

Analysis Tools for CLAS12

- Application of RooFit

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CLAS Joint Working Groups
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Goals

General

parameter estimation for all reaction cases → **Yields**

→ allow customised PDF classes

→ use of strings to define PDF

Polarisation Observables

Spin Density Matrix

Angular Moments

Amplitudes...

User friendly

→ Use ROOT Jupyter notebooks and/or PyROOT

→ Minimise amount of user code

Fast Efficient calculation

→ Cache data and integrals

→ parallelise with via RooFit multicore Likelihood splitting

→ parallelise bins with ROOT PROOF multicore

→ parallelise on farm

Reliable results

Systematic uncertainties in extraction procedure?

→ ToyMC fits

RooFit and RooStats

arXiv.org > physics > arXiv:physics/0306116

Search.

Help

Physics > Data Analysis, Statistics and Probability

The RooFit toolkit for data modeling

Wouter Verkerke, David Kirkby

(Submitted on 14 Jun 2003)

RooFit is a library of C++ classes that facilitate data modeling in the ROOT environment. Mathematical concepts such as variables, (probability density) functions and integrals are represented as C++ objects. The package provides a flexible framework for building complex fit models through classes that mimic math operators, and is straightforward to extend. For all

Used in many analysis in HEP

RooFit ~830 citations

RooStats ~670 citations

Optimised maximum likelihood
fitter + lots of features

arXiv.org > physics > arXiv:1009.1003

Search...

Help | Advan

Physics > Data Analysis, Statistics and Probability

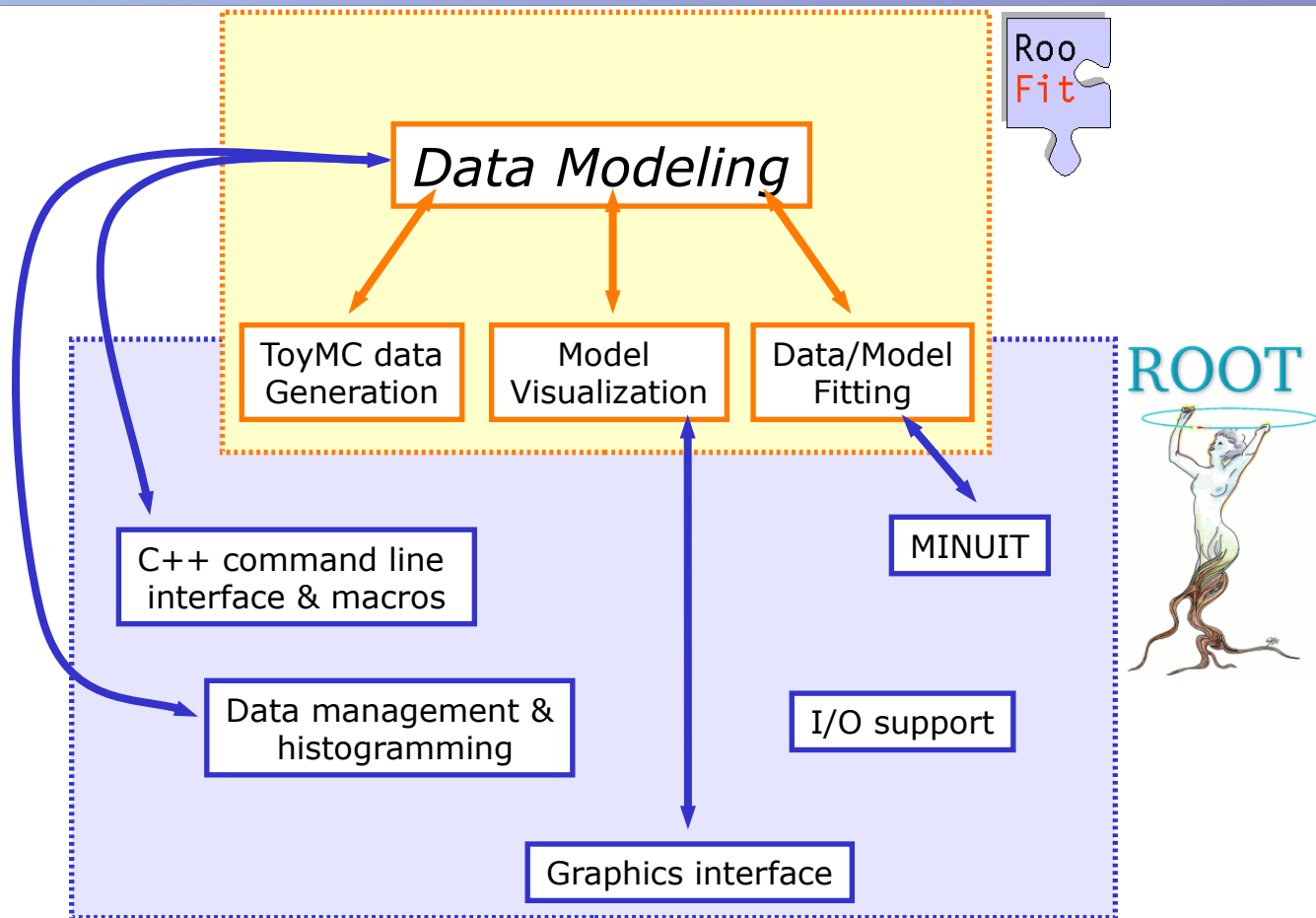
The RooStats Project

Lorenzo Moneta, Kevin Belasco, Kyle Cranmer, Sven Kreiss, Alfio Lazzaro, Danilo Piparo, Gregory Schott, Wouter Verkerke, Matthias Wolf

(Submitted on 6 Sep 2010 (v1), last revised 1 Feb 2011 (this version, v2))

RooStats is a project to create advanced statistical tools required for the analysis of LHC data, with emphasis on discoveries, confidence intervals, and combined measurements. The idea is to provide the major statistical techniques as a set of C++ classes with coherent interfaces, so that can be used on arbitrary model and datasets in a common way. The classes are built on top of the RooFit

Roofit



RooFit

- Mathematical objects are represented as C++ objects

Mathematical concept			RooFit class
variable	x	→	<code>RooRealVar</code>
function	$f(x)$	→	<code>RooAbsReal</code>
PDF	$f(x)$	→	<code>RooAbsPdf</code>
space point	$x_{\max}$	→	<code>RooArgSet</code>
integral	$\int_{x_{\min}} f(x) dx$	→	<code>RooRealIntegral</code>
list of space points		→	<code>RooAbsData</code>
PDF summation	$f_1(x) + f_2(x)$	→	<code>RooAddPdf</code>
PDF product	$f_1(x) \cdot f_2(x)$	→	<code>RooProdPdf</code>

Model building – (Re)using standard components

- List of most frequently used pdfs and their factory spec

Gaussian `Gaussian::g(x,mean,sigma)`

Breit-Wigner `BreitWigner::bw(x,mean,gamma)`

Landau `Landau::l(x,mean,sigma)`

Exponential `Exponential::e(x,alpha)`

Polynomial `Polynomial::p(x,{a0,a1,a2})`

Chebyshev `Chebyshev::p(x,{a0,a1,a2})`

Kernel Estimation `KeysPdf::k(x,dataSet)`

Poisson `Poisson::p(x,mu)`

Voigtian `Voigtian::v(x,mean,gamma,sigma)`

(=BW $\otimes$ G)

Application of RooFit

Is it useful for extracting parameters from fits to angular distributions of real experimental data ?

Arbitrary PDF

- Skeleton Pdf code generator
- Define evaluate() function
- Use TFormula string→function→pdf
- Analytic or Numerical integrator

Signal / background

- simultaneous fit to both
$$\mathbf{S}(\mathbf{x}) * \mathbf{f}_s(\mathbf{y}, \theta_s) + \mathbf{B}(\mathbf{x}) * \mathbf{f}_b(\mathbf{y}, \theta_b(\mathbf{x}))$$
- sPlot + weighted likelihood fits

Acceptance correction → **No, but could use weights**

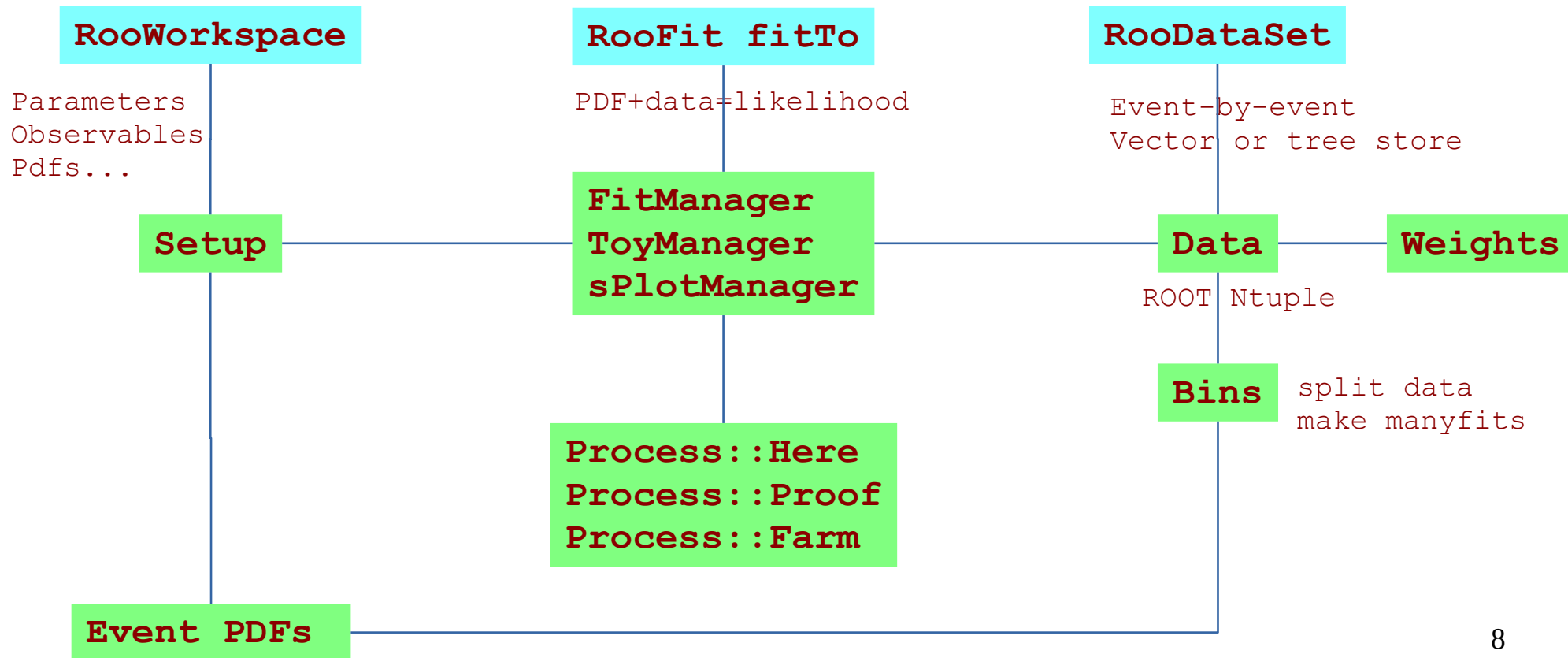
Systematic uncertainties → ToyMC studies

Many fits to different bins → Not trivially

Software Structure

RooFit

New



Maximum Likelihood with acceptance

$$L(p) = \prod_i \frac{f(\tau_i : p)}{\int f(\tau_i : p) d\tau}$$

Likelihood = product Prob.
function f with observables τ and pars p

$$L_{acc}(p) = \prod_i \frac{f(\tau_i : p) \eta(\tau_i)}{\int f(\tau_i : p) \eta(\tau_i) d\tau}$$

With acceptance need product
of f and acceptance function

$$A(p) = \int f(\tau_i : p) \eta(\tau_i) d\tau$$

Probability Normalisation $\int$

$$L_{acc}^{ext}(p) = \left[\frac{A(p)^N}{N!} e^{-A(p)} \right] \prod_i \frac{f(\tau_i : p) \eta(\tau_i)}{A(p)}$$

Extended ML with Poisson statistics

$$-\ln L_{acc}^{ext}(p) \propto -\sum_i \ln f(\tau_i : p) \eta(\tau_i) + A(p)$$

Minimise $-\ln(L)$
Drop terms independent of p

$$= -\sum_i \ln f(\tau_i : p) - \sum_i \ln \eta(\tau_i) + A(p)$$

Drop acceptance in sum over events

$$\propto -\sum_i \ln f(\tau_i : p) + A(p)$$

Need to minimise this

$$A(p) \simeq \sum_j^M f(\tau_j : p)$$

Approximating $A(p)$ as sum of
 f over M accepted Monte-Carlo events

Simulated MC integrals for RooFit : RooHSEventsPDF

RooFit PDF classes require normalisation

- This can be done by users own method
- Or RooFit performs its own numerical integration(vegas)

RooHSEventsPDF calculates its own integral by summing over MC events - includes partial integration for plotting

Requires ROOT tree of reconstructed MC events with fit variables

Can give weights to MC events (e.g. better match non-fitted data and MC distributions, importance sampling)

Additionally can generate events using true values if these are passed through in the tree

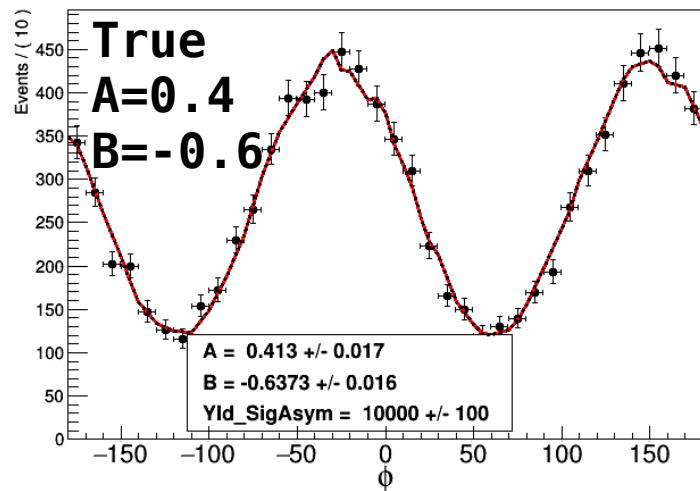
- Evaluate function with generated truth, produce events with simulated reconstructed

Works any number of fit observables or parameters

- i.e. given a Ndim Model it will fit for Mpar parameters accounting for detector acceptance effects

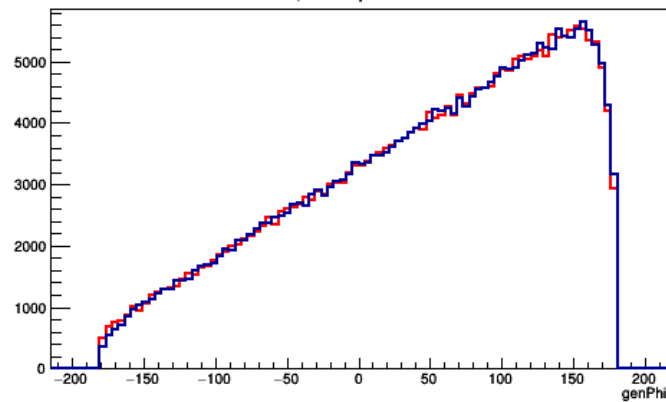
Fits with acceptance

Fit components for Phi

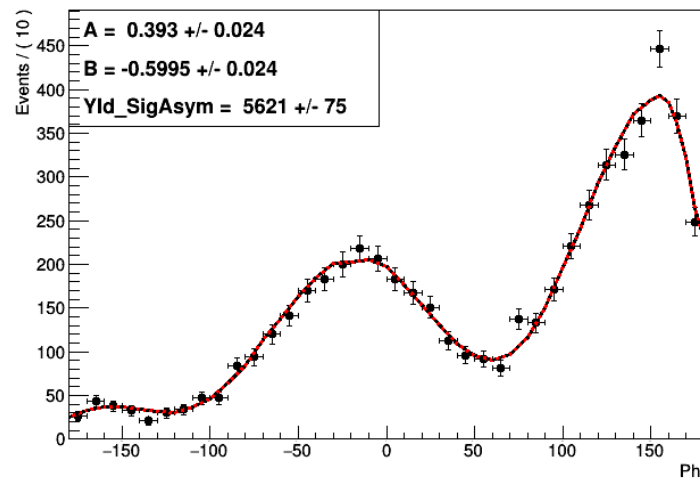


+

ϕ Acceptance



Fit components for Phi



=

100% acceptance

PDF : $1 + A\cos(2\phi) + B\sin(2\phi)$

Plotting of fit function
integrates over MC events

Example Fit in Jupyter

```
In [ ]: 1 FitManager RF;  
2 RF.SetUp().SetOutDir("outObs/");
```

Set the fit observables in the dataset

```
In [ ]: 1 RF.SetUp().LoadVariable("Phi[-180,180]");  
2 RF.SetUp().LoadVariable("Pol[0,1]");  
3 RF.SetUp().LoadCategory("PolState[Polp=1,Polm=-1]");
```

If I want to use any other variable, or example to apply a cut, load it as an AuxVar.

```
In [ ]: 1 RF.SetUp().LoadAuxVar("M1[0,10]"); //Load Aux Var, limits used as cut  
2 RF.SetUp().AddCut("M1>2"); //Additional cut based on vars or aux vars
```

Create and load into the fit manager my PDF class.

```
In [ ]: 1 Loader::Compile("PhiAsymmetry.cxx");  
2 RF.SetUp().FactoryPDF("PhiAsymmetry::SigAsym"  
3 " ( Phi,Pol,PolState,A[0,-1,1],B[0,-1,1] )");  
4 RF.SetUp().LoadSpeciesPDF("SigAsym",1);
```

Split the data into 4 energy bins to perform separate fits on.

```
In [ ]: 1 RF.Bins().LoadBinVar("Eg",4,3,4);
```

