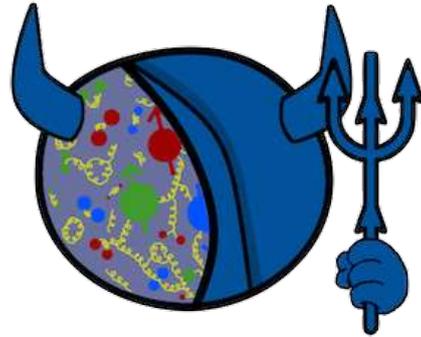


Exact Polarization in Relativistic Fluids at Global Equilibrium

Based on Eur.Phys.J.Plus 138 (2023) 6
In collaboration with F. Becattini



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Andrea Palermo



Stony Brook
University

Motivation

The global polarization of the Lambda particle is well described by the thermal vorticity

Becattini, Chandra, Del Zanna, Grossi, Annals Phys. 338 (2013) 32-49

$$S^\mu(p) = -\frac{1}{8m} \epsilon^{\mu\rho\sigma\tau} p_\tau \frac{\int d\Sigma \cdot p n_F (1 - n_F) \varpi_{\rho\sigma}(x)}{\int d\Sigma \cdot p n_F}$$

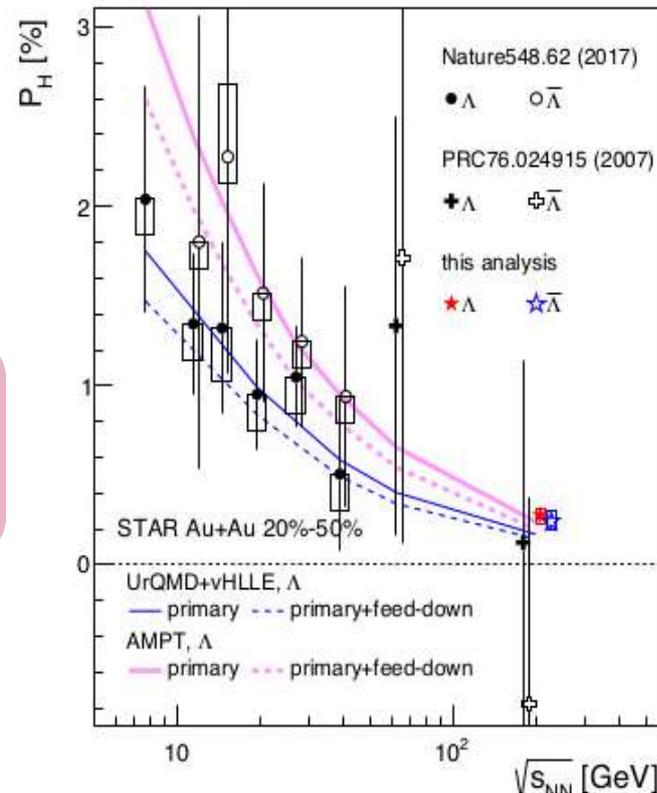
$$\varpi_{\mu\nu} = -\frac{1}{2} (\partial_\mu \beta_\nu - \partial_\nu \beta_\mu)$$

Thermal vorticity

$$\beta^\mu = \frac{u^\mu}{T}$$

Four-temperature

The formula is only a **first order approximation**. Higher order corrections in vorticity are unknown, even in thermal equilibrium.



T. Niida, Nucl.phys.A, 2019

Main results

Exact spin density matrix for spin-S particles at general global equilibrium:

$$\Theta(p) = \frac{\left[(-1)^{2S} \mathbb{I} - e^{-b \cdot p + \theta_0 \cdot D^S(\mathbf{J})} \right]^{-1} - (-1)^{2S} \mathbb{I}}{\sum_{k=-S}^S \left(e^{b \cdot p - k \sqrt{-\theta^2}} - (-1)^{2S} \right)^{-1}}$$

Exact spin vector for Dirac field at global equilibrium

$$\theta^\mu = -\frac{1}{2m} \epsilon^{\mu\nu\rho\sigma} \varpi_{\nu\rho} p_\sigma \quad S^\mu(p) = \frac{1}{2} \frac{\theta^\mu}{\sqrt{-\theta^2}} \frac{\sinh\left(\frac{\sqrt{-\theta^2}}{2}\right)}{\cosh\left(\frac{\sqrt{-\theta^2}}{2}\right) + e^{-b \cdot p + \zeta}}$$

