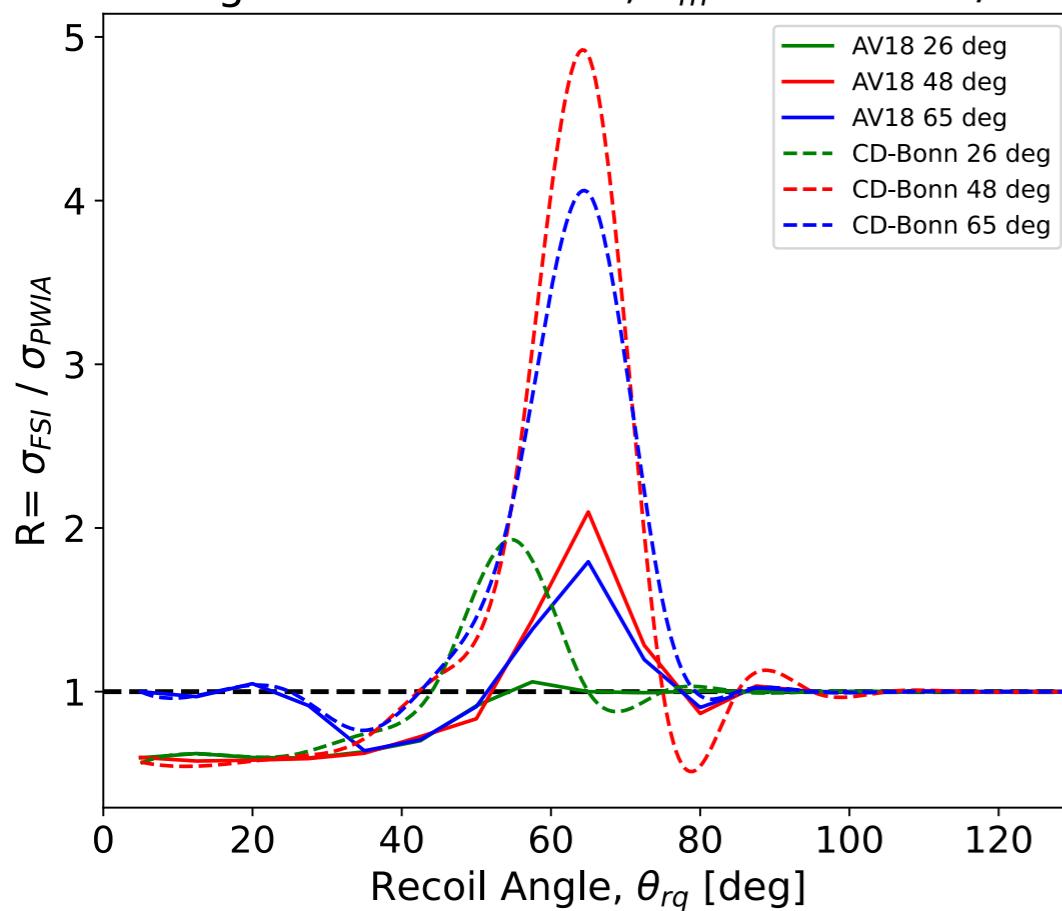
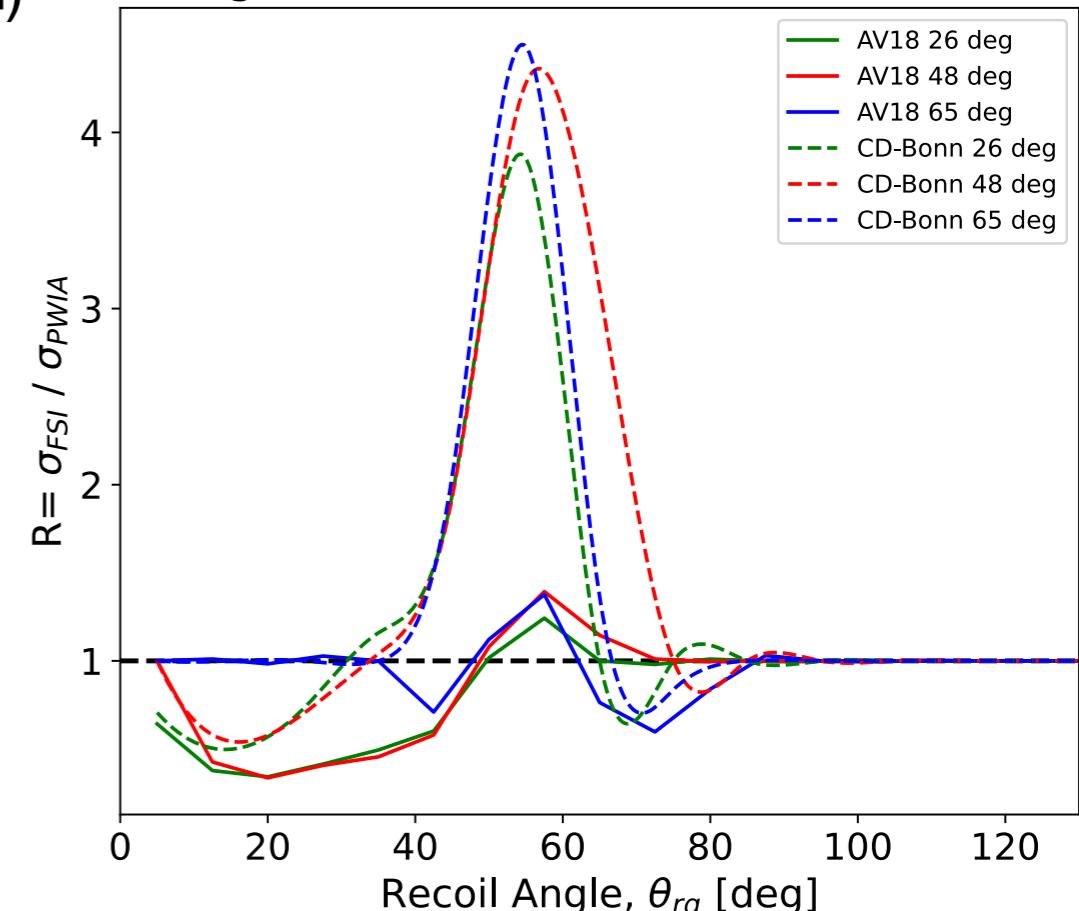
- optimum kinematics @ $E_b = 14 \text{ GeV}$, $Q^2 = 9.5 \text{ GeV}^2$
 (optimum \rightarrow widest angular coverage at highest possible beam energy)
 - ▶ increase in beam energy requires increase in Q^2
 in order to keep electron spectrometer within limits
 - ▶ **problem:** higher $Q^2 \rightarrow$ reduced count rates



