

Calibration Data Flow and Performance at Belle II



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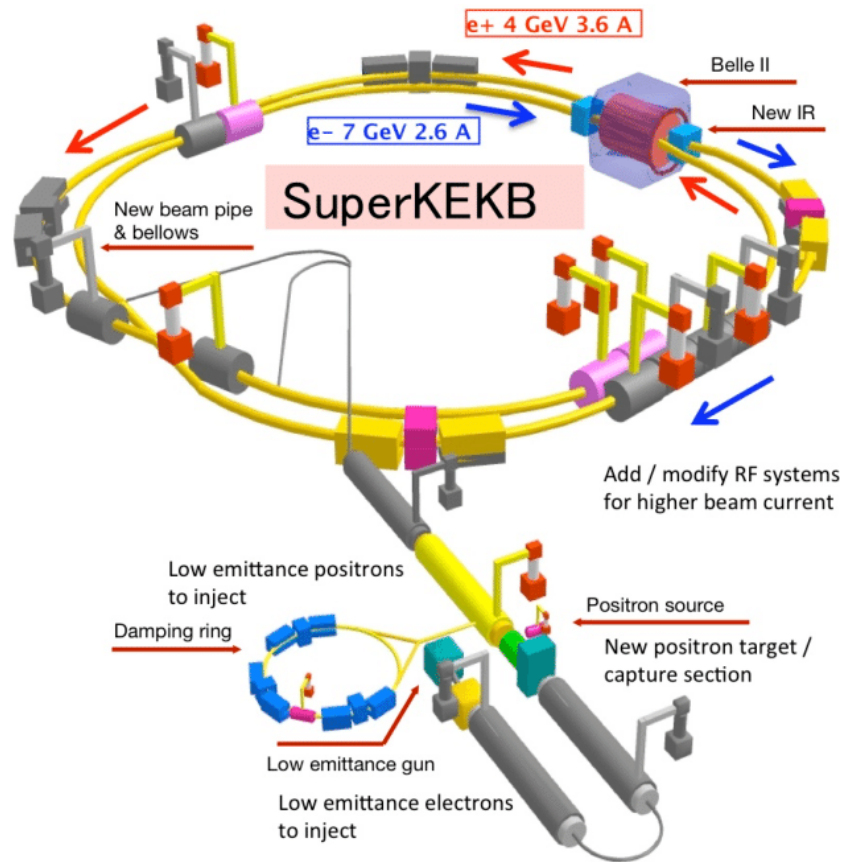


Bundesministerium
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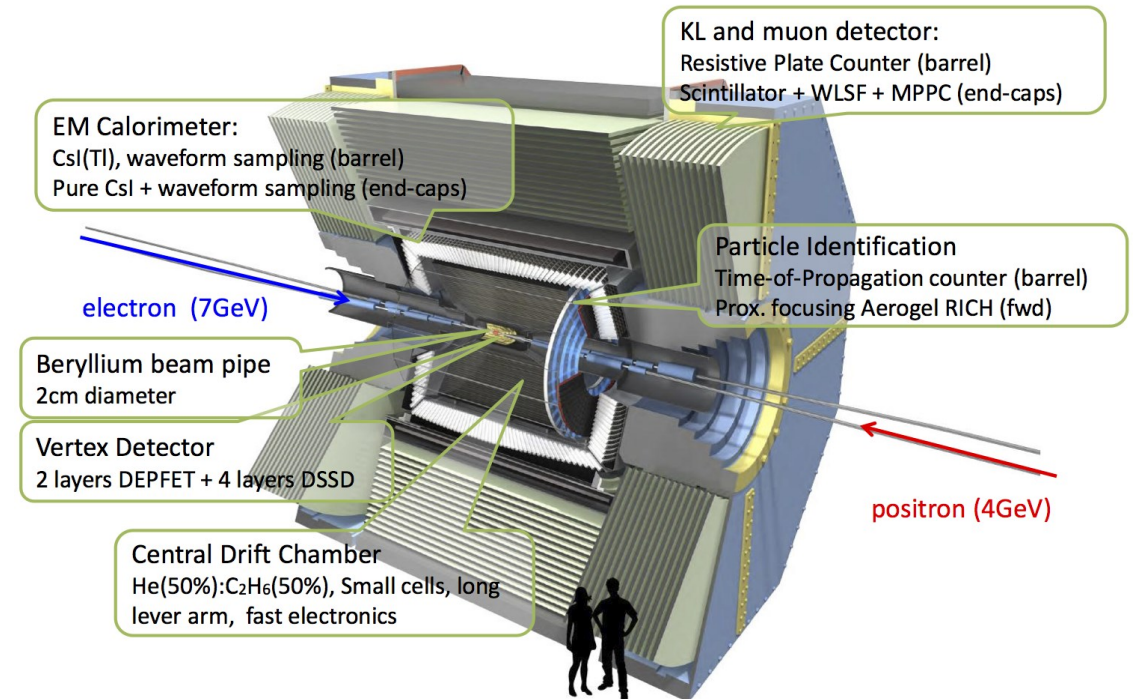
From e^+e^- collisions to raw data