Including **all quantum corrections** in vorticity

Expectation values

To compute mean values we need the spin density matrix:

$$\Theta_{sr}(p) = \frac{\text{Tr}(\hat{\rho}\hat{a}_r^\dagger(p)\hat{a}_s(p))}{\sum_l \text{Tr}(\hat{\rho}\hat{a}_l^\dagger(p)\hat{a}_l(p))} = \frac{\langle\hat{a}_r^\dagger(p)\hat{a}_s(p)\rangle}{\sum_l \langle\hat{a}_l^\dagger(p)\hat{a}_l(p)\rangle}$$

Once we are given a spin density matrix, the expectation value of the Pauli-Lubanski operator is:

[F. Becattini, Lect.Notes Phys. 987 (2021) 15-52, AP, F. Becattini Eur.Phys.J.Plus 138 (2023) 6, 547]

$$S^\mu(p) = \sum_{i=1}^3 [p]_i^\mu \text{tr}(\Theta(p)D^S(J^i)), \quad m \neq 0$$

Other methods to compute polarization rely on the Wigner function.

Global equilibrium

Density operator at global equilibrium:

$$\hat{\rho} = \frac{1}{Z} \exp \left[-b_{\mu} \hat{P}^{\mu} + \frac{1}{2} \varpi_{\mu\nu} \hat{J}^{\mu\nu} \right] \quad \langle \hat{O} \rangle = \text{Tr} \left(\hat{\rho} \hat{O} \right)$$

The vector b is constant and the thermal vorticity ϖ is a constant antisymmetric tensor. The four-temperature β vector is a Killing vector:

$$\beta^{\mu}(x) = b^{\mu} + \varpi^{\mu\nu} x_{\nu} \equiv \frac{u^{\mu}}{T}$$

At global equilibrium:

$$\frac{A^{\mu}}{T} = \varpi^{\mu\nu} u_{\nu}$$

Acceleration

$$\frac{\omega^{\mu}}{T} = -\frac{1}{2} \epsilon^{\mu\nu\rho\sigma} \varpi_{\nu\rho} u_{\sigma}$$

Angular velocity

The generators of the Poincaré group appear in the density operator.

Analytic continuation of the thermal vorticity: $\varpi \mapsto -i\phi$

$$\hat{\rho} = \frac{1}{Z} \exp \left[-b_\mu \hat{P}^\mu - \frac{i}{2} \phi_{\mu\nu} \hat{J}^{\mu\nu} \right]$$

$P \mapsto$ translations
 $J \mapsto$ Lorentz transformations

Factorization of the density operator:

$$\hat{\rho} = \frac{1}{Z} \exp \left[-\tilde{b}_\mu(\phi) \hat{P}^\mu \right] \exp \left[-i \frac{\phi_{\mu\nu}}{2} \hat{J}^{\mu\nu} \right] \equiv \frac{1}{Z} \exp \left[-\tilde{b}_\mu(\phi) \hat{P}^\mu \right] \hat{\Lambda}$$

$$\tilde{b}^\mu(\phi) = \sum_{k=0}^{\infty} \frac{1}{(k+1)!} \underbrace{(\phi_{\alpha_1}^{\mu} \phi_{\alpha_2}^{\alpha_1} \cdots \phi_{\alpha_k}^{\alpha_{k-1}})}_{k \text{ times}} b^{\alpha_k} \quad \hat{\Lambda} \equiv e^{-i \frac{\phi_{\mu\nu}}{2} \hat{J}^{\mu\nu}}$$

We can use **group theory** to calculate **thermal expectation values**.

The number operator at (imaginary) global equilibrium:

$$\langle \hat{a}_s^\dagger(\mathbf{p}) \hat{a}_t(\mathbf{p}') \rangle = 2\varepsilon' \sum_{n=1}^{\infty} (-1)^{2S(n+1)} \delta^3(\Lambda^n \mathbf{p} - \mathbf{p}') D^S(W(\Lambda^n, \mathbf{p}))_{ts} e^{-\tilde{b} \cdot \sum_{k=1}^n \Lambda^k \mathbf{p}}$$

$D(W) = [\Lambda p]^{-1} \Lambda[p]$ is the Wigner rotation in the spin-S representation of the rotation group.

