

JLab 20+ DVCS Data:

A look at the Benefits

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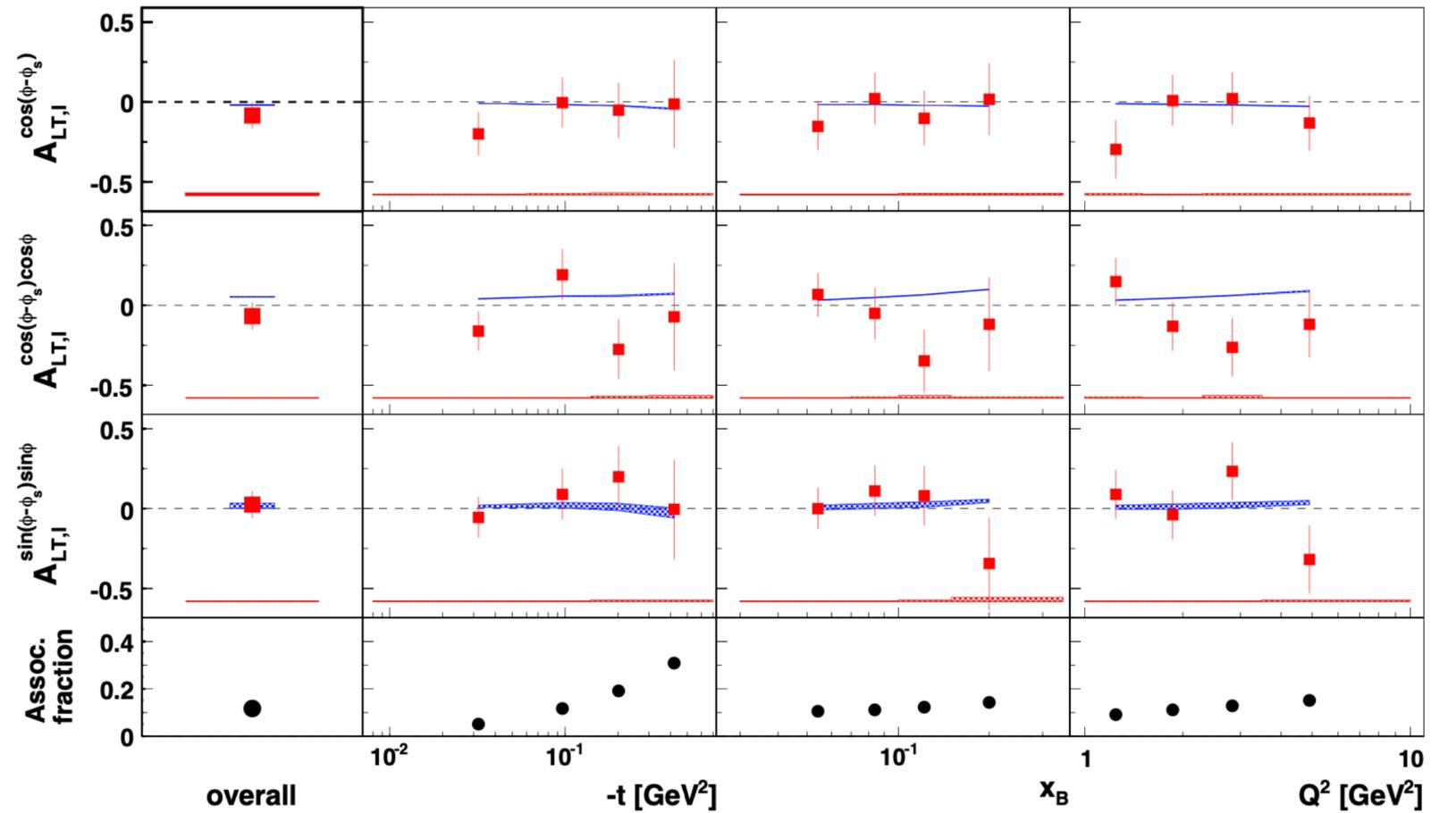
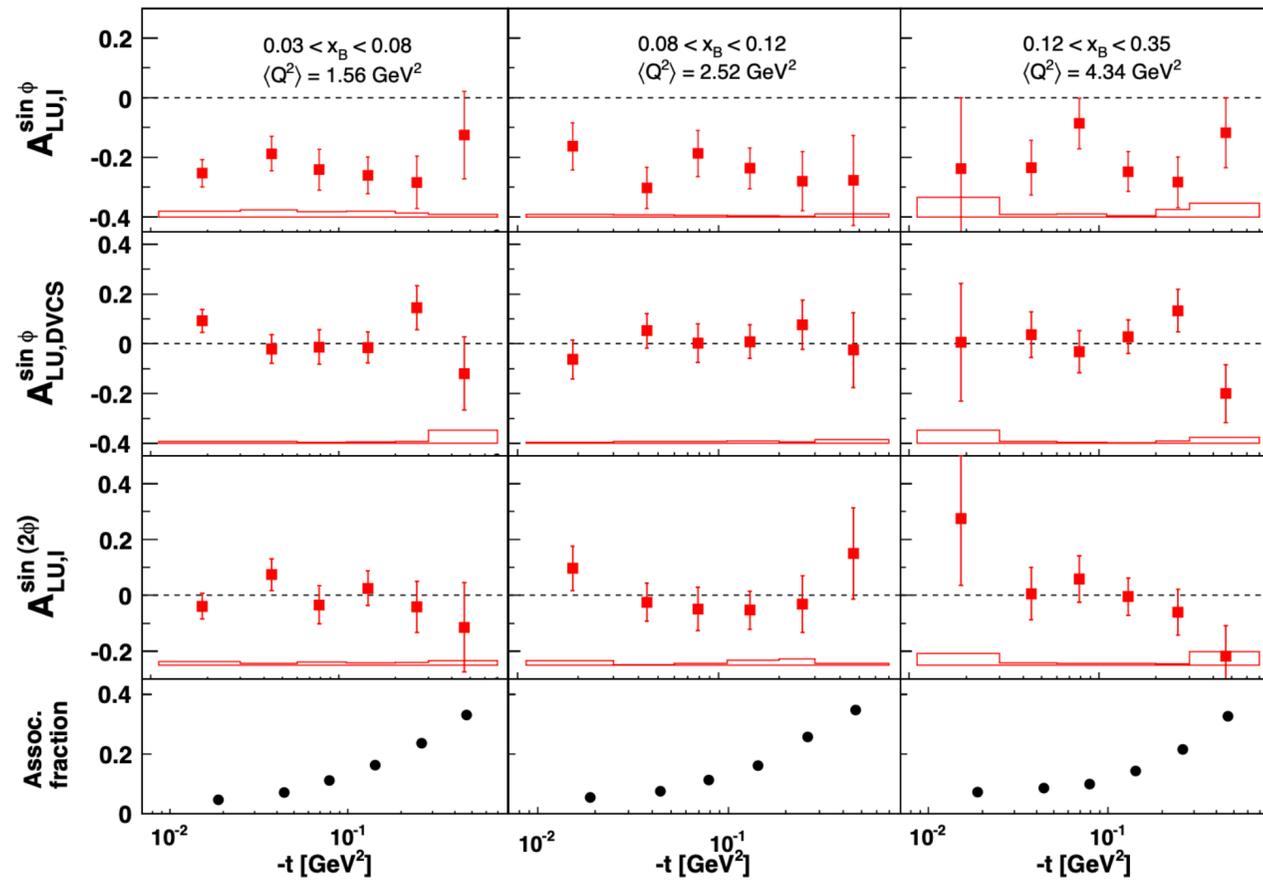
Overview

- **Review of HERMES data**
 1. Data Examples
 2. Extracted CFFs
- **JLab DVCS at 20+ GeV**
 1. Kinematics
 2. Comparison: HERMES to JLAB
 3. 20+ GeV Cross section predictions compared to 10 GeV
 4. Remarks, Compton Form Factor fitting
- **Conclusions**

First let's look at the HERMES data..