**Asymmetric e^+e^- collider
@ and around the $Y(4S)$ resonance**

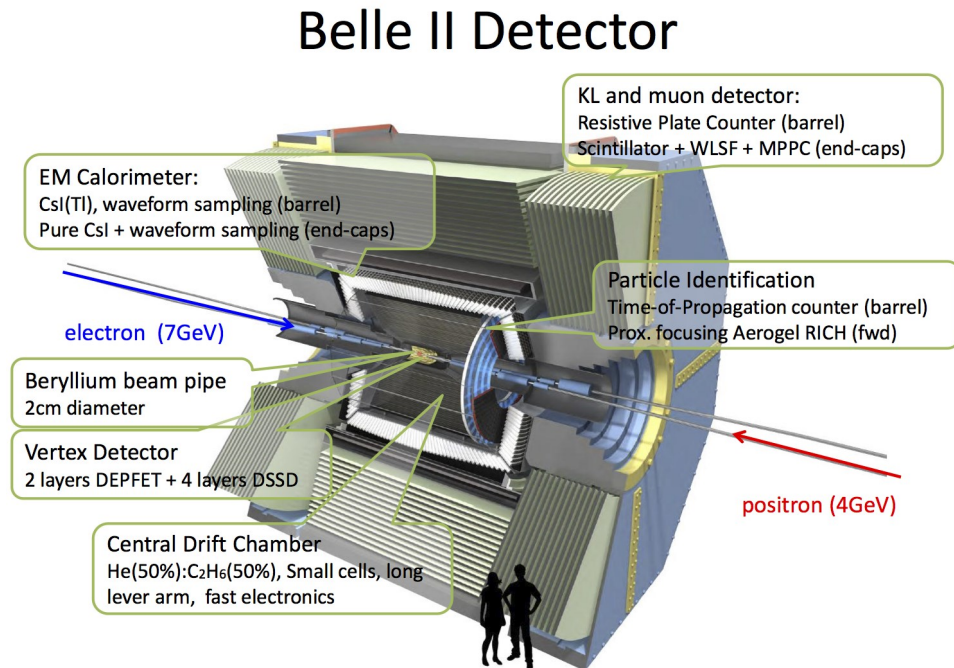


Belle II Detector

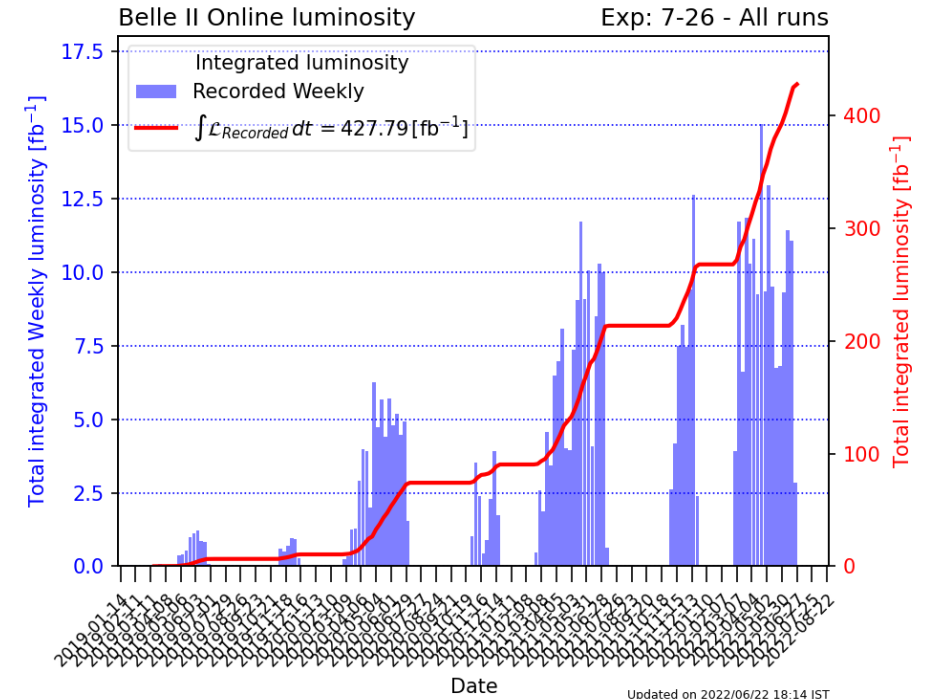


**General purpose
 4π detector**

From e^+e^- collisions to raw data

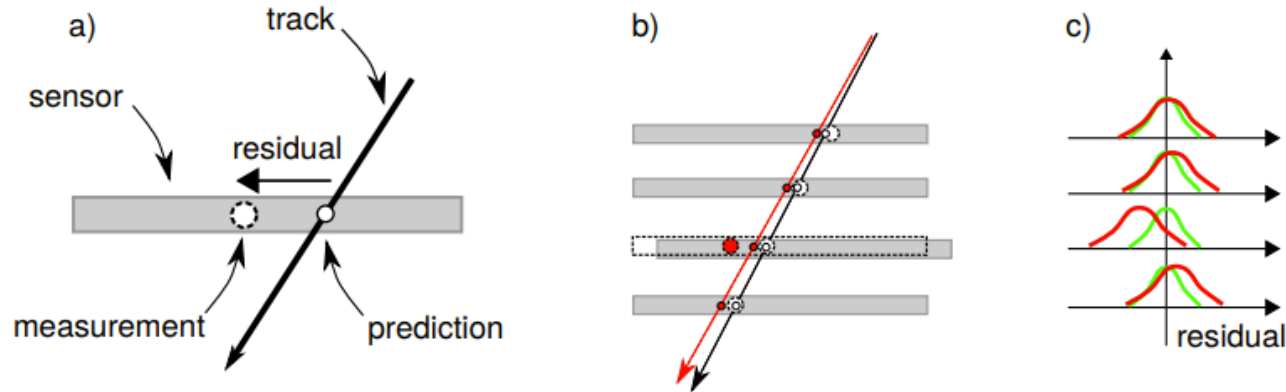


General purpose
 4π detector



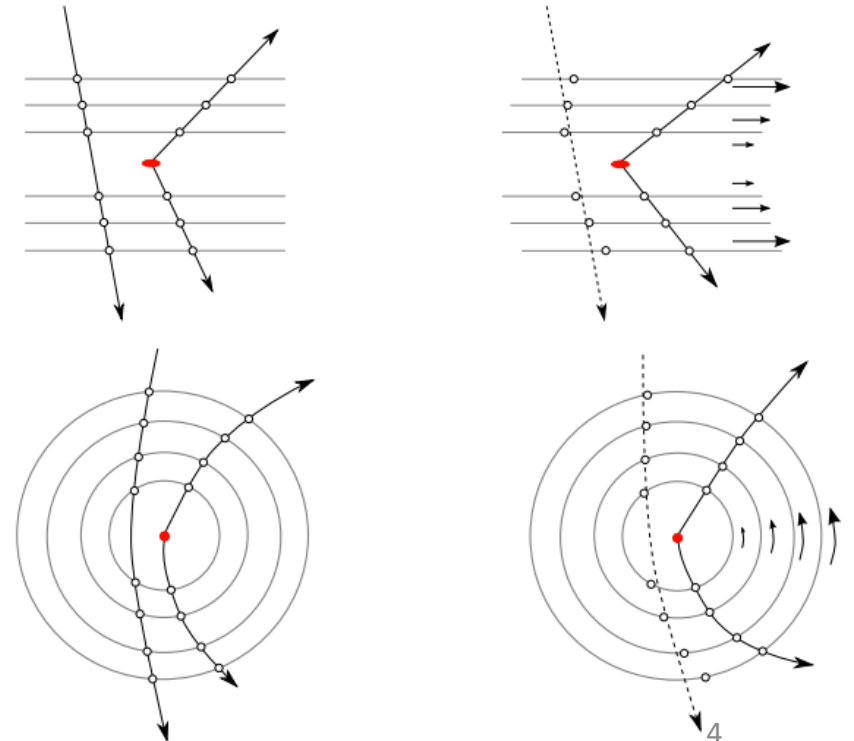
