



New experimental frontiers in the study of the hadronic interactions with ALICE at the LHC

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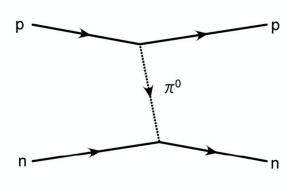
On behalf of the ALICE Collaboration

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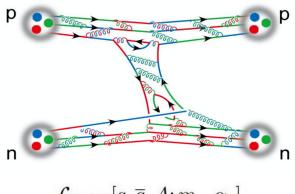
Residual strong interaction among hadrons



$$\mathcal{L}_{EFT}[\pi, N, \ldots; m_{\pi}, m_{N}, \ldots, C_{i}]$$

Non-perturbative region of QCD

- Hadrons as degrees of freedom
- Effective Field Theories (EFT) with low-energy coefficients constrained by data

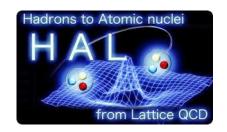


 $\mathcal{L}_{QCD}[q,\overline{q},A;m_q,\alpha_s]$

Lattice QCD

-Understanding of the interaction starting from quarks and gluons

Residual strong interaction from lattice

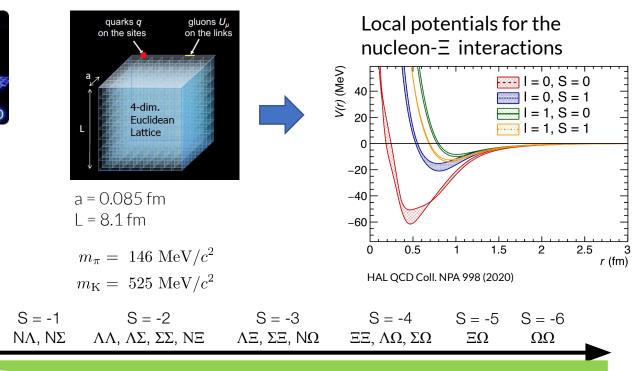


T. Hatsuda, K. Sasaki et al.

S = 0

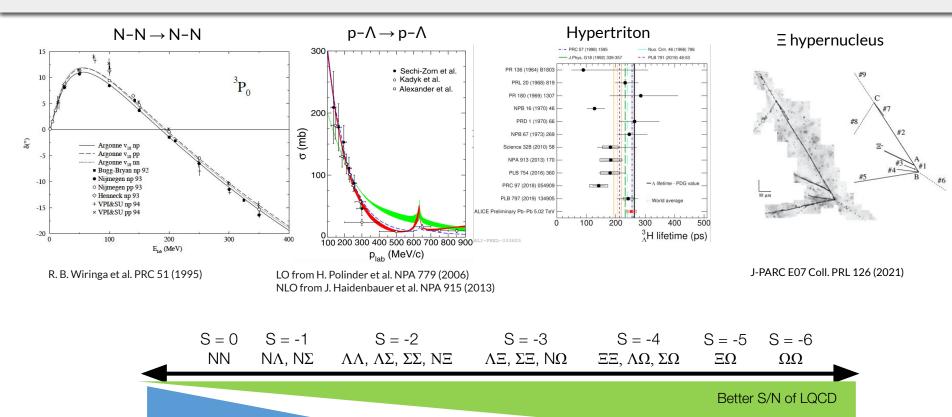
NN

HAL QCD Coll. PLB 792 (2019) HAL QCD Coll. NPA 998 (2020) HAL QCD Coll. PRD 99 (2019)