- HERMES raw data runs from 2001-2007, publications from ~2006-2012
- Exclusive study of ASYMMETRIES
- Electron and positron beams used
- Many observables measured: $A_C, A_{LU}, A_{UL}, A_{LL}, A_{UT}, A_{LT}$ totals over 14 harmonic coeff's
- Beam energy of 27.6 GeV
- Lower luminosity $\leq 10^{31} cm^{-2} s^{-1}$
- Most data within $0.05 \leq x_B \leq 0.2$ $2 GeV^2 \leq Q^2 \leq 6 GeV^2$

Examples of HERMES data

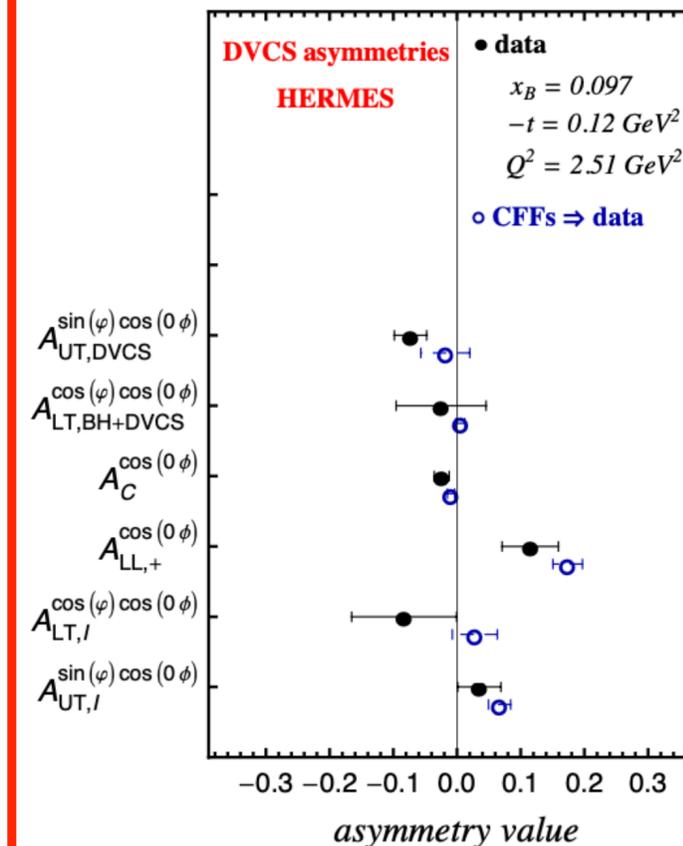
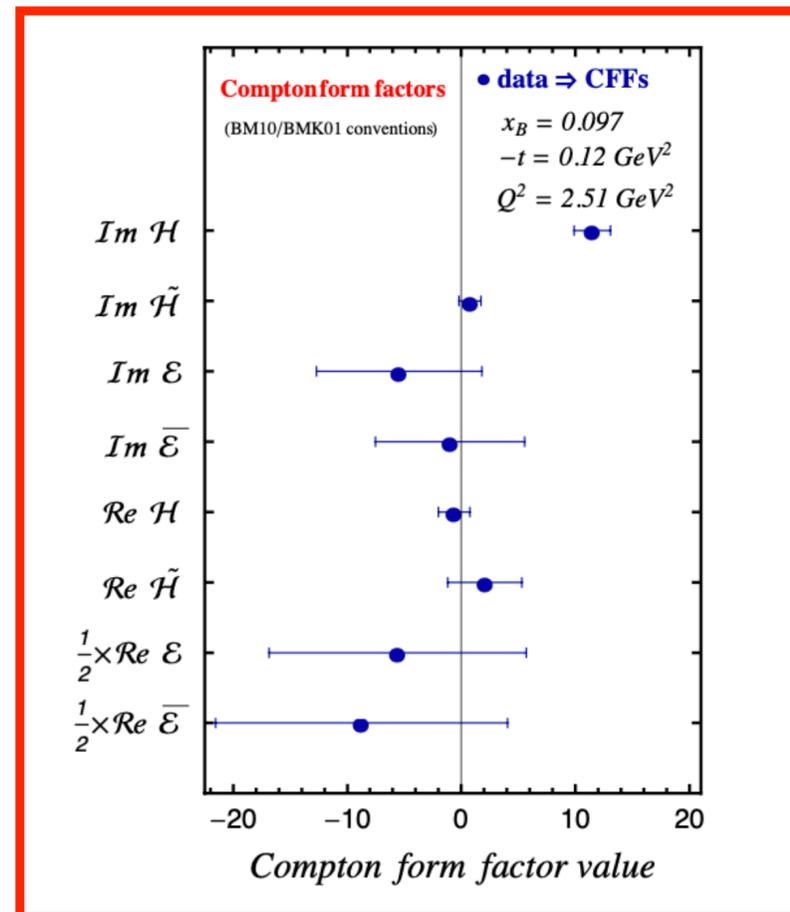
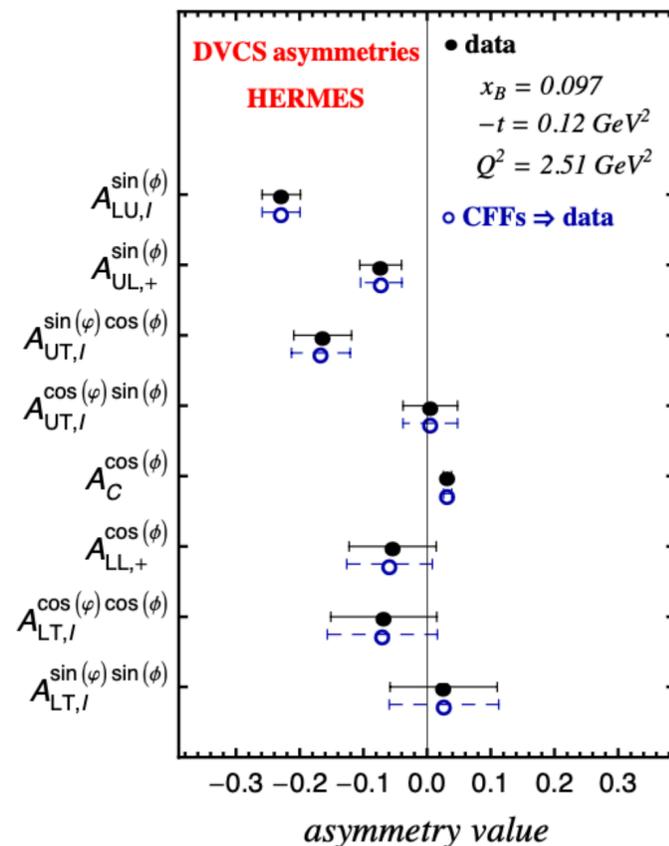


• Okay, so what does this data give us ??...

An analysis of HERMES data

e.g. Kumericki, Muller & Murray 2014:

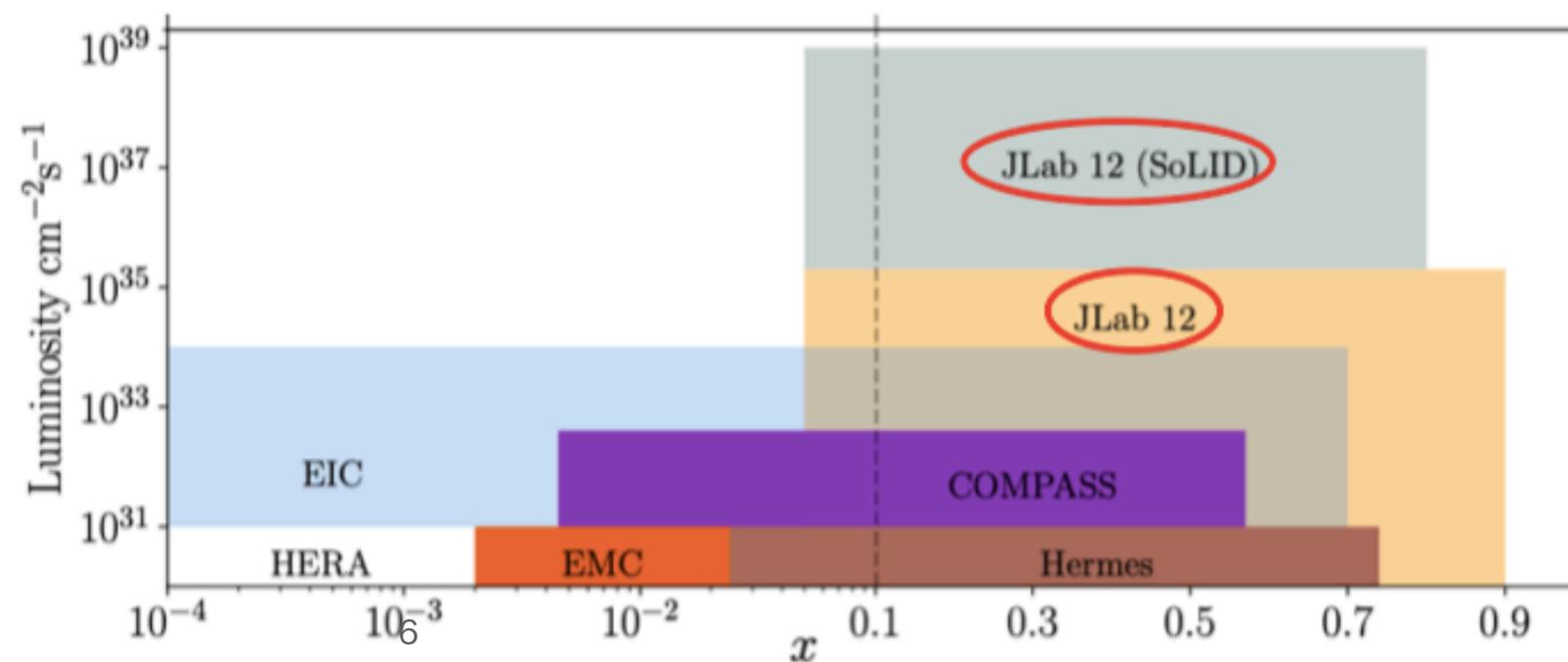
- Extracted all 8 twist-2 CFFs using linear approximations



- Other, more rigorous analyses have been done, which also includes JLab data, but for this case here is a “best-case scenario” given only HERMES data

JLab 20+ GeV Data

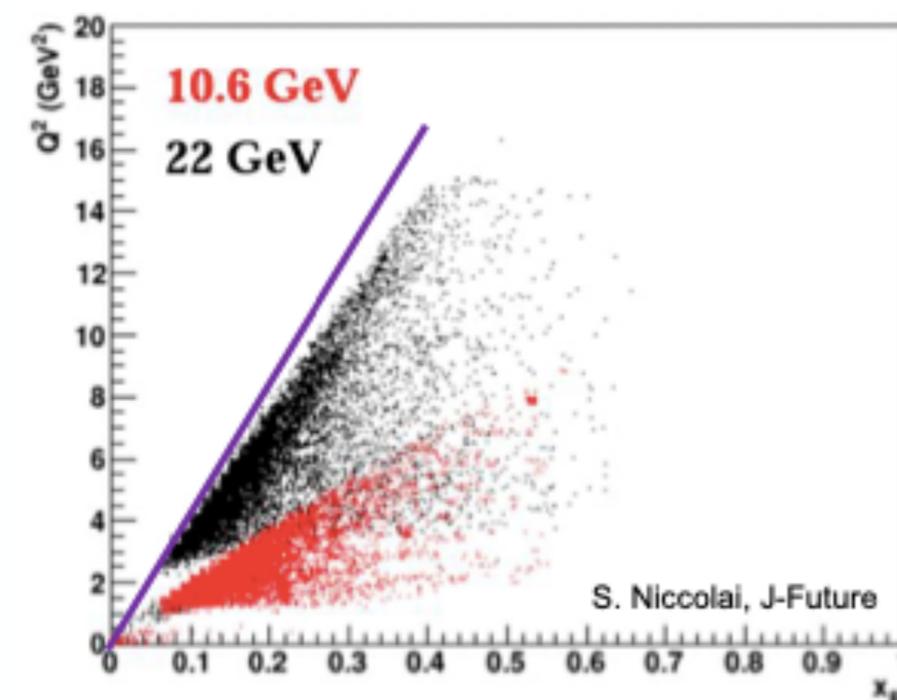
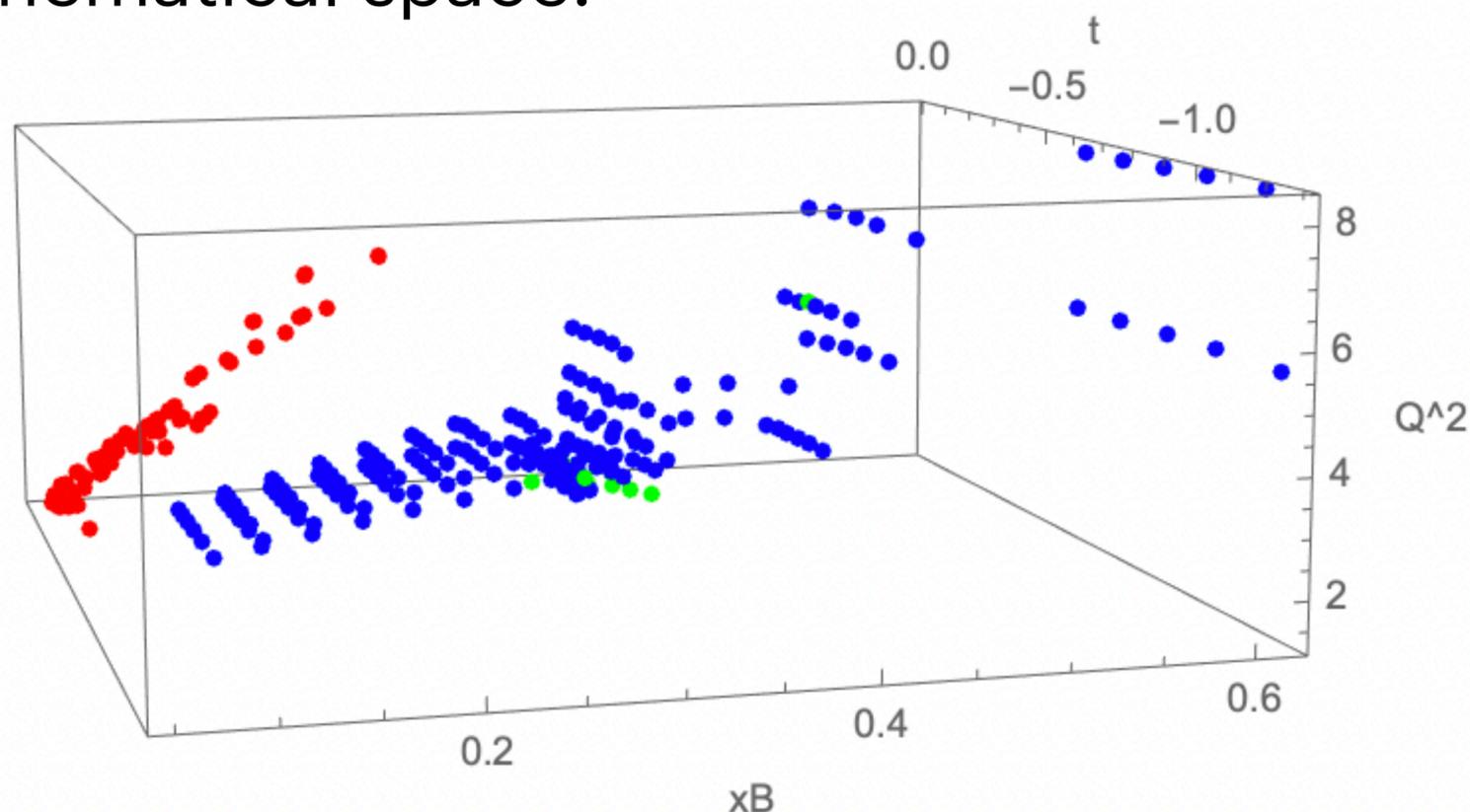
- JLab 20+ GeV can be thought of as a “super HERMES”
- Similar maximum energy (but variable!) but much higher luminosity
- Therefore, if we simply performed the same (or similar) observable measurement program, the improved statistics will allow much higher constraints on CFFs (& GPDs)