Total integrated luminosity
 $\mathcal{L} = 424 \text{ fb}^{-1}$

Alignment – A brief introduction



Challenges:

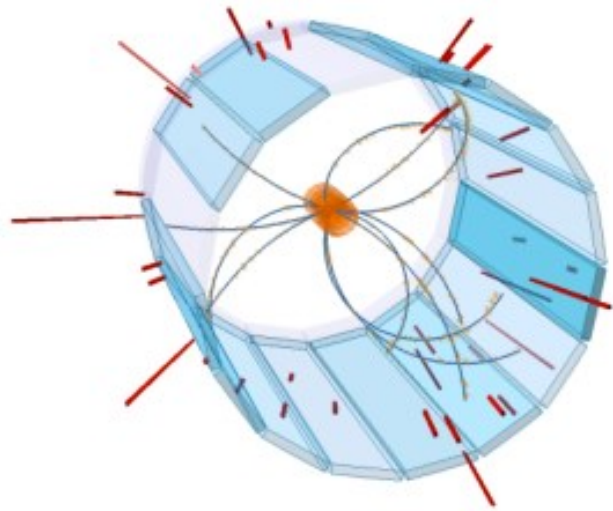
Weak modes due to lack of absolute reference
→ Complementary data samples



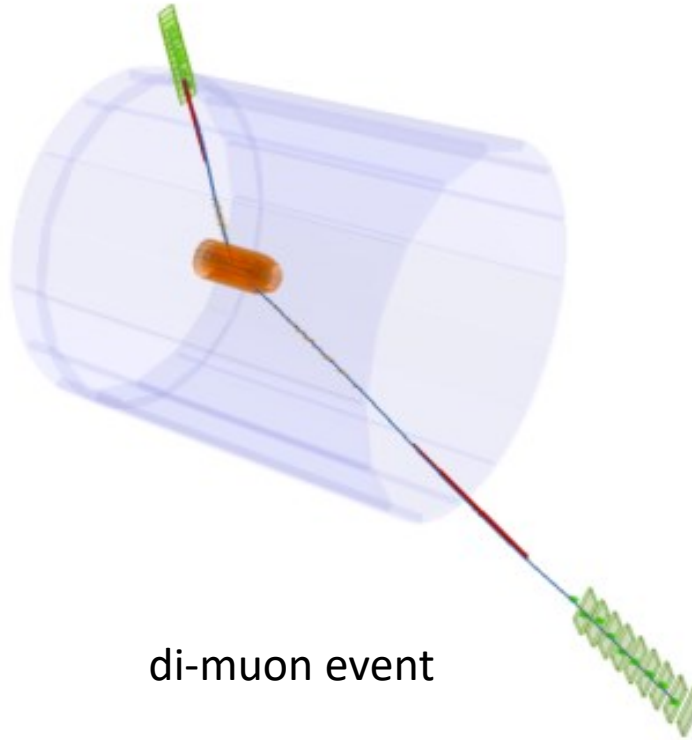
**Belle II: Global Alignment
with Millepede II**



