SHMS / HMS Detector System Summary / Updates (Mostly Cerenkov Updates)

B. Sawatzky (JLab)





Summary as of June 2018

- Generally Detectors/DAQ are working as planned
 - → Calibrations, cross-checks are in progress
 - » See upcoming Hall A/C collaboration meeting and Analysis workshop!
 - → SHMS gas Cerenkov yields are lower than predicted, investigations underway
 - » SHMS NGC has one very noisy PMT
 - Spare PMT available, will be replaced this summer
 - » SHMS NGC yield may be concern at high momentum settings (> 7 GeV/c)?
 - May want to be conservative with PID expectations at such settings?
 - » SHMS HGC yields are also lower than predicted
 - → HMS mirrors found to be badly damaged
 - » Seems probable it has been like this for years (since 2008/9?)
 - » Spare mirrors located and being installed now





HMS Gas Cherenkov





HMS Cerenkov Mirror Replacement

- Completed over Summer
 - → Kudos to Howard F. and Tech staff
- Went well
 - → Both mirrors replaced with good spares
 - → Focus checked and optimized

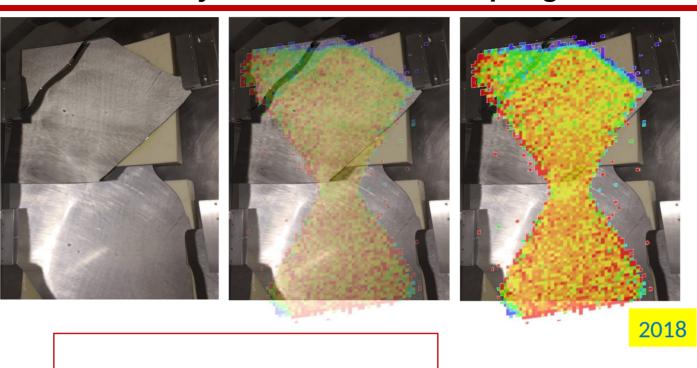
 HMS Cerenkov performance seems to be fully restored







HMS Cerenkov(OLD): Have we really been taking data that way, or did the mirrors just break after the "Spring-18 Run"?



Certainly the mirrors were broken during the (end of) "Spring 18", at least.

They may have been broken (breaking?) for a very long time.

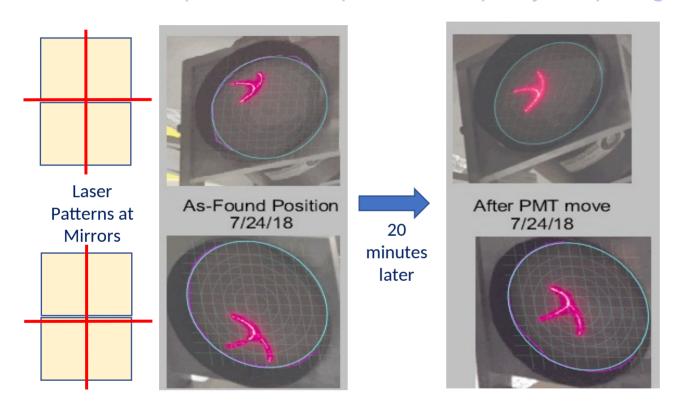
Slide from H. Fenker (June 2018)

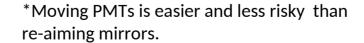


HMS gas Cerenkov

- Results guided the mirror alignment procedure
 - Laser lines projected on mirror _ Ideal pattern on PMT face

Yielded performance improvement beyond just replacing broken mirrors





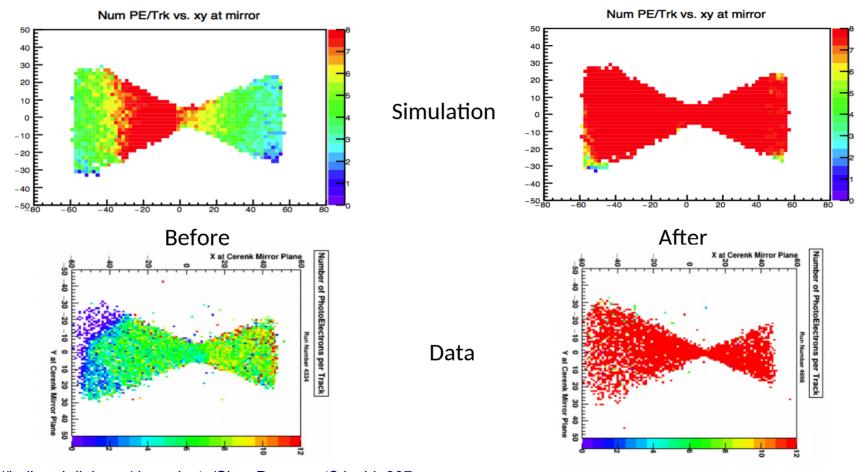
Calculated pattern that should be seen on PMT faces for crossed-lines laser. (Note that PMT circles shown are at X= ±32cm. Mirrors are nominally at X= ±30cm.)

https://hallcweb.jlab.org/doc-private/ShowDocument?docid=997

Fenker 29 Nov 2018

HMS gas Cerenkov

- Results guided the mirror alignment procedure
 - Laser lines projected on mirror _ Ideal pattern on PMT face
 - Yielded performance improvement beyond just replacing broken mirrors



https://hallcweb.jlab.org/doc-private/ShowDocument?docid=997

SHMS Heavy Gas Cherenkov





SHMS Heavy Gas Cerenkov (HGC)



- SHMS HGC
 - \rightarrow 4 mirrors
 - » ~90% reflection to 200 nm
 - → 4 UV glass PMTs
 - » Good QE down to ~200 nm
 - → Tank can be pumped to vacuum then filled to (partial-)pressure
- Typical fill
 - \rightarrow 1 atm of C_4F_80
 - → pion threshold = ~2.6 GeV/c
- Expected npe yield
 - > 25 photo-electrons
 - → Seeing less than that...