Load fit data and simulated MC data for normalisation integral

```
In [ ]: 1 RF.LoadData("MyModel",pwd+"Data.root");  
2 RF.LoadSimulated("MyModel",pwd+"MC.root","SigAsym");
```

Now I attach the weights from my sPlot fit. I want to use the signal weights which were given the name "Signal" in the sPlot notebook.

```
In [ ]: 1 RF.Data().LoadWeights("Signal","sPlotFit/Tweights.root");
```

Now run the fits. I use the Process classes which allow me to choose between running directly here on a single core or multicore via PROOF-lite. It doesn't make sense to run with PROOF unless multiple splits have been defined with LoadBinVar or you are using Bootstrap, in which case you should relate the number of cores requested to the number of splits or bootstraps.

```
In [ ]: 1 Here::Go(&RF);  
2 //OR run with PROOF-LITE on N=4 cores (you can change the 4)  
3 Proof::Go(&RF,4);
```

Output location

Fit Observables

Cuts to other variables

Define PDF (here load a class)

$1 + A \cdot \text{PolState} \cdot \text{Pol} \cdot \cos(2\phi) + B \cdot \text{PolState} \cdot \text{Pol} \cdot \sin(2\phi)$

Split data in bins of E_g

Set Fit and MC Integral data

Set Event weights

Run fit here or on PROOF¹²

sPlot

M. Pivk, F.R. Le Diberder, Nucl.Inst.Meth.A 555, 356-369, 2005

Given discriminatory PDF for signal and background calculates weight :

$${}_s\mathcal{P}_n(y_e) = \frac{\sum_{j=1}^{N_s} \mathbf{V}_{nj} f_j(y_e)}{\sum_{k=1}^{N_s} N_k f_k(y_e)}$$

N_s = Number of species

f_k = PDF for species k

N_k = Yield for species k

$\mathbf{V}$ = covariance matrix

Part of RooStats(used here)

Can include multiple signal
and background species

Can use directly in likelihood fits

In this package running sPlot, same as running any fit

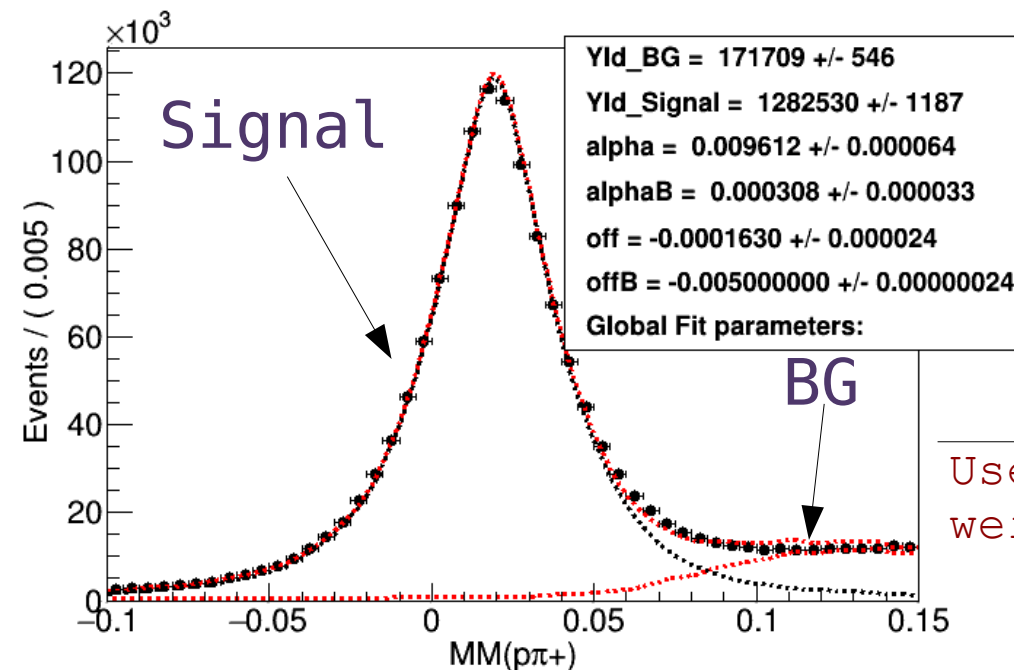
But only as good as fit model...

And Signal OR Background observable distributions cannot
correlate with discriminatory variables

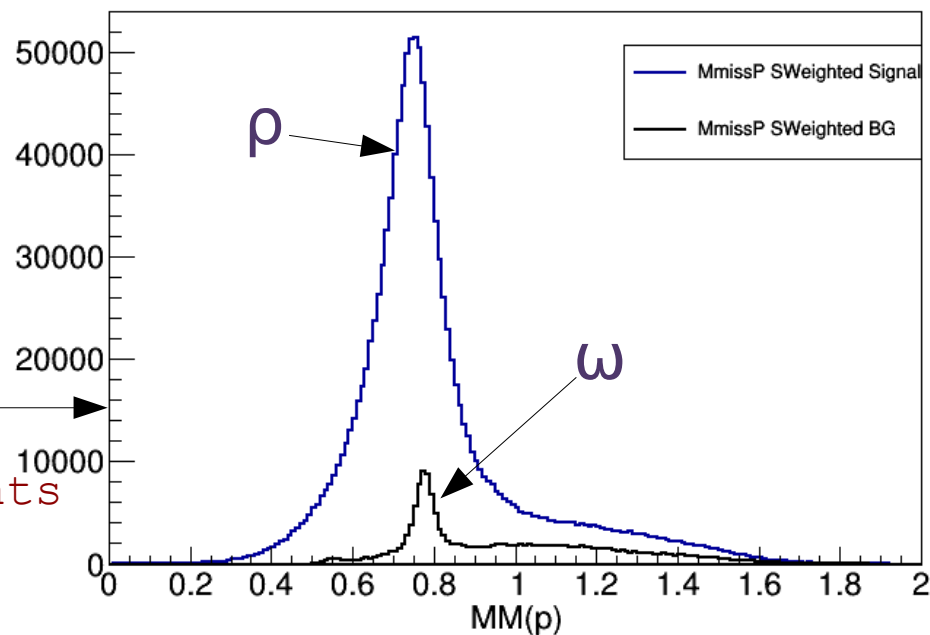
Example fit and disentangled distributions

$\gamma p \rightarrow p \pi^+ (\pi^-)$

Models from simulated $\pi^+ \pi^- p$ and $\pi^+ \pi^- \pi^0 p$ events



Use weights

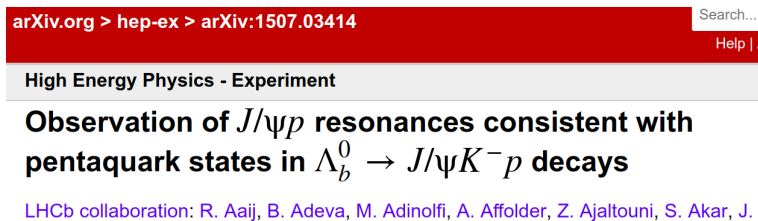


Note 2 fits required. First fixes alpha and off.
Second, only Yields free => Covariance matrix

Maximum likelihood with weights

Many analysis include additional factor to better approximate uncertainties

e.g



$$-2\ln(L(p)) = -2\alpha \left[\sum_i W_i \ln(f(\tau_i, p)) + A(p) \right]$$


Weights subtract off background contribution to likelihood

$$\alpha = \frac{\sum_i W_i}{\sum_i W_i^2}$$

α term reduces gradient
Therefore increase uncertainty

RooFit with weights

$$-2\ln(L(p)) = -2\left[\sum_i W_i \ln(f(\tau, p)) + A(p)\right] \quad (1) \quad \text{Minimise this}$$

$$-2\ln(L(p)) = -2\left[\sum_i W_i^2 \ln(f(\tau, p)) + A(p)\right] \quad (2) \quad \text{Also calculate this}$$

$$G_{\mu\nu} = \sum_{i=1}^N w_i \frac{dp_i}{d\alpha_\mu} \frac{dp_i}{d\alpha_\nu} \quad \text{covariance matrix for (1)}$$

$$F_{\mu\nu} = \sum_{i=1}^N (w_i)^2 \frac{dp_i}{d\alpha_\mu} \frac{dp_i}{d\alpha_\nu} \quad \text{covariance matrix for (2)}$$

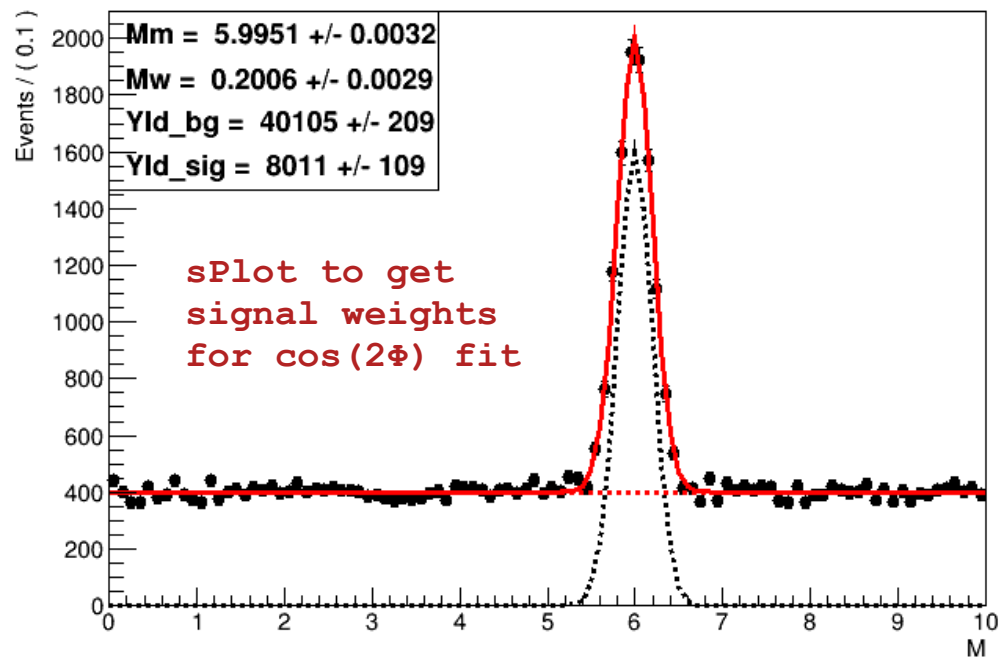
$$\text{COV}(\vec{\alpha}, \vec{\alpha}) = G^{-1} F G^{-1}$$

$$\text{errors scale with } \frac{\sum_i W_i^2}{\sum_i W_i}$$

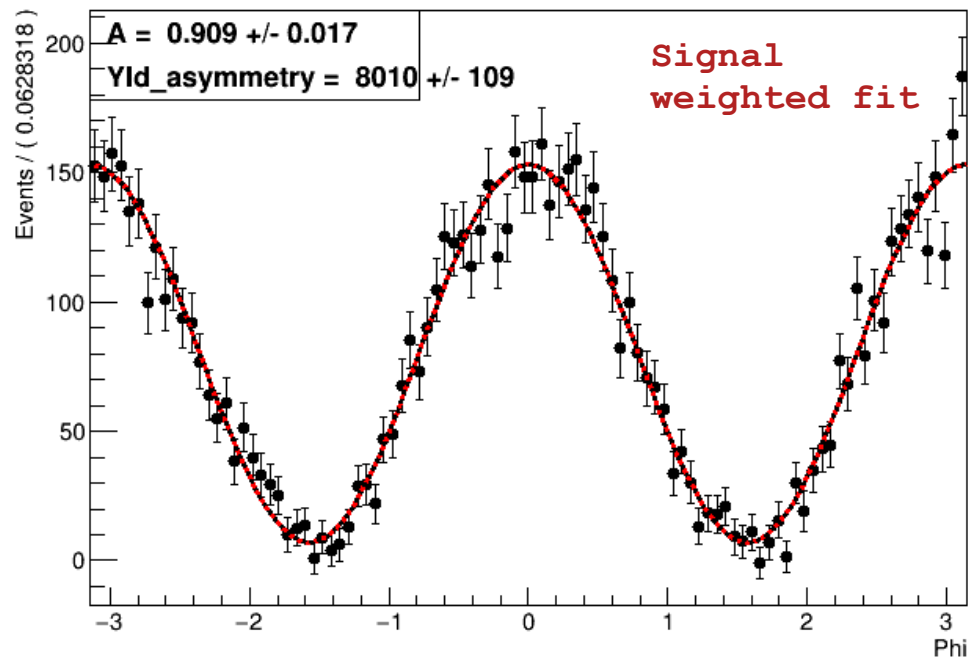
Weighted Fit uncertainties validation

Toy Fits : Signal 8k $f(M) = \text{Gaussian}(6, 0.2)$ $g(\Phi) = 1 + 0.92\cos(2\Phi)$
 Backgr 40k $f(M) = \text{Uniform}$ $g(\Phi) = \text{Uniform}$

Fit components for M



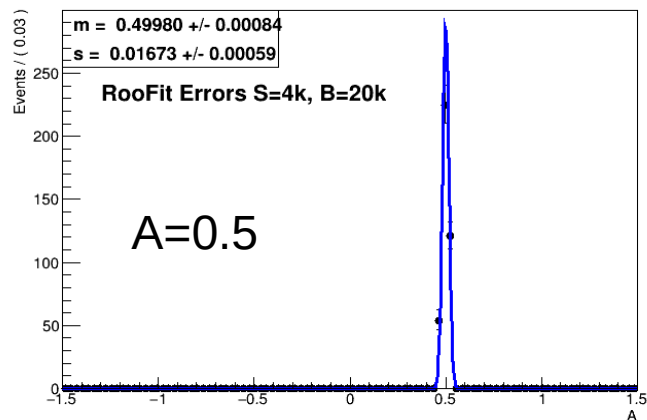
Fit components for Phi



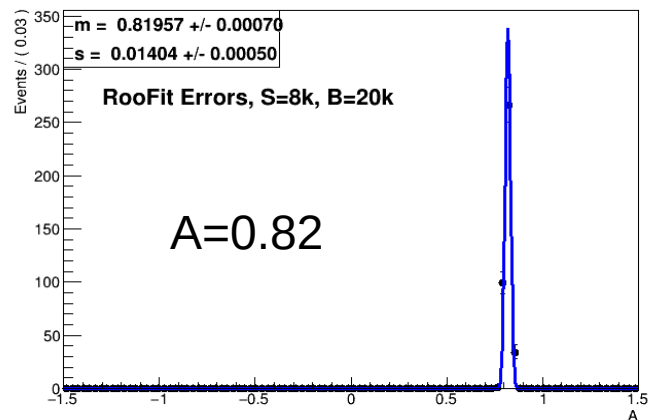
Weighted Fit uncertainties validation

RooFit errors, 400 fits, change A

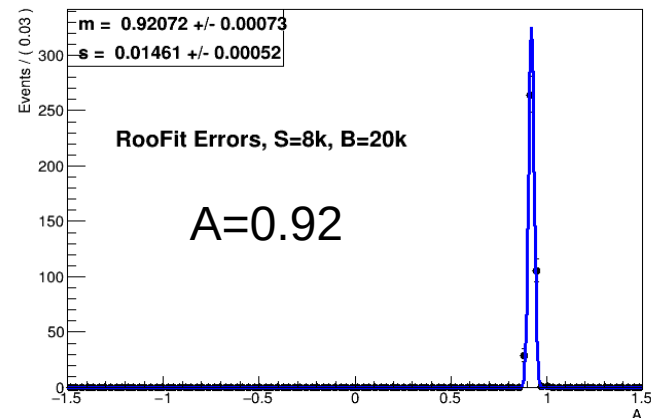
A RooPlot of "A"



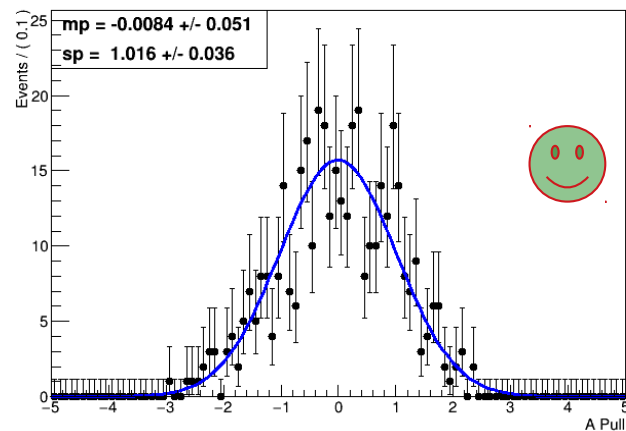
A RooPlot of "A"



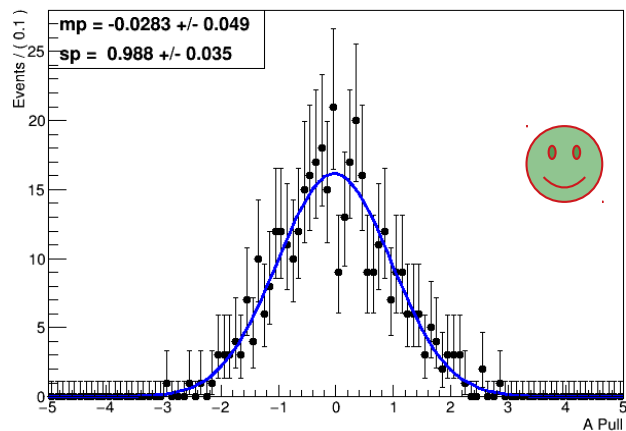
A RooPlot of "A"