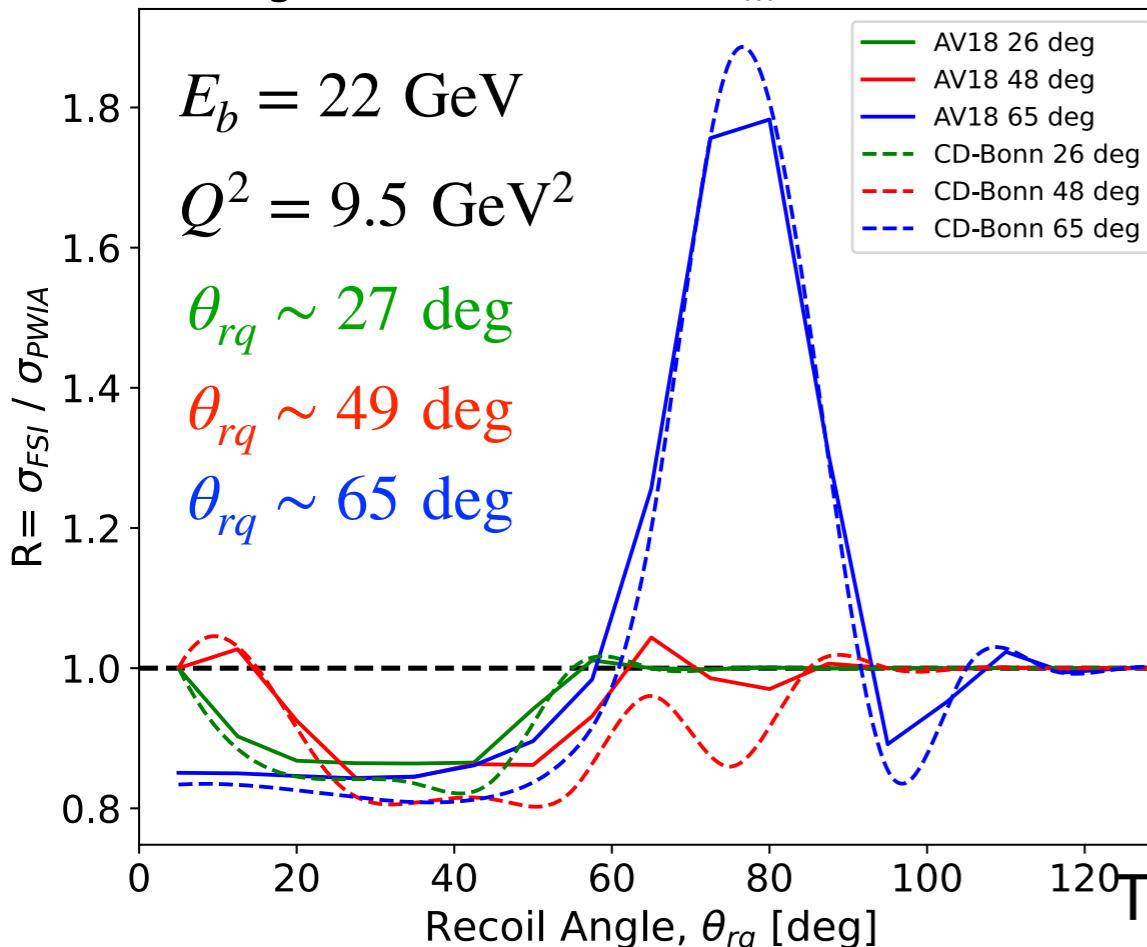
| | |
|---|------------------------|
| ● | $E_b = 11 \text{ GeV}$ |
| ● | $E_b = 12 \text{ GeV}$ |
| ● | $E_b = 13 \text{ GeV}$ |
| ● | $E_b = 14 \text{ GeV}$ |
| ● | $E_b = 15 \text{ GeV}$ |
| ● | $E_b = 16 \text{ GeV}$ |
| ● | $E_b = 17 \text{ GeV}$ |