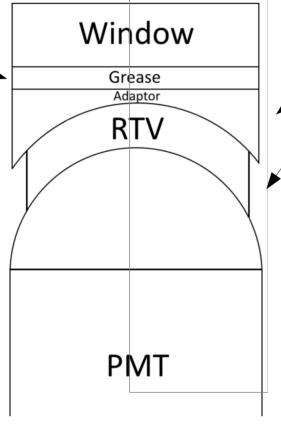




Original SHMS HGC PMT Optical Config

Removed from PMT1, 2

in Nov 2017



 These replaced with a stand-off ring in PMT1 in Spring 2018 run.

→ Same procedure applied to all 4 PMTs prior to Fall run

- RTV layer measured
 - → ~0.14 mm (PMT1)
 - \rightarrow ~0.06 mm (PMT2)
 - → Consistent with earlier finding that removal did not improve performance

SHMS HGC (Heavy Gas Cerenkov)

- SHMS HGC had lower npe then predicted during Fall 2017, Spring 2018 runs
 - \rightarrow 13–15 pe vs >25 pe at 1 atm C₄F₈O
- Low yield theories investigated and eliminated in last year
 - → Poor mirror reflectivity
 - → Poor mirror focus
 - → PMT optical coupling absorbing UV?
 - » Optical grease between PMT adapter and gas window?
 - » RTV joint between PMT and PMT adapter?
 - » PMT adapter not quartz?

- Last things to check...
 - → Directly measure PMT QE in the UV range
 - » Will also measure UV absorption of quartz gas window and quartz PMT adapter while we're at it...
 - → Completed in July 2018

* Ref: <u>Update on Dec/Jan Running</u>





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- Last things to check...
 - → Directly measure PMT QE in the UV range
 - » Will also measure UV absorption of quartz gas window and quartz PMT adapter while we're at it...
 - → Completed in July 2018
 - » No problems found, everything was within spec...

* Ref: <u>Update on Dec/Jan Running</u>





Conclusions and Follow-up Prior to Fall 2018

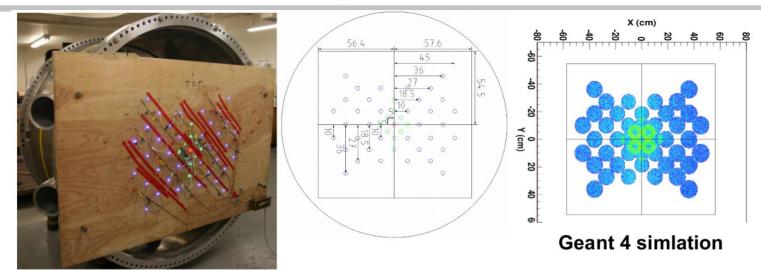
- SHMS HGC PMTs really are UV glass
- Quartz gas windows and quartz PMT adapters are transparent in UV
- RTV layer cuts enough that we should remove it, but is NOT a major problem
 - → Before and After npe response of PMT1 in production showed no/minimal improvement after removing RTV layer...

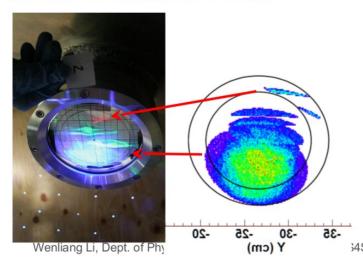
- Steps completed prior to Fall Run
 - → Removed RTV, quartz adapter, etc from upper PMTs
 - → Re-insert bottom PMTs
 - → Reinstall HGC in stack
- Regina group checked mirror alignment and made some small adjustments
 - → Hope was to improve collection efficiency overall with some minor impact on collection associated with central tracks





SHMS HGC Alignment Apparatus





- LED alignment Array (Xmas tree)
 - Colored LEDs: Red, Blue and Green
 - Replicate photon envelope
 - Compared to MC
- Significantly helped us aligning the detector

34S0A2, Canada.

https://hallcweb.jlab.org/doc-private/ShowDocument?docid=974

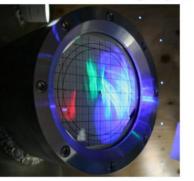




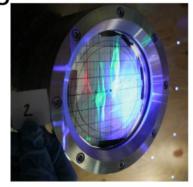
Before and After Fall 2018 Alignment

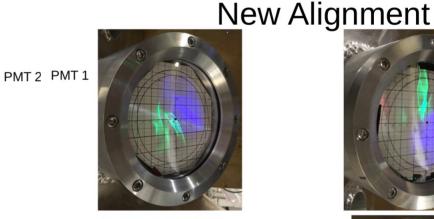
Goal was to ensure light from greater majority of HGC acceptance was guaranteed to hit PMT. This required shifting the single, central, red LED response toward the PMT edge.

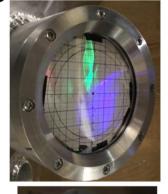
Old Alignment

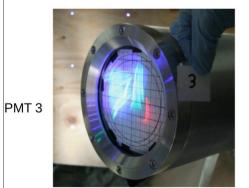


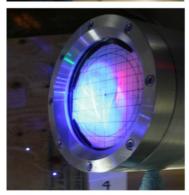
PMT 1

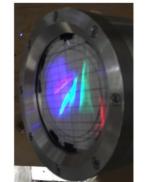


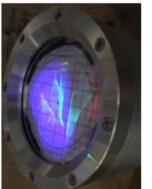












PMT 4

PMT 2

https://hallcweb.jlab.org/doc-private/ShowDocument?docid=974

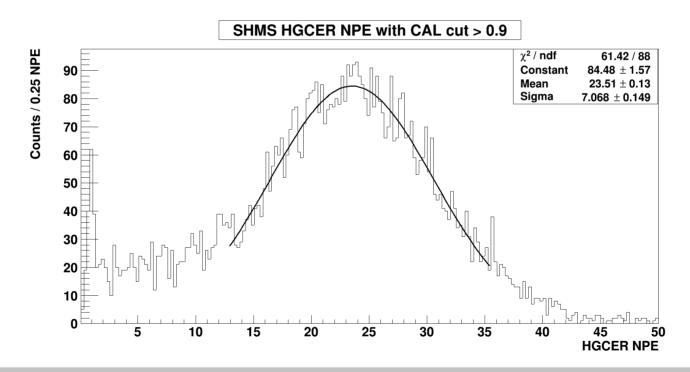
PMT 4 PMT 3



SHMS HGC Performance in Fall 2018

https://logbooks.jlab.org/entry/3599771

- New results puzzle me...
 - \rightarrow Early conclusion was big improvement of 15 \rightarrow 25 npe for electrons!
 - → Plot from Sept 29 analysis by Regina group