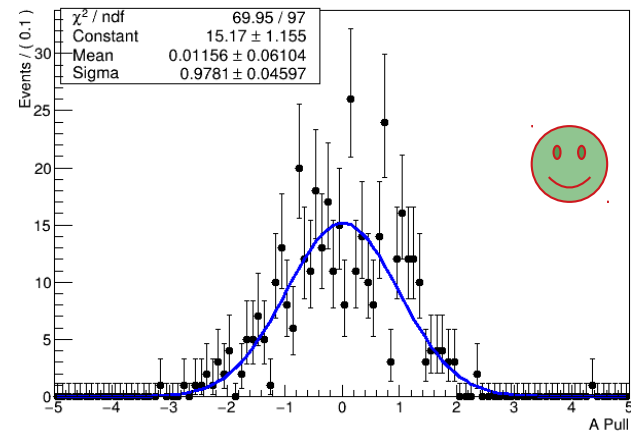
A RooPlot of "A Pull"



A RooPlot of "A Pull"



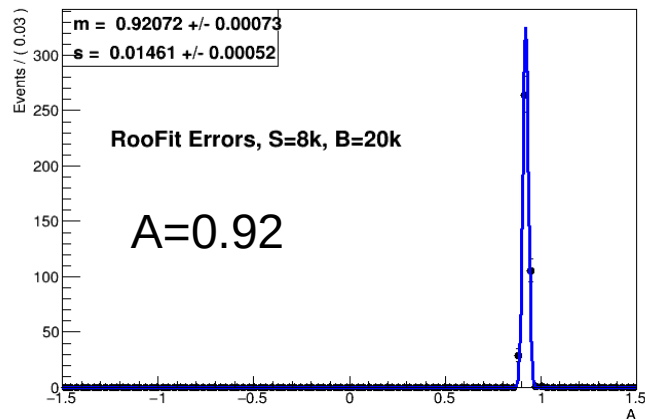
A RooPlot of "A Pull"



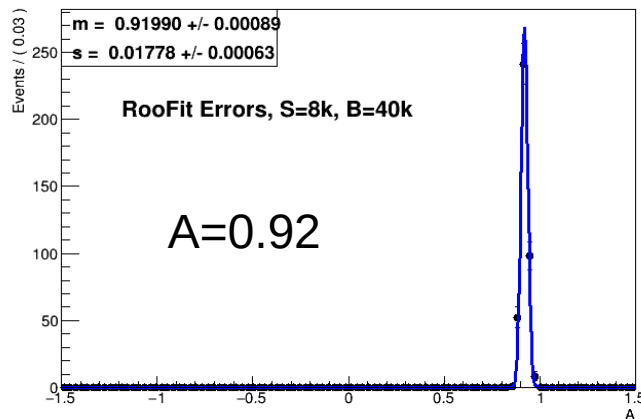
Weighted Fit uncertainties validation

RooFit errors, 400 fits, change BG

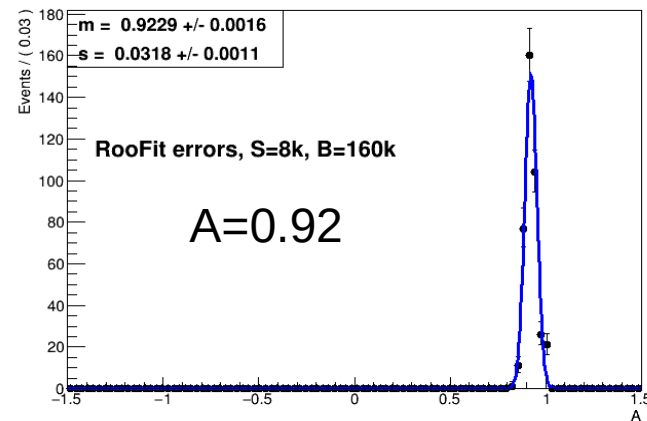
A RooPlot of "A"



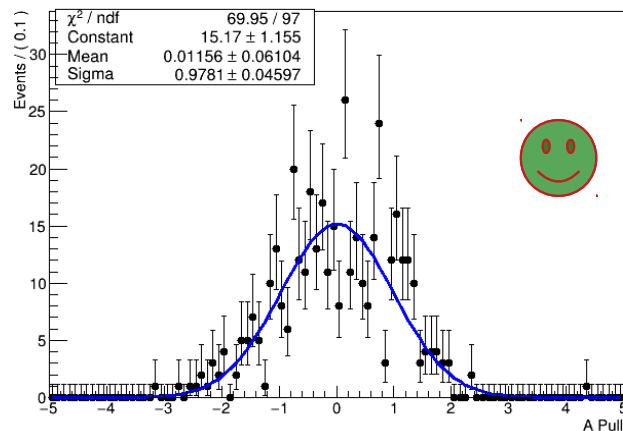
A RooPlot of "A"



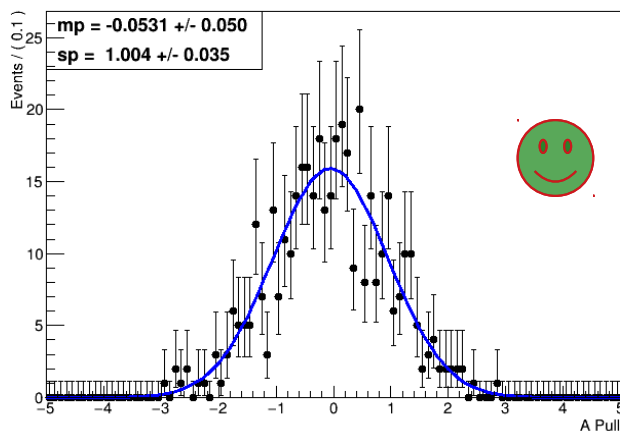
A RooPlot of "A"



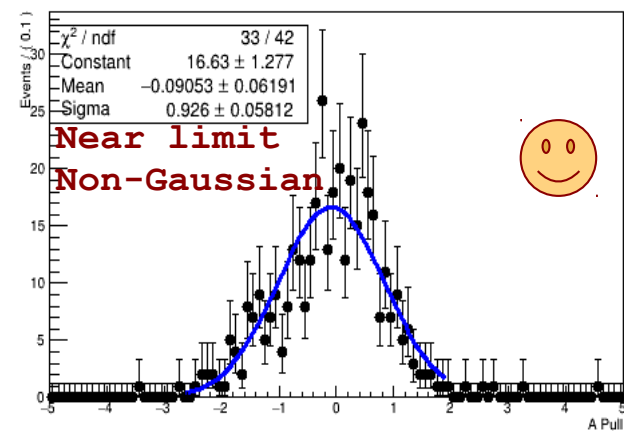
A RooPlot of "A Pull"



A RooPlot of "A Pull"



A RooPlot of "A Pull"



Uncertainties with acceptance

Uncertainties from simulated MC integral are not propagated to Minuit covariance matrix

Normally try to have $\times 10$ MC events $\sim 1/\sqrt{10} \sim 0.3$ additional factor on statistical uncertainty

Can be estimated via bootstrapping the MC events samples

For asymmetries MC statistical uncertainty can be removed by calculating

$$\sum_{i=0}^{N_{acc}} f(\tau, p, P=+1) + \sum_{i=0}^{N_{acc}} f(\tau, p, P=-1)$$

In the limit $N(+)=N(-)=N_{acc}/2$

As this will be constant, rather than

$$\sum_{i=0}^{N_{acc}/2} f(\tau, p, P=+1) + \sum_{i=N_{acc}/2}^{N_{acc}} f(\tau, p, P=-1)$$

With small corrections in the case of asymmetries in luminosity $N(+)\neq N(-)$

Measuring Asymmetries

C. Mullen, Glasgow → Genoa

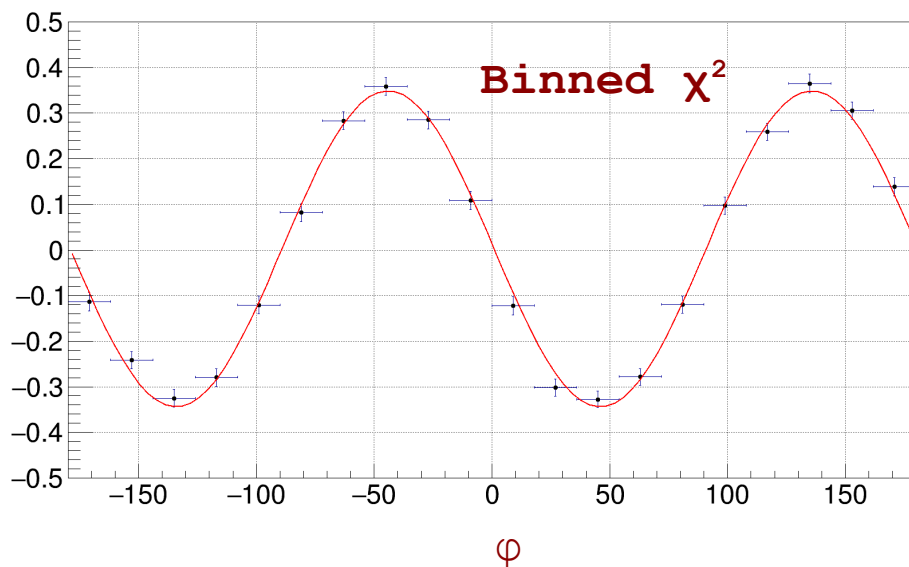
Example CB@MAMI $n\pi^0$ photoproduction beam asymmetry

$$\text{PDF : } f(\phi, P_y, P=\pm 1) = (1 + AP_y P \cos(2\phi))$$

Using histograms

$$\Sigma P_y^{mean} \cos(2\phi) = \frac{N(\phi, P=+1) - N(\phi, P=-1)}{N(\phi, P=+1) + N(\phi, P=-1)}$$

Do not require acceptance correction
(to first order)



Measuring Asymmetries with event based Likelihood

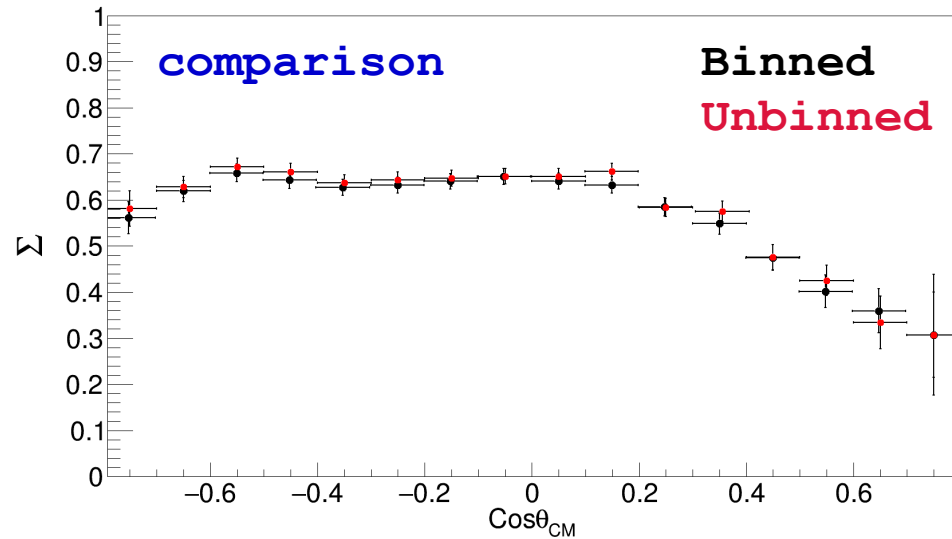
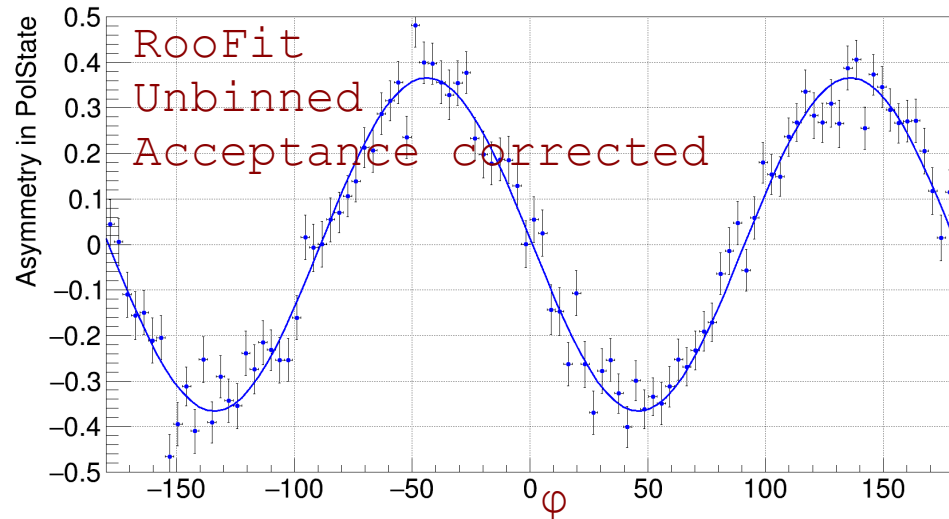
Normalisation Integral

$$\sum_{i=0}^{N_{acc}} f(\tau_i, p, P=+1) + \sum_{i=0}^{N_{acc}} f(\tau_i, p, P=-1) = \sum_{i=0}^{N_{acc}} (1 + AP_{y,i} \cos(2\phi_i)) + \sum_{i=0}^{N_{acc}} (1 - AP_{y,i} \cos(2\phi_i)) = 2 N_{acc}$$

= constant, when $N(P=+1)=N(P=-1)$ and $P_Y(P=+1)=P_Y(P=-1)$
 $\Rightarrow$ does not effect position of likelihood maximum

First order, ignore acceptance correction, do not need to calculate integral

Second order, use acceptance, corrects for experimental polarisation and luminosity asymmetries



Minimisers

It is straightforward to try different minimisers to fit the data

Minuit2 is used by default

To use Minuit :

```
Fitter.SetMinimiser(new RooMinuit());
```

Or an MCMC implementation :

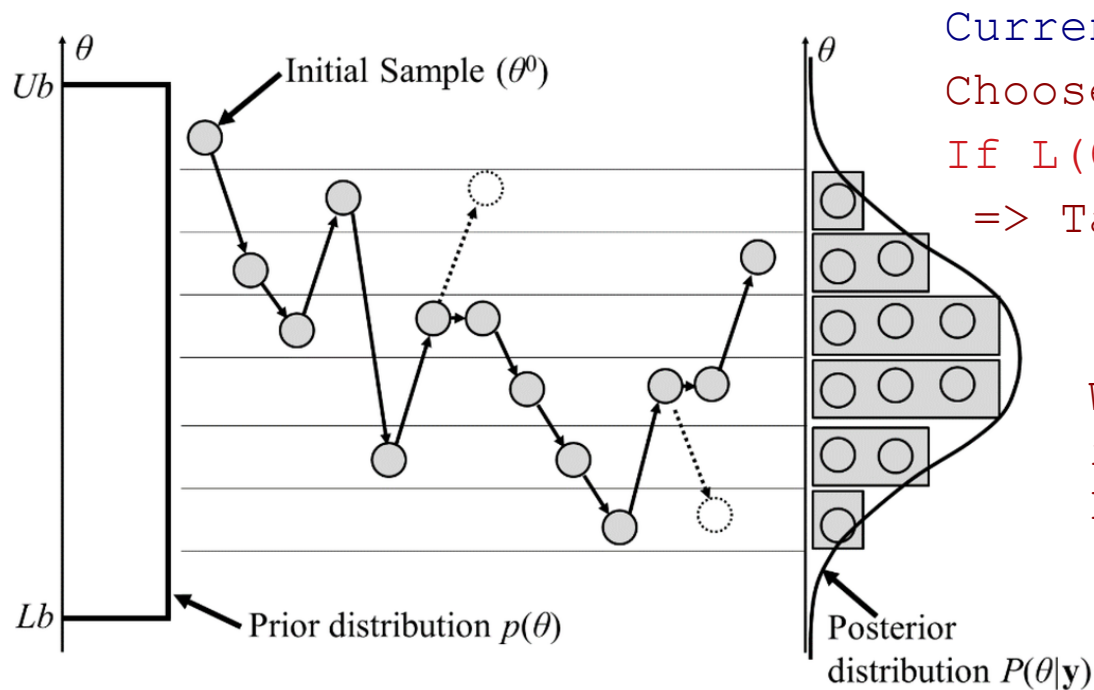
```
Fitter.SetMinimiser(new RooMcmcSeq(30000,10000,1000));
```

Number of samples

Number of burnin

1/StepSize

MCMC with Metropolis-Hastings, for finding parameter values



Current parameters = θ_k

Choose θ' from $q(\theta', \theta_k)$ (proposal func.)

