

Validation of neutrino energy estimation using electron scattering data

Student

Mariana Khachatryan

Supervisor

Lawrence Weinstein

Outline

- ❑ The importance of energy reconstruction in neutrino oscillation experiments
- ❑ Neutrino-nucleon Charged Current interactions
- ❑ Testing neutrino beam energy reconstruction methods with electron scattering CLAS e2a experiment data

Long Baseline Oscillations

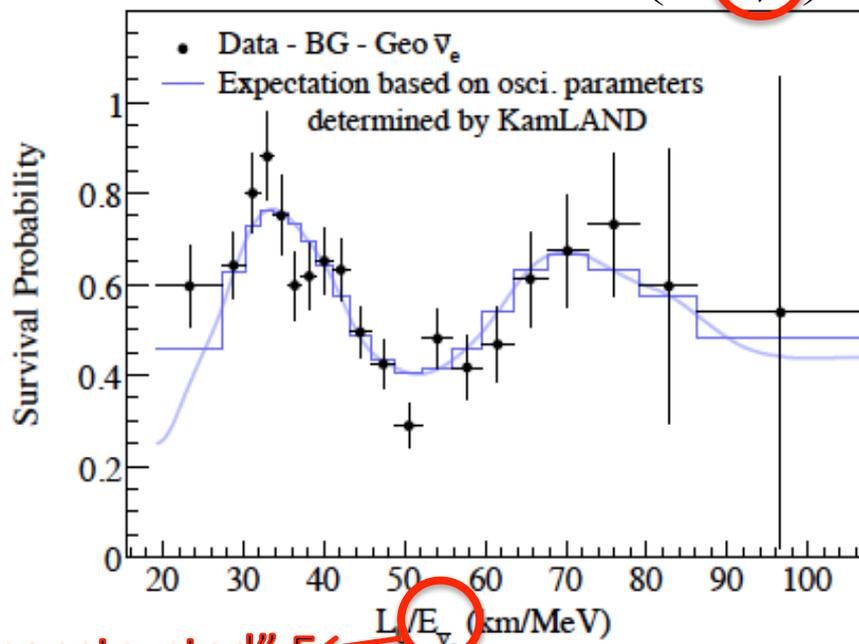


Types

1. Appearance experiments: Use neutrino beam of type A and search for neutrinos of type B after distance L

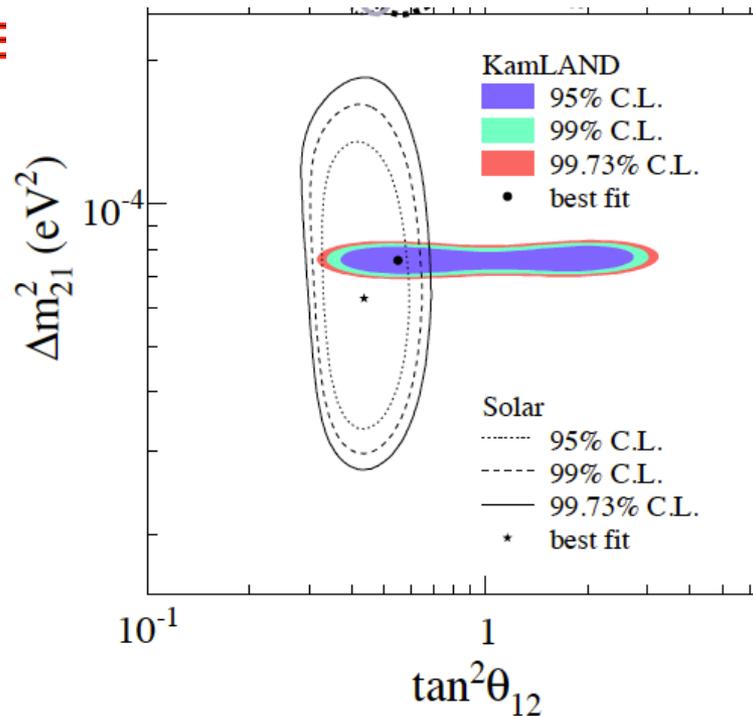
2. Disappearance experiments: Compare the fluxes of neutrinos of given type before and after oscillations

$$P(\nu_\mu \rightarrow \nu_x) = \sin^2(2\theta) \times \sin^2\left(\frac{\Delta m^2 L}{4E_\nu}\right) \rightarrow \text{Real E}$$



“Reconstructed” E

KamLAND, PRL 100, 221803 (2008)
(Reactor beam)