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SHMS HGC Performance in Fall 2018

https://logbooks.jlab.org/entry/3599771

- New results puzzle me...
 - \rightarrow Early conclusion was big improvement of 15 \rightarrow 25 npe for electrons!
 - → This would be great, but...
- All changes in optical path were tested in-situ during Spring run and found to have minimal impact (+1–2 npe, max)
 - → Removal of optical grease
 - → Removal of RTV layer
 - → Removal of quartz adapter
- Which leaves just the adjustment in mirror focus done prior to Fall 2018 run



SHMS HGC Performance in Fall 2018

https://logbooks.jlab.org/entry/3599771

- New results puzzle me...
 - \rightarrow Early conclusion was big improvement of 15 \rightarrow 25 npe for electrons!
 - → This would be great, but...
- The change in mirror focus was fairly minor...
 - → Studies were done in Spring placing cuts to constrain electron tracks to pass through selected positions on mirror(s), could never get more than ~15 npe
 - → LED 'spot' images moved few cm on PMT face in Fall 2018 refocusing effort
 - » I don't understand how both observations can both be true...
 - Cerenkov ring diam is significantly smaller than the mirror dimensions
 - There had to be some region of a mirror that focused 100% of its photons on its PMT in the Spring, why were we unable to find it and see a high npe yield?
 - If no region of any mirror could focus all rays onto a PMT, then a cm level shift of the mirror focal point on the PMT face shouldn't make a huge difference...





Mirror focus "by eye"

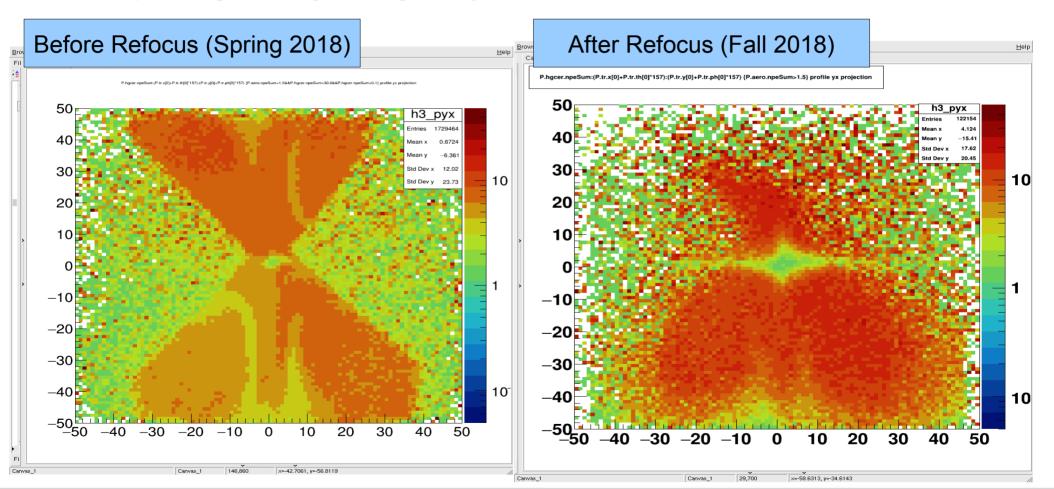


- Grid pattern seen in mirrors is reflection of gridded paper target on PMT surfaces
 - → No obvious dead zones
 - → Most of the light reflecting off the mirror seems to intersect with the PMT surface.
- Not sure how shifting mirror focus could generate 2x the npe yield overall without us being able to see the same maximal yields by placing cuts on mirror regions.



Hole in SHMS HGC Acceptance Grew

https://logbooks.jlab.org/entry/3608378

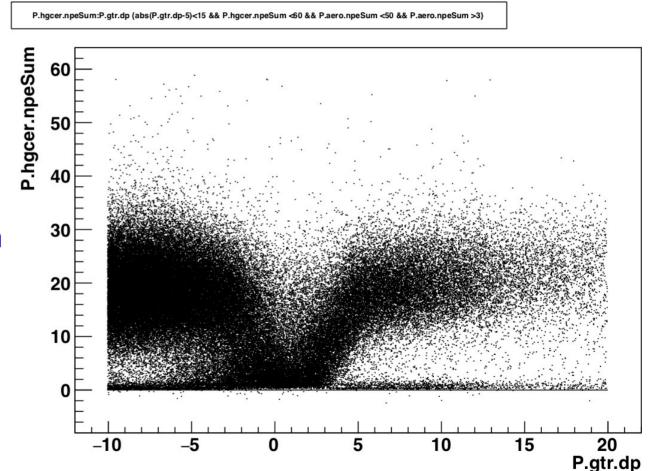






Hole in SHMS HGC Acceptance Grew

- Also seen in larger 'dip' in npe yield in the central delta region
 - → Will clearly impact PID in that region.
- Unfortunately this does seem to be 'real' and not an artifact of cuts, kinematics, etc...