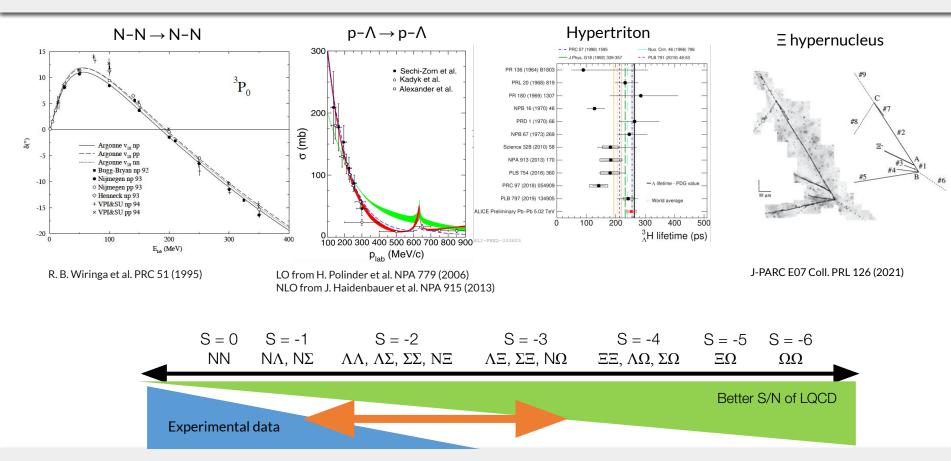
Better S/N of LQCD

Experimental data for two-body interactions



Experimental data

Experimental data for two-body interactions



Impact on the Equation of State of neutron stars

Neutron stars

Dimensions

R ~ 10 - 15 km

 $M \sim 1.5 - 2.2 M_{\odot}$

Outer Crust

Ions, electron gas, Neutrons

Inner Core

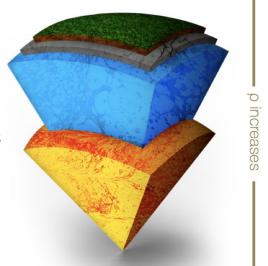
Neutrons?

Protons?

Hyperons?

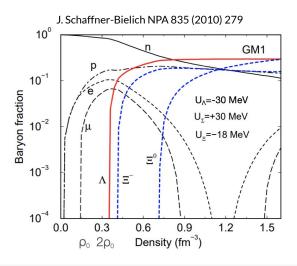
Kaon condensate?

Quark Matter?

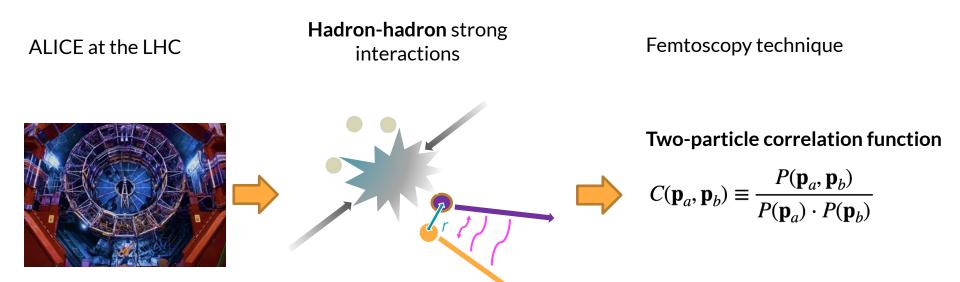


Neutron stars are very dense, compact objects

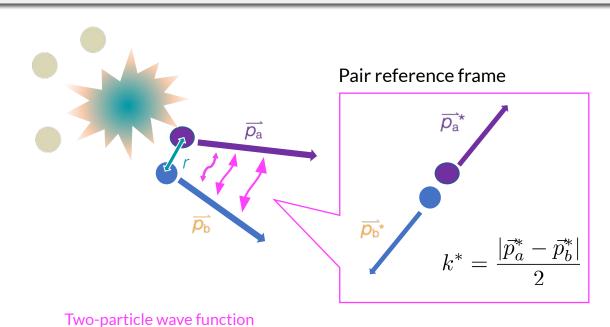
What is the Equation of State?
What are the constituents to consider?
How do they interact?



Investigating hadronic interactions at LHC



Femtoscopy technique



Correlation function:

$$C(k^*) = \mathcal{N} \cdot \frac{N_{\text{same}}(k^*)}{N_{\text{mixed}}(k^*)} = \int S(r) \left| \psi(k^*, r) \right|^2 d^3r$$

Emission source

Femtoscopy technique

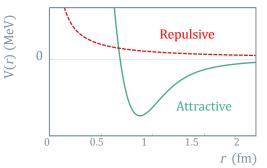
Source parameterization



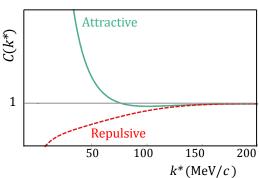


Gaussian source

Interacting potential



Correlation function



Schrödinger equation

CATS (<u>C</u>orrelation <u>A</u>nalysis <u>T</u>ool using the <u>S</u>chrödinger equation) D. Mihaylov et al. EPJC 78 (2018)