For vanishing vorticity (i.e. $\Lambda=I$) we recover Bose and Fermi statistics:

$$\langle \hat{a}_s^\dagger(\mathbf{p}) \hat{a}_t(\mathbf{p}') \rangle = 2\varepsilon' \sum_{n=1}^{\infty} (-1)^{2S(n+1)} \delta^3(\mathbf{p} - \mathbf{p}') \delta_{ts} e^{-nb \cdot \mathbf{p}} = \frac{2\varepsilon \delta^3(\mathbf{p} - \mathbf{p}') \delta_{ts}}{e^{b \cdot \mathbf{p}} + (-1)^{2S+1}}$$

We want to use this result to compute the spin density matrix:

$$\Theta(p)_{rs} = \frac{\langle \hat{a}_s^\dagger(p) \hat{a}_r(p) \rangle}{\sum_t \langle \hat{a}_t^\dagger(p) \hat{a}_t(p) \rangle}$$

From the **analytic continuation** of the density operator:

$$\langle \hat{a}_s^\dagger(p) \hat{a}_t(p) \rangle = 2\varepsilon \sum_{n=1}^{\infty} (-1)^{2S(n+1)} \delta^3(\Lambda^n \mathbf{p} - \mathbf{p}) D^S(W(\Lambda^n, p))_{ts} e^{-\tilde{b} \cdot \sum_{k=1}^n \Lambda^k p}$$

The spin density matrix is singular unless $\Lambda \mathbf{p} = \mathbf{p}$. The analytic continuation of the density operator is forced to be in the little group of \mathbf{p} .

Exact spin physics

The constraint equation:

$$\Lambda p = p \implies \phi^{\mu\nu} p_\nu = 0$$

$$\phi^{\mu\nu} = \epsilon^{\mu\nu\rho\sigma} \xi_\rho \frac{p_\sigma}{m} \quad \xi^\rho = -\frac{1}{2m} \epsilon^{\rho\mu\nu\sigma} \phi_{\mu\nu} p_\sigma$$

Thanks to the constraint, it is possible to compute the spin density matrix for any spin

$$\xi^\mu \mapsto \theta^\mu = -\frac{1}{2m} \epsilon^{\mu\nu\rho\sigma} \varpi_{\nu\rho} p_\sigma \quad \theta_0^\mu = [p]^{-1\mu}{}_\nu \theta^\nu$$

$$\Theta(p) = \frac{\left[(-1)^{2S} \mathbb{I} - e^{-b \cdot p + \theta_0 \cdot D^S(\mathbf{J})} \right]^{-1} - (-1)^{2S} \mathbb{I}}{\sum_{k=-S}^S \left(e^{b \cdot p - k \sqrt{-\theta^2}} - (-1)^{2S} \right)^{-1}}$$

We can compute the spin vector for Dirac fermions:

$$S^\mu(p) = \frac{1}{2} \frac{\theta^\mu}{\sqrt{-\theta^2}} \frac{\sinh\left(\frac{\sqrt{-\theta^2}}{2}\right)}{\cosh\left(\frac{\sqrt{-\theta^2}}{2}\right) + e^{-b \cdot p + \zeta}}$$

And for generic spin-S particles:

$$S^\mu(p) = \frac{\theta^\mu}{\sqrt{-\theta^2}} \frac{\sum_{k=-S}^S k \left[e^{b \cdot p - \zeta - k\sqrt{-\theta^2}} - (-1)^{2S} \right]^{-1}}{\sum_{k=-S}^S \left[e^{b \cdot p - \zeta - k\sqrt{-\theta^2}} - (-1)^{2S} \right]^{-1}}$$

These formulae account for corrections to **all orders in vorticity**

The formula:

- Reproduces linear results

$$S^\mu(p)_{\theta \rightarrow 0} \sim \theta^\mu \frac{S(S+1)}{3} (1 + (-1)^{2S} n_{F/B}(b \cdot p - \zeta))$$