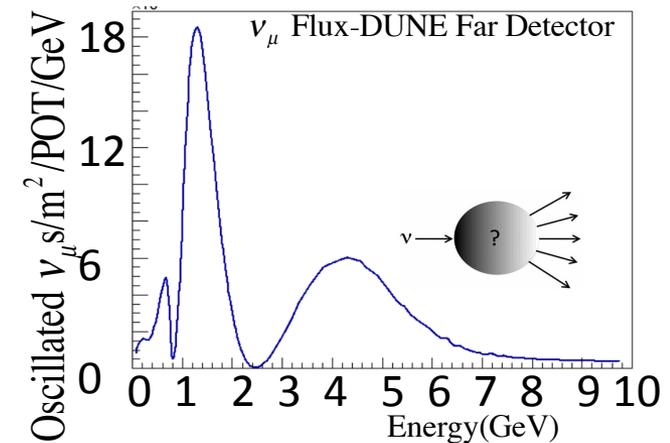
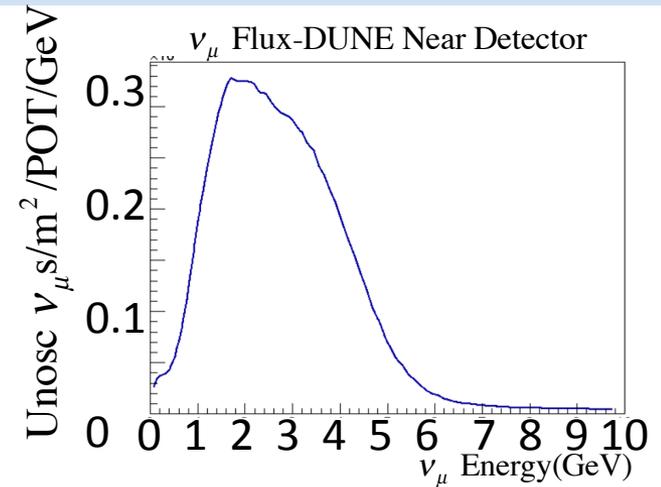


Allowed region for neutrino oscillation parameters from KamLAND and solar neutrino experiments

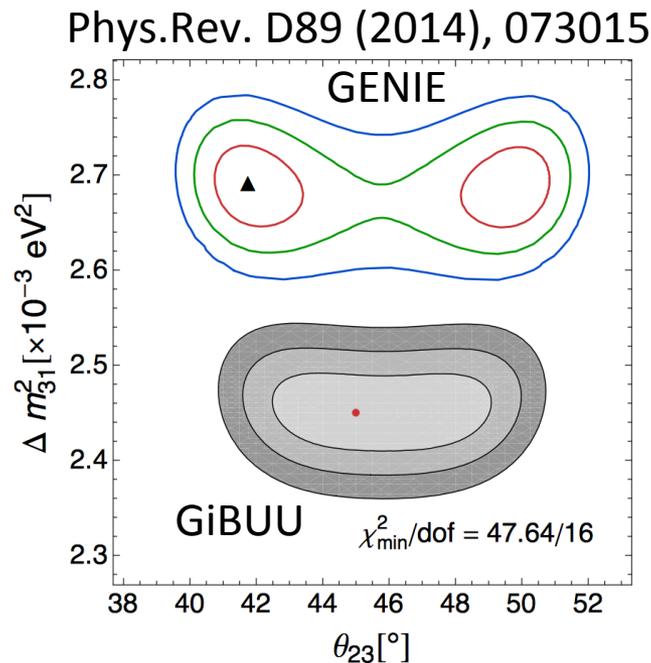
(Long Baseline) Oscillation Challenge

Oscillations are basically ratios of reconstructed ν energy spectra:

- Energy (x-axis): Reconstructed from the measured final state.
- Flux (y-axis): Corrected using reaction model



=> Incorrect neutrino-nucleus interaction modeling can bias the extracted oscillation parameters



Testing energy reconstruction methods in neutrino experiments with e- scattering data

$\nu - N$ and $e^- - N$ scattering
Weak Interactions and **Electromagnetic Interaction**

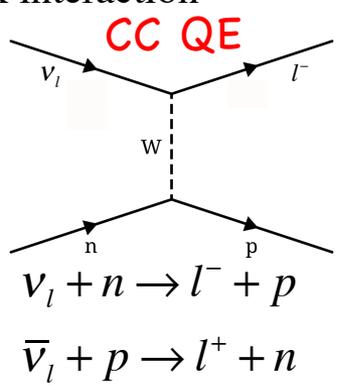
CC Weak Interactions

NC Weak Interactions

Charged Current (CC) Weak interaction
 mediated by W^\pm bosons

$$j_\mu^\pm = \bar{u} \frac{-ig_W}{2\sqrt{2}} (\gamma^\mu - \gamma^\mu \gamma^5) u$$