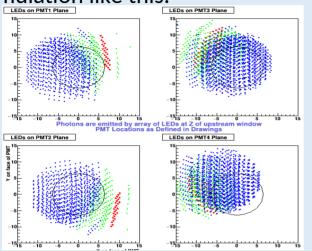
(Plot by D. Gaskell)

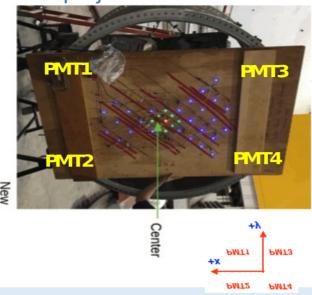


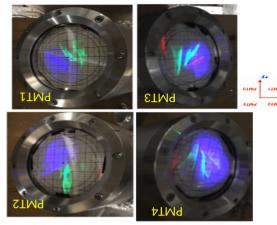
SHMS Heavy-Gas Cerenkov

Compare colored LED patterns as projected onto PMTs

With the ideal mirrors 'aligned' in the simulation like this:

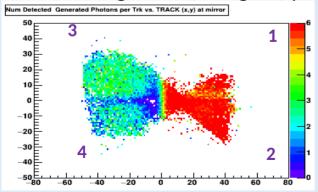




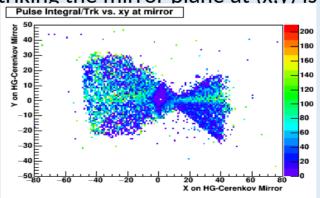


Note: PMT layout above matches the scatterplot layout, not the physical arrangement of PMTs.

when I generate 6 Cerenkov photons along each 'real' track path, the average NPE (or signal) per track striking the mirror plane at (x.v) is



SIMULATION: #photons/track



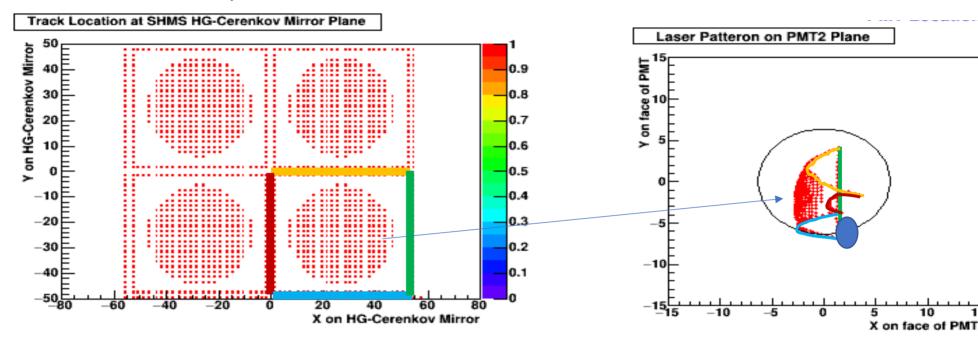
can simulate loss of signal along cracks, but not shadows or enhancements.

Fenker 29 Nov 2018

DATA: goodAdcPulseInt/track https://hallcweb.jlab.org/doc-private/ShowDocument?docid=997

Fenker 29 Nov 2018

Laser lines map onto PMT faces



Mapping of light rays from mirror to PMT face. Rays originate 10m in front of mirror plane. Colored lines along mirror perimeter above map onto similarly colored lines on face of PMT. PMT2 is used in this example. The blue oval approximates the projection of its photocathode onto the mirror plane.

This study could be used to attempt to align mirrors/pmts just as was done in the HMS.

https://hallcweb.ilab.org/doc-private/ShowDocument?docid=997

We would need to open the HGC outside the SHMS and set up the laser beam to see the pattern on PMT faces.

Near Term and Future Plans...

- Restoring the 'old' focus is relatively straight forward, BUT will require <u>at least</u> 5 days
 - → Remove HGC from Stack
 - → Pull (heavy) gas windows
 - → Mount alignment 'xmas tree'
 - → Match the old alignment images (documentation is great, and helps a lot here)
 - → Button up
 - → Leak check tank
 - → If all good, reinstall in Stack
- Next window for this: January?

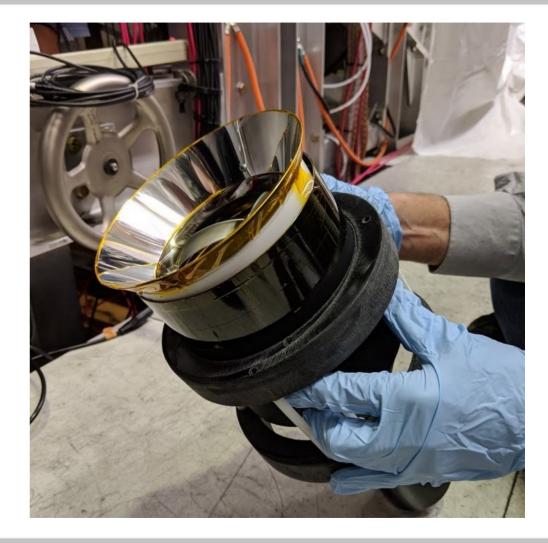
- "Hail Mary" fix in November
 - → Regina group and I installed some 'Winston Cones' between PMT and gas window
 - » Installation time ~ 1 shift
 - » HGC stays in stack
 - » Gas/vacuum seals not touched
 - » Increase 'effective' PMT diameter 5" → 6"
 - » Hopefully catch photons being lost at perimeter?





SHMS HGC "Winston Cone" Patch

- "Hail Mary" fix in November
 - → Regina group and I installed some 'Winston Cones' between PMT and gas window (all 4 PMTs)
 - » HGC stays in stack
 - » Gas/vacuum seals not touched
 - » Increase 'effective' PMT diameter 5" → 6"
 - » Hopefully catch photons being lost at perimeter of PMT?





SHMS HGC "hole" still present...

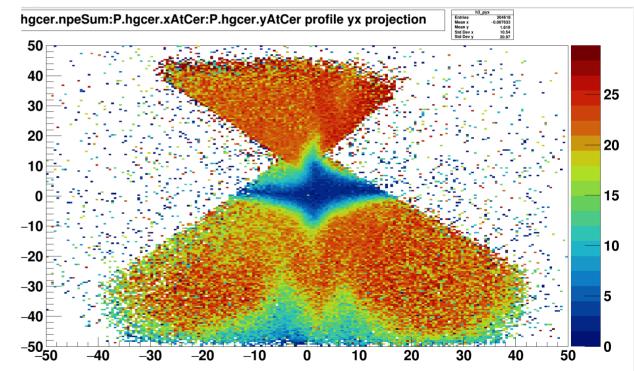
 Unfortunately no (significant) improvement

> → Direct comparison is tricky (different kinematics)

→ No worse than before, at least...