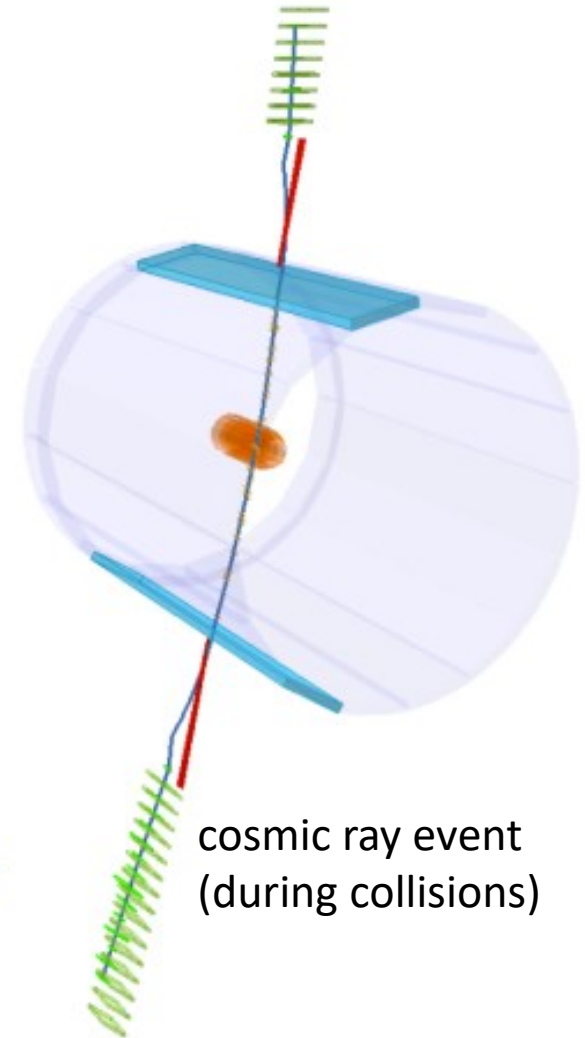
Alignment – data samples



hadron event

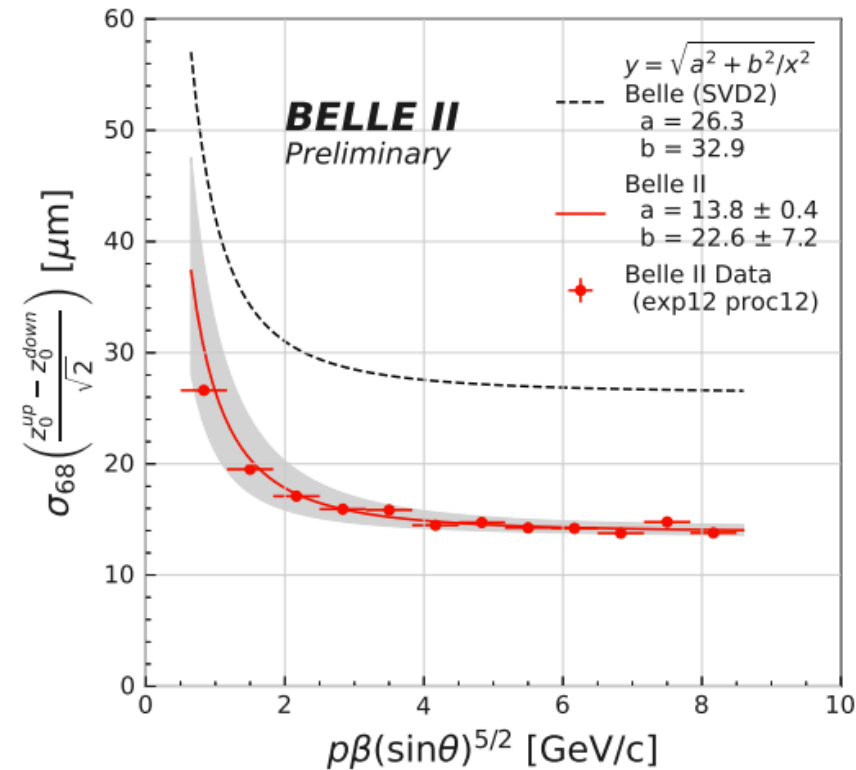
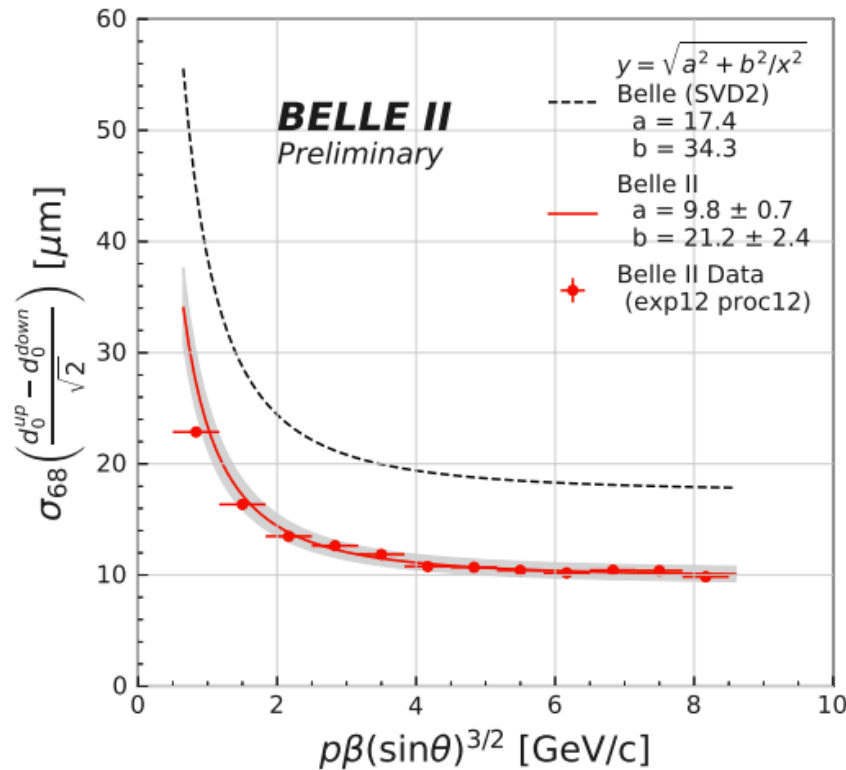


di-muon event



cosmic ray event
(during collisions)