- Exact calculation with Boltzmann statistics

$$S_B^\mu = \frac{\theta^\mu}{\sqrt{-\theta^2}} \frac{\chi'(\sqrt{-\theta^2})}{\chi(\sqrt{-\theta^2})} \quad \chi(\sqrt{-\theta^2}) = \sum_{k=-S}^S e^{k\sqrt{-\theta^2}}$$

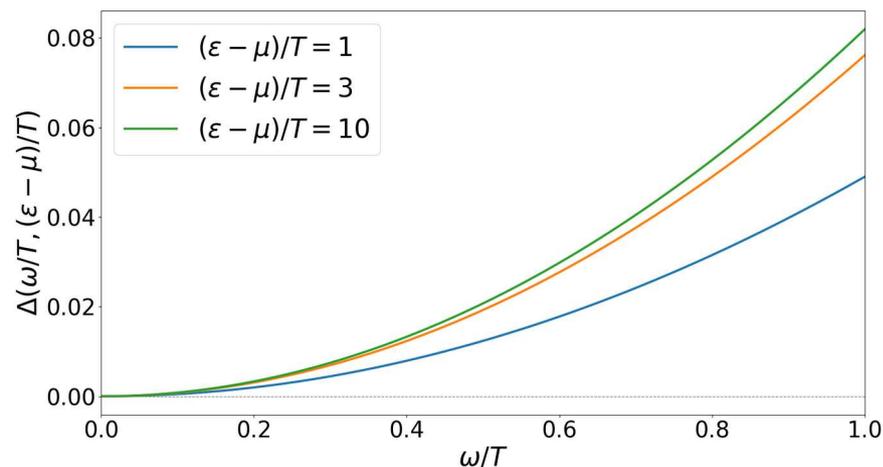
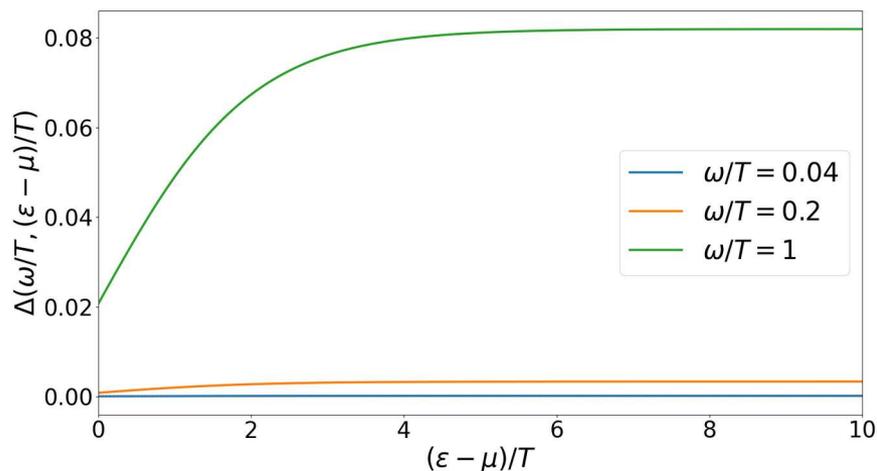
- Is unitary

$$S^\mu(p)_{\theta \rightarrow \infty} \sim S \frac{\theta^\mu}{\sqrt{-\theta^2}}$$

Exact polarization in heavy-ion collisions

Consider the Λ polarization in relativistic heavy ion collisions $\omega \sim 10^{22} \text{ s}^{-1}$ and $\omega/T \sim 0.04$.