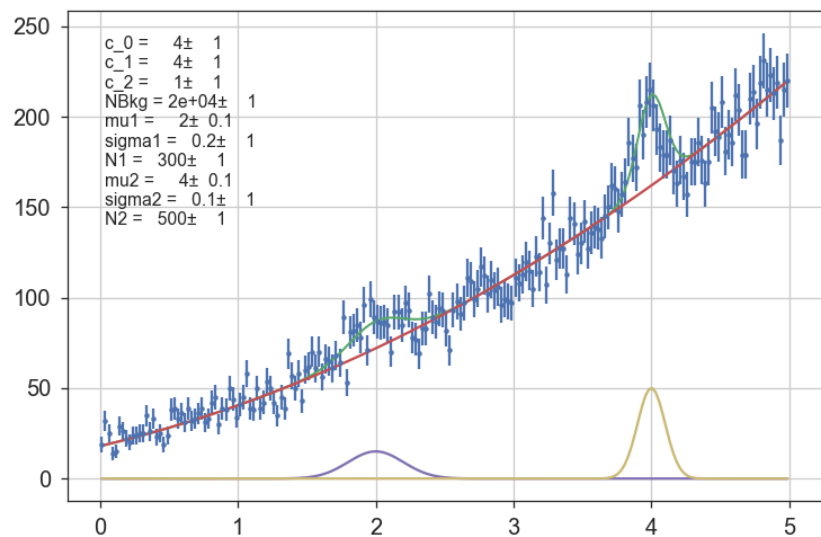
If $L(\theta')/L(\theta_k) > \text{Uniform}(0,1)$

=> Take θ' as next parameters sample

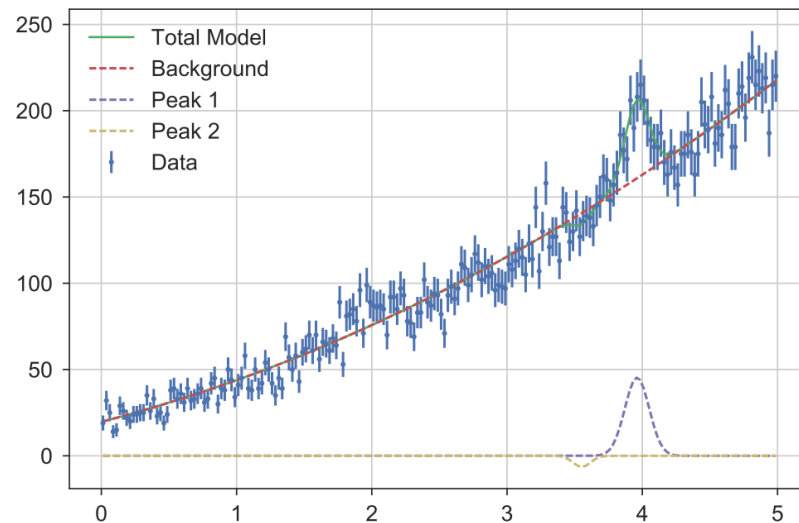
We use sequential proposal
i.e. change 1 parameter to
 $P += \text{RandGaussian}(P, P_{\text{range}} * \text{StepSize})$

MCMC Test Minuit fit

Generated Data with true PDFs

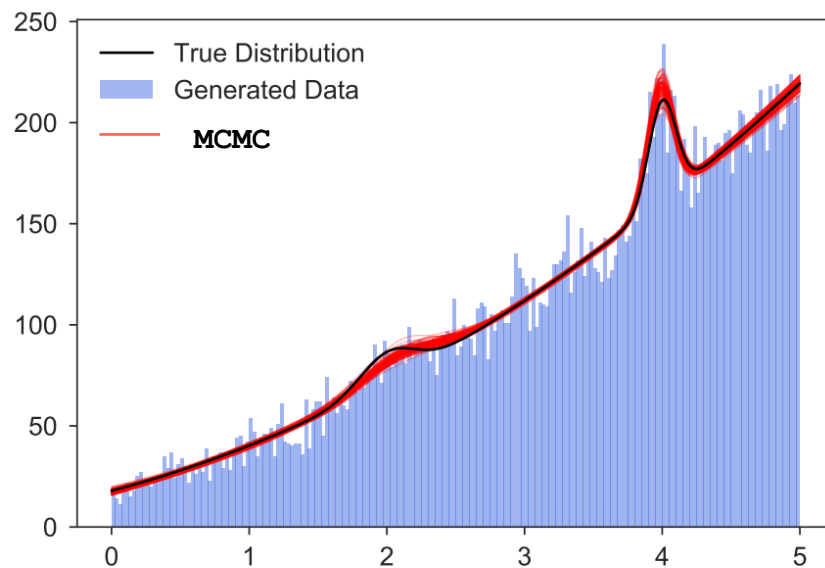
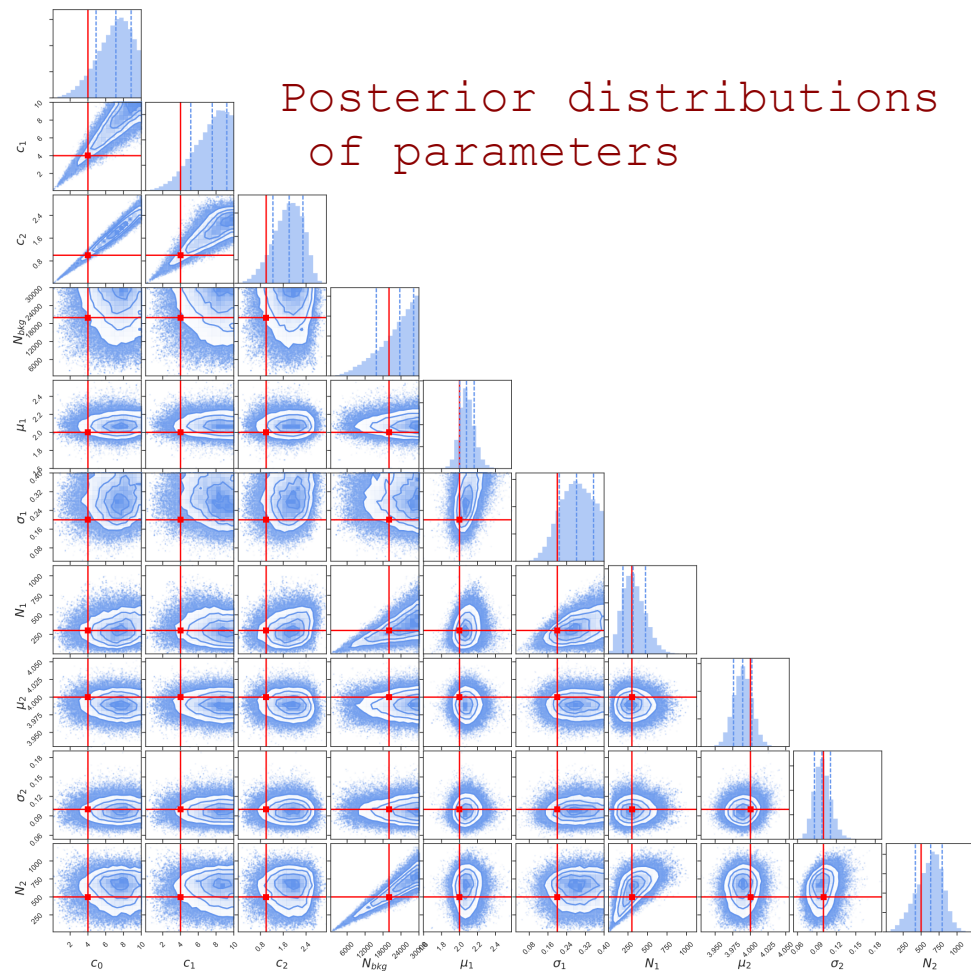


Minuit Fit results
Random initial parameters



MCMC Test fit

Posterior distributions
of parameters



Systematic Studies: ToyMC in Jupyter

 SimpleGenerateToyFit Last Checkpoint: 16 hours ago (autosaved) Logout Terminal

File Edit View Insert Cell Kernel Widgets Help Trusted Python 3



Example event generation, fitting and ToyMC study

Load the ROOT and fitting modules. Turn on javascript ROOT for nice interactive plots

```
In [1]: import ROOT
ROOT.gROOT.ProcessLine(".x $HSCODE/hsfit/LoadFit.C")
%jsroot
Welcome to JupyROOT 6.16/00
```

Construct a Toy manager for generating initial data set. This would be equivalent to your real data. The argument 1 tells it to only create one data set per bin.

```
In [2]: toy = ROOT.ToyManager(1)
```

Give an output directory for storing the "data"

```
In [3]: toy.SetUp().SetOutDir("outSimpleToys/");
toy.SetUp().SetIDBranchName("UID");
```

Declare your fit variable and its range

```
In [4]: toy.SetUp().LoadVariable("Mmiss[0,10]");
```

Declare your PDF to generate from. Here a Signal is Gaussian with mean 6 (with range 4-7) and width 0.2 (with range 0.0001-3).

LoadSpecies adds this PDF to the total PDF, while the 100 is the typical number of events to generate. Actual number will include Poisson statistics fluctuation.

```
In [5]: toy.SetUp().FactoryPDF("Gaussian::Signal( Mmiss, Gmean[6,4,7], Gsigma[0.2,0.0001,3] )");
toy.SetUp().LoadSpeciesPDF("Signal",100);
```

The Background BG, is a Chebychev polynomial, which is going to have twice as many (200) events as the signal contribution.

```
In [6]: toy.SetUp().FactoryPDF("Chebychev::BG(Mmiss,{a0[-0.1,-1,1],a1[0.1,-1,1]}");
toy.SetUp().LoadSpeciesPDF("BG",200);
```

Generate the data!

```
In [8]: ROOT.Here.Go(toy);
```

Fitting the generated data

A ToyManager that has been used to generate data can be directly used to create a fitter with the same model. Alternately you can define a completely new fit model here.

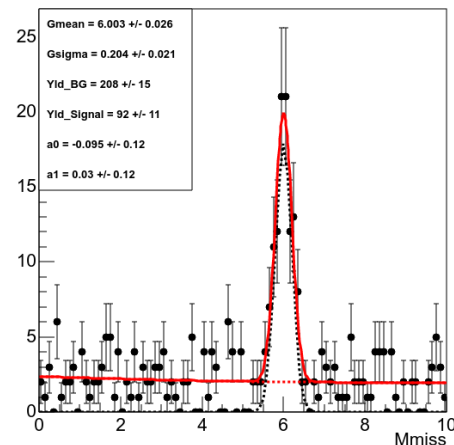
```
In [9]: fit0=toy.Fitter();
DataEvents::Load ToyData 1
```

Fit the data on one local CPU.

We should see the minimiser output with the final fit plots at the end.

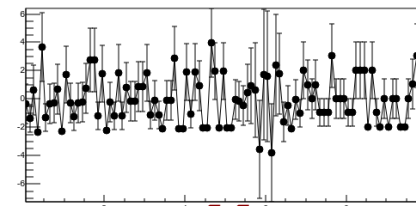
```
In [10]: ROOT.Here.Go(fit0);
```

Fit components for Mmiss



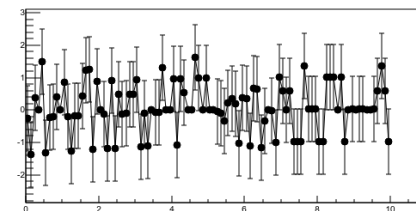
FIT residuals

Residual Histogram of DataEvents_plot_Mmiss and Projection of total model



FIT pulls

Pull of Histogram of DataEvents_plot_Mmiss and Projection of total model



Example : ToyMC study in Jupyter

Toy MC study

Now I have successful fit results I want to study the fit for bias etc. I can do this by generating many data sets from my fit results and fitting them to make sure the extracted parameters are consistent each time.

The ToyMC model with the fit results found in the previous cell can be combined into a ToyManager with 1 line of code using the fit model (fit0 here), which also specifies the number of toy datasets (400) to generate. We should then set a new output directory and generate the toy datasets.

```
In [11]: toy2=ROOT.ToyManager.GetFromFit(400,fit0,"ResultsToy0H5Minuit2.root")
toy2.SetUp().SetOutDir("outSimpleToy2");
ROOT.Here.Go(toy2.get());
```

There are going to be 400 toy fits so lets run them in parallel with PROOF. The next cell is need to initialise PROOF.

```
In [12]: from ROOT import TProof
```

Get my fitter from the new ToyManager and run the fits on PROOF with 4 workers.

```
In [ ]: fit2=toy2.Fitter()
ROOT.Proof.Go(fit2,4)
```

Collect all the fits and create parameter distributions and pulls. This is automated by the ToyManager Summarise function.

```
In [14]: toy2.Summarise()
```

Summarise ResultTree /work/Dropbox/Haspect/dev/HASPECT6/tutorials/RooFitExamples/Generators/outSimpleToy2/

```
ToyManager::Summarise() Initial Parameters
1) 0x56492f5c9900 RooRealVar:: Gmean = 5.96906 L(4 - 7) "Gmean"
2) 0x56492f5d2120 RooRealVar:: Gsigma = 0.206953 L(0.0001 - 3) "Gsigma"
3) 0x56492f5950d0 RooRealVar:: a0 = -0.0433 L(-1 - 1) "a0"
4) 0x56492f52e730 RooRealVar:: a1 = 0.0877073 L(-1 - 1) "a1"
5) 0x56492f5ca1a0 RooRealVar:: Yld_Signal = 109.655 L(0 - 1e+12) "Yld_Signal"
6) 0x56492f5ca9a0 RooRealVar:: Yld_BG = 171.461 L(0 - 1e+12) "Yld_BG"
```

```
Gmean 5.96715 +- 0.0237717 sigma 0.0241361 meanPull 0.00223759 sigmaPull 1.02776
bias -0.00190544 bias Pull -0.0790702 sigma 1.02767
```

```
Gsigma 0.207819 +- 0.0204853 sigma 0.0225982 meanPull -0.115595 sigmaPull 1.1047
bias 0.00086601 bias Pull -0.0722243 sigma 1.10013
```

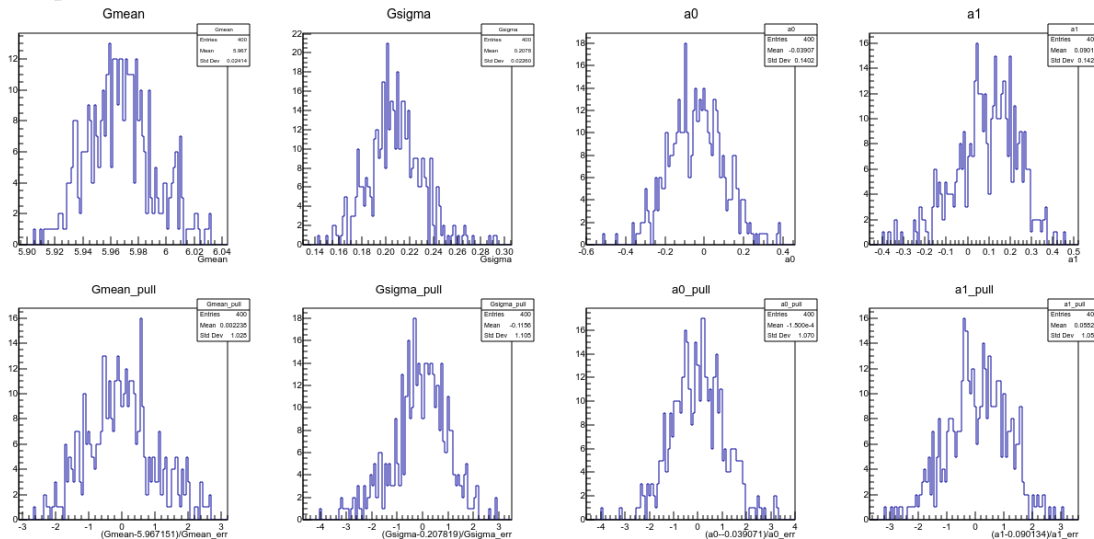
```
a0 -0.0390713 +- 0.132536 sigma 0.140173 meanPull -0.000150034 sigmaPull 1.06967
bias 0.00422866 bias Pull 0.0318165 sigma 1.06967
```

```
a1 0.0901345 +- 0.132916 sigma 0.142643 meanPull 0.058276 sigmaPull 1.05851
bias 0.00242745 bias Pull 0.0766501 sigma 1.05951
```

Get new ToyManager from previous fit

Run multicore on PROOF-lite
(or could use ROOT.Farm.Go(fit2))

Plot parameter distributions and pulls for the 400 toy fits



Speeding up likelihood calculation

Intrinsic RooFit optimisations implemented in RooNLLVar

Variables not used by PDF are dropped

PDF normalisation integrals only recalculated when range or parameter of that PDF changes

Components with no or constant parameters are precalculated and cached

→ but you must construct your PDF carefully to benefit

Fit with several parameters but 1 component

Generate toy data (100k):

$$f(\varphi) = 1 + 0.5 \cdot \cos(2\varphi)$$

Fit with function :

$$f(\varphi) = 1 + A \cdot \cos(2\varphi) + B \cdot \sin(2\varphi) + C \cdot \cos(\varphi) + D \cdot \cos(4\varphi)$$

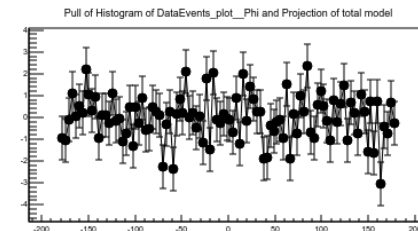
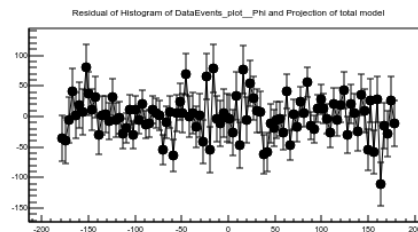
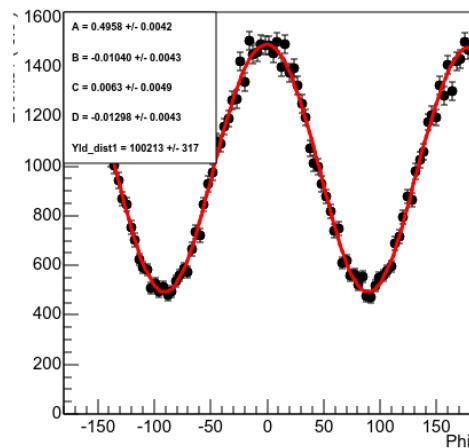
Use RooFit Factory Expression Pdf, takes general formula string

Only has 1 component, so any time A,B,C,D change must recalculate for every event

This uses a numerical integration

```
fit1.SetUp().LoadVariable("Phi[-180,180]");
fit1.SetUp().LoadParameter("A[0.,-1,1]");
fit1.SetUp().LoadParameter("B[0.5,-1,1]");
fit1.SetUp().LoadParameter("C[0.1,-1,1]");
fit1.SetUp().LoadParameter("D[-0.2,-1,1]");
fit1.SetUp().FactoryPDF("EXPR::dist1('1+A*cos(2*Phi/57.29578)'+B*sin(2*Phi/57.29578)+C*cos(Phi/57.29578)+'D*cos(4*Phi/57.29578)',Phi,A,B,C,D)");
fit1.SetUp().LoadSpeciesPDF("dist1",1);
```