Alignment – Resolution of track parameters



$$\sigma^2(\tilde{p}) = a^2 + \frac{b^2}{\tilde{p}^2}$$

$$\tilde{p}_{d_0} = p\beta \sin(\theta)^{3/2}$$

$$\tilde{p}_{z_0} = p\beta \sin(\theta)^{5/2}$$

a : detector resolution

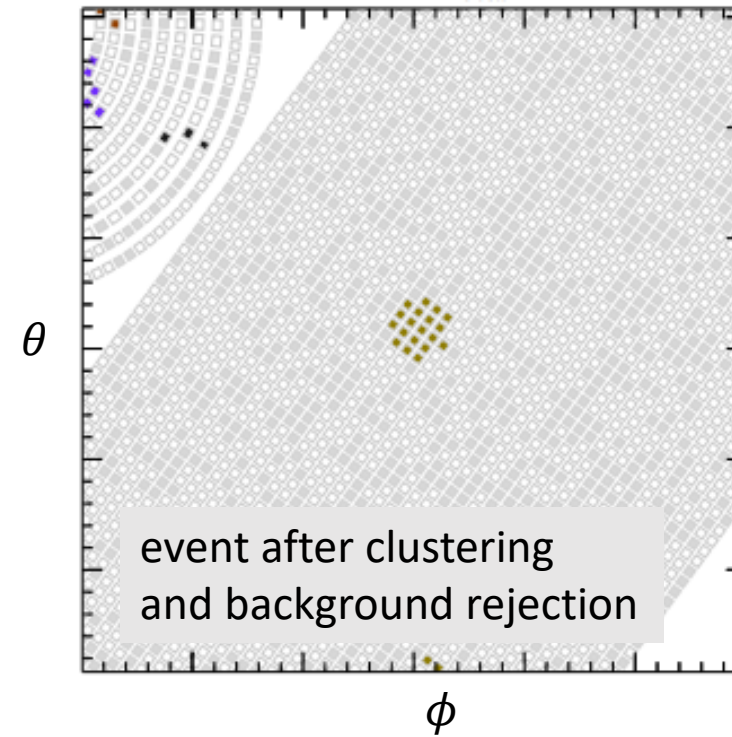
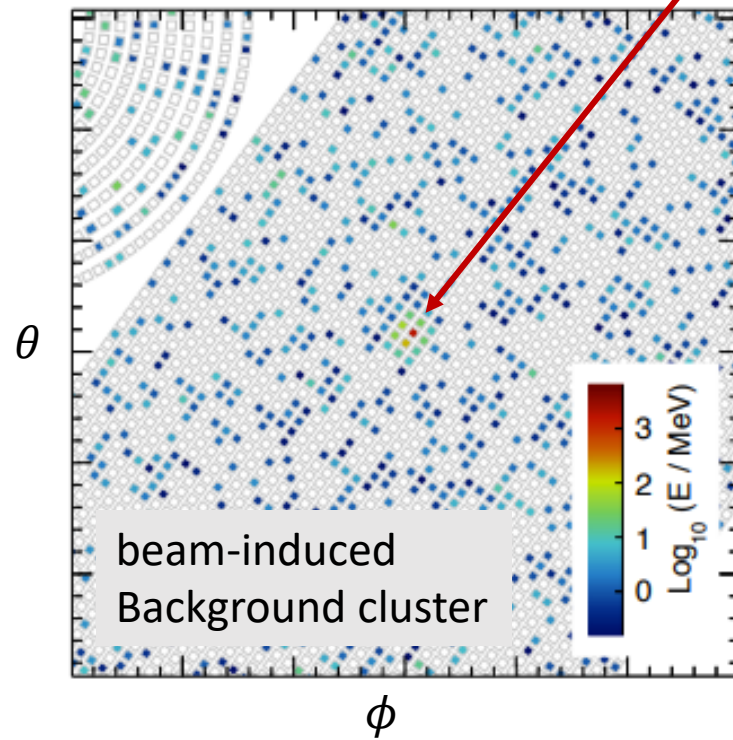
b : material budget affecting multiple scattering

\tilde{p} : pseudo-momentum

ECL reconstruction

Simulated single photon
and beam background

signal
photon



ECL calibration

ECL leakage

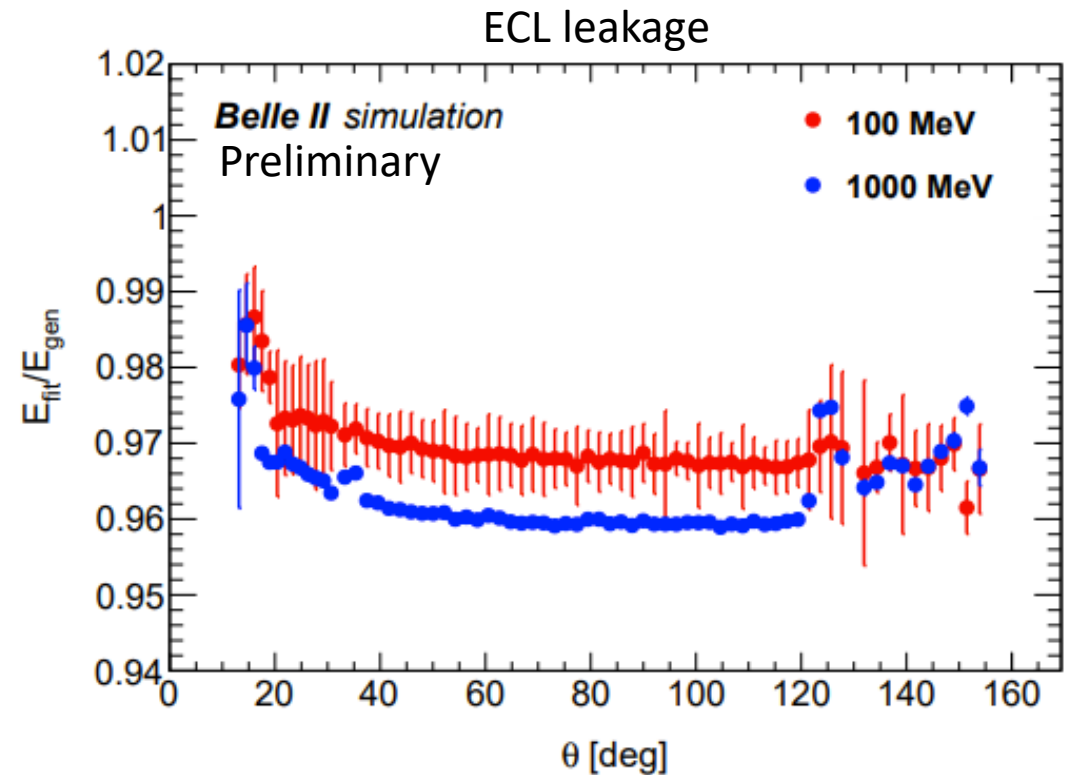
- Reconstructed energy smaller than true energy through leakage
- Derived from simulated single photon + beam background

Absolute energy

- $ee \rightarrow \gamma\gamma$ or cosmic ray events

Final photon energy

- symmetric π^0 and η candidates



Calibration setup

- Small background rates allow for loose trigger lines
 - The high-level trigger does not require precise calibrations
- Offline calibrations feasible