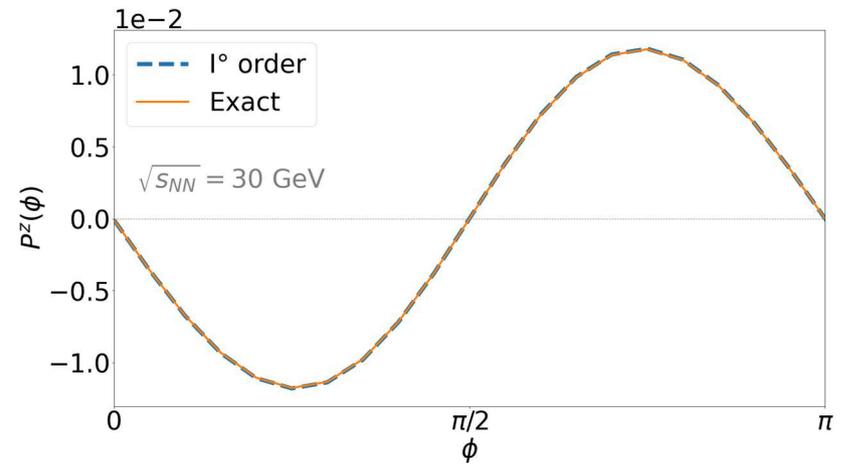
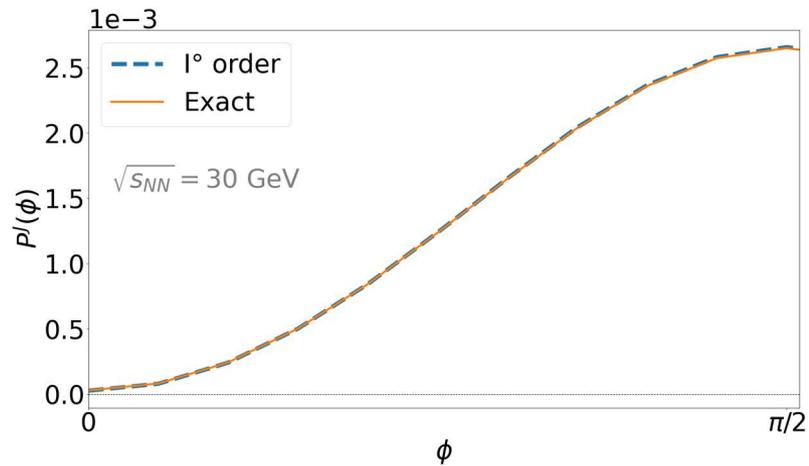
$$\Delta = \left| \frac{S_E - S_L}{S_E} \right|$$



The difference is very small in most physical cases.

Extending the formula to local equilibrium with $\varpi(\mathbf{x})$

$$S^\mu(p) = -\frac{1}{4m} \epsilon^{\mu\nu\rho\sigma} p_\sigma \frac{\int d\Sigma \cdot p n_F \frac{\varpi_{\nu\rho}}{\sqrt{-\theta^2}} \frac{\sinh(\sqrt{-\theta^2}/2)}{\cosh(\sqrt{-\theta^2}/2) + e^{-b \cdot p + \zeta}}}{\int d\Sigma \cdot p n_F}$$



Conclusions

Exact expectation values can be computed using the **analytic continuation of the density operator**.

Exact spin polarization vector and **spin density matrix** for particles at global equilibrium.

Higher order corrections in vorticity are negligible, at least in high energy collisions...

**Thank you for the
attention!**

Backup

Pauli-Lubanski Vector

The Hilbert space of relativistic particles is built using the four-momentum and the Pauli-Lubanski vector

$$\hat{\Pi}^\mu = -\frac{1}{2}\epsilon^{\mu\nu\rho\sigma}\hat{J}_{\nu\rho}\hat{P}_\sigma \quad [\hat{\Pi}^\mu, \hat{P}^\nu] = 0$$

For states with definite momentum p

$$\hat{\Pi}^\mu|p\rangle = \hat{\Pi}^\mu(p)|p\rangle \quad \hat{\Pi}^\mu(p) = -\frac{1}{2}\epsilon^{\mu\nu\rho\sigma}\hat{J}_{\nu\rho}p_\sigma$$

One has

$$p_\mu\hat{\Pi}^\mu(p) = 0$$

If we consider **massive** particles the Pauli-Lubanski vector is connected to the generators of rotations

$$\widehat{S}^\mu(p) = \frac{\widehat{\Pi}^\mu(p)}{m} = -\frac{1}{2m} \epsilon^{\mu\nu\rho\sigma} \widehat{J}_{\nu\rho} p_\sigma$$

$$\widehat{S}^\mu(p) = \sum_{i=1}^3 \widehat{S}_i(p) n_i^\mu(p) \quad [\widehat{S}_i(p), \widehat{S}_j(p)] = i\epsilon_{ijk} \widehat{S}_k(p)$$