- Need to proceed with "Plan A"
 - → Alignment check
 - → Alignment restoration

R. Ambrose (U of Regina) https://logbooks.jlab.org/entry/3633305



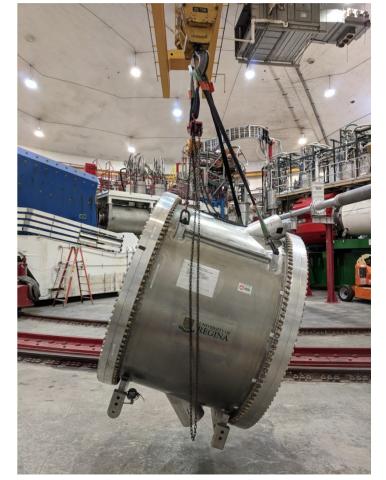


SHMS HGC Future Plans...

Restoring the 'old' focus is relatively straight forward, BUT will require

at least 5 days

- → Remove HGC from Stack
- → Pull (heavy) gas windows
- → Mount alignment 'xmas tree'
 - » Match the old alignment images
- → Do "Laser Check" based on Howard's sim (worked for HMS)
- → Button up
- → Leak check tank
- → If all good, reinstall in detector stack
 - » refill with C₄F₈O
- Next window for this: January?





SHMS / HMS Summary as of December 2018

Generally Detectors/DAQ are working as planned

- → Calibrations, cross-checks continue
 - » See upcoming Hall C collaboration meeting in January
- → HMS Cerenkov fully restored

SHMS Gas Cerenkov are not performing to design yet...

- → SHMS Noble Gas Cerenkov (NGC) yields are still a question mark...
 - » NGC is idle on the Hall C floor (noisy PMT will be replaced Summer 2019)
 - » NGC yield *may* be concern at high momentum settings (> 7 GeV/c)?
 - May want to be conservative with PID expectations at such settings?
- → SHMS Heavy Gas Cerenkov (HGC) yields are also an issue...
 - » Overall npe yields are not well understood
 - Better now than Spring 2018 (but I don't understand why)
 - "Hole" in acceptance needs to be addressed
 - Next opportunity is Jan 2019 (check alignment, realign)
 - » Verify models, improve understanding, and move forward from there...





Questions?





BACKUP SLIDES





"Band" structure in SHMS HGC npe yields vs position



"Banding" in SHMS HGC npe Yields

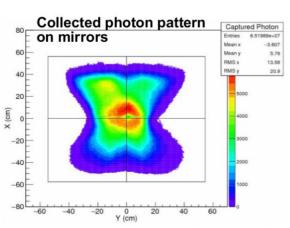
Also seen in Simulation

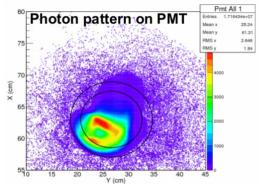
Optics Simulations - 7 GeV/c π @ 0.95 atm

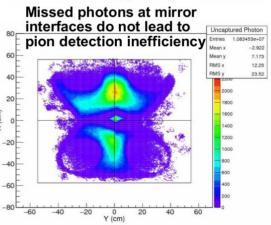
Cherenkov cone angle largest at 7 GeV/c $(2.84^{\circ}) \rightarrow$ light hardest to focus.

GEANT4 Simulation includes:

- -C₄F₁₀ properties vs wavelength
- -Measured 4 mirror oblateness
- -Measured AL reflectivity vs wavelength
- -Quartz adapter losses vs wavelength
- -PMT efficiency vs wavelength







Details in JLAB-PHY-12-1697 (M.Sc. thesis)



Garth Huber, huberg@uregina.c

"Banding" in SHMS HGC npe Yields

Data (Spring 2018)

NPE from Particle in Focal Plane profile xy projection

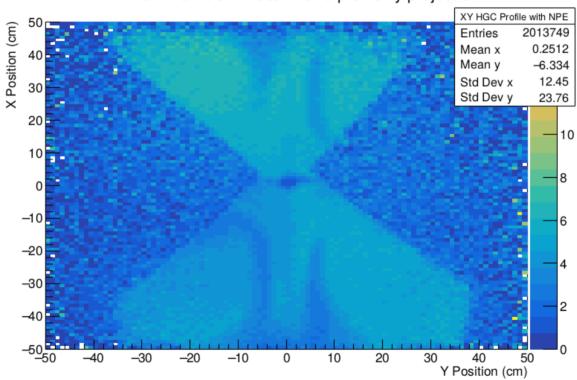


Figure 6: Number of photoelectrons received in the HGC across the mirror plane. Color axis refers to the number of photoelelectrons. Run 1583.

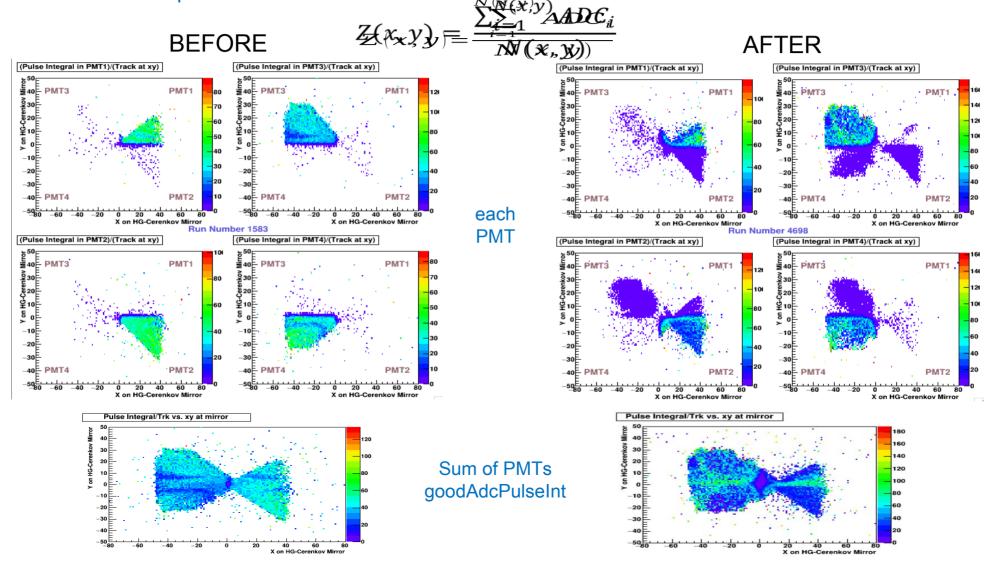




Crosstalk????

