#### Parallelization of Air Shower Simulation with IceCube

Kevin Meagher University of Wisconsin May 8, 2023 CHEP2023 Norfolk, VA













### Overview

- Introduction to IceCube Neutrino Observatory
- IceCube's old simulation chain for cosmic-ray air showers
- Improvements to the simulation chain
- Results









### IceCube Laboratory

Data is collected here and sent by satellite to the data warehouse at UW-Madison

#### 1450 m

50 m

IceTop-



#### **Digital Optical** Module (DOM) 5,160 DOMs deployed in the ice

2450 m









### **IceCube Needs a Lot of Background Simulation**

- Cosmic-ray air showers occur at a rate of about 3kHz
- Neutrinos at a rate of about 2mHz
- An event selection that keeps 1 in 10<sup>6</sup> means we need a lot of background simulation
- Old simulation chain suffered from low CPU utilization, IO bottlenecks, and memory spikes
- Trend in Computing centers less Memory per CPU core



#### **Atmospheric Background**











- Simulates the development of Cosmic-ray air showers
- The energy distribution of daughter particles is highly stochastic
- Muons are the only daughter particle which is relevant for simulating IceCube's deep ice detectors
- CORSIKA is a single 100k fortran file that has to be compiled as a separate binary



#### **OpenCL Detector Simulation** GPU CPU **I3MCTree I3Photon Propagate Photons Detector Response** Primary Cosmic Ray 11 1 1 1

IceTray - IceCube's processing framework







- "PROPOSAL" is the name of a software package not an actual proposal
- Propagates muons as they travel through ice
- Muons energy loss is also highly stochastic and • non-parallelizable
- All muons will lose a significant fraction of their energy before making it to the deep ice detector
- Only muons with very high energy will make it to the detector









- Light emitted by muons is propagated through the Ice
- Layers of different concentrations of dust cause the scattering and absorption of the ice to be different at different heights
- Photon propagation is highly parallelizable and can only be effectively performed on GPUs with OpenCL/CUDA













#### **Detector Response/Reconstruction**

- Simulates DOM Photomultiplier tube and front-end electronics
- Simulates noise
- Simulates the data acquisition system and trigger
- Performs all reconstructions the same as if it were experimental data





## **Problems with Old Simulation Chain**



Most air showers don't produce muons that have enough energy to make it to the detector: Low energy showers have very small chance of producing a high energy muon Inclined showers have a much higher threshold to make it to the detector









- Move to a Client/Server model for Air Shower, Muon, and Photon propagators
  - Each client has a work queue can span multiple events  $\bullet$
- 1 instance of CORSIKA Server for air showers propagation
- 1 instance OpenCL Server collects work units and sends them to the GPU
- 8 instances of PROPOSAL for muon propagation
- 8 instances of the simulation chain are used in parallel to connect to the propagation servers
  - Keeps the GPU fully utilized





#### **I3PrimaryInjector**

- Primaries are now sampled from new IceTray module rather than hard coded  $\bullet$ functions in CORSIKA
- Primaries are generated on directly on the detector surface with arbitrary energy  $\bullet$ and spatial distributions
- Allows for future schemes for oversampling regions that produce more background







#### **CORSIKA Server**

- CORSIKA was modified to run as a server which can be configured with individual primaries (thanks to D. Baack)
- Different CORSIKA cards for different showers
  - The energy needed to reach the detector is highly dependent on the inclination of the shower
  - The energy at which CORSIKA will stop propagating particles is now higher for inclined showers
- Showers with higher energy leading edge muons are undersampled:
  - Showers with low energy muons are killed before the rest of the shower is calculated
- Since individual particles are sent over IPC no files are written avoiding IO bottleneck





muons to reach the detector



New CORSIKA Server will set the minimum muon energy higher for more inclined showers





#### **Memory Resources**

- Memory is saved with more efficient dataclass for storing photons: I3MCPE  $\bullet$
- The Propagation Client's work queue doesn't have to store intermediate data for the entire event reducing memory usage
- The memory is amortized across the 8 instances of icetray lacksquare
  - 8 instance sharing 8GB are much less likely to be taken down by a memory spike than 8 jobs with 1GB each  $\bullet$

![](_page_12_Picture_7.jpeg)

![](_page_12_Picture_12.jpeg)

## **Optimizing Leading Muon Undersampling**

![](_page_13_Figure_2.jpeg)

![](_page_13_Picture_4.jpeg)

#### Air showers are undersampled based on the energy of the first muon to appear in the shower

![](_page_13_Picture_7.jpeg)

### **Comparison Between New and Old Simulation**

![](_page_14_Figure_1.jpeg)

Good agreement is seen between old and new

![](_page_14_Picture_3.jpeg)

![](_page_14_Picture_5.jpeg)

#### Performance

![](_page_15_Figure_1.jpeg)

![](_page_15_Picture_2.jpeg)

	Old Simulation Chain	New Simulation Chain
	101 µs	3.70 µs
	7GB	1GB
obs/	60%	84%
	45%	80%

![](_page_15_Picture_5.jpeg)

![](_page_16_Picture_0.jpeg)

# Questions

![](_page_16_Picture_2.jpeg)