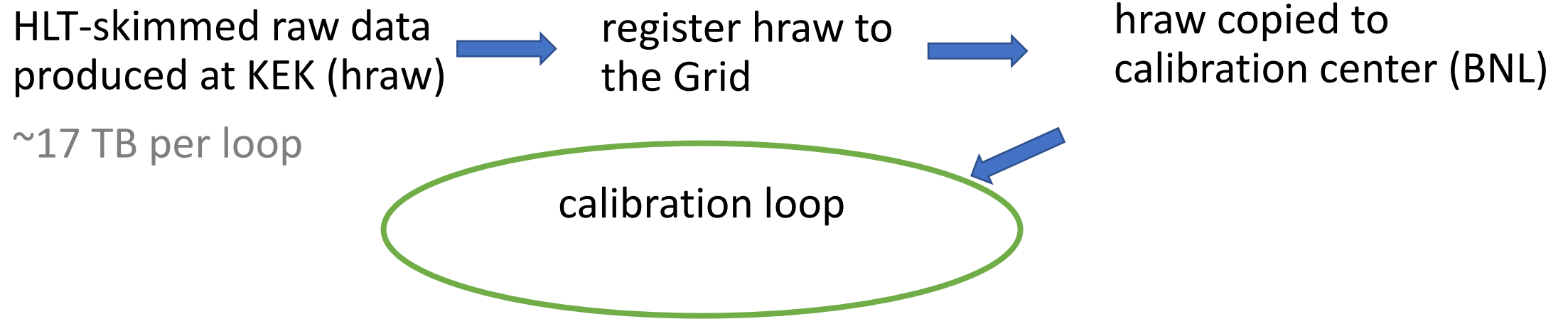
Advantage:

Calibration of the bucket with the current bucket data instead of the previous bucket data

Disadvantage:

Data “lagging” behind, but prompt calibration done within weeks of data taking

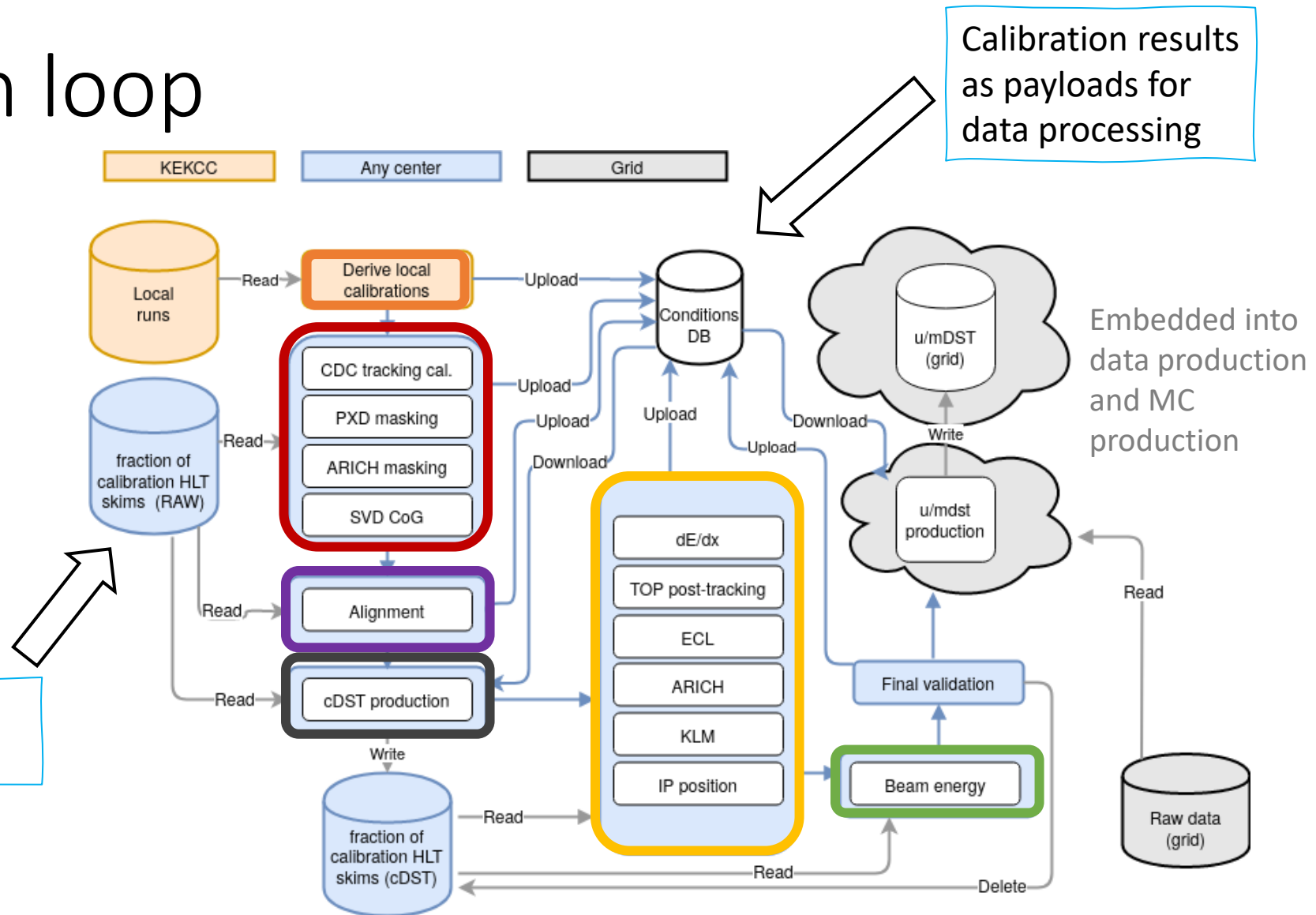
Data Flow



The calibration loop

1. Local calibrations
2. Pre-tracking/raw-data based calibrations
3. Alignment
4. cDST production
5. Post-tracking/data based calibration
6. Analysis-based calibrations