Angular Distributions, $P_m = 0.40 \text{ GeV}/c$ Angular Distributions, $P_m = 0.60 \text{ GeV}/c$ 

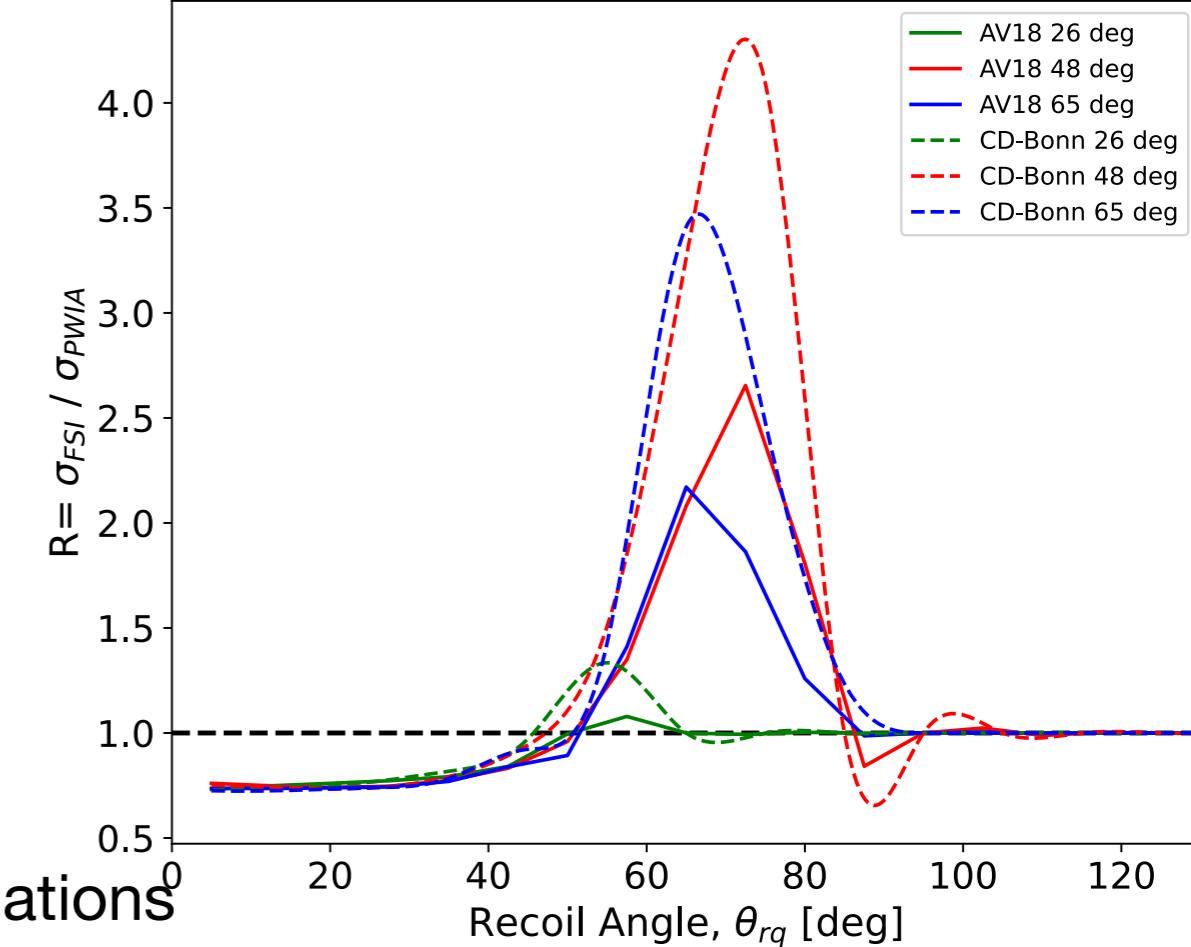
Theory calculations
of angular distributions
(by. M. Sargsian)

Angular Distributions, $P_m = 0.80 \text{ GeV}/c$ Angular Distributions, $P_m = 1.00 \text{ GeV}/c$ 