HERMES vs JLAB data

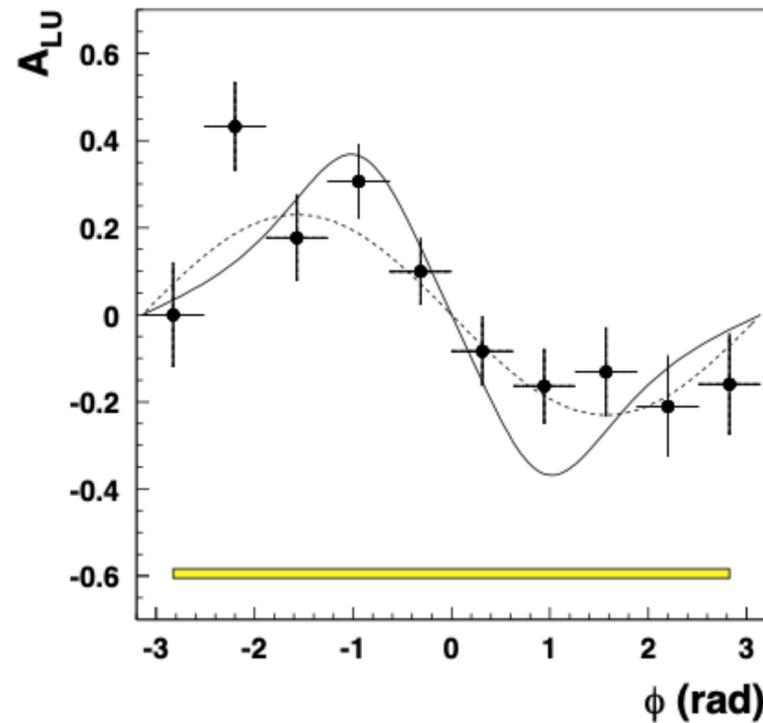
- Current HERMES data greater Q^2 coverage, but less t coverage than current JLab data
- Upgrading to 20^+ GeV will bring all that wonderful JLab precision to the HERMES kinematical space!

- HERMES data
- current JLAB data



Case Comparison: A_{LU}

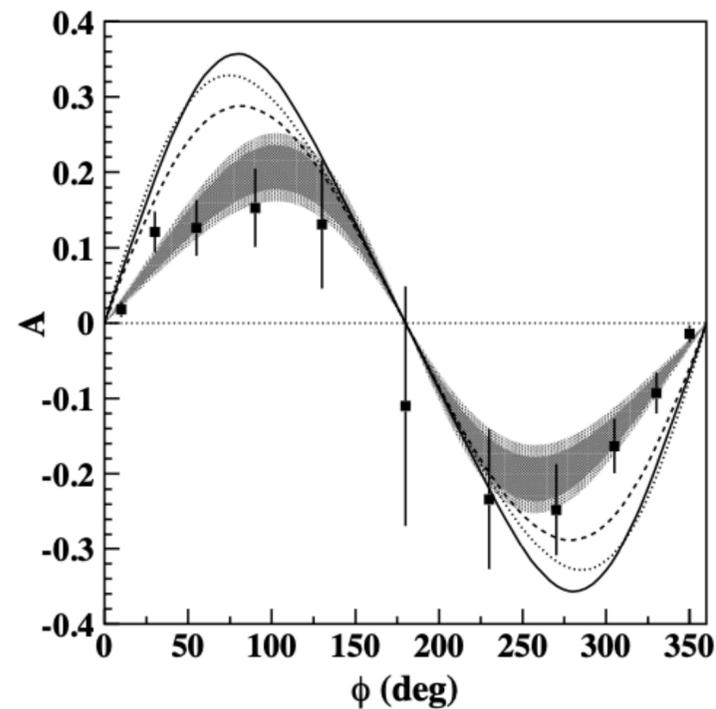
HERMES 2001



By combining the $A_{LU}^{\sin\phi}$ data in the same M_x region as was used for Fig. 3 ($-1.5 < M_x < 1.7$ GeV), an average value of -0.23 ± 0.04 (stat) ± 0.03 (syst) is obtained. The average values of the kinematic variables corresponding to this measurement are: $\langle x \rangle = 0.11$, $\langle Q^2 \rangle = 2.6$ GeV² and $\langle -t \rangle = 0.27$ GeV². Since the BH pro-

$$A_{LU}^{\sin(\phi)} = 0.23 \pm .04$$

CLAS 2001



a $\sin\phi$ modulation. Curves in Figure 4 show the results of theoretical calculations from Refs. [5,11,13] at fixed values of $Q^2 = 1.25$ (GeV/c)², $x_B = 0.19$, and $-t = 0.19$ (GeV/c)². The limited experimental informa-

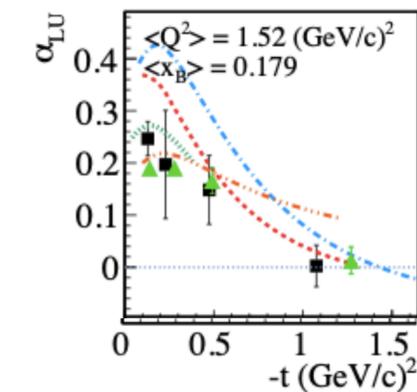
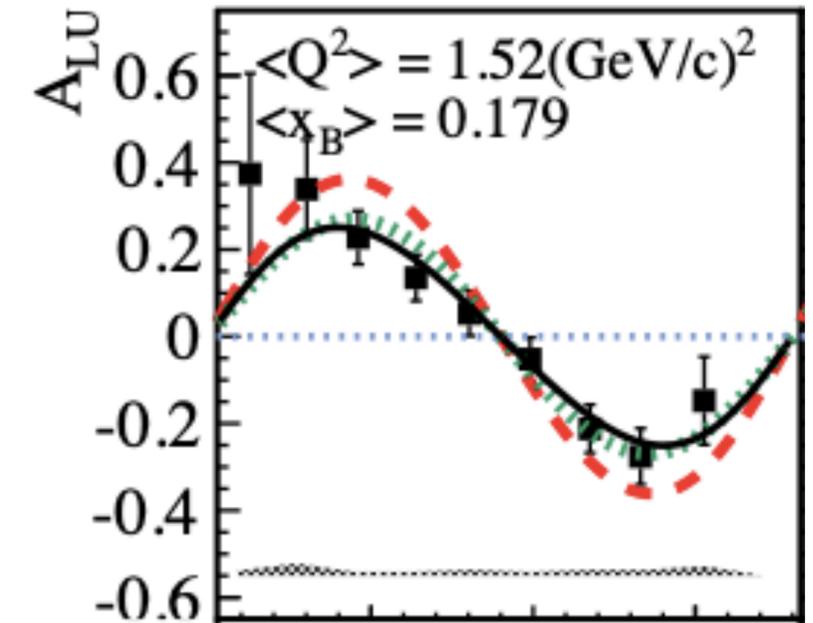
The data points are fitted with the function $A(\phi) = \alpha \sin\phi + \beta \sin 2\phi$. The fitted parameters are $\alpha = 0.202 \pm 0.028^{stat} \pm 0.013^{sys}$ and $\beta = -0.024 \pm 0.021^{stat} \pm 0.009^{sys}$.