Correlation function:

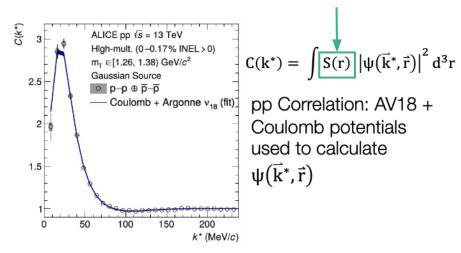
$$C(k^*) = \mathcal{N} \cdot \frac{N_{\text{same}}(k^*)}{N_{\text{mixed}}(k^*)} = \int S(r) \left| \psi(k^*, r) \right|^2 d^3r$$

Emission source

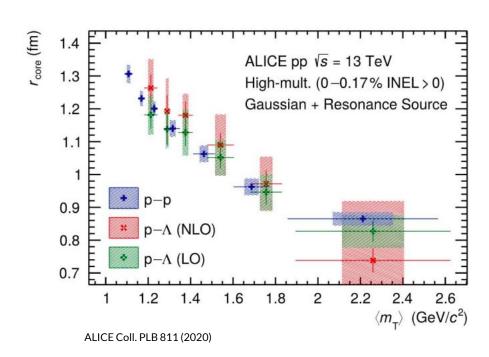
Two-particle wave function

- >1 if the interaction is attractive
- = 1 if there is no interaction
- <1 if the interaction is repulsive

Source model

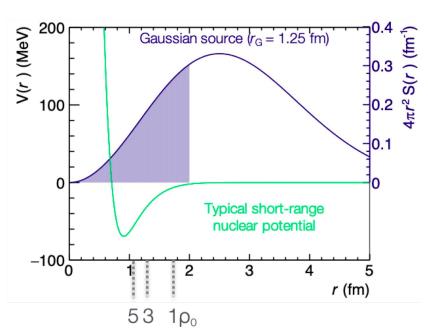


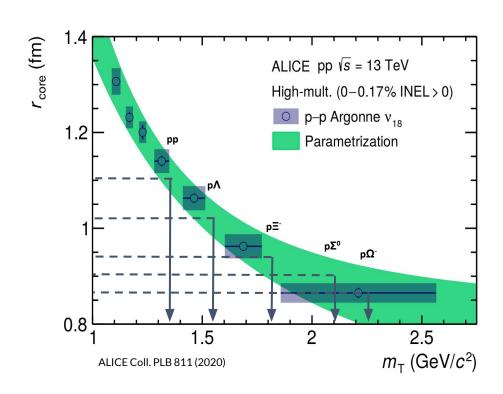
One universal source for all hadrons with strong resonance decays considered for each pair of interest



Source model

Small particle-emitting source created in pp and p-Pb collisions at the LHC.

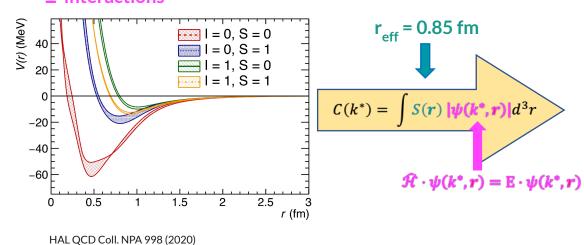




|S|=2 sector: p≡ interaction and first test of LQCD

Lattice QCD potentials from HAL QCD collaboration available

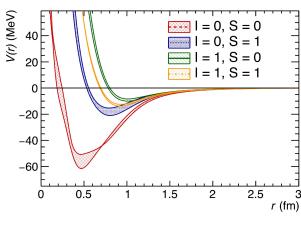
Local potentials for the nucleon-≡ interactions