Angular Distributions, $P_m = 0.40 \text{ GeV}/c$

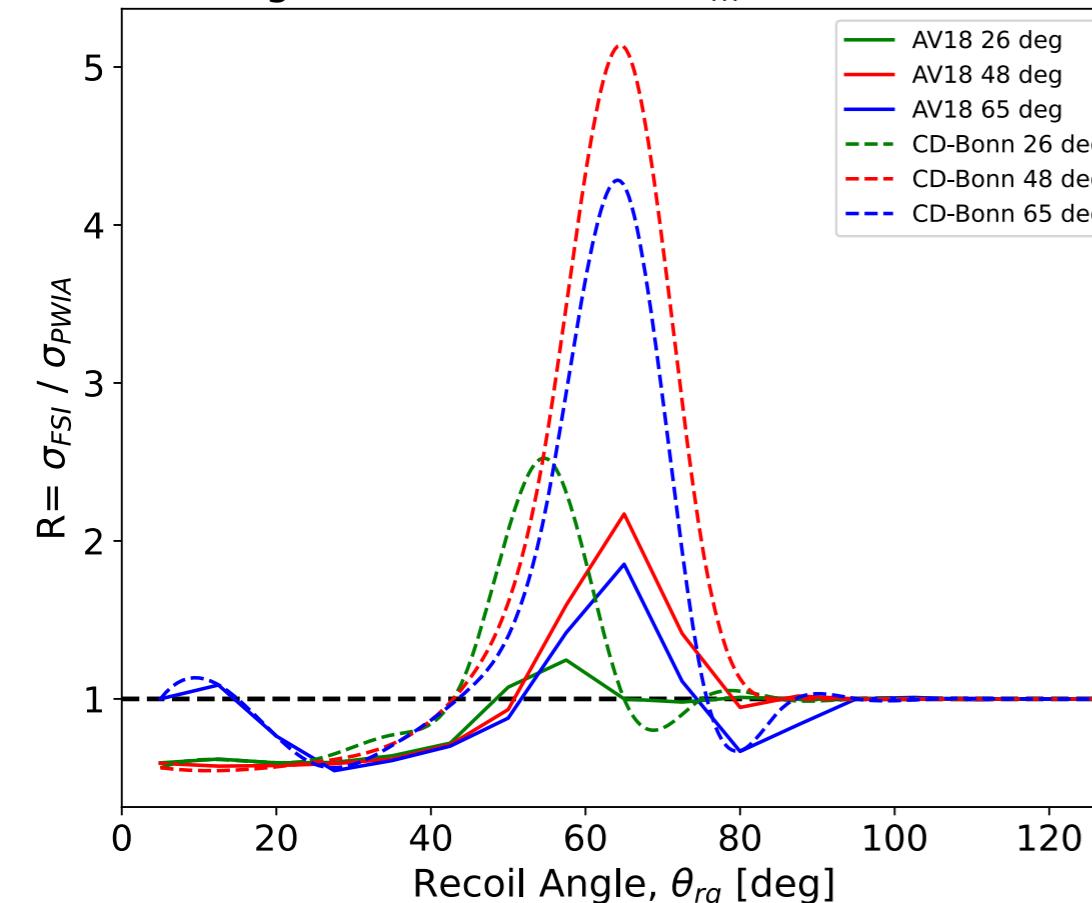


Angular Distributions, $P_m = 0.60 \text{ GeV}/c$

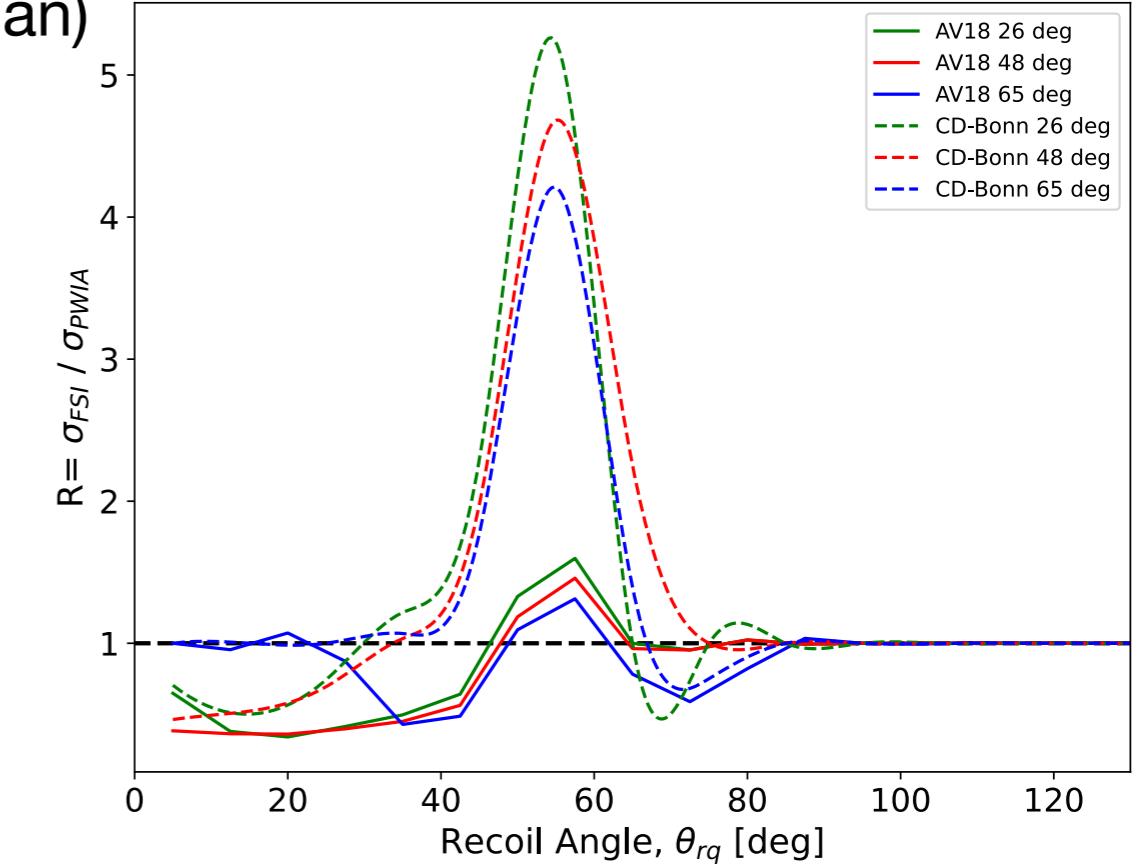


Theory calculations
of angular distributions
(by. M. Sargsian)