$$A_{LU}^{\sin(\phi)} = 0.20 \pm .03$$

$A_{LU}^{\sin(\phi)}$ fractional uncertainty improved by factor of 2

This improvement directly translates to CFFs

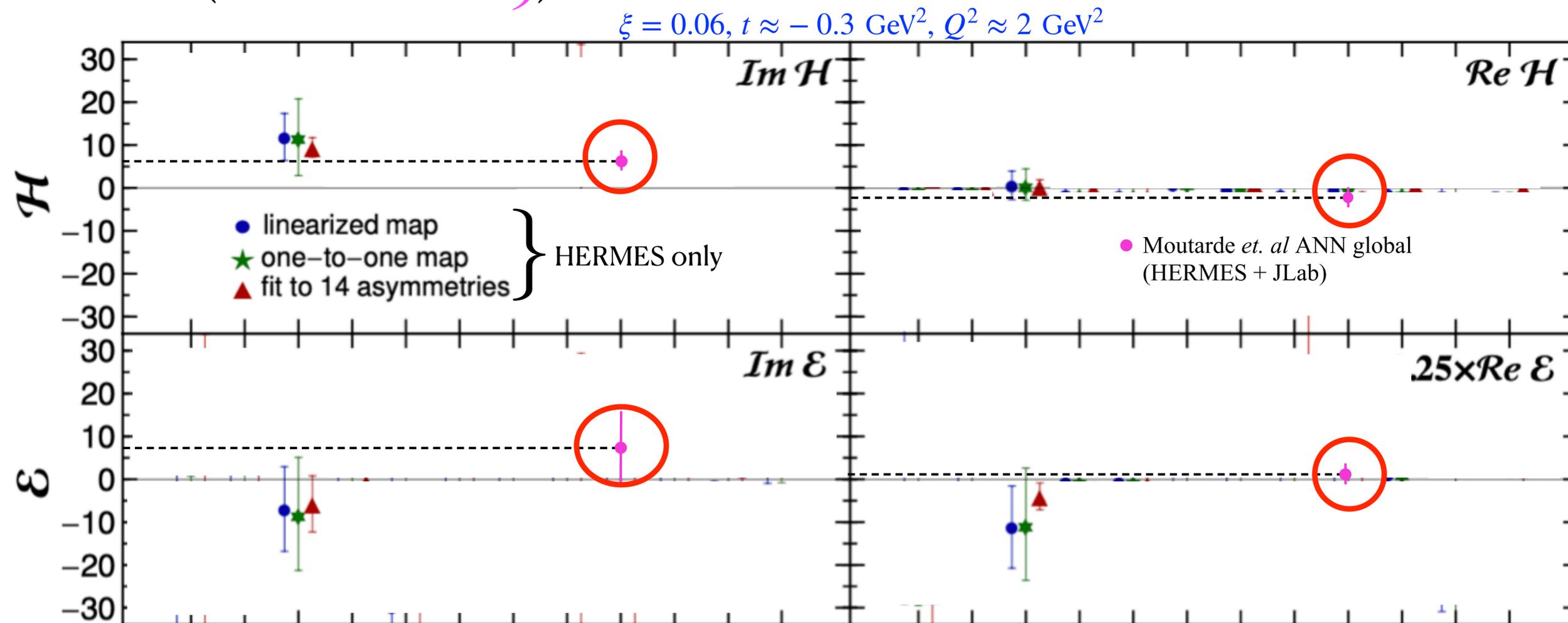
CLAS 2015



$$A_{LU}^{\sin(\phi)} = 0.25 \pm .02$$

Global Analyses of DVCS data

- A general global analysis should not use linear approximations
- They should either use numerical χ^2 fitting method (Guidal & Boer 2015) or NNs (Moutarde 2019)

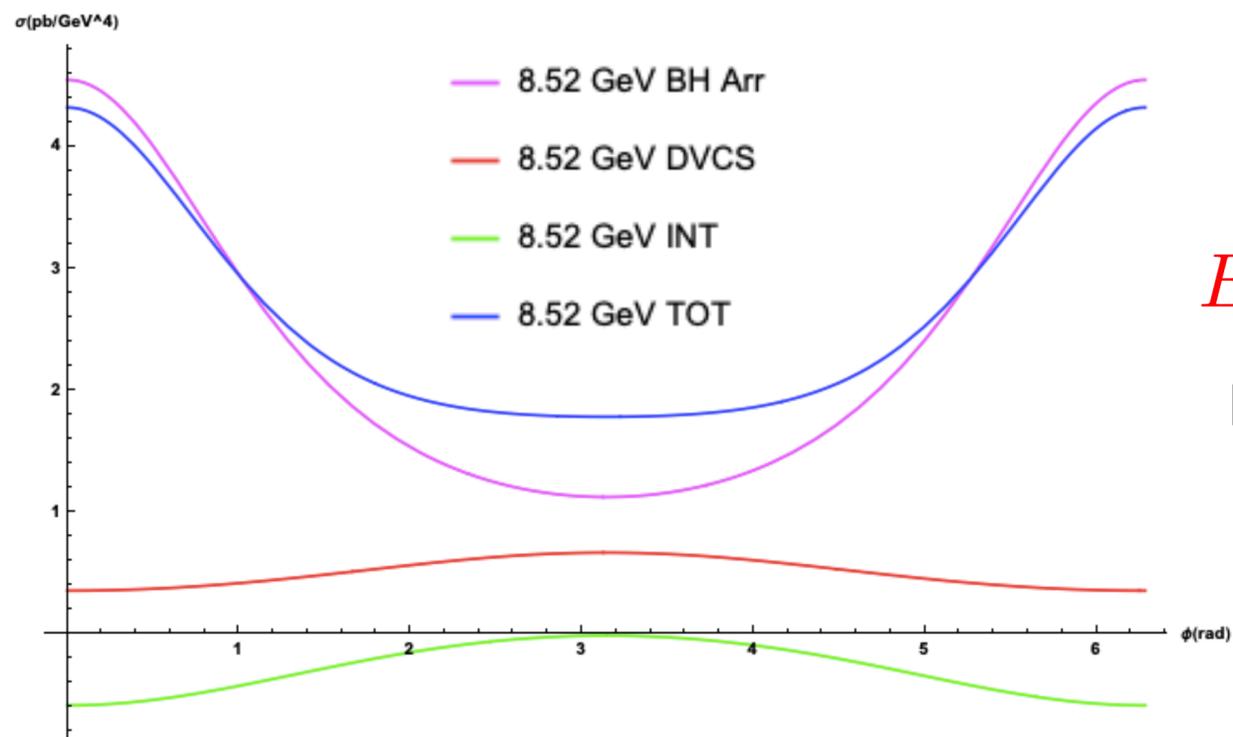
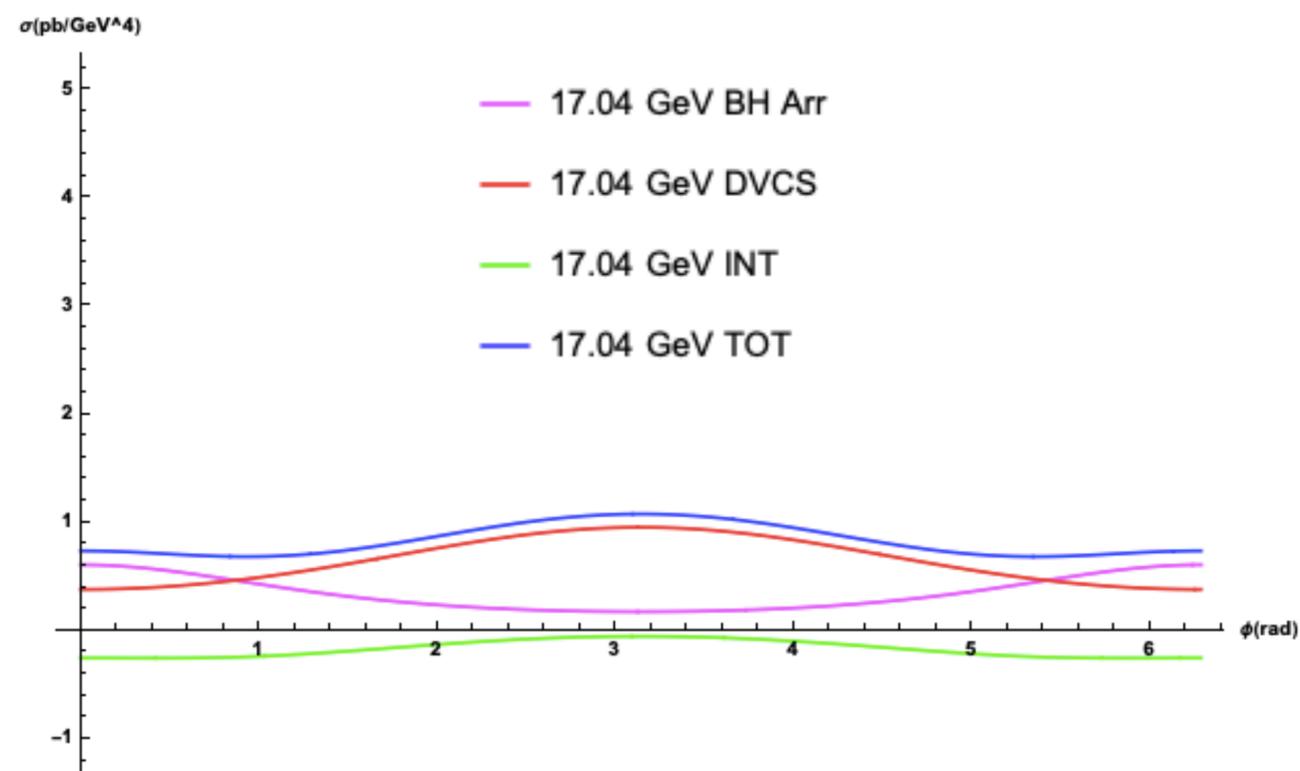
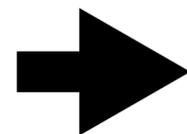