Fit components for Phi



Fit in 10.5 seconds

Fit with many components

Generate toy data (100k):

$$f(\varphi) = 1 + 0.5 \cdot \cos(2\varphi)$$

Fit with function :

$$f(\varphi) = 1 + A \cdot \cos(2\varphi) + B \cdot \sin(2\varphi) + C \cdot \cos(\varphi) + D \cdot \cos(4\varphi)$$

Use RooFit Factory Expression Pdf,
takes general formula string

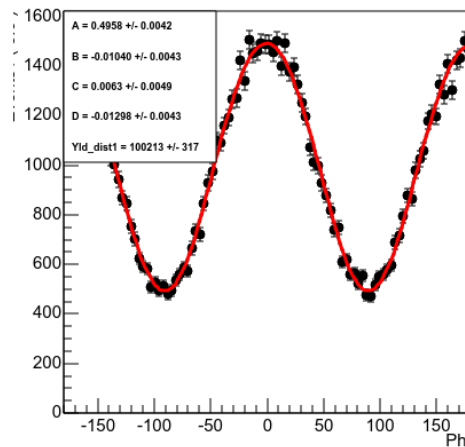
Define 4 separate terms which are
Combined into 1 PDF

COS2, SIN2 etc do not depend on any
parameters => precalculated and
cached

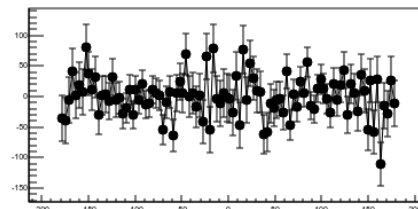
This uses a numerical integration

```
fit2.SetUp().LoadVariable("Phi[-180,180]");  
fit2.SetUp().LoadParameter("A[0.,-1,1]");  
fit2.SetUp().LoadParameter("B[0.5,-1,1]");  
fit2.SetUp().LoadParameter("C[0.1,-1,1]");  
fit2.SetUp().LoadParameter("D[-0.2,-1,1]");  
fit2.SetUp().LoadFormula("COS2=cos(2*@Phi[]/57.29578)");  
fit2.SetUp().LoadFormula("COS=cos(@Phi[]/57.29578)");  
fit2.SetUp().LoadFormula("COS4=cos(4*@Phi[]/57.29578)");  
fit2.SetUp().LoadFormula("SIN2=sin(2*@Phi[]/57.29578)");  
fit2.SetUp().FactoryPDF("EXPR::dist2('1+A*COS2+B*SIN2-C*COS+D*COS4',  
                                Phi,A,B,C,D,COS2,COS,COS4,SIN2)");  
fit2.SetUp().LoadSpeciesPDF("dist2",1);
```

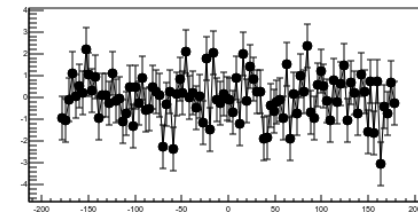
Fit components for Phi



Residual of Histogram of DataEvents_plot_Phi and Projection of total model



Pull of Histogram of DataEvents_plot_Phi and Projection of total model



Fit in 4.5 seconds

Speeding up MC integrals

$$A(p) = \sum_{i=0}^{N_{acc}} f(\tau_i, p)$$

No factorisation, must recalculate every time one parameter changes

$$A(p) = \sum_{i=0}^{N_{acc}} P_0(p) * f_0(\tau_i) + P_1(p) * f_1(\tau_i) + P_2(p) * f_2(\tau_i) + \dots$$

$$A(p) = P_0(p) \sum_{i=0}^{N_{acc}} f_0(\tau_i) + \dots$$

Summations over $\mathbf{f}_j$ s only needs done once and then cached = $\mathbf{F}_j$
i.e just 1 loop of MC events

Calculation of integral each step becomes very fast to calculate

$$A(p) = P_0(p) * F_0 + P_1(p) * F_1 + P_2(p) * F_2 + \dots$$

Already done in dedicated Amplitude Analysis software

RooComponentsPDF

Inherits from RooHSEventsPDF

- loads MC integral events tree

Takes any number of parameters, functions, formula, ...

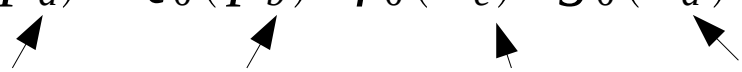
- anything that is RooAbsReal

Sorts each component into parameter dependent/independent terms

Precalculates normalisation integral terms at start of fit

Can be used for any PDF that can be written as a sum of products with fit parameters and observable parts factorised :

$$P_0(p_a)*Q_0(p_b)*f_0(\tau_c)*g_0(\tau_d) + P_1(p_a)*Q_1(p_b)*f_1(\tau_c)*g_1(\tau_d) + \dots$$



Some subsets of $\mathbf{p}$ Some subsets of $\boldsymbol{\tau}$

Fit with RooComponentsPDF

Generate toy data (100k):

$$f(\varphi) = 1 + 0.5 \cdot \cos(2\varphi)$$

Fit with function :

$$f(\varphi) = 1 + A \cdot \cos(2\varphi) + B \cdot \sin(2\varphi) + C \cdot \cos(\varphi) + D \cdot \cos(4\varphi)$$

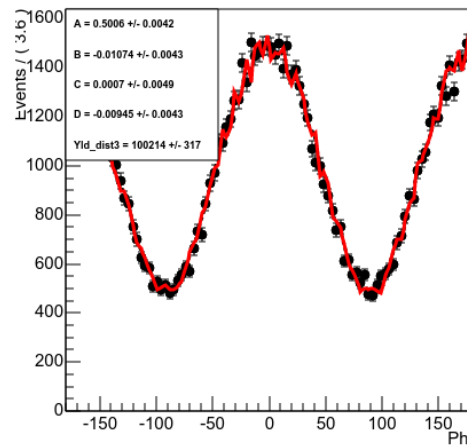
Use **RooComponentsPDF**,
give 100k MC events for integral

Define 4 separate terms which are
Combined into 1 PDF

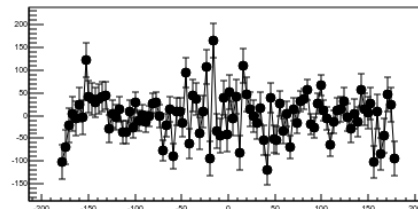
COS2, SIN2 etc do not depend on any
parameters => precalculated and
cached

```
fit3.SetUp().LoadVariable("Phi[-180,180]");  
fit3.SetUp().LoadParameter("A[0.,-1,1]");  
fit3.SetUp().LoadParameter("B[0.5,-1,1]");  
fit3.SetUp().LoadParameter("C[0.1,-1,1]");  
fit3.SetUp().LoadParameter("D[-0.2,-1,1]");  
fit3.SetUp().LoadFormula("COS2=cos(2*@Phi[]/57.29578)");  
fit3.SetUp().LoadFormula("COS=cos(@Phi[]/57.29578)");  
fit3.SetUp().LoadFormula("COS4=cos(4*@Phi[]/57.29578)");  
fit3.SetUp().LoadFormula("SIN2=sin(2*@Phi[]/57.29578)");  
  
fit3.SetUp().FactoryPDF("RooComponentsPDF::dist3  
|(1,{Phi},=A;COS2:B;SIN2:C;COS:D;COS4)");
```

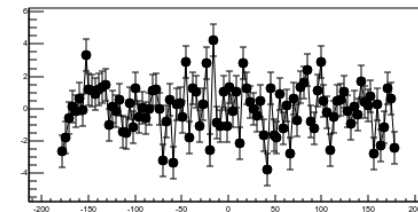
Fit components for Phi



Residual of Histogram of DataEvents_plot_Phi and Projection of total model



Pull of Histogram of DataEvents_plot_Phi and Projection of total model



Fit in 2.2 seconds

Measuring hadron SDMEs

P. Pauli, Glasgow

Toy Example $\vec{\gamma}p \rightarrow XY^+$ with $X \rightarrow P^+P^-$

Use decay angles of X : θ, φ

Planar production angle relative to Linear polarisation Φ

Degree of linear polarisation P_γ

Fit function:

$$W_0 = \frac{1}{4\pi} \left[3 \left(\frac{1}{2} - \rho_{11}^0 \right) \sin^2(\theta) + \rho_{11}^0 (1 + 3 \cos^2(\theta)) - 2\sqrt{3} \left(\text{Re}(\rho_{31}^0) \cos(\varphi) \sin(2\theta) + \text{Re}(\rho_{3-1}^0) \cos(2\varphi) \sin^2(\theta) \right) \right]$$

$$W_1 = \frac{1}{4\pi} \left[3\rho_{33}^1 \sin^2(\theta) + \rho_{11}^1 (1 + 3 \cos^2(\theta)) - 2\sqrt{3} \left(\text{Re}(\rho_{31}^1) \cos(\varphi) \sin(2\theta) + \text{Re}(\rho_{3-1}^1) \cos(2\varphi) \sin^2(\theta) \right) \right]$$

$$W_2 = \frac{1}{4\pi} \left[2\sqrt{3} \left(\text{Im}(\rho_{31}^2) \sin(\varphi) \sin(2\theta) + \text{Im}(\rho_{3-1}^2) \sin(2\varphi) \sin^2(\theta) \right) \right]$$

$$W = W_0 - P_\gamma \cos(2\Phi) W_1 - P_\gamma \sin(2\Phi) W_2$$

with Φ being the angle between production plane and polarisation plane.

SDME code

```
RF.SetUp().LoadVariable("theta[0,3.2]");
RF.SetUp().LoadVariable("phi[-3.2,3.2]");
RF.SetUp().LoadVariable("PHI[-1.5,3.2]");
RF.SetUp().LoadVariable("P[-0.2,1]");

RF.SetUp().LoadFormula("A=3./(8.*Pi)*sin(@theta[])*sin(@theta[])");
RF.SetUp().LoadFormula("B=1./(2.*Pi)*(3.*cos(@theta[])*cos(@theta[])-1.)");
RF.SetUp().LoadFormula("C=(-1.)*sqrt(3.)/(2.*Pi)*cos(@phi[])*sin(2.*@theta[])");
RF.SetUp().LoadFormula("D=(-1.)*sqrt(3.)/(2.*Pi)*cos(2.*@phi[])*sin(@theta[])*sin(@theta[])");
RF.SetUp().LoadFormula("E=(-1.)*3./(4.*Pi)*@P[]*cos(2*@PHI[])*sin(@theta[])*sin(@theta[])");
RF.SetUp().LoadFormula("F=(-1.)*3./(4.*Pi)*@P[]*cos(2*@PHI[])*(1./3.+cos(@theta[])*cos(@theta[]))");
RF.SetUp().LoadFormula("G=sqrt(3.)/(2.*Pi)*@P[]*cos(2*@PHI[])*cos(@phi[])*sin(2.*@theta[])");
RF.SetUp().LoadFormula("H=sqrt(3.)/(2.*Pi)*@P[]*cos(2*@PHI[])*cos(2.*@phi[])*sin(@theta[])*sin(@theta[])");
RF.SetUp().LoadFormula("I=(-1.)*sqrt(3.)/(2.*Pi)*@P[]*sin(2*@PHI[])*sin(@phi[])*sin(2.*@theta[])");
RF.SetUp().LoadFormula("J=(-1.)*sqrt(3.)/(2.*Pi)*@P[]*sin(2*@PHI[])*sin(2.*@phi[])*sin(@theta[])*sin(@theta[])");

RF.SetUp().LoadParameter("Rho011[0,-1,1]");
RF.SetUp().LoadParameter("Rho031[0,-1,1]");
RF.SetUp().LoadParameter("Rho03m1[0,-1,1]");
RF.SetUp().LoadParameter("Rho111[0,-1,1]");
RF.SetUp().LoadParameter("Rho133[0,-1,1]");
RF.SetUp().LoadParameter("Rho131[0,-1,1]");
RF.SetUp().LoadParameter("Rho13m1[0,-1,1]");
RF.SetUp().LoadParameter("Rho231[0,-1,1]");
RF.SetUp().LoadParameter("Rho23m1[0,-1,1]");

RF.SetUp().FactoryPDF("RooComponentsPDF::SDMES(0,{theta,phi,phiAngle_KPlus,PHI,P},
    =A:B;Rho011:C;Rho031:D;Rho03m1:E;Rho133:F;Rho111:G;Rho131:H;Rho13m1:I;Rho231:J;Rho23m1)");
```

Formula independent of parameters
=> precalculated and cached

RooComponentsPDF used for fast integration

SDME Pseudo data Fits

Acceptance : Holes in Lab θ

Reduced acceptance in decay θ

Number data events =13k MC Integral 100k

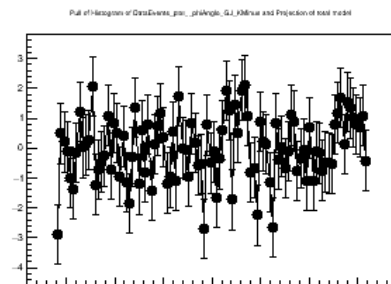
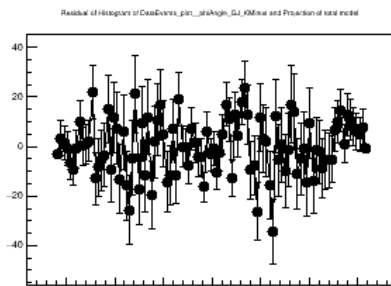
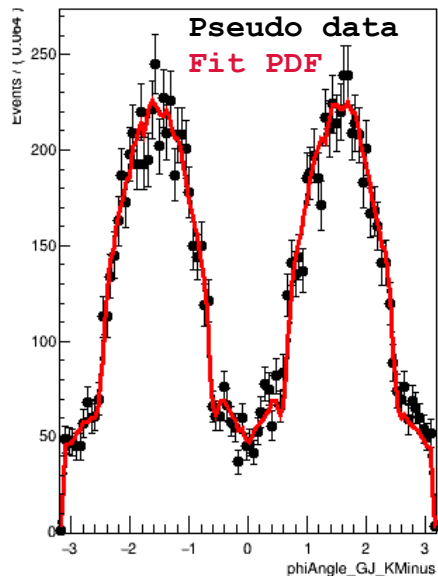
Run MCMC for 5000 posterior points (efficiency was ~3%)

Old Time = 64,000 s RooComponentsPDF = 1,400 s ~factor 50 speed-up

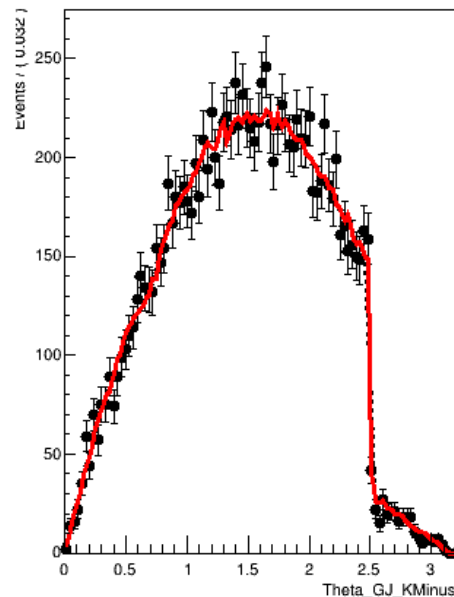
~23% is optimal so
Could be ~7x faster
Just change step size!