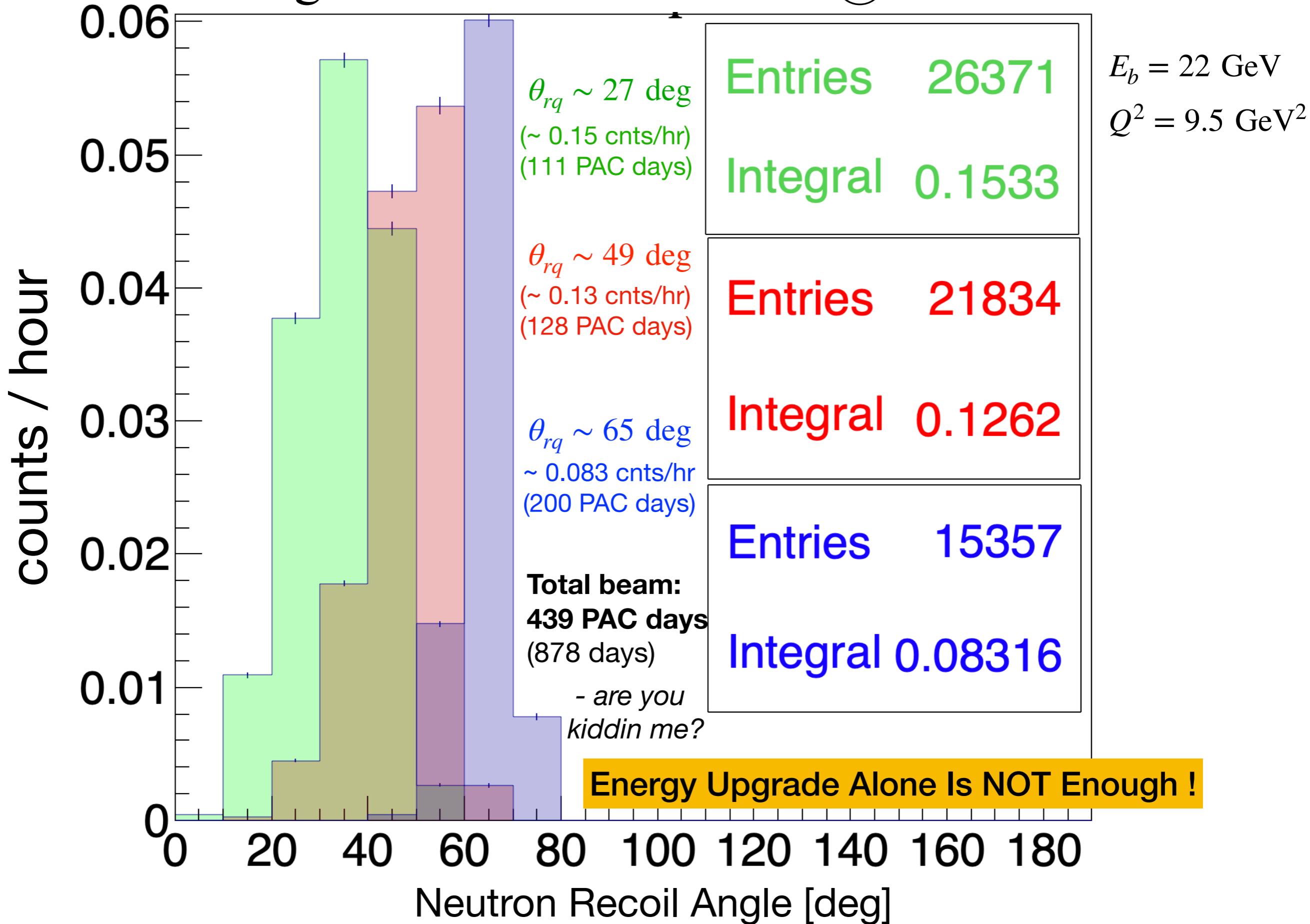
Angular Distributions, $P_m = 0.80 \text{ GeV}/c$



