# Beyond hadronic physics with secondary beams @JLAB:

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## New research lines using secondary beams

- The interaction of high-current (O(100µA)), medium-energy (O(10GeV)) electron beam with a dump produces intense secondary beams:
  - Light Dark Matter (if exists)
    - BDX
  - > Neutrinos
    - CEvNS
  - > Muons
    - Proton radius



## Light Dark matter beam: production

In the "dark sector" scenario, 3 main LDM production mechanisms in fixed-target, lepton-beam experiments



### A) A'-strahlung:

- > Radiative A' emission in nucleus EM field followed by A'->XX
- > Scales as  $Z^2 \alpha^3$
- > Forward-boosted, high-energy A' emission

### **B)** Non-resonant e+e- annhilation:

- >  $e+e- \rightarrow A'\gamma$  followed by A'->XX
- > Scales as  $Z\alpha^2$
- > Forward-backward A' emission in the CM

### C) Resonant e+e- annhilation:

- > e+e- -> A'->XX
- Scales as Zα
- $\succ$  Breit-Wigner like cross section with  $M_{A'}=\sqrt{2m_eE}$

## BDX @ 20 GeV e- beam

- The Beam Dump Experiment is optimized to run
  @ 12 GeV e- beam
  - > Fully parasitic wrt Moller experiment
  - New facility (Civil construction + passive shielding) downstream HallA beam dump
  - Detector : EM calorimeter + Veto systems



 $10^{-7}$  $10^{-8}$ BABAR  $\epsilon^2 \alpha_D (m_{\chi}/m_{A'})^4$  $10^{-9}$  $(\mathbf{II})$ BDX - MINI $10^{-10}$  $10^{-11}$ COHERI Ш  $10^{-12}$ S BDX + BDX@24(I) Pseudo-Dirac Fermion Relic  $10^{-13}$ (II) Majorana Relic (III) Scalar Relic  $10^{-14}$  $10^{0}$  $10^{1}$  $10^{2}$  $m_{\chi}$  [MeV]

- BDX can benefit of 24 GeV e- beam extending the reach
  - Pro: Increase number of secondary particles in the dump -> enhanced DM production
  - WARNING: beam-related background need to be studied

#### see M. Spreafico talk

### Neutrino production in the dump

- Neutrinos production in the dump due to muon and pion decays
  - The majority (low-energy v) come from pion and muon decay at rest
    - π decay produces a prompt 28.5 MeV v<sub>µ</sub> along with
      a µ which subsequently decay producing v<sub>p</sub> e v<sub>m</sub>
    - Weak angular dependence
  - High-energy v from in-flight pion and muon decay



#### Detector 5

## Neutrino secondary beam

- Neutrinos flux estimated through MC simulations based on FLUKA
  - Used Hall-A beam dump description implemented in FLUKA by the JLAB Rad Con
  - Neutrino flux evaluated both for 10Gev e- beam and 20GeV ebeam
    - used current dump (not optimized) for 20GeV simulation
  - Flux computed on several planes located both downstream and on top of Hall-A beam dump



Credit to A. Fulci



#### Detector 1 Credit to A. Fulci

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  - Flux computed on several planes located both downstream and on top of Hall-A beam dump
  - Flux comparable to SNS@ Oak Ridge National Lab one





## Physics case for a Neutrino beam @ JLAB

### Coherent Elastic v-Nucleus Scattering

- Integrated cross section as a function of v energy Low energy neutrinos (< 100 MeV) coherently scatter on >entire target nucleus  $q < (1/R_{N})$
- Cross section scaling with N<sup>2</sup>  $\succ$
- The largest cross section among neutrino scattering >channels for Ev < 100 MeV
- Low recoil energy due to kinematics O(10 keV) >
- First measurement in 2017 on CsI by COHERENT >collaboration (~134 events)
- why interesting?: \*
  - weak parameters -> mixing angle
  - nuclear properties -> neutrons distribution radius >
  - sterile neutrino >
  - neutrino magnetic moment >
  - non standard interaction mediated by exotic particles >

#### **Requirements**:

10-35

10-3

10-37

10-38

10-39

10-40

10-41

7 [cm<sup>2</sup>]

High-intense v-flux >

100

 $\succ$ v-flux energy range: few MeV - few 100 MeV

SNS

10<sup>1</sup>

Ev [MeV]

LBNF

10<sup>2</sup>

detector has be sensitive to small energy  $\succ$ depositions

Reactor

backgrounds need to be sufficiently small to  $\succ$ observe the signal.