Plot goodAdcPulseInt from each PMT vs. (x,y) of TRACK at mirror plane...

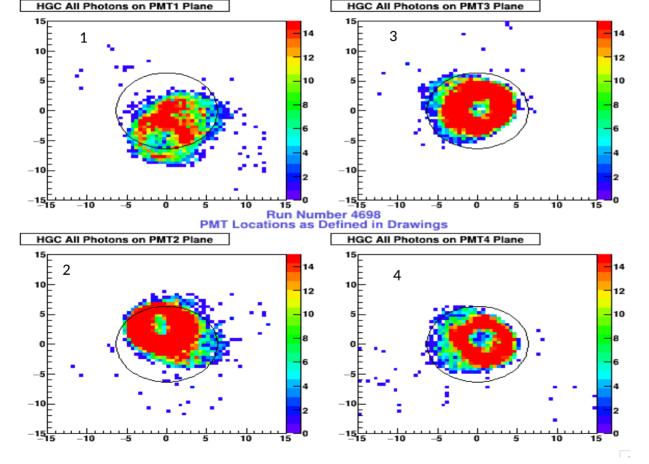
... and see more crosstalk <u>after new alignment</u>. I assume it is optical, but maybe it is electronic.



SHMS Heavy-Gas Cerenkov

- Determine 'Image' size of Physics Tracks on PMT face:
 - For each reconstructed particle track (real data) radiate 6 Cerenkov photons along length of each track through Cerenkov gas. Propagate them to the PMTs.

Resulting image size on PMT face is consistent with the 5-inch PMT



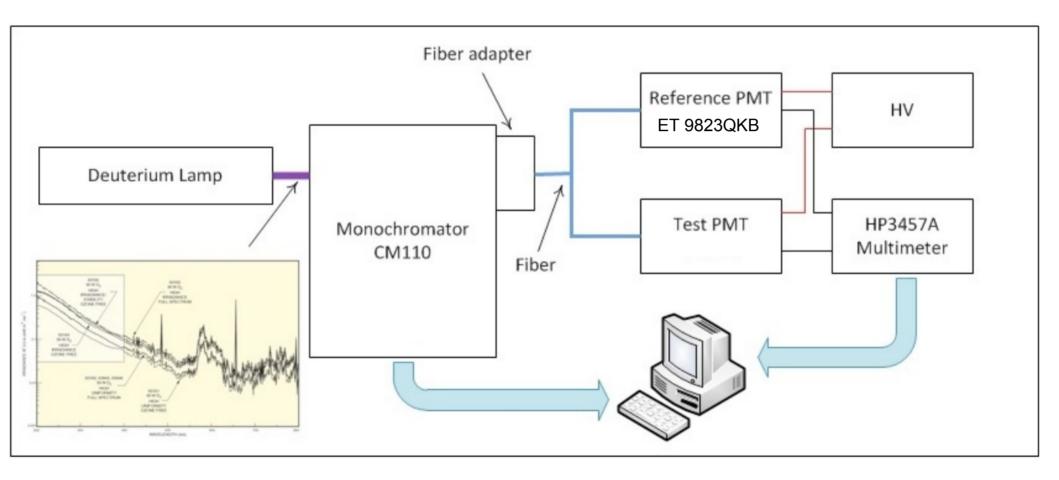
Note: mirror shapes are taken as perfect spherical sections. Alignment is arbitrary in this run.