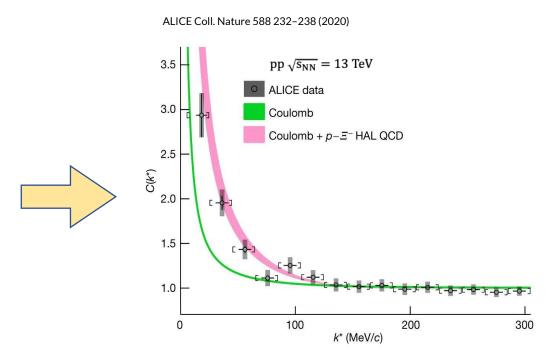
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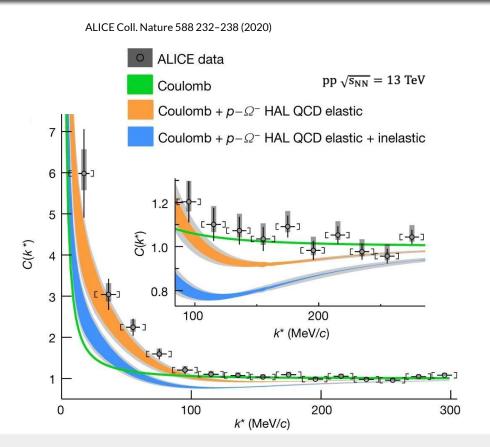


HAL QCD Coll. NPA 998 (2020)



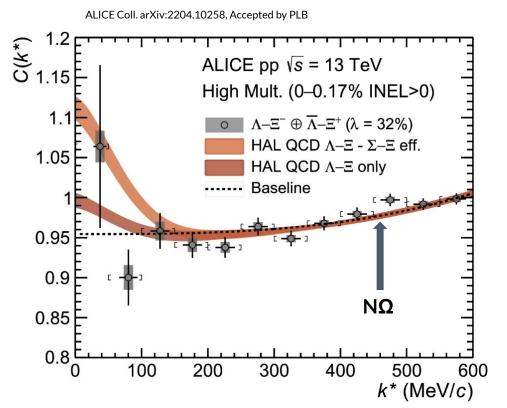
Observation of a strong attractive interaction beyond Coulomb in agreement with lattice predictions

|S|=3: p- Ω^{-} correlation function in pp at 13 TeV



- Enhancement above Coulomb
 - → Observation of the strong interaction
- Attraction in ⁵S₂ results in the prediction of a bound state (Binding Energy = 1.54 MeV)
- Missing potential of the ³S₁ channel
 - \rightarrow Test of two cases:
 - Inelastic channels dominated by absorption
 - Neglecting inelastic channels
- Data more precise than lattice calculations
- So far, no indication of a bound state

$|S|=3: \Lambda-\Xi^-$ interaction with femtoscopy



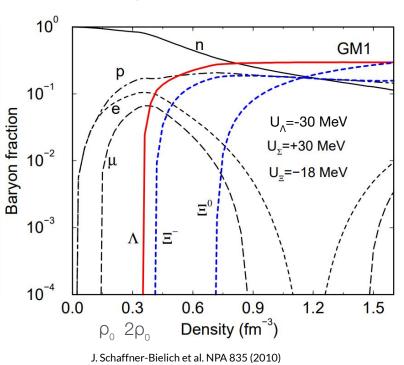
- Unknown contribution from coupled channels in Lattice QCD calculations
 - \rightarrow Coupling $\Lambda\Xi$ - $\Sigma\Xi$ sizable in HAL QCD calculation
 - \rightarrow No sensitivity yet

 ("No coupling" 0.64 n σ VS "Coupling" 1.43 n σ)
- No $N\Omega$ cusp visible
 - \rightarrow Hint to negligible N Ω - $\Lambda\Xi$ coupling

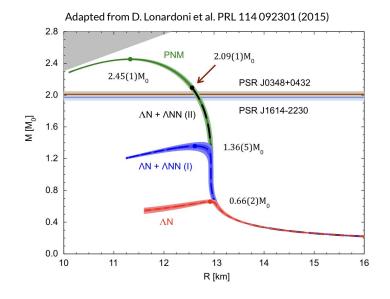
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Hyperon appearance in neutron stars?