Data generated with random SDME values

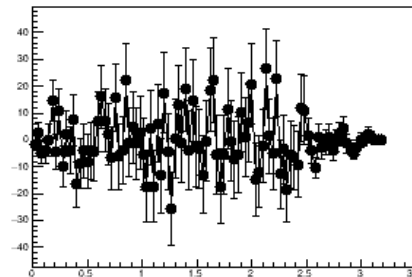
Fit components for $\phi_{\text{Angle_GJ_KMinus}}$



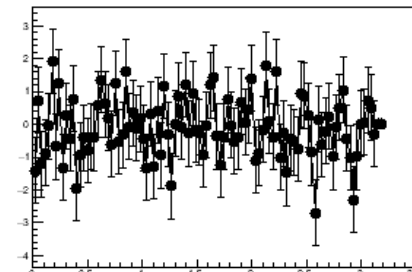
Fit components for $\Theta_{\text{GJ_KMinus}}$



Residual of histogram of DataEvents_hist_Theta_GJ_KMinus and Projection of total model



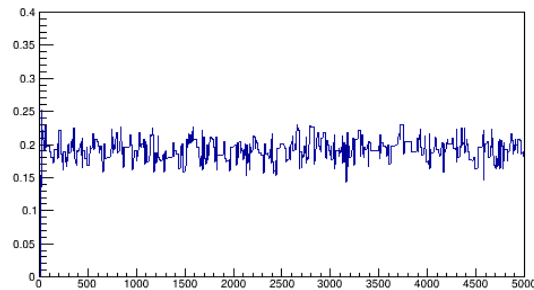
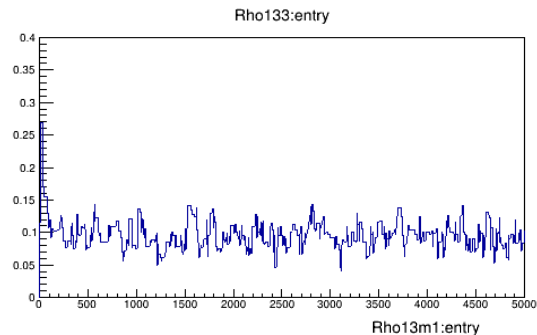
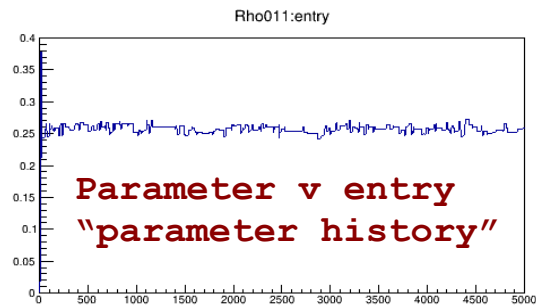
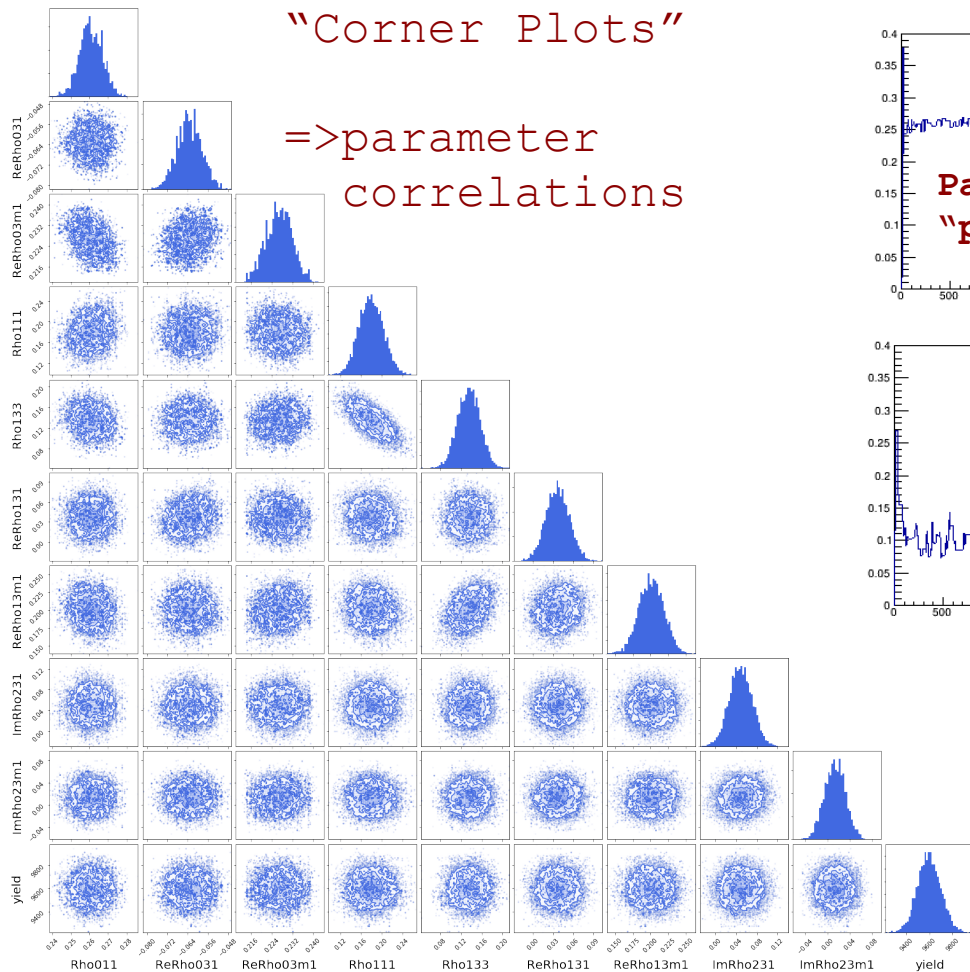
Pull of histogram of DataEvents_hist_Theta_GJ_KMinus and Projection of total model



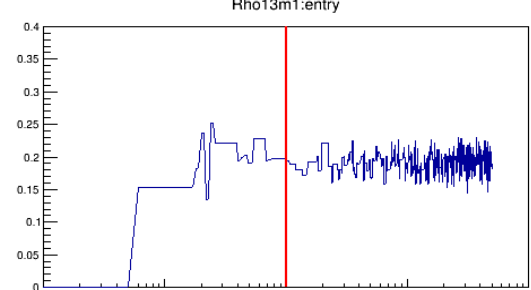
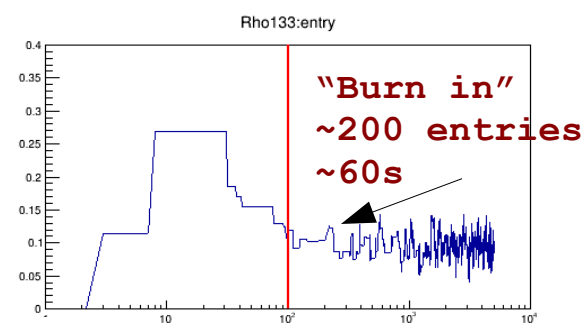
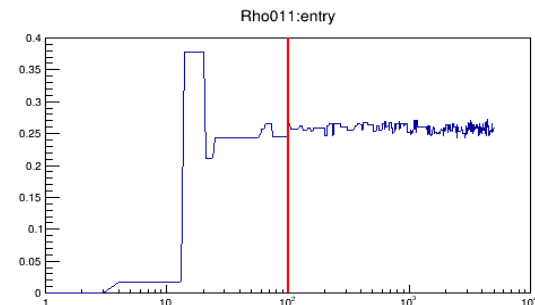
SDME MCMC diagnostics

"Corner Plots"

=>parameter
correlations



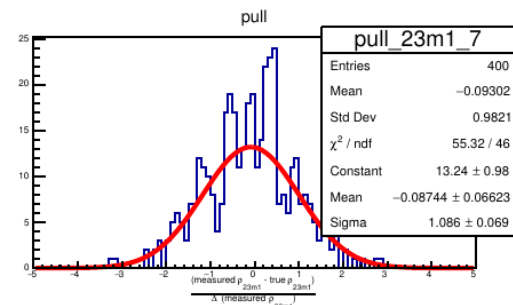
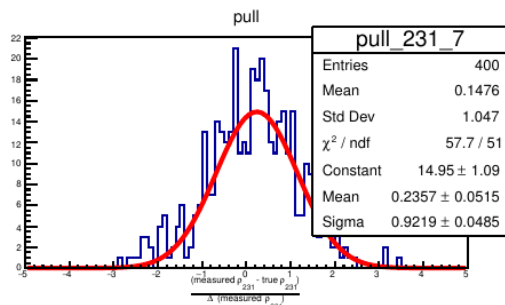
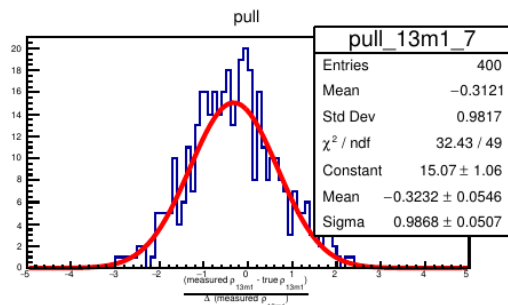
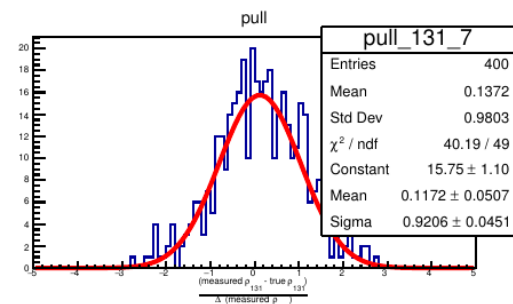
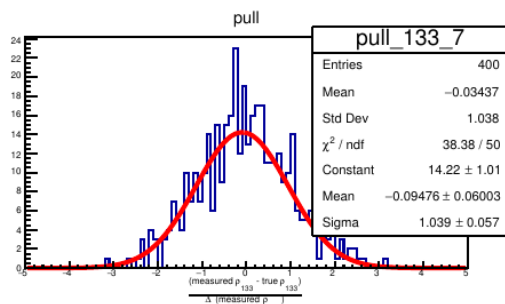
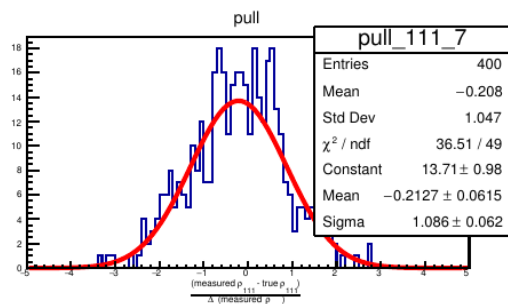
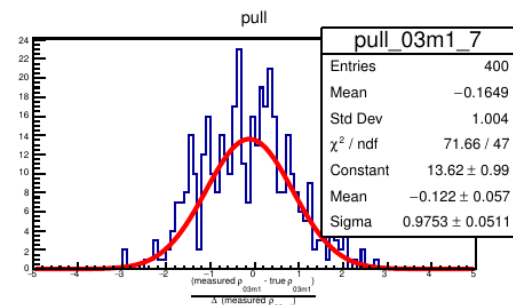
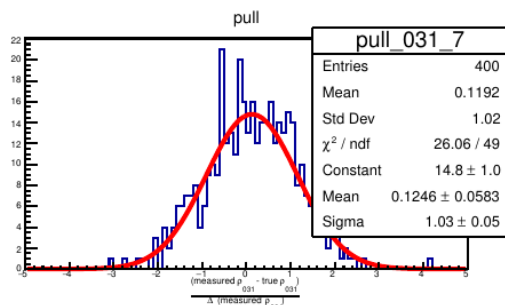
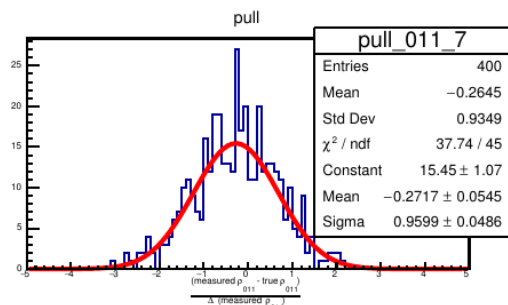
MCMC entry #



Log x scale

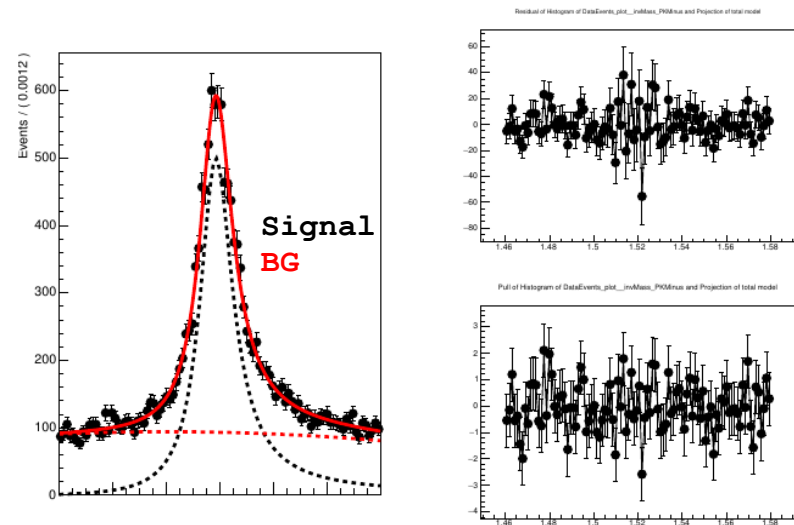
SDME Generate and MCMC 400 Toydatasets

- Look for bias, correct uncertainties



SDME Generate and MCMC 400 Toydatasets

- Include background and sWeights



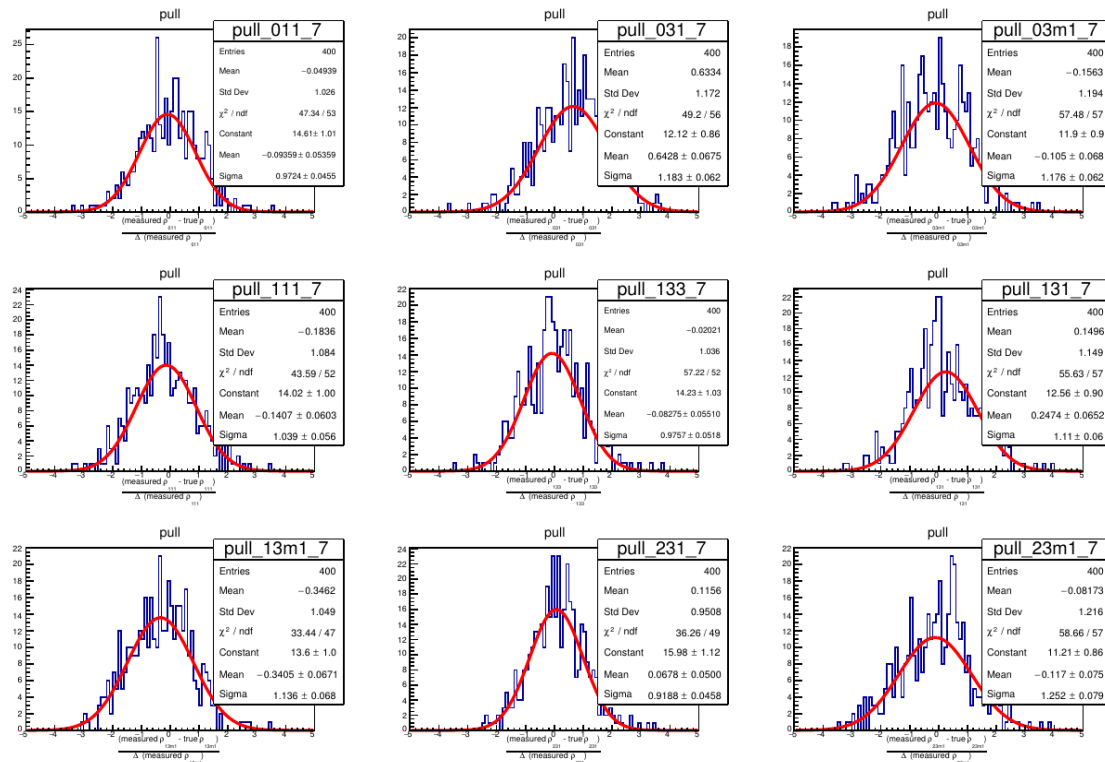
Fit to pseudo discriminatory variable

i.e. different PDF for sig and BG

Use weights in toy fits

Look for distortions due to background subtraction

MCMC more robust than Minuit with weights



Would like this to be standard procedure

JPAC, Two Meson Photoproduction : Moments

arXiv.org > hep-ph > arXiv:1906.04841

Search...

Help | Advanced

High Energy Physics - Phenomenology

Moments of angular distribution and beam asymmetries in $\eta\pi^0$ photoproduction at GlueX

V. Mathieu, M. Albaladejo, C. Fernández-Ramírez, A. W. Jackura, M. Mikhasenko, A.

Pilloni, A. P. Szczepaniak (JPAC collaboration)

(Submitted on 11 Jun 2019)

Directly relate Y_L^M moments to partial waves :

$$H^0(00) = H^1(00) + 2 \left[|P_1^{(+)}|^2 + |D_1^{(+)}|^2 + |D_2^{(+)}|^2 \right] ,$$

$$H^1(00) = 2 \left[|S_0^{(+)}|^2 + |P_0^{(+)}|^2 + |D_0^{(+)}|^2 \right] ,$$

$$H^0(10) = H^1(10) + \frac{4}{\sqrt{5}} \operatorname{Re}(P_1^{(+)} D_1^{(+)*}) ,$$

$$H^1(10) = \frac{4}{5\sqrt{3}} \left[2\sqrt{5} \operatorname{Re}(P_0^{(+)} D_0^{(+)*}) + 5 \operatorname{Re}(S_0^{(+)} P_0^{(+)*}) \right] ,$$

$$I(\Omega, \Phi) = I^0(\Omega) - P_y I^1(\Omega) \cos(2\Phi) - P_y I^2(\Omega) \sin(2\Phi)$$

$$I^0(\Omega) = \sum_L \sum_{M=0}^{M \leq L} \sqrt{\left(\frac{2L+1}{4\pi}\right)} (2 - \delta_{M,0}) H^0(L, M) \Re[Y_L^M(\Omega)]$$

$$I^1(\Omega) = - \sum_L \sum_{M=0}^{M \leq L} \sqrt{\left(\frac{2L+1}{4\pi}\right)} (2 - \delta_{M,0}) H^1(L, M) \Re[Y_L^M(\Omega)]$$

$$I^2(\Omega) = 2 \sum_L \sum_{M=0}^{M \leq L} \sqrt{\left(\frac{2L+1}{4\pi}\right)} \Im[H^2(L, M)] \Im[Y_L^M(\Omega)]$$

Fit Function
= sum of products
=> RooComponentsPDF

Just requires
RooHSSphHarmonic
class

Experimental : Parsers

Parsers → Create string for configuring RooComponentsPDF

```
auto parser=HS::PARSER::PolarisedSphHarmonicMoments("Moments","CosTh","Phi","PolPhi","Pol",4,-2,2);
string sum;
sum+="H_0_0_0[1]*ReY_0_0(CosTh,Phi,Y_0_0)";
sum+="SUM(L[1|3],M[0|2<L+1]){H_0_L_M[0,-1,1]*ReY_L_M(CosTh,Phi,Y_L_M)}";
sum+="SUM(L[0|3],M[0|2<L+1]){H_1_L_M[0,-1,1]*ReY_L_M(CosTh,Phi,Y_L_M)*COS2PHI}";
sum+="SUM(L[1|3],M[1|2<L+1]){H_2_L_M[0,-1,1]*ImY_L_M(CosTh,Phi,Y_L_M)*SIN2PHI}";

Fitter.SetUp().ParserPDF(sum,parser);

Fitter.SetUp().LoadSpeciesPDF("Moments",1);
```