Test PMTs at UVa





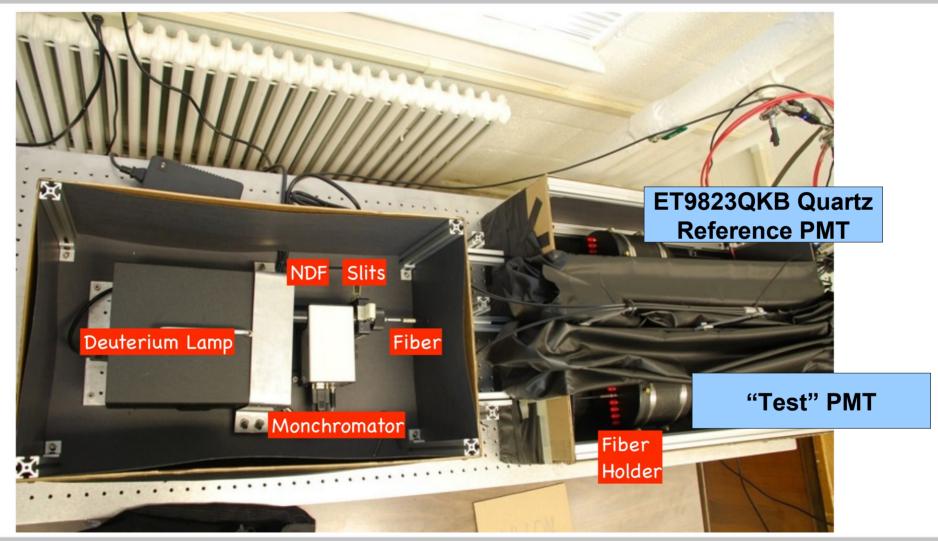
UVa PMT Test Stand







UVa PMT Test Stand



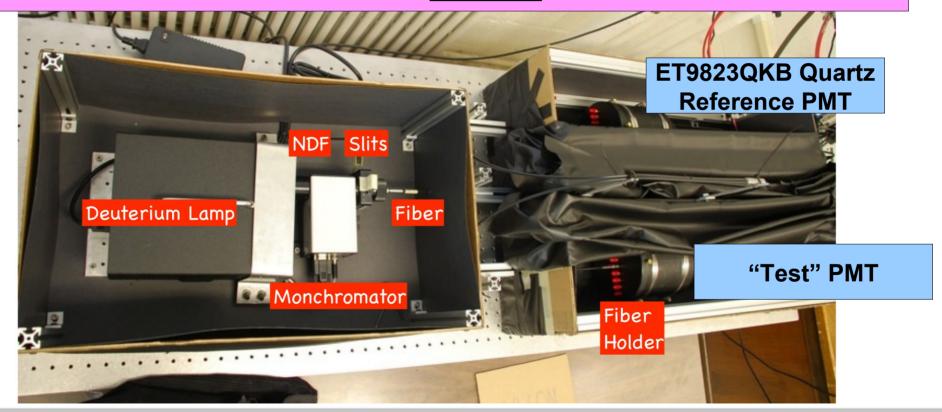




UVa PMT Test Stand

Normalization required for absolute QE measurement - Difficult, not needed, so we did not do it.

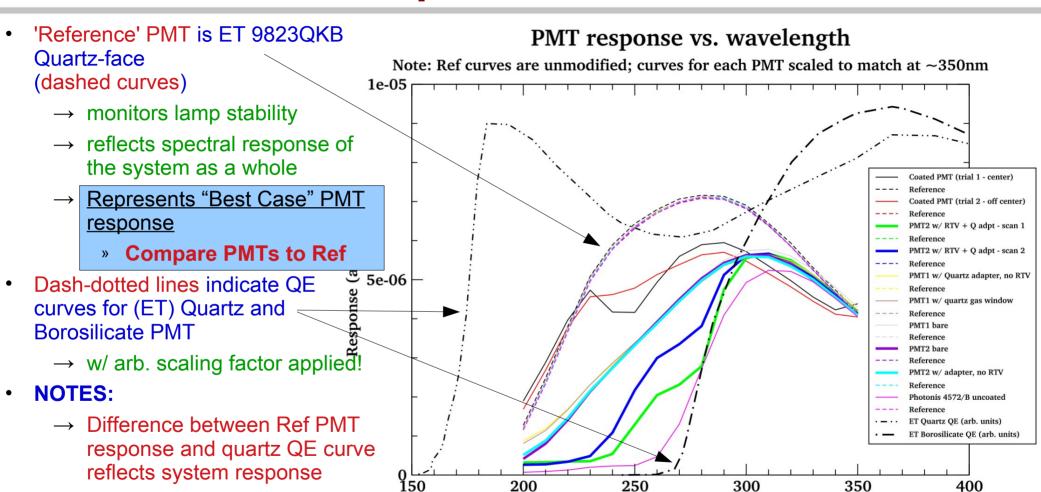
Measurements are relative not absolute







Interpretation Notes





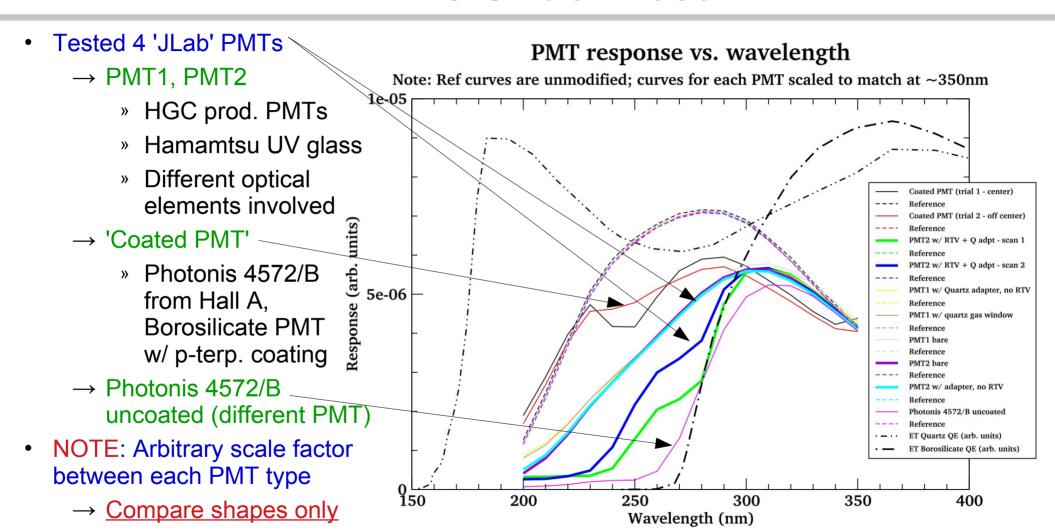
Lamp intensity, absorption

in fibers, etc



Wavelength (nm)

PMTs Under Test





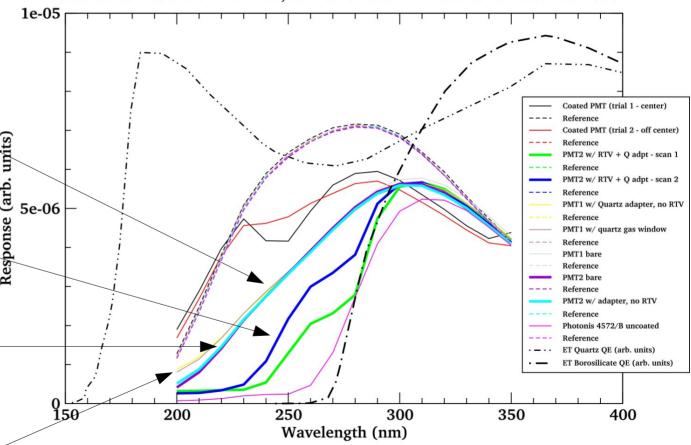


'Production' PMT responses

- PMT1 (light lines)
 - → Prod. PMT that had all 'interface' parts removed
 - » no adapter, no RTV, no opt. grease
 - → As expected
- PMT2 (heavy lines)
 - → Production PMT
 - » w/ adapter + RTV
 - degraded
 - » adapter + RTV layer (0.06 mm) removed
 - same responseas PMT1
- Quartz glass PMT adapter and gas window tested
 - → transparent (no impact)

PMT response vs. wavelength

Note: Ref curves are unmodified; curves for each PMT scaled to match at ~350nm







SHMS HGC Optical Configuration





PMT with Quartz Adapter Attached



Side view of quartz adapter on PMT 1 (now removed)

Note: Yellow color is due to kapton tape around perimeter of quartz adapter, NOT radiation damage.