- The inclusion of higher luminosity, higher precision data results in determining CFFs (and thus GPDs) with smaller uncertainties

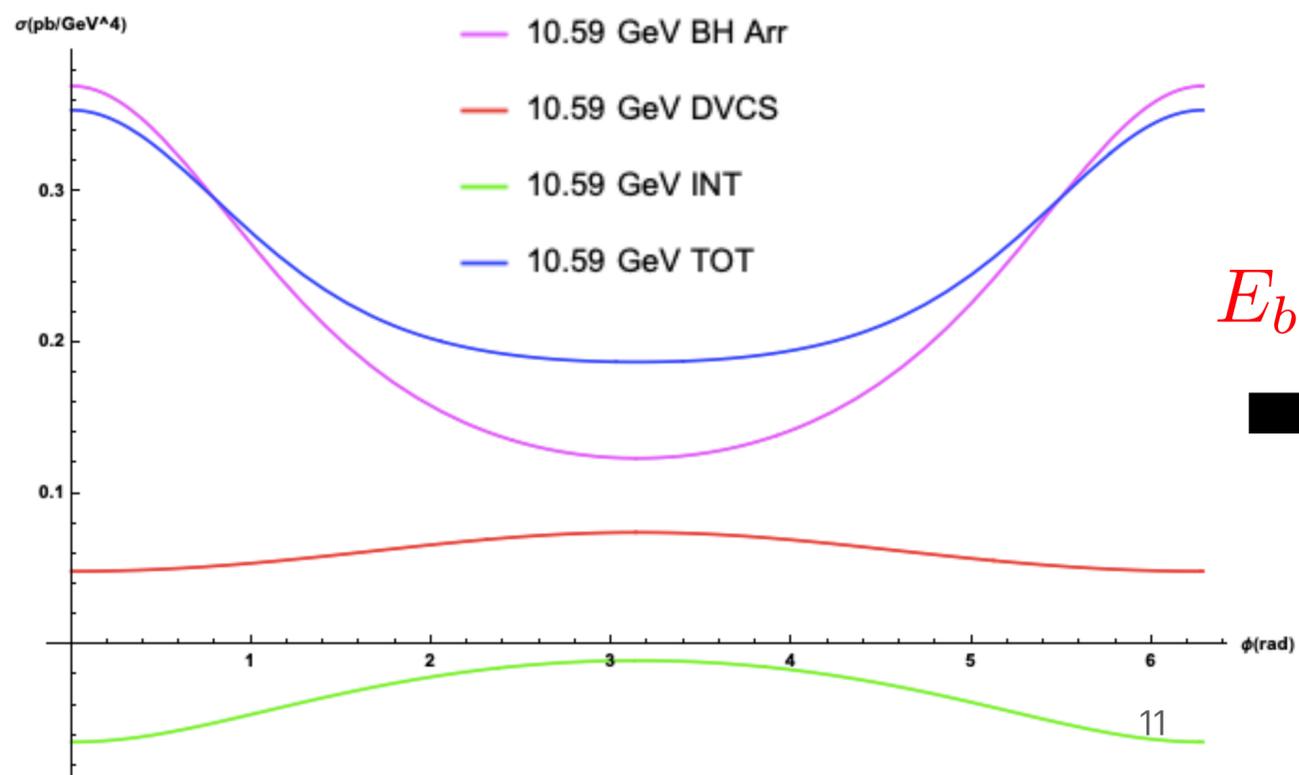
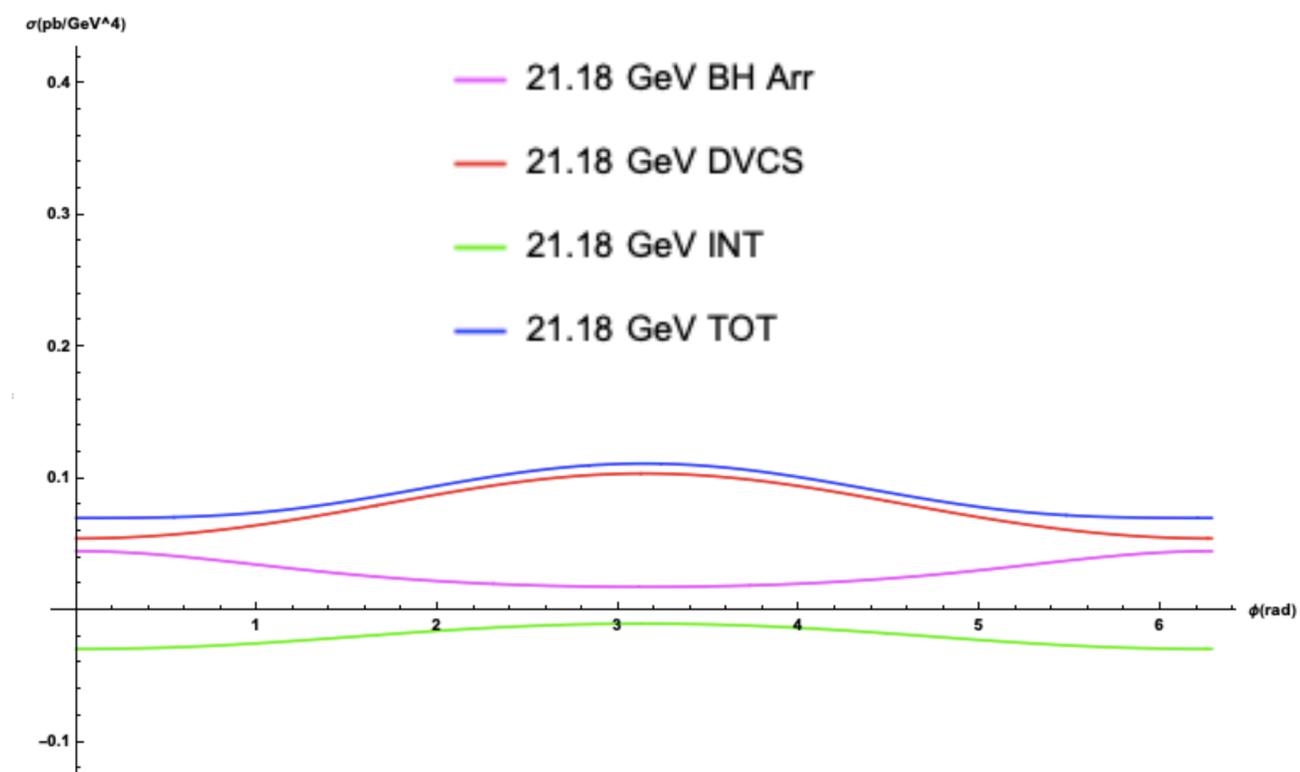
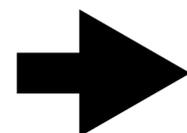
20 GeV DVCS Cross Section Predictions