In the **massless** case:

$$\widehat{\Pi}^\mu(p) = \widehat{h}(p) p^\mu + \widehat{\Pi}_1(p) n_1^\mu(p) + \widehat{\Pi}_2(p) n_2^\mu(p) \quad \widehat{\Pi}_{1,2}(p) |p, h\rangle = 0$$

$$\widehat{h}(p) = \frac{\widehat{\Pi}(p) \cdot q}{p \cdot q} = -\frac{1}{2p \cdot q} \epsilon^{\mu\nu\rho\sigma} \widehat{J}_{\nu\rho} p_\sigma q_\mu$$

$$q \cdot p \neq 0, \quad q^2 = 0, \quad q \cdot n_i(p) = 0$$

Helicity is the only physical degree of freedom

Polarization and the Wigner function

To compute expectation values it is useful to use the Wigner function.

$$W(x, k)_{ab} = -\frac{1}{(2\pi)^4} \int d^4y e^{-ik \cdot y} \langle : \bar{\Psi}_b(x + y/2) \Psi_a(x - y/2) : \rangle$$

Massive Dirac fermions [F. Becattini, Lect.Notes Phys. 987 (2021) 15-52]:

$$S^\mu(p) = \frac{1}{2} \frac{\int d\Sigma \cdot p \operatorname{tr}(\gamma^\mu \gamma_5 W_+(x, p))}{\int d\Sigma \cdot p \operatorname{tr}(W_+(x, p))}$$

Massless Dirac fermions

$$\Pi^\mu(p) = p^\mu \sum_{h=\pm S} h \Theta_{hh}(p) \qquad \Pi^\mu(p) = \frac{p^\mu}{2} \frac{\int d\Sigma \cdot p \operatorname{tr}(\not{d} \gamma^5 W_+(x, p))}{\int d\Sigma \cdot p \operatorname{tr}(W_+(x, p) \not{d})}$$

In the massless case the mean spin always points in the direction of momentum.

See also [Y.-C. Liu, K. Mameda, X.-G. Huang, *Chin. Phys. C* 44(9), 094101]

Exact Wigner function for free fermions at global equilibrium:

$$W(x, k) = \frac{1}{(2\pi)^3} \int \frac{d^3 p}{2\varepsilon} \sum_{n=1}^{\infty} (-1)^{n+1} e^{-n\tilde{\beta}(in\phi)\cdot p} \times$$

$$\left[e^{-in\frac{\phi:\Sigma}{2}} (m + \not{p}) \delta^4(k - (\Lambda^n p + p)/2) + (m - \not{p}) e^{in\frac{\phi:\Sigma}{2}} \delta^4(k + (\Lambda^n p + p)/2) \right]$$

Where $\Lambda = e^{-i\frac{\phi}{2}:J}$ is in the four-vector representation.

Solves the Wigner equation! Full summation of the “ \hbar expansion”.

Can be used to compute expectation values:

$$\langle : \{ \bar{\Psi}\Psi, j^\mu, T^{\mu\nu} \} : \rangle = \int d^4 k \{ \text{tr}(W), \text{tr}(\gamma^\mu W), k^\mu \text{tr}(\gamma^\nu W) \}$$

The Dirac delta is integrated out and results are expressed as series of functions that can be regularized using the **analytic distillation**.

Spin vector

Spin vector of massive Dirac fermions:

$$S^\mu(p) = \frac{1}{2} \frac{\int d\Sigma \cdot p \operatorname{tr}(\gamma^\mu \gamma_5 W_+(x, p))}{\int d\Sigma \cdot p \operatorname{tr}(W_+(x, p))}$$

Exact spin vector at global equilibrium:

$$S^\mu(p) = \frac{1}{2m} \frac{\sum_{n=1}^{\infty} (-1)^{n+1} e^{-n\tilde{b}(in\phi) \cdot p} \operatorname{tr}\left(\gamma^\mu \gamma_5 e^{-in\frac{\phi \cdot \Sigma}{2}} \not{p}\right) \delta^3(\Lambda^n p - p)}{\sum_{n=1}^{\infty} (-1)^{n+1} e^{-n\tilde{b}(in\phi) \cdot p} \operatorname{tr}\left(e^{-in\frac{\phi \cdot \Sigma}{2}}\right) \delta^3(\Lambda^n p - p)}$$

How to handle a ratio of series of δ -functions? Where does it come from?