SHMS HGC PMT 'Ring'





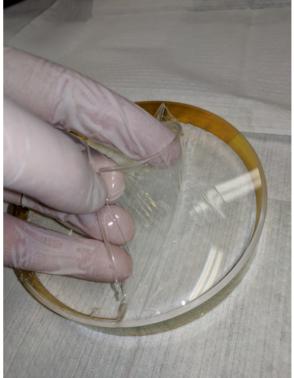
PMT2 Quartz Adapter Removed

 Quartz adapter on PMT2 removed at UVa

PMT2 RTV layer thickness

 \rightarrow 0.06 mm









Both SHMS and HMS running since Fall 2017

Hall C production running began in January 2018

Detectors/DAQ (generally) performing well

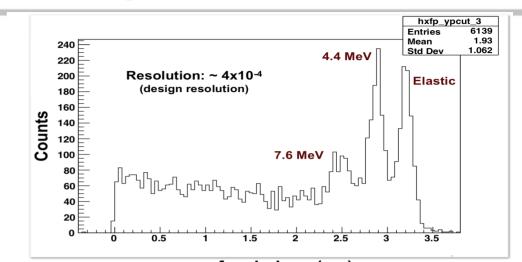
- Hodoscopes < 98% efficient (4/4 planes)
- Wire chambers < 99% efficient
- Good tracking, Good energy resolution Analysis, Calibrations underway!

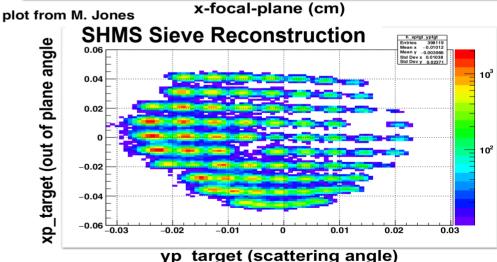
But...

- SHMS HGC has lower npe yield than predicted (13 pe vs >25 pe), investigate over Summer 2018
- SHMS NGC yield is lower than predicted, may want to be conservative with PID performance at high momenta (ie. large neon fraction)?
- HMS HGC mirrors found to be damaged
 - Spares available(!), repair underway now

Plots on right from Commissioning data taken last December, January

- •See also: <u>Update on Dec/Jan Running</u> (D. Dutta, Jan 2017 Hall C Collab. Meeting)
- •See also: Upcoming Hall A/C Meeting / Analysis Workshop!



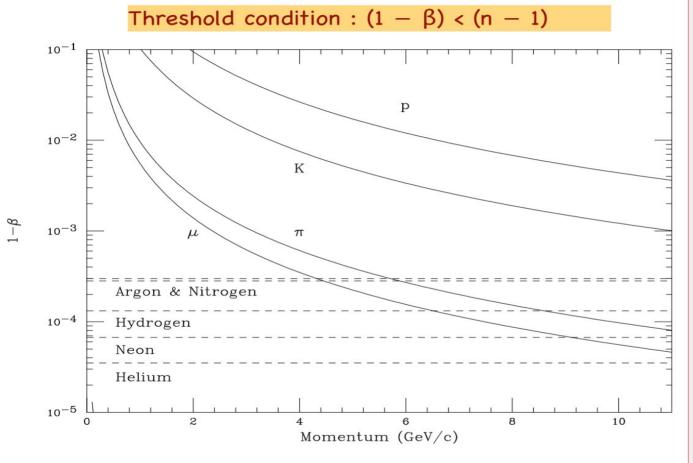


using COSY matrix elements without any optimization





SHMS NGC (Noble Gas Cerenkov)



Concern:

As neon fraction rises, the npe yield drops. At 100% neon, it is down by a factor of ~3–4 from 100% Argon.

Simulation:

~40 p.e. (100% Ar)

~ 9 p.e. (100% Ne)

[\sim 55 p.e. (CO₂) estimated]

Data:

~ 17 (100% Ar)

~ 25 (100% CO₂)

 \rightarrow ~ 3.8 p.e. (Ne)

NOTE: VERY PRELIMINARY!

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HClog 3519060

Slide from D. Day (Hall C ERR)

Choice of gases

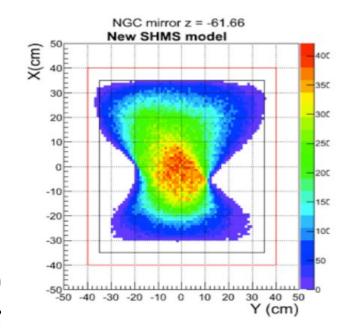
Argon up to 6 GeV and a mixture of Argon and Neon up to 11 GeV





SHMS PID Requirements: negative polarity

	•				
Experiment	P (GeV/c)	Req'd e⁻:π⁻ Disc.	Spec'd NG Cerenkov	Spec'd Calorimeter	Total Expected
E12-06-101 (Fpi-3)	2.2 - 8.1	4.5•10³:1			
E12-06-104 (σ_{L}/σ_{T})	5.4 - 5.8	1O ³ :1	50:1 (HMS Cerenkov gives up to 300:1 now)	>200:1 (1000:1 above 6 GeV/c)	
E12-07-103 (pion factorization) (d)	2.4 - 8.5	10³:1			
E12-06-105 (x>1)	4.8 -10.6	5•10³:1			>104:1
E12-06-110 (c)	2.2 - 6.8	10³:1			
E12-06-121 (g ₂ ^{n,} d ₂ ⁿ)	6.3 - 7.5	>10 ² :1	Joon Howy		



4 overlapping spherical mirrors

R = 135 cm, 43 by 43 cm

2 m of active length

Noble gas at 1 Atm

Slide from D. Day (Hall C ERR)