Becomes function with 22 $H^i(L,M)$ fit parameters,

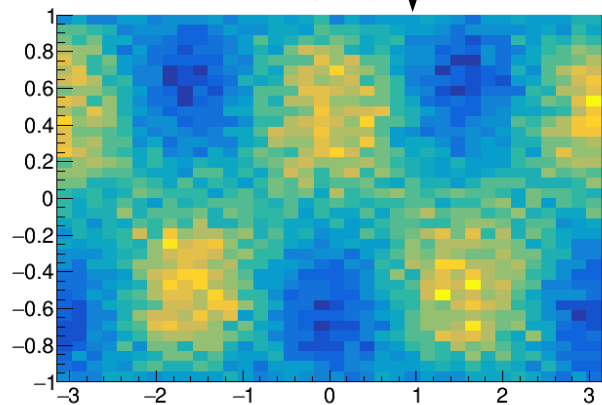
Objects of
RooHSSphHarmonicRe

```
RooComponentsPDF::Moments(0,{CosTh,Phi,PolPhi,Pol},=H_0_0_0[1];ReY_0_0:
H_0_1_0[0,-1,1];ReY_1_0:H_0_1_1[0,-1,1];ReY_1_1:H_0_2_0[0,-1,1];ReY_2_0:
H_0_2_1[0,-1,1];ReY_2_1:H_0_2_2[0,-1,1]; ReY_2_2:H_0_3_0[0,-1,1];ReY_3_0:
H_0_3_1[0,-1,1];ReY_3_1:H_0_3_2[0,-1,1];ReY_3_2:H_1_0_0[0,-1,1]; ReY_0_0;COS2PHI:
H_1_1_0[0,-1,1];ReY_1_0;COS2PHI:H_1_1_1[0,-1,1];ReY_1_1;COS2PHI:
H_1_2_0[0,1,1];ReY_2_0;COS2PHI:H_1_2_1[0,-1,1];ReY_2_1;COS2PHI:
H_1_2_2[0,-1,1];ReY_2_2;COS2PHI:H_1_3_0[0,-1,1]; ReY_3_0;COS2PHI:
H_1_3_1[0,1,1];ReY_3_1;COS2PHI:H_1_3_2[0,1,1];ReY_3_2;COS2PHI:
H_2_1_1[0,-1,1];ImY_1_1;SIN2PHI:H_2_2_1[0,-1,1];ImY_2_1;SIN2PHI:
H_2_2_2[0,-1,1];ImY_2_2;SIN2PHI:H_2_3_1[0,1,1];ImY_3_1;SIN2PHI:H_2_3_2[0,-1,1];ImY_3_2;SIN2PHI)
```

Fit Pseudo Data with Moments PDF

Pseudo Data 100K events
+ 100K MC events

Decay θ V ϕ



L<4, M<3 22 parameters

Fit with Minuit

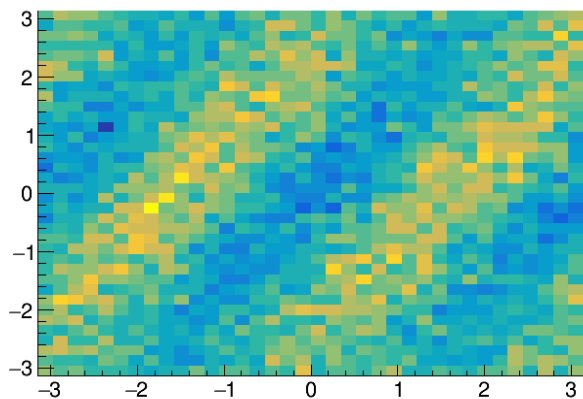
→ 40s

Fit with MCMC

2000 samples

→ 260s 18%accept

Decay ϕ V polarised Φ



L<5, M<4 36 params.

Fit with Minuit

→ 130s

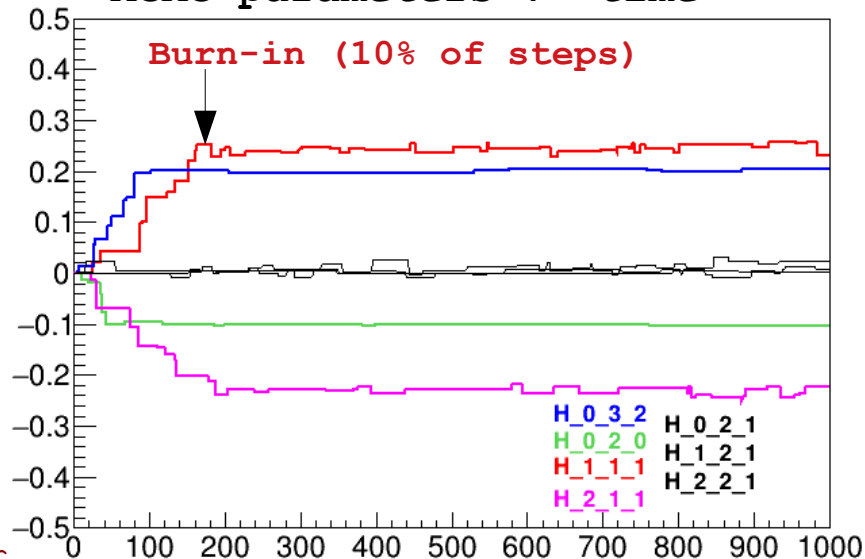
Fit with MCMC

2000 samples

→ 360s 18%accept

MCMC scales better
with parameter #

MCMC parameters V "time"



Proof of principal : Fitting Amplitudes

RooFit does not directly support complex numbers

- Try expanding intensities with amplitudes so only deal with real valued numbers and functions

$$(\sum AB)^2 = \sum_i F(A_i, A_i) F(B_i, B_i) + 2 \sum_i \sum_{j, i < j} [F(A_i, A_j) F(B_i, B_j) - G(A_i, A_j) G(B_i, B_j)]$$

$$F(C, D) = \Re(C) \Re(D) + \Im(C) \Im(D) \quad G(C, D) = \Im(C) \Re(D) - \Re(C) \Im(D)$$

Define RooFit function for F and G which take 4 real numbers

These numbers can be real and imaginary parts of complex parameters or functions (e.g. Y_L^M)

$$\begin{aligned} (\sum h_L^M Y_L^M)^2 &= F(h_0^0, h_0^0) F(Y_0^0, Y_0^0) + F(h_1^0, h_1^0) F(Y_1^0, Y_1^0) \dots \\ &+ 2 F(h_0^0, h_1^0) F(Y_0^0, Y_1^0) - 2 G(h_0^0, h_1^0) G(Y_0^0, Y_1^0) + \dots \end{aligned}$$

Sum of products of
F and G functions
→ RooComponentsPDF 44

JPAC, Two Meson Photoproduction: Amps

Amplitudes given as

$$U_k^{(\epsilon)}(\Omega) = \sum_{\ell, m} [\ell]_{m; k}^{(\epsilon)} Y_{\ell}^m(\Omega) ,$$

$$\tilde{U}_k^{(\epsilon)}(\Omega) = \sum_{\ell, m} [\ell]_{m; k}^{(\epsilon)} [Y_{\ell}^m(\Omega)]^* .$$

$$I^0(\Omega) = \kappa \sum_{\epsilon, k} |U_k^{(\epsilon)}(\Omega)|^2 + |\tilde{U}_k^{(\epsilon)}(\Omega)|^2 ,$$

Use "Parser" to code as

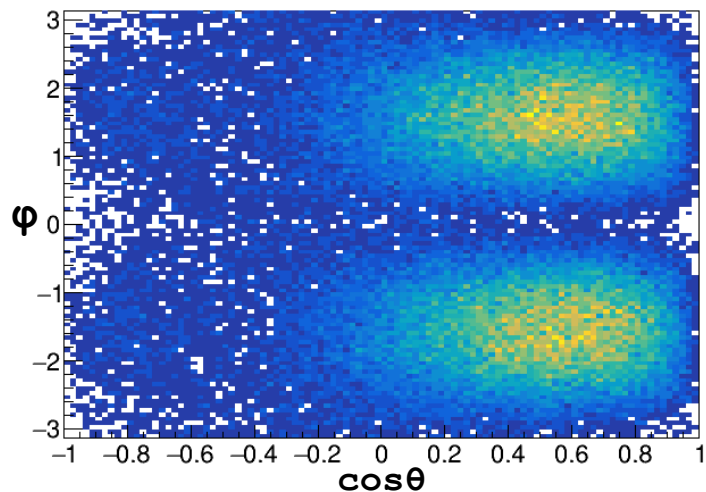
```

                                L<2          M=-1,0,1 <L
string sum = "SUM(L[0|2],M[-1|1<L+1>-L-1])
              {h_L_M[0,-1,1][0,-1,1]*Y_L_M(CosTh,Phi,L,M)}^2
+ SUM(L[0|2],M[-1|1<L+1>-L-1])
  {h_L_M[0,-1,1][0,-1,1]*Y_L_M^CONJ(CosTh,Phi,L,M)}^2 ";
```

Real and imaginary fit parameters

Complex spherical harmonics

Pseudo data, Spherical Harmonic Moments Fit to data generated with amplitudes

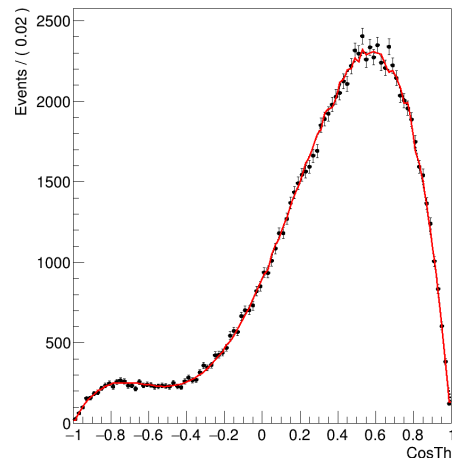


Analytical Moments
from Amplitudes

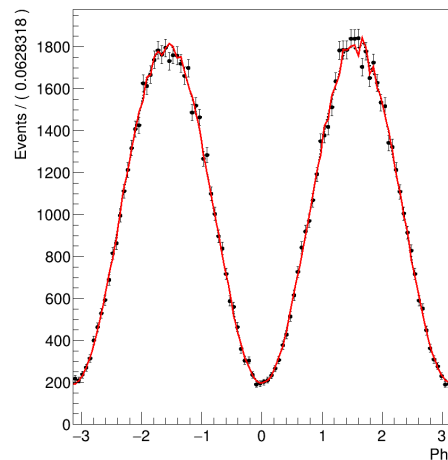
Fit Results

(L,M) = (1,0): $H_0 = 0.3578$;	$H_{0\ 1\ 0} = 0.358112 \pm 0.00797507$
(L,M) = (1,1): $H_0 = 0.0000$;	$H_{0\ 1\ 1} = -0.000740184 \pm 0.000722244$
(L,M) = (2,0): $H_0 = -0.0629$;	$H_{0\ 2\ 0} = -0.0623866 \pm 0.000474004$
(L,M) = (2,1): $H_0 = 0.0000$;	$H_{0\ 2\ 1} = -0.000551358 \pm 0.000478794$
(L,M) = (2,2): $H_0 = -0.1680$;	$H_{0\ 2\ 2} = -0.169541 \pm 0.000885905$
(L,M) = (3,0): $H_0 = -0.1533$;	$H_{0\ 3\ 0} = -0.153232 \pm 0.00448687$
(L,M) = (3,1): $H_0 = 0.0000$;	$H_{0\ 3\ 1} = -0.000258408 \pm 0.000334288$
(L,M) = (3,2): $H_0 = -0.1400$;	$H_{0\ 3\ 2} = -0.140019 \pm 0.00137683$
(L,M) = (3,3): $H_0 = 0.0000$;	$H_{0\ 3\ 3} = 0.000338522 \pm 0.000556094$
(L,M) = (4,0): $H_0 = -0.0762$;	$H_{0\ 4\ 0} = -0.0765479 \pm 0.000712912$
(L,M) = (4,1): $H_0 = 0.0000$;	$H_{0\ 4\ 1} = 0.00024877 \pm 0.00030687$
(L,M) = (4,2): $H_0 = -0.0602$;	$H_{0\ 4\ 2} = -0.0605708 \pm 0.000410417$
(L,M) = (4,3): $H_0 = 0.0000$;	$H_{0\ 4\ 3} = -0.000109336 \pm 0.000434857$
(L,M) = (4,4): $H_0 = 0.0000$;	$H_{0\ 4\ 4} = 0.000138865 \pm 0.000480976$

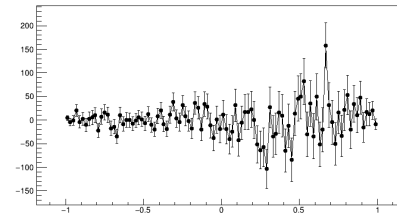
Fit components for CosTh



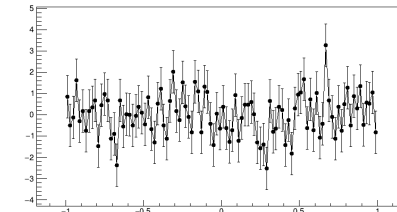
Fit components for Phi



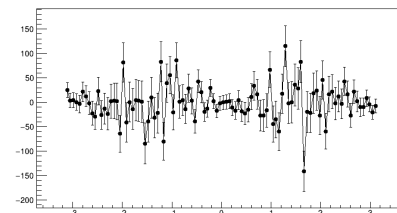
Residual of Histogram of DataEvents_plot_CosTh and Projection of total model



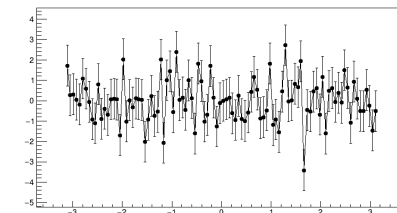
Pull of Histogram of DataEvents_plot_CosTh and Projection of total model



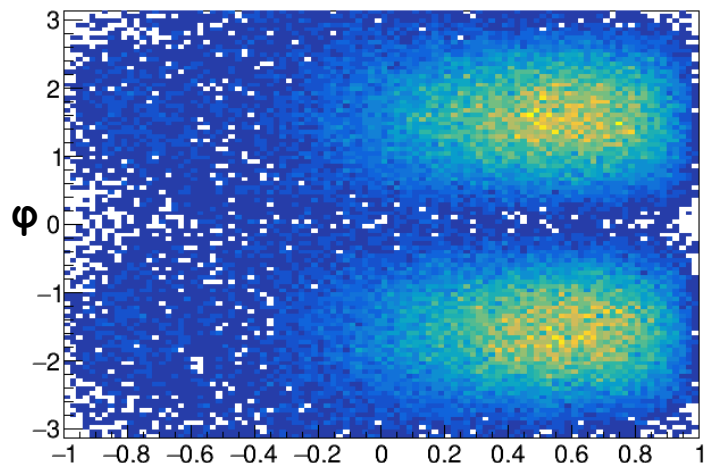
Residual of Histogram of DataEvents_plot_Phi and Projection of total model



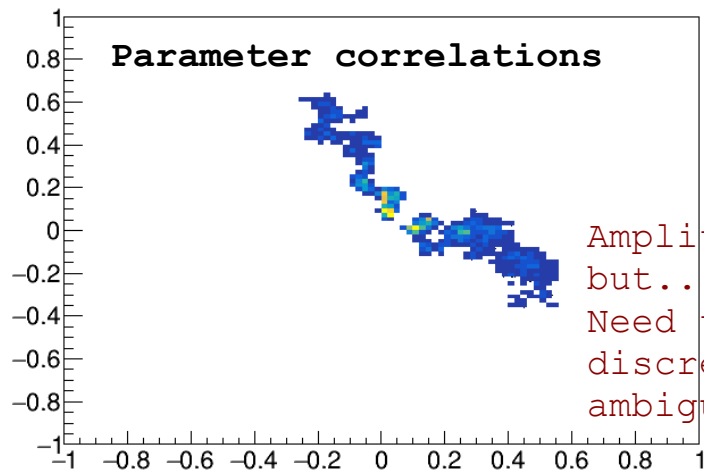
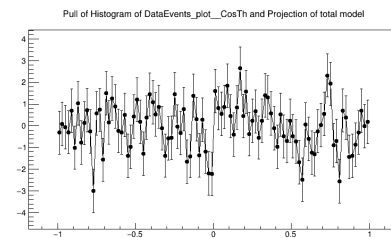
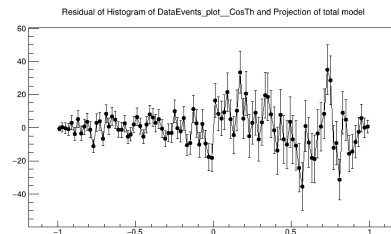
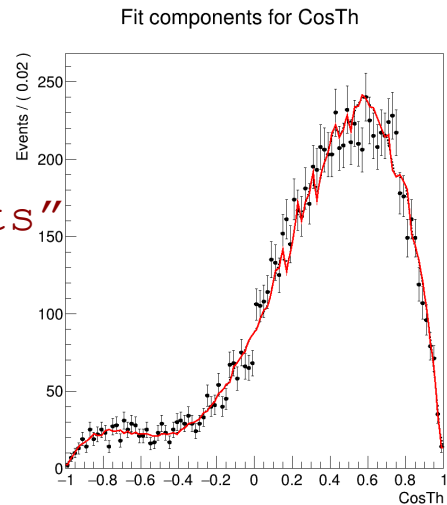
Pull of Histogram of DataEvents_plot_Phi and Projection of total model



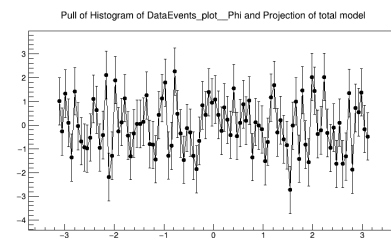
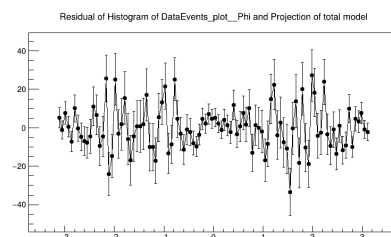
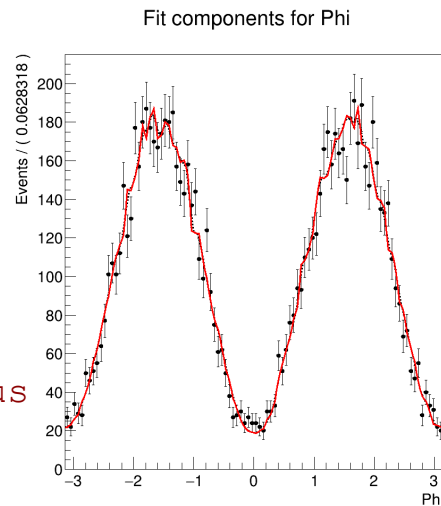
Pseudo data, Amplitudes Fit to data generated with amplitudes



MCMC "Fits"



Amplitude fit works,
but...
Need to account for
discrete and continuous
ambiguities



BruFit

<https://github.com/dglazier/brufit>

BruFit

A RooFit based event based maximum likelihood fitting package

The purpose of this package is to add to the RooFit package to allow analysis of hadronic physics scattering reactions.