Angular Distributions, $P_m = 1.00 \text{ GeV}/c$



Angular Distribution Rates @ 80 uA

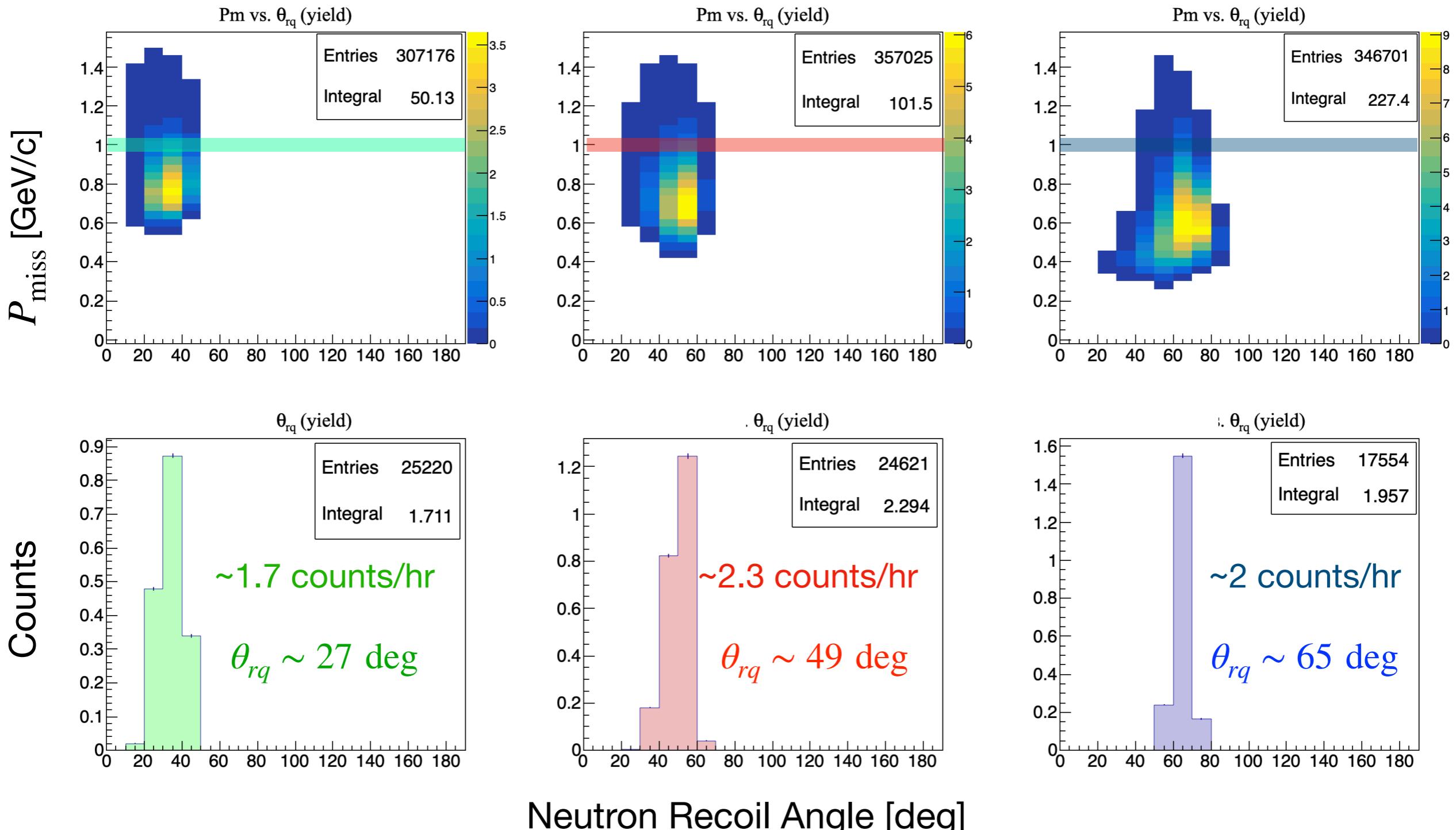


Baseline Integrated Count Rates Estimates @ 80 uA

10 deg bins

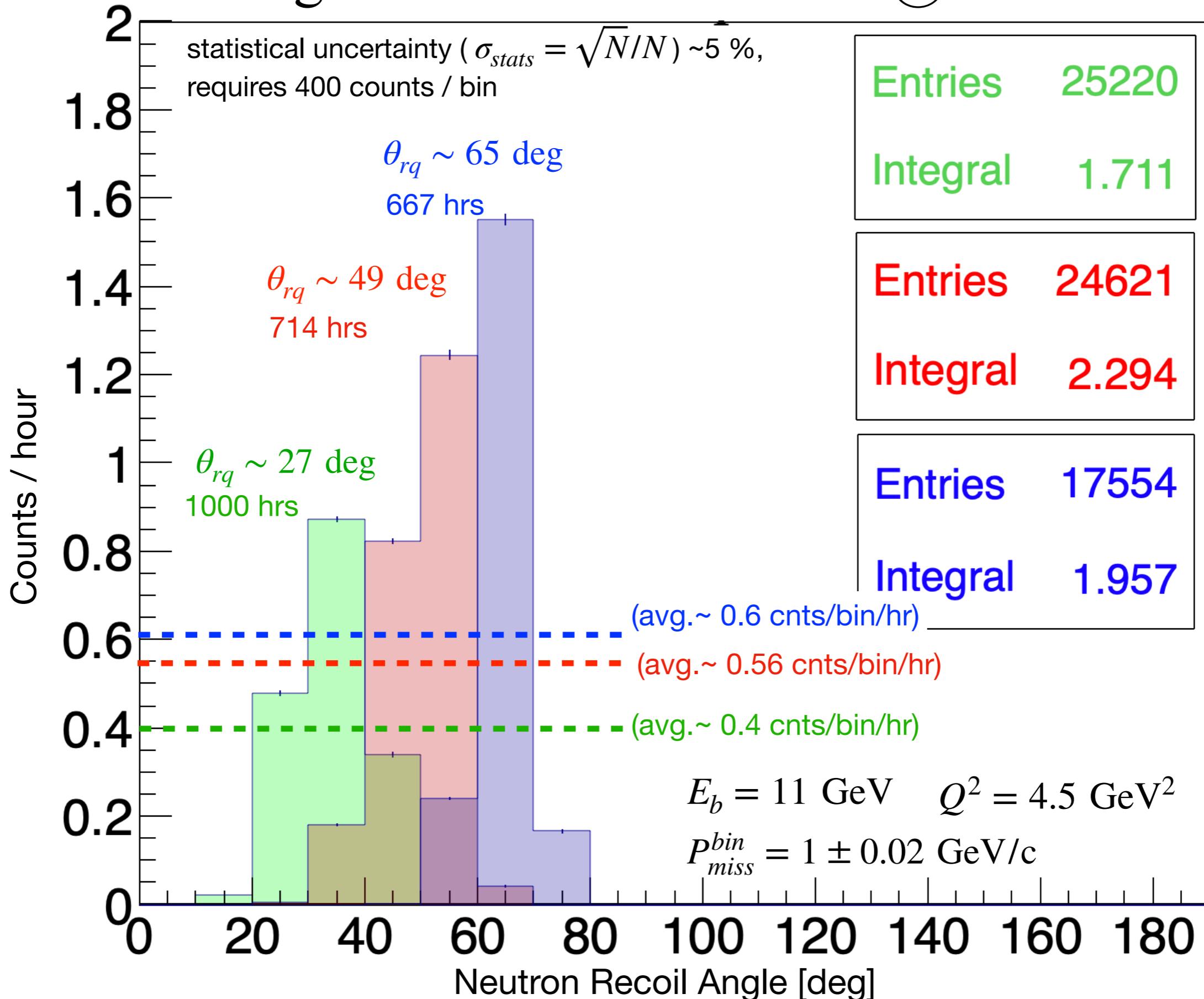
Baseline

- no spectrometer upgrade or energy upgrade required
- Plotted 2D Missing Momentum vs. Recoil Angle
- Projected Angular Distribution for Missing Momentum bin $P_{miss}^{bin} = 1 \pm 0.02$ GeV/c