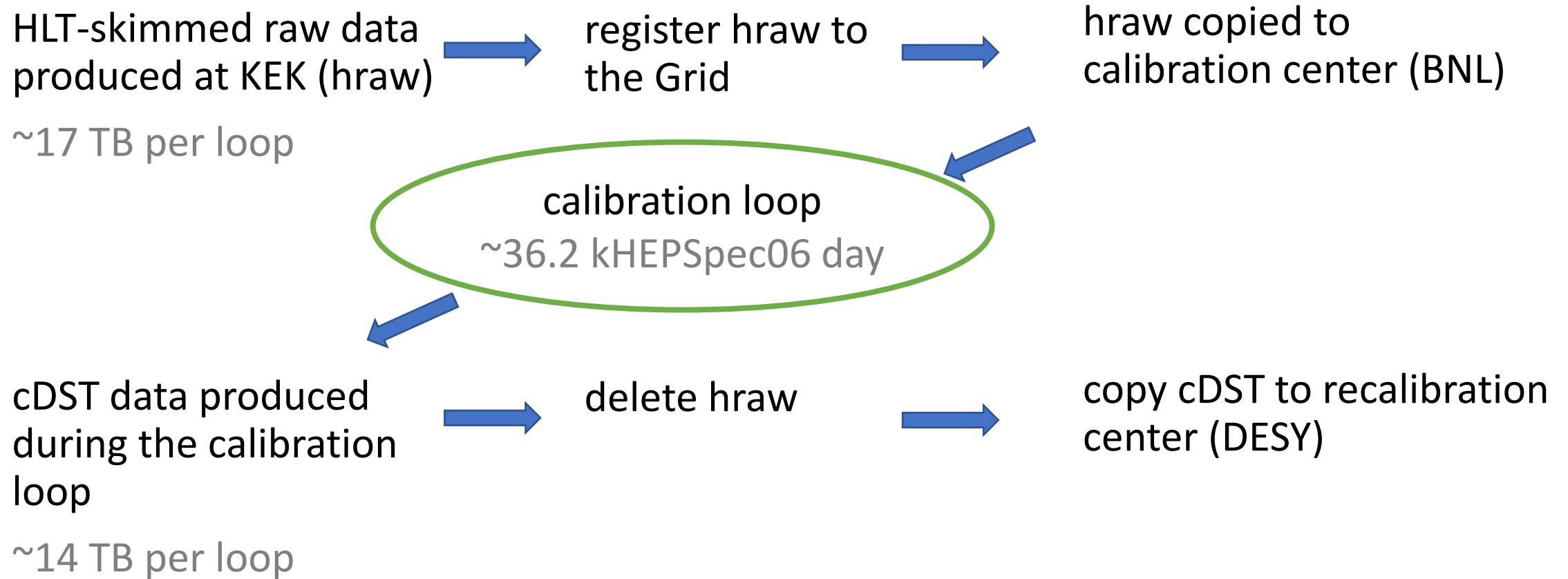
Calibration skims [hraw]
(adaptive prescaling to 9 fb^{-1})



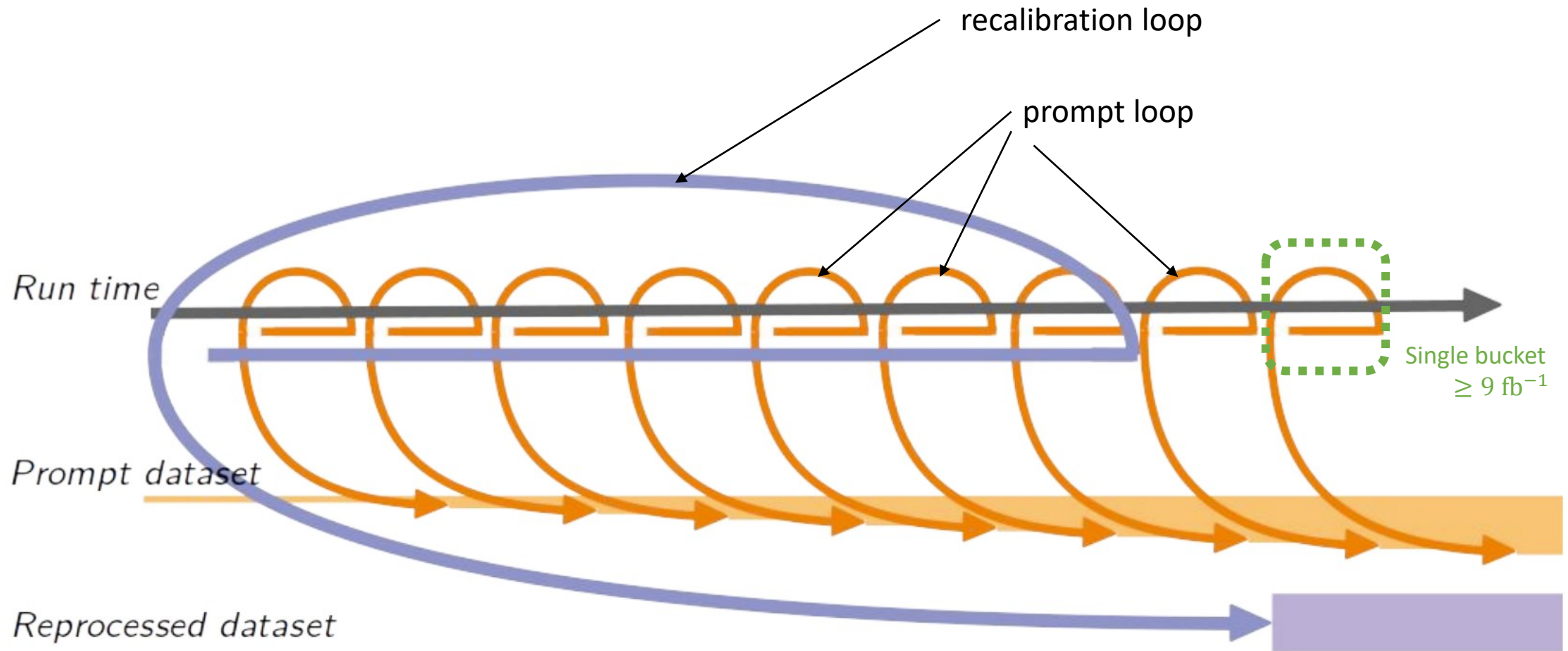
**The calibration loop is fully automated
and provides physics data within ~2 weeks of data taking**