The main feature is a PDF class (RooHSEventsPDF) that allows calculation of normalisation integrals from Monte Carlo detector simulations which allow the acceptance of detector systems to be corrected for when extracting observables.

In addition the (RooComponentsPDF) class provides caching of these integrals for fast evaluation when the PDF is a sum of products

```
> brufit
root [1] FitManager fm
root [2] fm.SetUp().SetOutDir("out/"); //Put results files in out/
root [3] fm.SetUp().LoadVariable("phi[-3.1416,3.1416]"); //phi is a variable in the data tree
root [4] fm.SetUp().FactoryPDF("EXPR::amplitude('1+A[0,-1,1]*cos(2*phi)',phi,A)"); //Fit a cos2phi distr
root [5] fm.SetUp().LoadSpeciesPDF("amplitude");//add to the total fit PDF
root [6] fm.LoadData("treeName","fileName.root"); //set data (ROOT tree)
root [7] Here::Go(&fm); //run the fit
```

Try the jupyter python notebooks....

Summary

We are developing general data fitting tools based on RooFit

Extend RooFit by adding PDF class with normalisation integral calculated from simulated MC events

Extend this by adding further PDF with cached MC integrals

Fits can be performed via "user friendly" Jupyter notebooks

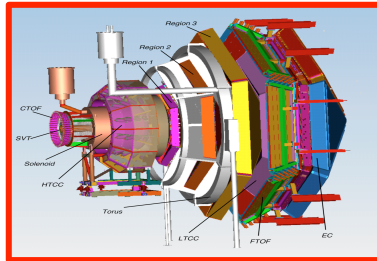
Applying ToyMC studies to fits with signal and background is a good way to investigate systematic effects in the parameter extraction

Tools will be used with MesonEx experiment

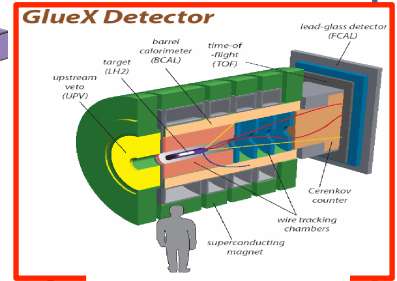
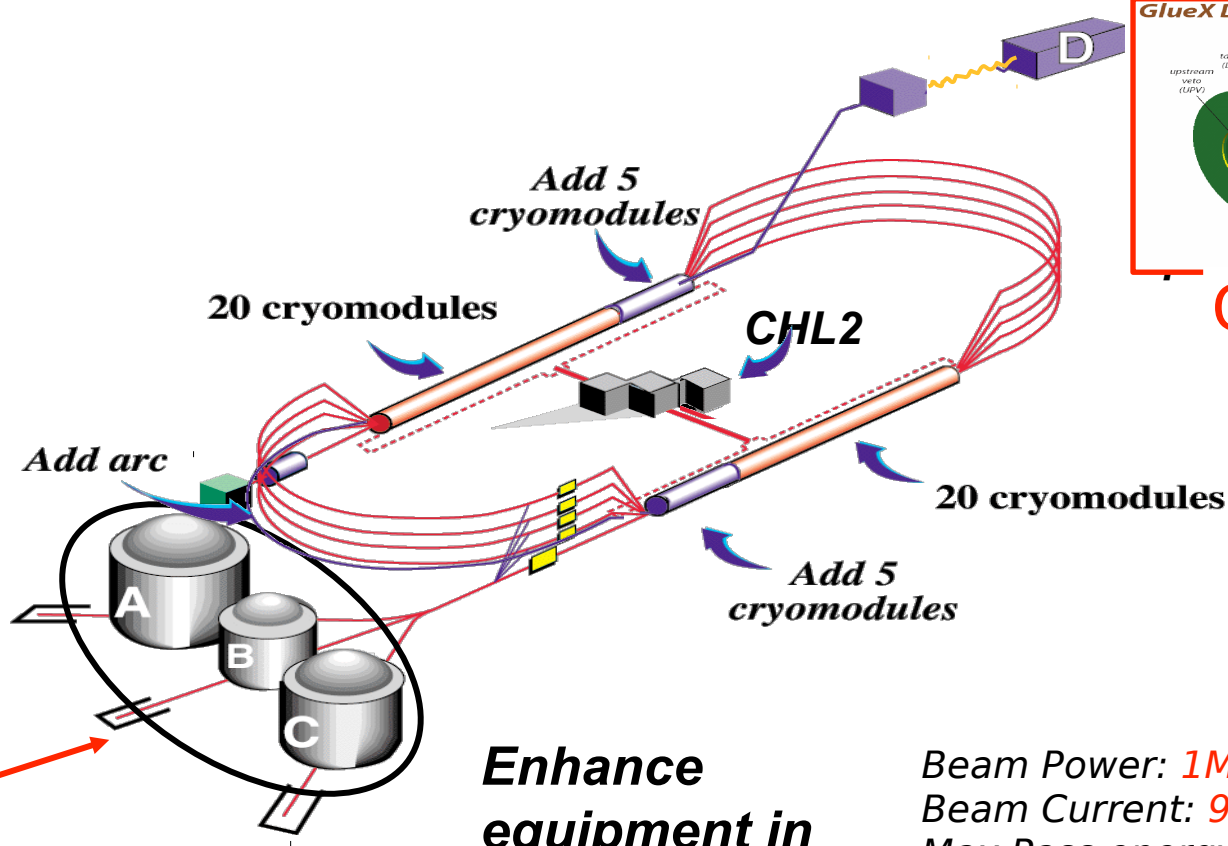
Already being applied to CLAS, GLUEX, CB@MAMI data

1: CSST3_h 2 0 h 1 0; CSST3_Y 2 0 Y 1 0: CSST3_h 2 0 h 1 1; CSST3_Y 2 0 Y 1 1: CSST3_h 2 0 h 2 -1; CSST3_Y 2 0 Y 2 -
1: CSST3_h 2 0 h 2 1; CSST3_Y 2 0 Y 2 1: CSST3_h 2 1 h 0 0; CSST3_Y 2 1 Y 0 0: CSST3_h 2 1 h 1 -1; CSST3_Y 2 1 Y 1 -
1: CSST3_h 2 1 h 1 0; CSST3_Y 2 1 Y 1 0: CSST3_h 2 1 h 1 1; CSST3_Y 2 1 Y 1 1: CSST3_h 2 1 h 2 -1; CSST3_Y 2 1 Y 2 -
1: CSST3_h 2 1 h 2 0; CSST3_Y 2 1 Y 2 0: CSST_h 0 0 h 0 0; CSST_Y 0 0 Y 0 0 CONJ: CSST_h 0 0 h 1 -1; CSST_Y 0 0 Y 1 -
1 CONJ: CSST_h 0 0 h 1 0; CSST_Y 0 0 Y 1 0 CONJ: CSST_h 0 0 h 1 1; CSST_Y 0 0 Y 1 1 CONJ: CSST_h 0 0 h 2 -1; CSST_Y 0 0 Y 2 -
1 CONJ: CSST_h 0 0 h 2 0; CSST_Y 0 0 Y 2 0 CONJ: CSST_h 0 0 h 2 1; CSST_Y 0 0 Y 2 1 CONJ: CSST_h 1 -1 h 0 0; CSST_Y 1 -1 Y 0 0 CONJ: CSST_h 1 -
1 h 1 -1; CSST_Y 1 -1 Y 1 -1 CONJ: CSST_h 1 -1 h 1 0; CSST_Y 1 -1 Y 1 0 CONJ: CSST_h 1 -1 h 1 1; CSST_Y 1 -1 Y 1 1 CONJ: CSST_h 1 -1 h 2 -
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1 CONJ: CSST_h 1 0 h 2 0; CSST_Y 1 0 Y 2 0 CONJ: CSST_h 1 0 h 2 1; CSST_Y 1 0 Y 2 1 CONJ: CSST_h 1 1 h 0 0; CSST_Y 1 1 Y 0 0 CONJ: CSST_h 1 1 h 1
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1 CONJ: CSST_h 2 0 h 1 0; CSST_Y 2 0 Y 1 0 CONJ: CSST_h 2 0 h 1 1; CSST_Y 2 0 Y 1 1 CONJ: CSST_h 2 0 h 2 -1; CSST_Y 2 0 Y 2 -
1 CONJ: CSST_h 2 0 h 2 0; CSST_Y 2 0 Y 2 0 CONJ: CSST_h 2 0 h 2 1; CSST_Y 2 0 Y 2 1 CONJ: CSST_h 2 1 h 0 0; CSST_Y 2 1 Y 0 0 CONJ: CSST_h 2 1 h 1
-1; CSST_Y 2 1 Y 1 -1 CONJ: CSST_h 2 1 h 1 0; CSST_Y 2 1 Y 1 0 CONJ: CSST_h 2 1 h 1 1; CSST_Y 2 1 Y 1 1 CONJ: CSST_h 2 1 h 2 -1; CSST_Y 2 1 Y 2 -
1 CONJ: CSST_h 2 1 h 2 0; CSST_Y 2 1 Y 2 0 CONJ: CSST_h 2 1 h 2 1; CSST_Y 2 1 Y 2 1 CONJ: CSST3_h 0 0 h 1 -1; CSST3_Y 0 0 Y 1 -
1 CONJ: CSST3_h 0 0 h 1 0; CSST3_Y 0 0 Y 1 0 CONJ: CSST3_h 0 0 h 1 1; CSST3_Y 0 0 Y 1 1 CONJ: CSST3_h 0 0 h 2 -1; CSST3_Y 0 0 Y 2 -
1 CONJ: CSST3_h 0 0 h 2 0; CSST3_Y 0 0 Y 2 0 CONJ: CSST3_h 0 0 h 2 1; CSST3_Y 0 0 Y 2 1 CONJ: CSST3_h 1 -1 h 0 0; CSST3_Y 1 -
1 Y 0 0 CONJ: CSST3_h 1 -1 h 1 0; CSST3_Y 1 -1 Y 1 0 CONJ: CSST3_h 1 -1 h 1 1; CSST3_Y 1 -1 Y 1 1 CONJ: CSST3_h 1 -1 h 2 -1; CSST3_Y 1 -1 Y 2 -
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1 Y 2 1 CONJ: CSST3_h 1 0 h 0 0; CSST3_Y 1 0 Y 0 0 CONJ: CSST3_h 1 0 h 1 -1; CSST3_Y 1 0 Y 1 -
1 CONJ: CSST3_h 1 0 h 1 1; CSST3_Y 1 0 Y 1 1 CONJ: CSST3_h 1 0 h 2 -1; CSST3_Y 1 0 Y 2 -
1 CONJ: CSST3_h 1 0 h 2 0; CSST3_Y 1 0 Y 2 0 CONJ: CSST3_h 1 0 h 2 1; CSST3_Y 1 0 Y 2 1 CONJ: CSST3_h 1 1 h 0 0; CSST3_Y 1 1 Y 0 0 CONJ: CSST3_h 1
1 h 1 -1; CSST3_Y 1 1 Y 1 -1 CONJ: CSST3_h 1 1 h 1 0; CSST3_Y 1 1 Y 1 0 CONJ: CSST3_h 1 1 h 2 -1; CSST3_Y 1 1 Y 2 -
1 CONJ: CSST3_h 1 1 h 2 0; CSST3_Y 1 1 Y 2 0 CONJ: CSST3_h 1 1 h 2 1; CSST3_Y 1 1 Y 2 1 CONJ: CSST3_h 2 -1 h 0 0; CSST3_Y 2 -
1 Y 0 0 CONJ: CSST3_h 2 -1 h 1 -1; CSST3_Y 2 -1 Y 1 -1 CONJ: CSST3_h 2 -1 h 1 0; CSST3_Y 2 -1 Y 1 0 CONJ: CSST3_h 2 -1 h 1 1; CSST3_Y 2 -
1 Y 1 1 CONJ: CSST3_h 2 -1 h 2 0; CSST3_Y 2 -1 Y 2 0 CONJ: CSST3_h 2 -1 h 2 1; CSST3_Y 2 -
1 Y 2 1 CONJ: CSST3_h 2 0 h 0 0; CSST3_Y 2 0 Y 0 0 CONJ: CSST3_h 2 0 h 1 -1; CSST3_Y 2 0 Y 1 -
1 CONJ: CSST3_h 2 0 h 1 0; CSST3_Y 2 0 Y 1 0 CONJ: CSST3_h 2 0 h 1 1; CSST3_Y 2 0 Y 1 1 CONJ: CSST3_h 2 0 h 2 -1; CSST3_Y 2 0 Y 2 -
1 CONJ: CSST3_h 2 0 h 2 1; CSST3_Y 2 0 Y 2 1 CONJ: CSST3_h 2 1 h 0 0; CSST3_Y 2 1 Y 0 0 CONJ: CSST3_h 2 1 h 1 -1; CSST3_Y 2 1 Y 1 -
1 CONJ: CSST3_h 2 1 h 1 0; CSST3_Y 2 1 Y 1 0 CONJ: CSST3_h 2 1 h 1 1; CSST3_Y 2 1 Y 1 1 CONJ: CSST3_h 2 1 h 2 -1; CSST3_Y 2 1 Y 2 -
1 CONJ: CSST3_h 2 1 h 2 0; CSST3_Y 2 1 Y 2 0 CONJ)

JLAB12



CLAS12



GLUEX

**Enhance
equipment in
existing halls**

Beam Power: **1MW**
 Beam Current: **90 μ A**
 Max Pass energy: **2.2 GeV**
 Max Energy Hall A-C: **10.9 GeV**
 Max Energy Hall D: **12 GeV**

Detect electrons at small angle to perform quasi-real photo-production experiments.

Calorimeter: electron energy/momentum

Photon energy ($\nu = E - E'$)

Polarization $\epsilon^{-1} \approx 1 + \nu^2/2EE'$

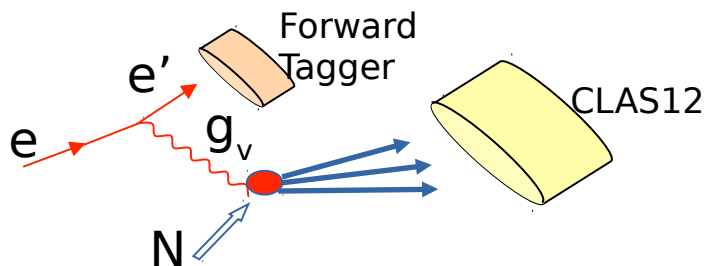
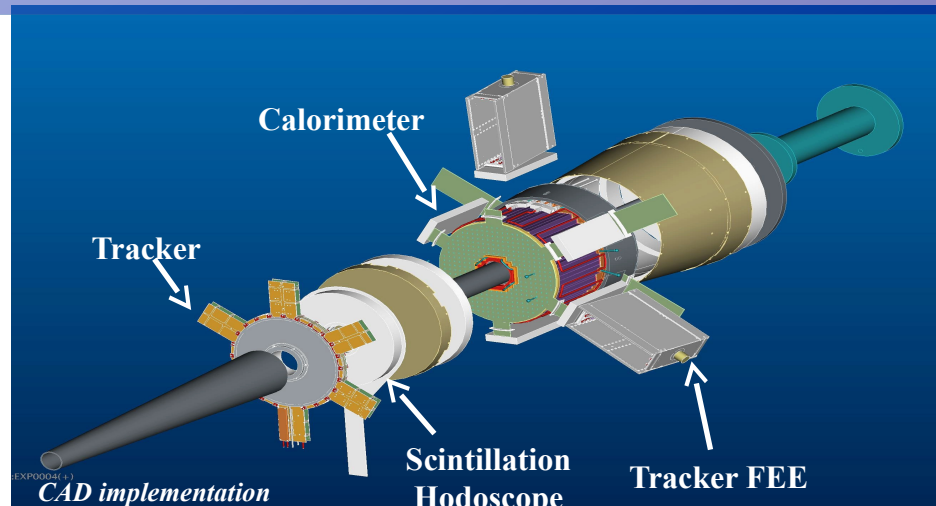
PbWO₄ crystals with APD/SiPM readout

Scintillation Hodoscope: veto for photons

Scintillator tiles with WLS readout

Tracker: electron angles, polarization plane

MicroMegas detectors



$E_{\text{scattered}}$	0.5 - 4.5 GeV
θ	2.5° - 4.5°
ϕ	0° - 360°
ν	6.5 - 10.5 GeV
Q^2	0.01 - 0.3 GeV ² ($< Q^2 > 0.1$ GeV ²)
W	3.6 - 4.5 GeV

Fitting discriminatory variables

Signal shapes are not always well described by parametric functions

⇒ Simulated PDFs

Systematic uncertainty in shape accounted for via morphing with additional nuisance parameters