## **CEvNS measurement @ JLAB**

### v-beam @ JLAB

- Produced by the interaction between e- beam and Hall A dump in a energy range: 10MeV - 300 MeV
- > Neutrino flux @ 10 GeV : ~10<sup>18</sup> v/m<sup>2</sup> at ~10 m above the dump for  $10^{22}EOT$
- Neutrino flux @ 20 GeV : ~2x10<sup>18</sup> v/m<sup>2</sup> at ~10 m above the dump for 10<sup>22</sup>EOT

### Detector technologies:

- > Detector located 10m on top of the dump
- > Two detection technology are under study:
  - Csl
  - LAr-TPC
- Veto system: active (plastic ...) and passive (lead, water, borate silicone and/or cadmium sheet layers...)



- Backgrounds:
  - beam-related background: neutron
  - beam-unrelated background: cosmic, radioactive detector contamination, environmental radioactivity
  - Extensive Monte-Carlo simulations and measurements campaign in situ are necessary

## **CEVNS** measurement @ JLAB

- Two targets under study:  $\otimes$ 
  - CsI(TI) crystal:  $\succ$ 
    - Pro: High-density, high LY, heavy nuclei
    - Cons: radioactive material, afterglow
  - Liquid Argon based detector:  $\succ$ 
    - Pro: Low threshold, directionality
    - Cons: depleted Ar

### A state of a state of a

Expected yield:			· · · · · · · · · · · · · · · · · · ·	LAr - 1m3		
Detector	e- @ 10 GeV v flux: 1E8 v/m² (*)	e- @ 20 GeV v flux: 2E8 v/m <sup>2</sup> (*)				z
CsI (1m <sup>3</sup> ) [thr : 10 keV]	8000		10		PRELIMINA	
LAr (1m <sup>3</sup> ) [thr: 10 keV]	2500				Energy threshold	Er_th(ke'
(*) for 10 <sup>22</sup> EOT	1	1	Cre	edit to A. Fulci,	S. Grazzi, A. Pillor	ו <sup>10</sup>

S

CEBAF @ 10 GeV

10

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Csl detector - 1m3

Energ<sup>1</sup>/<sup>o</sup>threshold (keV)

Integrated events

(\*) for 10<sup>22</sup> EOT

### vBDX-CEvNS @ 20 GeV

- **Two targets under study:** 
  - ➤ CsI(TI) crystal:
    - Pro: High-density, high LY, heavy nuclei
    - Cons: radioactive material, afterglow
  - Liquid Argon based detector:
    - Pro: Low threshold, directionality
    - Cons: depleted Ar

### Expected yield:

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Detector	e- @ 10 GeV v flux: 1E8 v/m <sup>2</sup> (*)	e- @ 20 GeV v flux: 2E8 v/m² (*)	grated events	LAr - 1m3	
CsI (1m <sup>3</sup> ) [thr : 10 keV]	~8000	~15000		PRELIMINARY	
LAr (1m³) [thr: 10 keV]	~2500	~4500	10	<sup>10</sup> Energy threshold (Kev	
				$\mathbf{A} = \{\mathbf{x}, \mathbf{y}, \mathbf{x}, \mathbf{y}, $	

CEBAF @ 20 GeV

10

Credit to A. Fulci, S. Grazzi, A. Pilloni <sup>11</sup>

10<sup>2</sup>

Er th(keV)

(\*) for 10<sup>22</sup> EOT

### Muon production in the dump



#### Credit to A. Fulci

# Muon secondary beam Muon flux estimated using the same FLUKA setup used for neutrinos

- \*
  - Flux scored on a plane to exit of the concrete dump  $\succ$
  - Flux is larger than CERN's M2 beam line one (Eµ>100GeV)  $\succ$

 $10^{-}$ 

5 6 8 9 10

E(GeV)



10-12



0

Y(cm

-40-20

14 E (GeV)

### Proton radius measurement with muon beam

- Proton radius can be measured with 2 techniques:
  - > Leptonic scattering
  - Spettroscopi measurement
- Persistent discrepancy between methods -> Proton radius puzzle



- Secondary muons produced in HALLA beam dump can be extracted and directed toward a new Hall
  - A magnet-based system to focus and to measure µ momentum -> study on going



### Summary

- High intensity extracted electron beams are a precious source of secondary beams:
  - Light Dark Matter (if exists)
  - > Neutrinos
  - > Muons
- A 24 GeV primary electron beam impinged on Hall-A dump can produce higher intensity secondary beams then the 12 GeV one
  - Neutrino beam with a DAR spectrum : flux of 2E8 v/m<sup>2</sup> for 10<sup>22</sup> EOT e-@ 20 GeV
  - Muon beam with a Bremsstrahlung-like spectrum. Energy range : O(10 MeV) O(10 GeV). Flux @ 20 GeV : 5E-6 µ/EOT.
- Secondary beams can be exploited to explore "hot" physics scenario