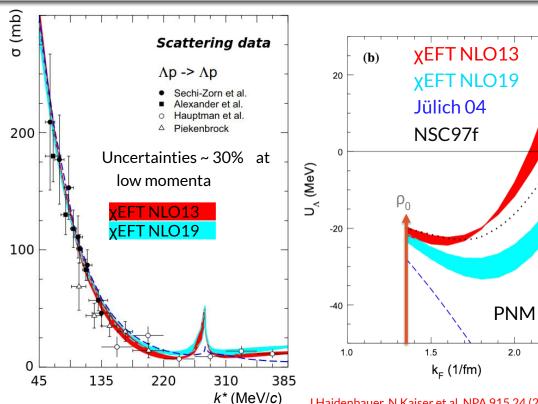


- Hyperons might appear in neutron stars since it is energetically favourable
- But the resulting equation of state might be too soft to explain heavy neutron stars



$|S| = 1: \Lambda$ -p interaction

- Low statistics and not available at low momenta
- ΛN - ΣN coupled system \rightarrow two-body coupling to ΣN is not (yet) measured
- ΣN coupling strength relevant for EoS
 - Strongly affects the behaviour of Λ at finite density
 - Implications for ΛNN interactions
- NLO19 predicts weak coupling NΛ-NΣ
 - Attractive ∧ interaction in neutron matter

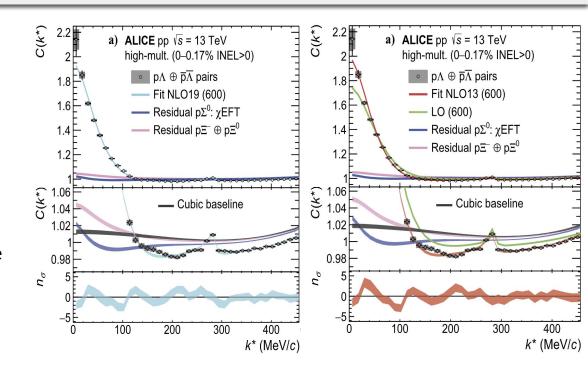


J.Haidenbauer, N.Kaiser et al. NPA 915 24 (2013)

J.Haidenbauer, U. Mei β ner EPJA 56 (2020)

$|S| = 1: \Lambda$ -p interaction

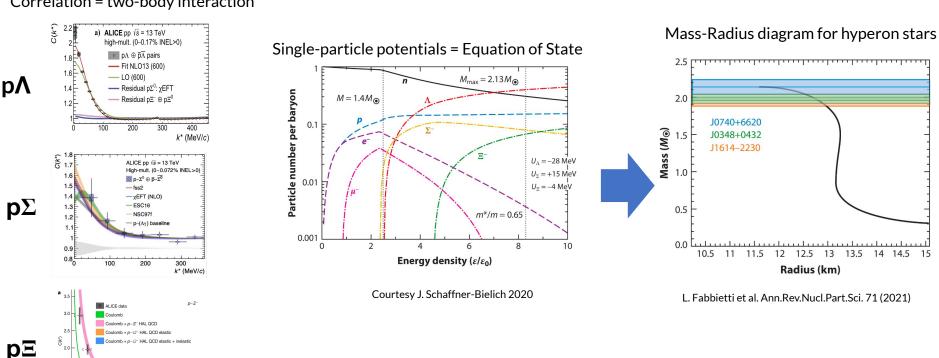
- Comparison with χEFT potentials
 - Sensitivity to different ΣN coupling strength
 - NLO19 favoured (n_σ = 3.7) → attractive interaction of Λ at large densities
 - Larger ΛNN repulsion required to stiffen the Equation of State at large densities



ALICE Coll. PLB 833 137272 (2022)

An example of Equation of State for neutron stars

Correlation = two-body interaction

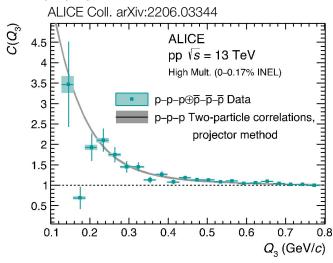


What about the three-body strong interaction?