- Assume the KM15 GPD model at kin. points where it agrees with JLab 12 GeV Hall A data, then predict the Bethe-Heitler (BH), DVCS and Interference cross sections at those kinematics in (E_b, ϕ, x_B, t, Q^2)
- As E_b increases (y decreases) the BH cross section decreases and so does the total cross section
- In general, the different harmonic coefficients of σ_{DVCS} & σ_{INT} behave differently with E_b , and since those cross sections make up a larger portion of the total, one can get a different ϕ dependence of the total cross section than at low energies

$$x_B = 0.48$$
$$t = -0.51 \text{ GeV}^2$$
$$Q^2 = 5.36 \text{ GeV}^2$$

 $E_b \times 2$ 

$$x_B = 0.6$$
$$t = -0.91 \text{ GeV}^2$$
$$Q^2 = 8.45 \text{ GeV}^2$$

 $E_b \times 2$ 

Additional Remarks about 20 GeV

- Bethe-Heitler cross section decreases while its uncertainty remains fixed
- However, the total cross section decreases too so the benefit appears to diminish
- Having multiple beam energies (doubled by upgrade) is a great benefit for CFF extraction, as they provide additional unique constraints on CFFs

$$\sigma_{\text{DVCS}}^{UU} = \frac{2\pi\Gamma}{Q^4} \sum_n h_n^U(\mathbf{E}_b; x_B, t, Q^2) \mathcal{D}(\mathcal{F}^2) \cos(n\phi) \quad \mathcal{F} = \mathcal{H}, \mathcal{E}, \dots(x_B, t, Q^2)$$

$$\#constr \approx \sum_{\text{pol.}} (\#E_b) \times (\#harm/pol.)$$

For stable CFF extraction (unique solution): $\#constr \geq \#param$

twist 2 CFFs \Rightarrow 8 param

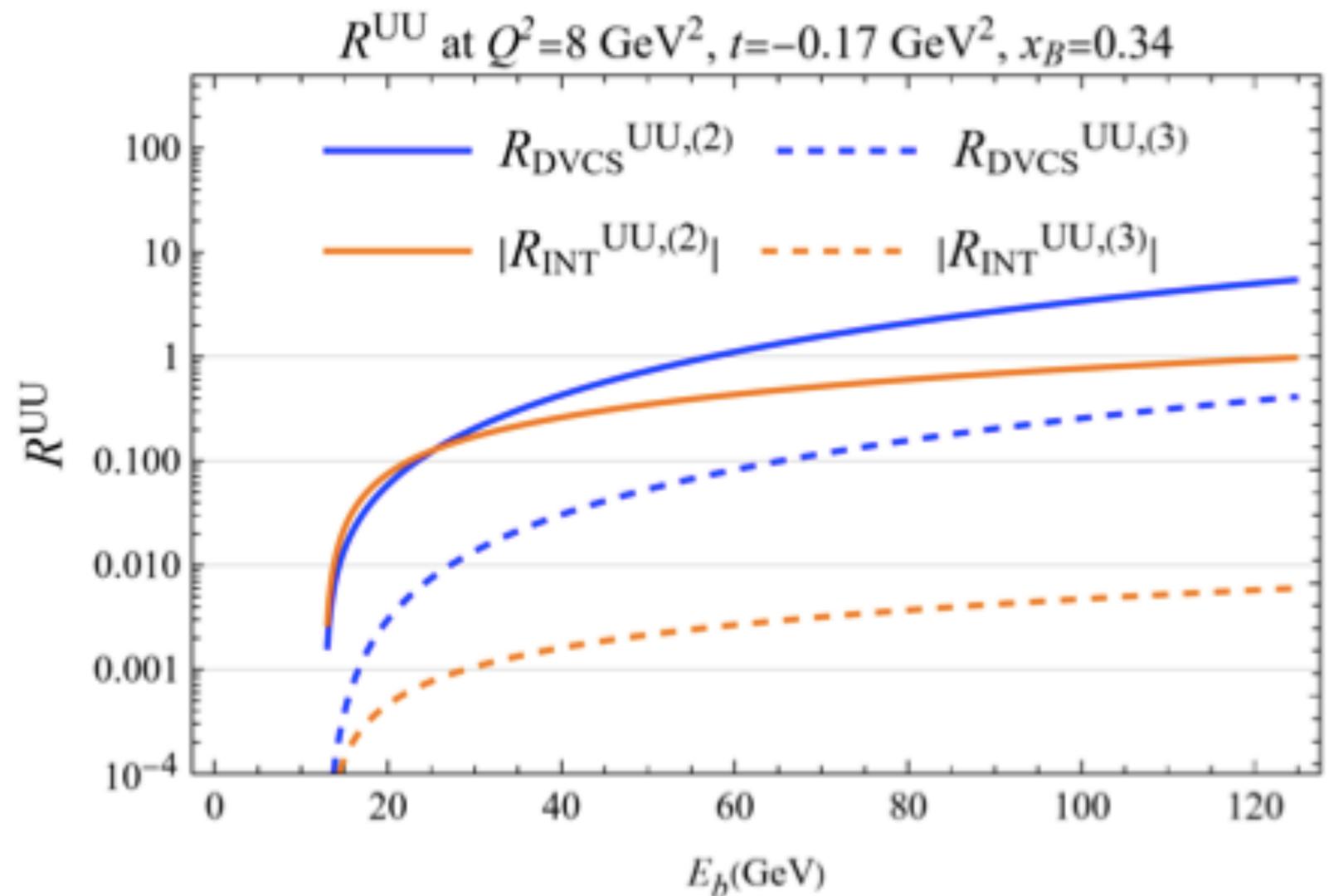
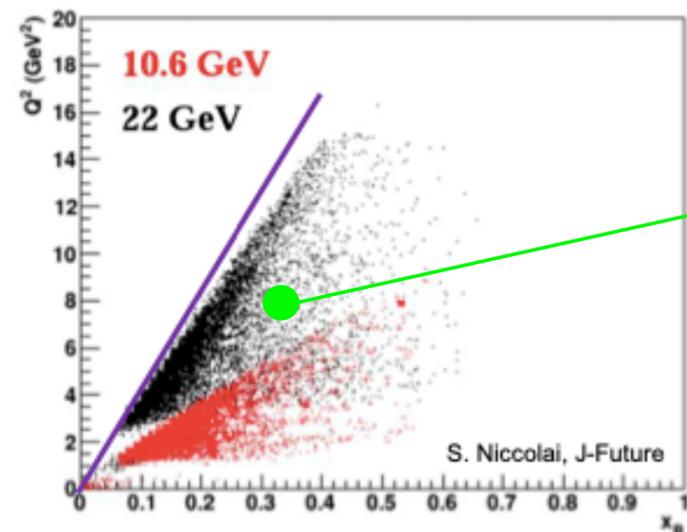
Higher Twist Suppression

- Higher beam energy allows higher Q^2 (at a fixed x_B), which allows one to get events where higher twist effects are very small
- Define ϕ -integrated cross section ratio:

$$R_i^{UU}(x_B, Q^2, t) = \frac{d^3\sigma_i^{UU}}{dx_B dQ^2 d|t|} \left(\frac{d^3\sigma_{BH}^{UU}}{dx_B dQ^2 d|t|} \right)^{-1}$$