Technical Details:
Talk by David Dosset
11.05.2023, 14:45 11

Data Flow



The Belle II calibration procedure



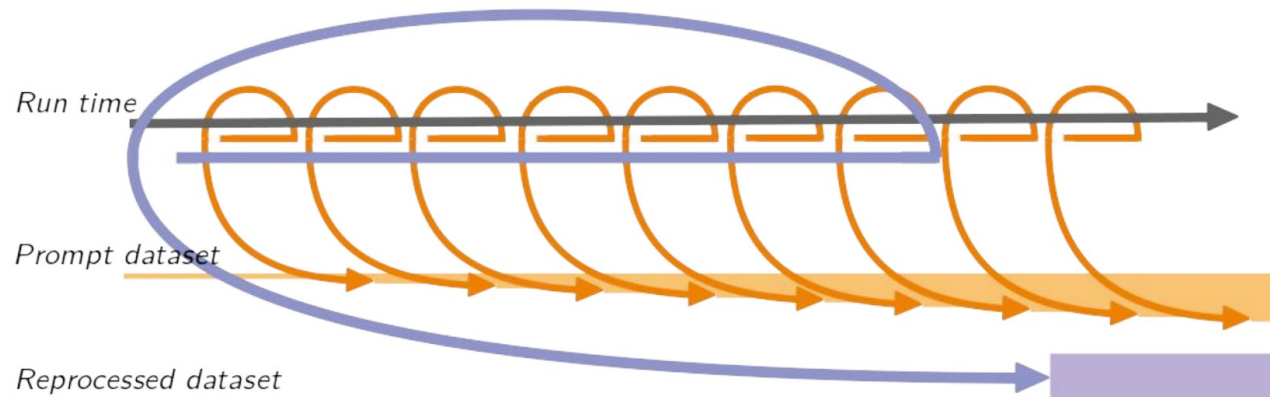
The Belle II calibration procedure

Promptly after data taking: prompt calibration @ BNL

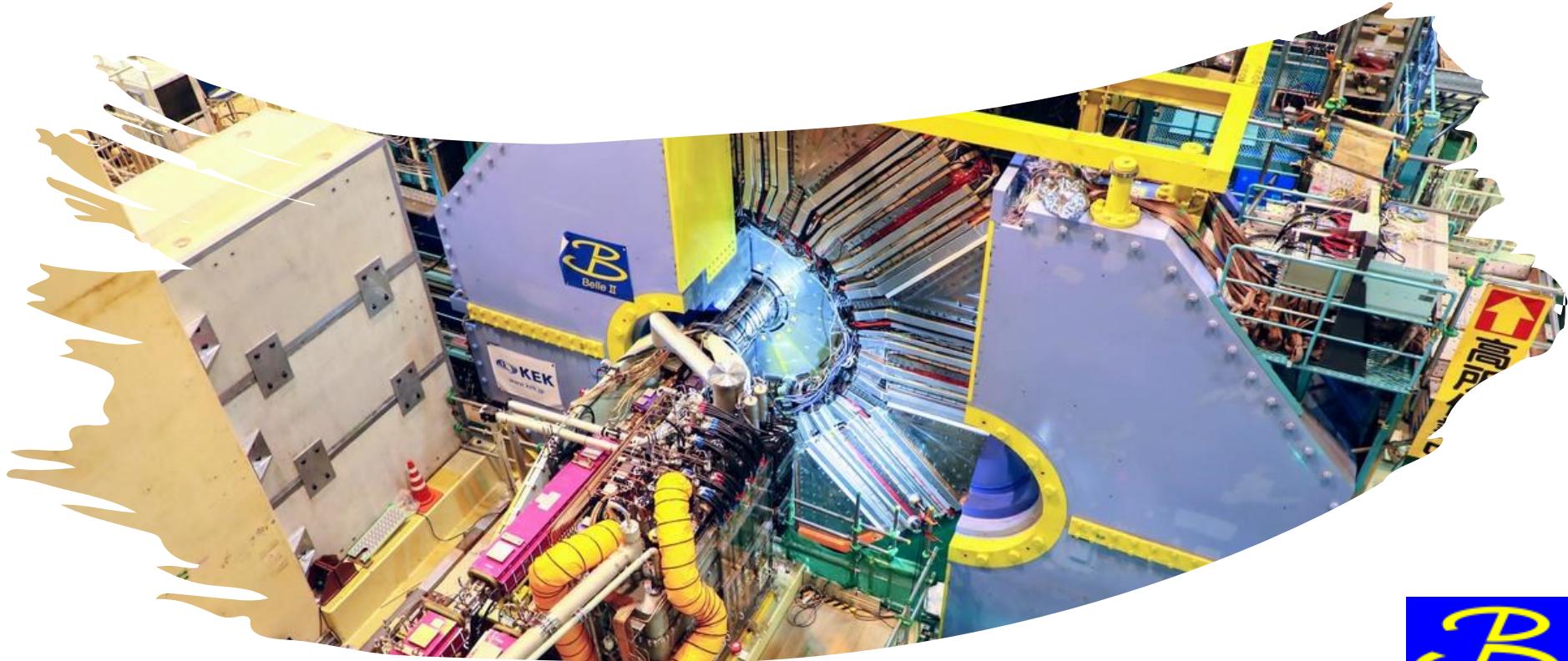
- Minimum luminosity (*) per prompt calibration “bucket”: 9 fb^{-1}
- All calibrations included
- Uses hraw as input
- ~one bucket every 2 weeks of data taking
- Ideally already the final calibration

Before data reprocessing: recalibration @ DESY

- To update older data with the latest reconstruction software
- Only calibrations with expected improvement
- Uses cDST as input
- ~once/year until 2025, every other year starting from 2025



(*) Data of each individual bucket is adaptively prescaled to correspond to 9 fb^{-1}



Summary

- Automation of the Belle II calibration
- Calibrations ready for high precision physics