p-p-p and p-p-Λ correlation functions

$$\begin{split} C(\mathbf{p}_1, \mathbf{p}_2, \mathbf{p}_3) &\equiv \frac{P(\mathbf{p}_1, \mathbf{p}_2, \mathbf{p}_3)}{P(\mathbf{p}_1) \cdot P(\mathbf{p}_2) \cdot P(\mathbf{p}_3)} = \\ &= \mathcal{N} \cdot \frac{N_{\text{same}}(Q_3)}{N_{\text{mixed}}(Q_3)} \end{split}$$

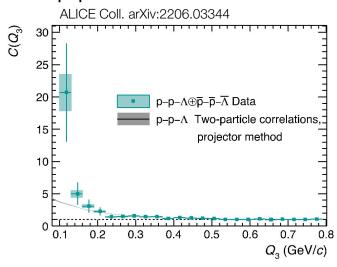
p-p-p correlation function



$$Q_{3} = \sqrt{-q_{12}^{2} - q_{23}^{2} - q_{31}^{2}}$$

$$q_{ij}^{\mu} = (p_{i} - p_{j})^{\mu} - \frac{(p_{i} - p_{j}) \cdot P_{ij}}{P_{ij}^{2}} P_{ij}^{\mu} \qquad P_{ij} \equiv p_{i} + p_{j}$$

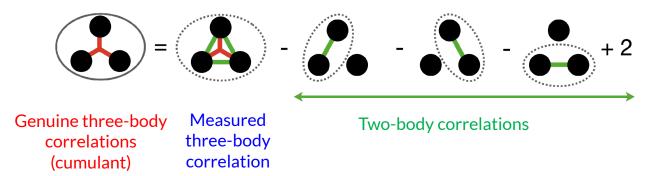
p-p- Λ correlation function



Cumulants

Genuine three-particle correlations isolated using the Kubo's cumulant expansion method:

R. Kubo, J. Phys. Soc. Jpn. 177 (1962)



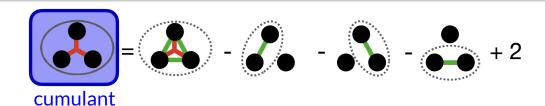
In terms of correlation functions:

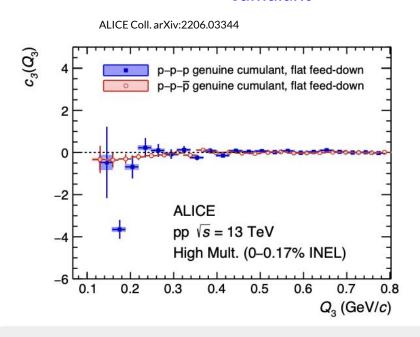
$$c_3(Q_3) = C(Q_3) - C_{12}(Q_3) - C_{23}(Q_3) - C_{31}(Q_3) + 2$$

Projector method

R.D.G. et al. EPJC 82 (2022) 244

p-p-p cumulant





Statistical significance

$$\rightarrow$$
 n_{σ} = 6.7 for Q_3 < 0.4 GeV/c

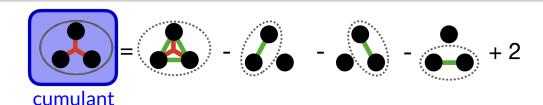
Conclusion

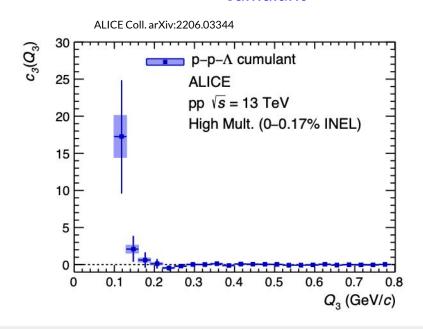
 \rightarrow Evidence of a genuine three-body effect in the p-p-p system at the LHC

Possible interpretations

- → Pauli blocking at the three-particle level
- → Long-range Coulomb interaction effects
- → Three-body strong interaction
- Test with mixed charge particles, cumulant negligible

p-p-∧ cumulant





Statistical significance

$$\rightarrow$$
 n_{σ} = 0.8 for Q_3 < 0.4 GeV/ c

Conclusion

→ No significant deviation from the null hypothesis

A factor 500 in statistics from the Run 3 data taking

- \rightarrow Non-zero cumulant can be directly linked to the three-body strong interaction
- → Important measurement for neutron stars

Summary