$i = \text{DVCS, INT}$

$$R_i^{UU} = R_i^{UU,(2)} + R_i^{UU,(3)} + \dots$$

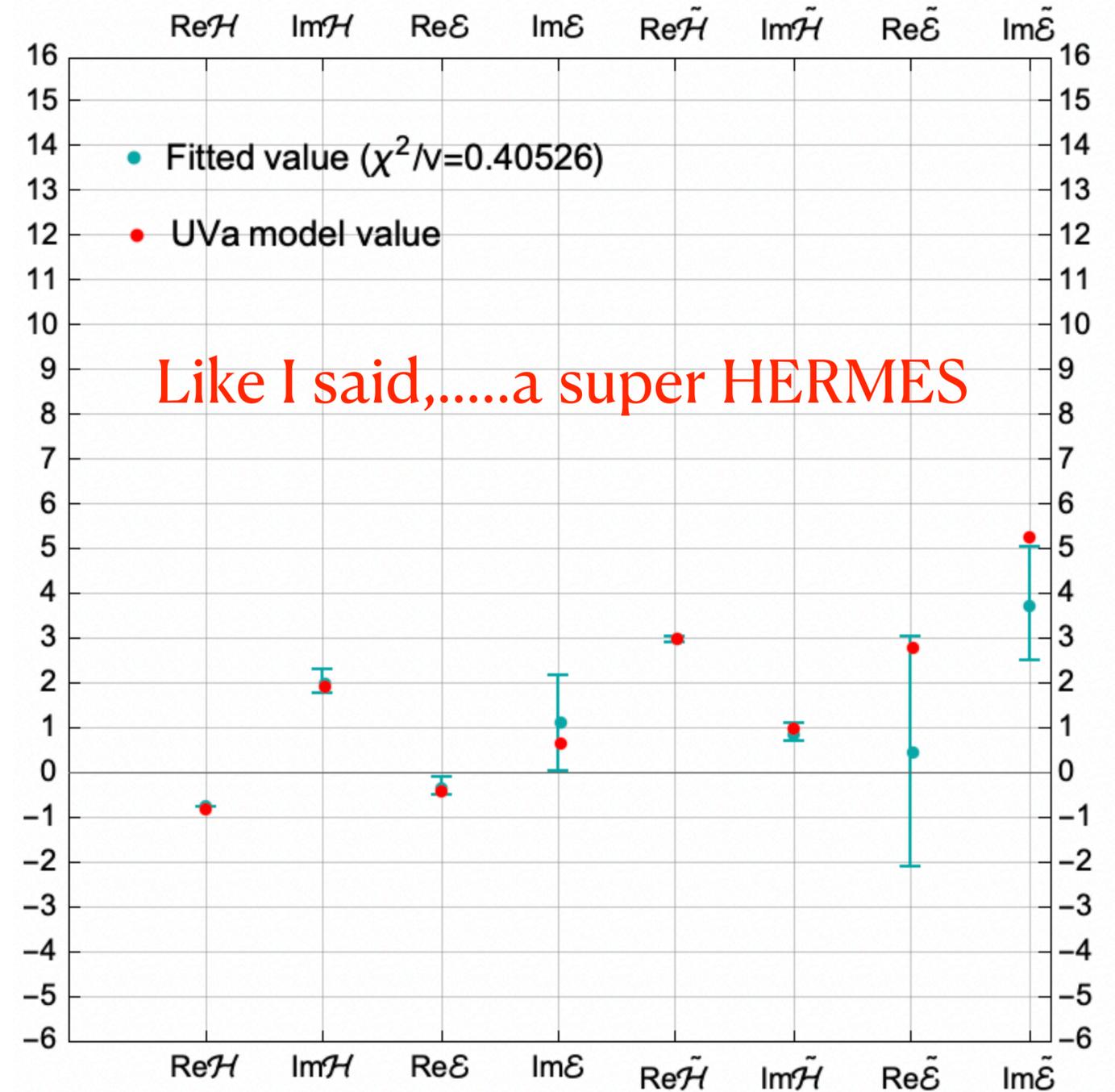
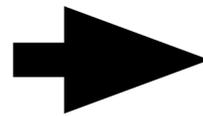
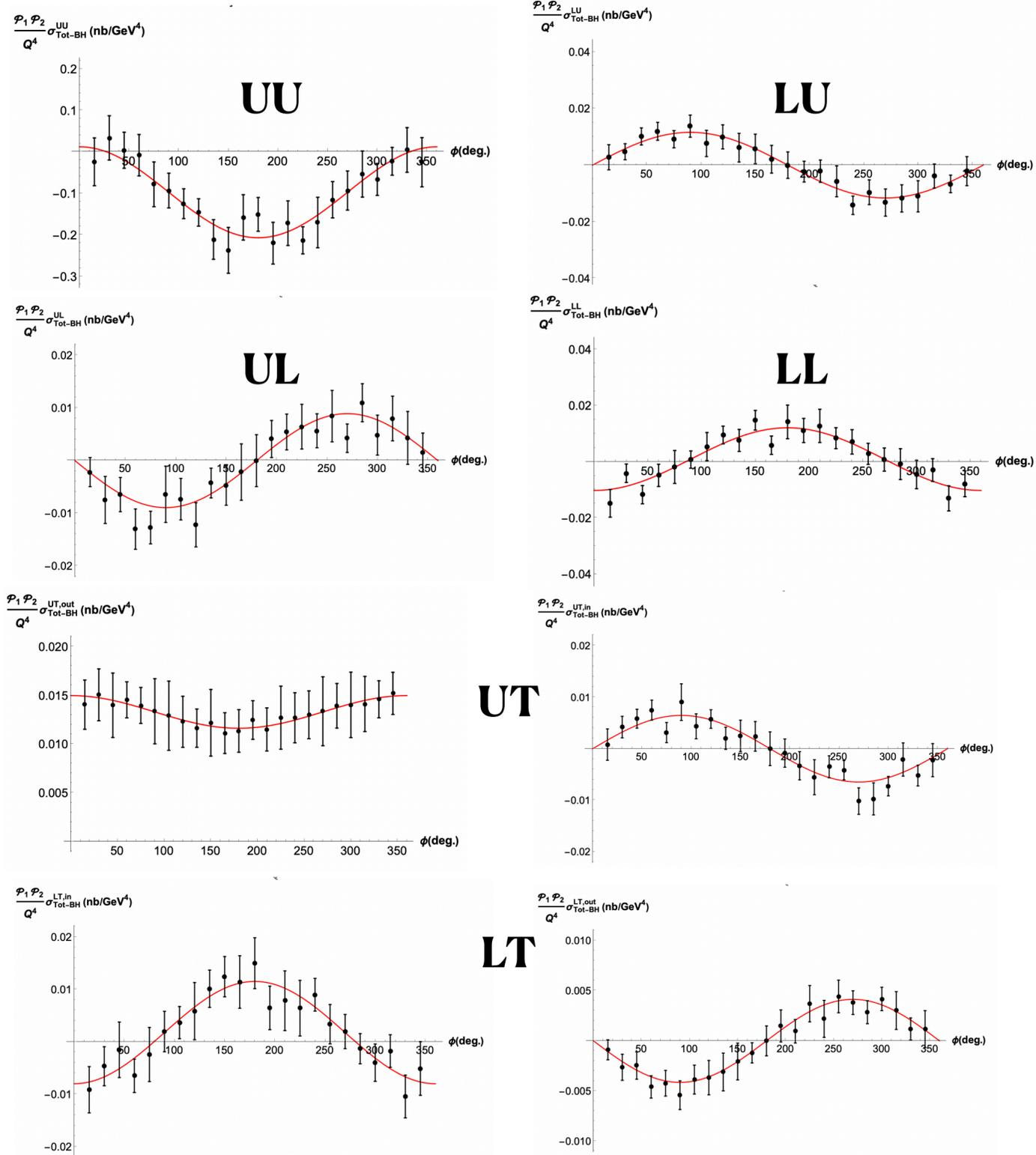


What we need next:

- We need MORE data:
 - A. More data points in general but also...
 - B. More kinematical coverage in: (Q^2, x_B, t)
 - C. More observable variety: UU, LU, **UL, LL, UT, LT**
 - D. Even extend that to charge asymmetries (positrons) and even other processes like DDVCS and TCS
 - E. Spread over full range of ϕ (and ϕ_S for Transverse targets)
- With that understood, going to **24 GeV gives additional advantages**

Hall B asymmetries

Observable Breadth: CFF Extraction



Final Remarks

- Higher CBAF energies allows greater kinematical coverage
- This together with a broad DVEP measurement program will truly make JLAB a “super HERMES”
- Additional discrete beam energy measurements **can help one better constrain CFFs** with limited number of observables
- The higher Q^2 reach allows one to **more cleanly extract pure twist-2 CFFs** with suppressed higher twist (≥ 3) contributions
- Evolution effects in GPDs may be studied more fully at the higher Q^2 as well