Any thermal expectation value in a free quantum field theory is obtained from:

$$\langle \hat{a}_s^\dagger(\mathbf{p}) \hat{a}_t(\mathbf{p}') \rangle = \frac{1}{Z} \text{Tr} \left(\exp[-\tilde{b}_\mu(\phi) \hat{P}^\mu] \hat{\Lambda} \hat{a}_s^\dagger(\mathbf{p}) \hat{a}_t(\mathbf{p}') \right)$$

$$[\hat{a}_s^\dagger(\mathbf{p}), \hat{a}_t(\mathbf{p}')]_{\pm} = 2\varepsilon \delta^3(\mathbf{p} - \mathbf{p}') \delta_{st}$$

Using Poincaré transformation rules and (anti)commutation relations (particle with spin S):

$$\begin{aligned} \langle \hat{a}_s^\dagger(\mathbf{p}) \hat{a}_t(\mathbf{p}') \rangle &= (-1)^{2S} \sum_r D^S(W(\Lambda, \mathbf{p}))_{rs} e^{-\tilde{b} \cdot \Lambda \mathbf{p}} \langle \hat{a}_r^\dagger(\Lambda \mathbf{p}) \hat{a}_t(\mathbf{p}') \rangle + \\ &+ 2\varepsilon e^{-\tilde{b} \cdot \Lambda \mathbf{p}} D^S(W(\Lambda, \mathbf{p}))_{ts} \delta^3(\Lambda \mathbf{p} - \mathbf{p}') \end{aligned}$$

$D(W) = [\Lambda p]^{-1} \Lambda[p]$ is the “Wigner rotation” in the S-spin representation.

We find a solution by iteration:

I $\langle \hat{a}_s^\dagger(p) \hat{a}_t(p') \rangle \sim 2\varepsilon e^{-\tilde{b} \cdot \Lambda p} D^S(W(\Lambda, p))_{ts} \delta^3(\Lambda \mathbf{p} - \mathbf{p}')$

⋮

II $\langle \hat{a}_s^\dagger(p) \hat{a}_t(p') \rangle \sim 2\varepsilon (-1)^{2S} D^S(W(\Lambda^2, p))_{ts} e^{-\tilde{b} \cdot (\Lambda p + \Lambda^2 p)} \delta^3(\Lambda^2 \mathbf{p} - \mathbf{p}') +$
 $+ 2\varepsilon e^{-\tilde{b} \cdot \Lambda p} D^S(W(\Lambda, p))_{ts} \delta^3(\Lambda \mathbf{p} - \mathbf{p}')$

⋮

∞ $\langle \hat{a}_s^\dagger(p) \hat{a}_t(p') \rangle = 2\varepsilon' \sum_{n=1}^{\infty} (-1)^{2S(n+1)} \delta^3(\Lambda^n \mathbf{p} - \mathbf{p}') D^S(W(\Lambda^n, p))_{ts} e^{-\tilde{b} \cdot \sum_{k=1}^n \Lambda^k p}$

Energy density for massless fermions, equilibrium with acceleration ($\phi=ia/T$)

$$\rho = \frac{3T^4}{8\pi^2} \sum_{n=1}^{\infty} (-1)^{n+1} \phi^4 \frac{\sinh n\phi}{\sinh^5(n\phi/2)}$$

The series is finite as long as ϕ is real. For real thermal vorticity it diverges!

The series includes terms which are non analytic at $\phi=0$.

Analytic distillation:



The series boils down to polynomials: $\alpha^\mu = \frac{A^\mu}{T}$ $w^\mu = \frac{\omega^\mu}{T}$

$$\rho = \frac{7\pi^2}{60\beta^4} - \frac{\alpha^2}{24\beta^4} - \frac{17\alpha^4}{960\pi^2\beta^4}$$

Expectation values vanish at the Unruh temperature $T_U = \sqrt{-A \cdot A}/2\pi$
 [G.Prokhorov, O. Teryaev, V. Zakharov, JHEP03(2020)137]