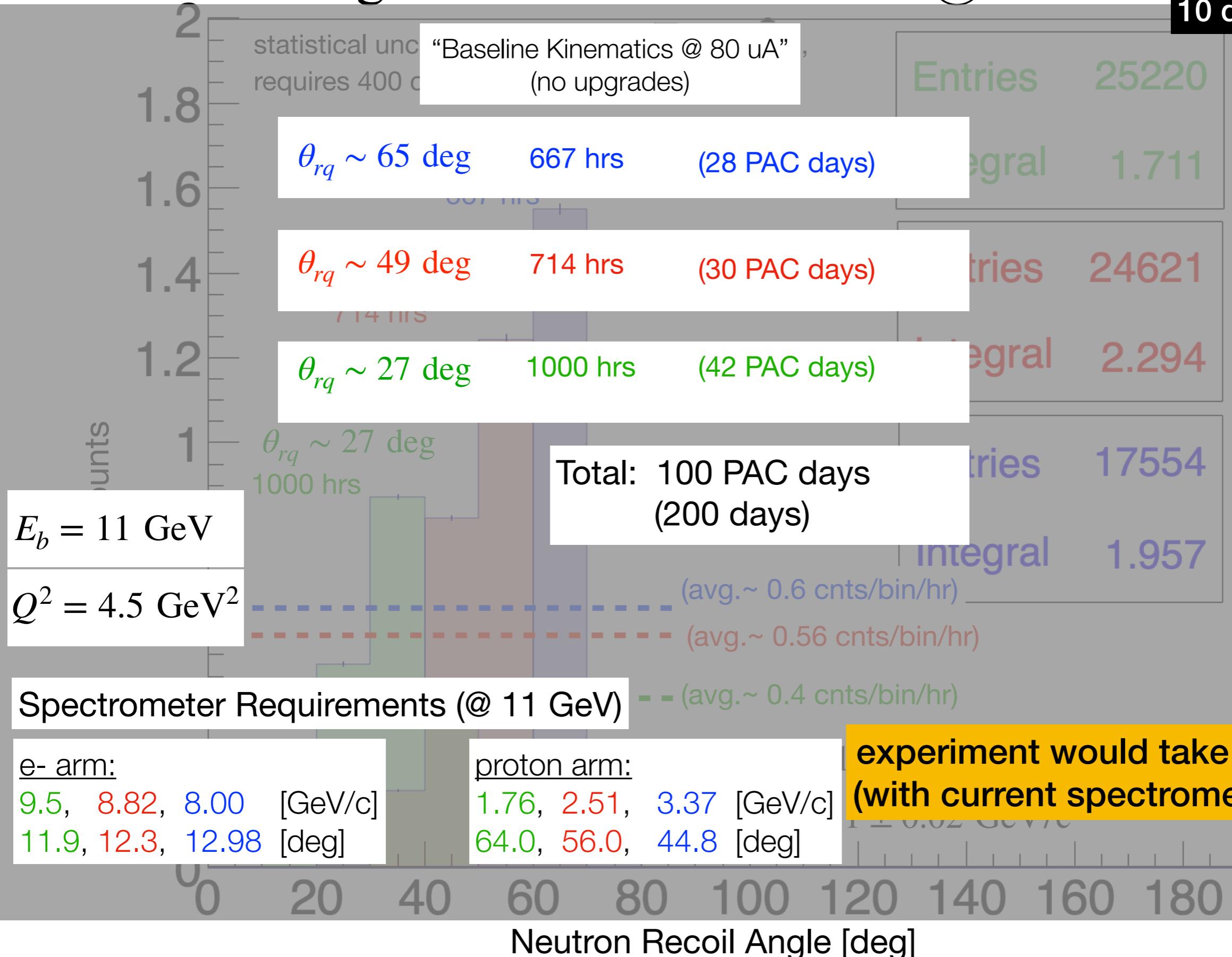
Angular Distribution Rates @ 80 uA

10 deg bins



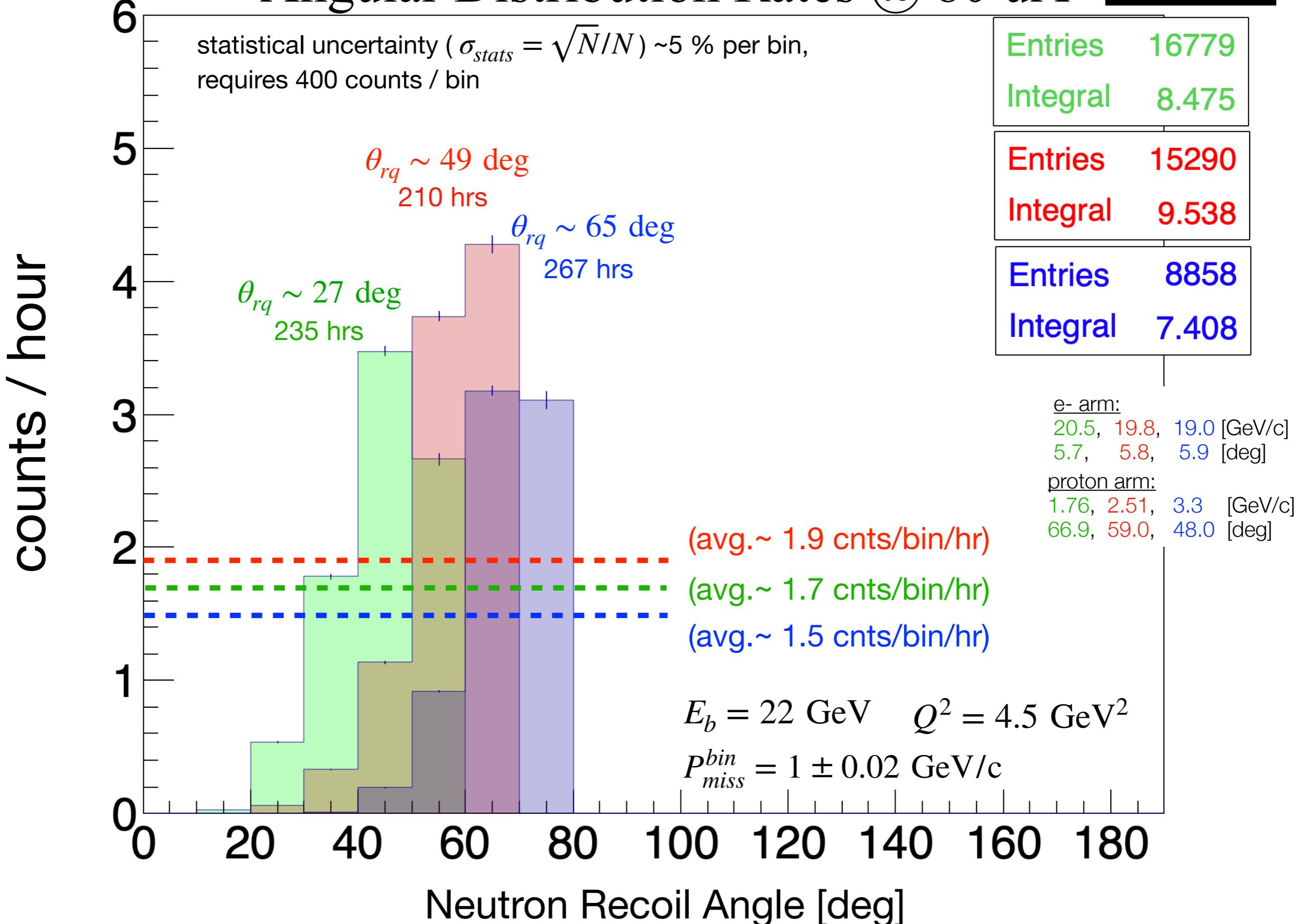
Angular Distribution Rates @ 80 uA

10 deg bins



Angular Distribution Rates @ 80 uA

10 deg bins



Angular Distribution Rates @ 80 uA