Axial current under rotation: [V. Ambrus, E. Winstanley, 1908.10244]

$$j_A^\mu = T^2 \left(\frac{1}{6} - \frac{w^2}{24\pi^2} - \frac{\alpha^2}{8\pi^2} \right) \frac{w^\mu}{\sqrt{\beta^2}}$$

First exact results at equilibrium with **both rotation and acceleration.**
 [V. Ambrus, E. Winstanley Symmetry 2021, 13(11)]

$$\rho = T^4 \left(\frac{7\pi^2}{60} - \frac{\alpha^2}{24} - \frac{w^2}{8} - \frac{17\alpha^4}{960\pi^2} + \frac{w^4}{64\pi^2} + \frac{23\alpha^2 w^2}{1440\pi^2} + \frac{11(\alpha \cdot w)^2}{720\pi^2} \right)$$

Massless Dirac field

$$\Pi^\mu = \frac{p^\mu \sum_{n=1}^{\infty} (-1)^{n+1} e^{-\tilde{b}(\phi) \cdot \sum_{k=1}^n \Lambda^k p} \text{tr} (\not{q} \gamma_5 \exp[-in\phi : \Sigma/2] \not{p}) \delta^3(\Lambda^n p - p)}{2 \sum_{n=1}^{\infty} (-1)^{n+1} e^{-\tilde{b}(\phi) \cdot \sum_{k=1}^n \Lambda^k p} \text{tr} (\not{q} \exp[-in\phi : \Sigma/2] \not{p}) \delta^3(\Lambda^n p - p)}$$

We can deal with the series as we did before

$$\phi^{\mu\nu} = \epsilon^{\mu\nu\rho\sigma} \frac{h_\rho p_\sigma}{p \cdot q} \quad h^\mu = -\frac{1}{2} \epsilon^{\mu\nu\rho\sigma} \phi_{\nu\rho} q_\sigma$$

The series reduces to

$$\eta = \frac{h \cdot p}{q \cdot p} = \frac{1}{2(p \cdot q)} \epsilon^{\mu\nu\alpha\beta} \phi_{\alpha\beta} p_\nu q_\mu$$

$$\Pi^\mu(p) = \frac{p^\mu}{2} \frac{i \sin(\eta/2)}{\cos(\eta/2) + e^{-b \cdot p}}$$

Dirac fermions:

$$\Pi^\mu(p) = -\frac{p^\mu}{2} \frac{\sinh(H/2)}{\cosh(H/2) + e^{-b \cdot p}}$$

$$H = \frac{1}{2(p \cdot q)} \epsilon^{\mu\nu\alpha\beta} \varpi_{\alpha\beta} p_\nu q_\mu = -\frac{1}{2\varepsilon} \epsilon^{0\nu\alpha\beta} \varpi_{\alpha\beta} p_\nu$$

The mean Pauli-Lubanski vector depends on the orientation of the momentum

$$\varphi^\nu = -\frac{1}{2} \epsilon^{\nu\alpha\beta 0} \varpi_{\alpha\beta} \quad \hat{\mathbf{p}} = \frac{\mathbf{p}}{\varepsilon}$$

Massless particles

For massless particles of arbitrary helicity S :

$$\Theta_{hk}(p) = \frac{\sum_{n=1}^{\infty} (-1)^{2S(n+1)} e^{-nb \cdot p} e^{n\eta h} \delta_{hk}}{\sum_{n=1}^{\infty} (-1)^{2S(n+1)} e^{-nb \cdot p} 2 \cos n\eta}$$

$$\Pi^{\mu}(p) = -p^{\mu} S \frac{\sinh(S H)}{\cosh(S H) - (-1)^{2S} e^{-b \cdot p}}$$

$$H = \frac{1}{2(p \cdot q)} \epsilon^{\mu\nu\alpha\beta} \varpi_{\alpha\beta} p_{\nu} q_{\mu} = -\frac{1}{2\epsilon} \epsilon^{0\nu\alpha\beta} \varpi_{\alpha\beta} p_{\nu}$$

The reason for this simplified structure is that massless particles only have $+S$ and $-S$ helicity and that

$$D^S(W(\Lambda^n, p))_{ts} = e^{i\eta r} \delta_{hk} \quad \hat{\Pi}_{1,2}(p)|p, h\rangle = 0$$