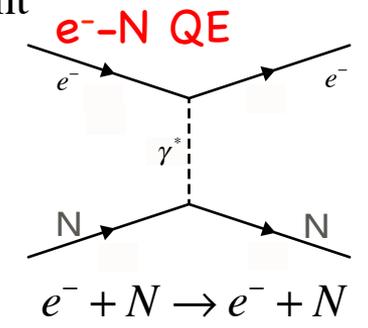
g_W - coupling strength



$e^- - N$ scattering
Electromagnetic Interaction

Electromagnetic current

$$j_\mu^{em} = \bar{u} \gamma_\mu u$$

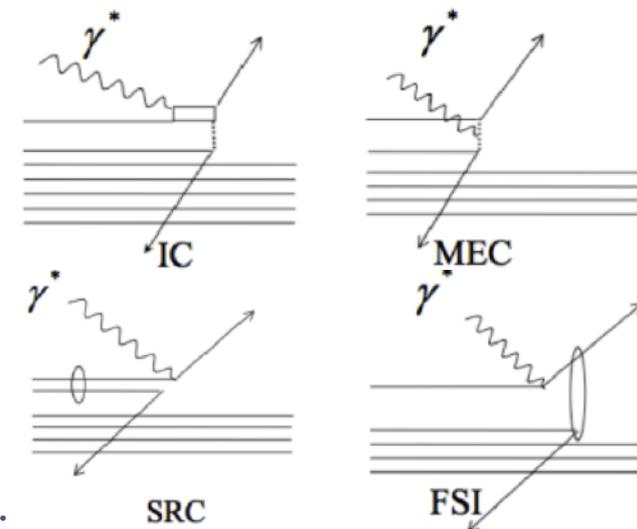


What can we learn from e- scattering studies?

e^- and neutrino interactions with matter have many similarities
 e^- beam energy is known \rightarrow can test energy reconstruction in selective kinematics

Goal:

- Analyze electron scattering data to study neutrino beam energy reconstruction methods for different energies and nuclei.
- Study nuclear responses (FSI, Resonance production, Multinucleon effects, etc.)
- Compare to Genie results and identify regions of phase space where simulation and data agree well



e^- -nucleus two body diagrams that lead to the same final state as that of e^- -N QE

Energy reconstruction methods in neutrino experiments

E_ν Reconstruction from lepton kinematics

$[(e, e') \text{ or } (\nu, l)]$ (assumes QE)

$$E_\nu^{\text{kin}} = \frac{2M\varepsilon + 2ME_1 - m_l^2}{2(M - E_1 + |k_1| \cos \theta)}$$

$\varepsilon \approx 20$ MeV single nucleon separation energy

M-nucleon mass

m_l outgoing lepton mass

k_1 - lepton three momentum

θ - lepton scattering angle

Problem: assumes QE

Cherenkov detectors:

- Electrons & Pions
- No protons / neutrons

E_ν Reconstruction from 'full' final state

$[(e, e' pX) \text{ or } (\nu, lX)]$

$$E_{\text{Calorimetric}} = E_e' + \sum T_p + E_{\text{Binding}} + \sum E_\pi$$

E_{Binding} - Binding energy

T_p - kinetic energy of knock out proton

E_e' - energy of scattered electron

E_π - energy of produced meson

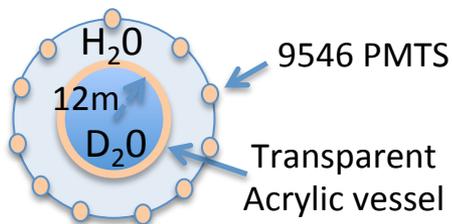
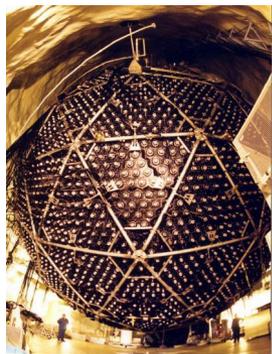
We ignore the kinetic energy of A-1 system.

Tracking detectors:

- Charged particles + π^0
- Neutron detection is challenging

SNO (Sudbury Neutrino Observatory, Canada, Ontario)

- 1000 ton heavy water D_2O and 3000 ton normal water
- Detect neutrinos via CCQE, NCQE and CC and NC e^- -neutrino elastic scattering

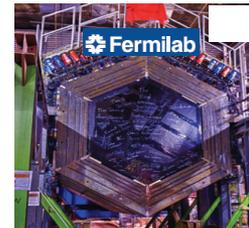


Water Cherenkov detector

Study: Solar neutrino problem

Art McDonald was co-awarded Nobel prize in 2015

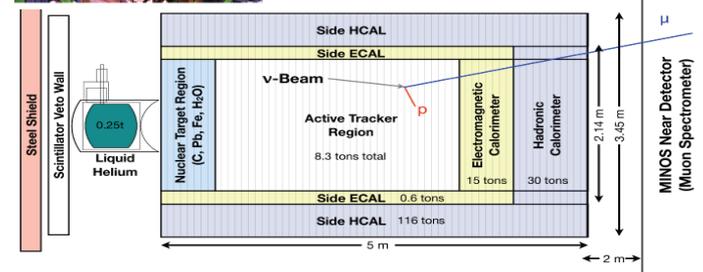
MINERvA (Fermilab)



Scintillator based detector

Study:

- Neutrino interaction cross sections
- nuclear effects



Electron scattering data

Have analyzed ^3He , ^4He , ^{12}C , ^{56}Fe 4.461, 2.261 GeV
e2a experiment data

Other data available ^3He , ^4He , C, Fe 1.1 GeV

E2a target properties

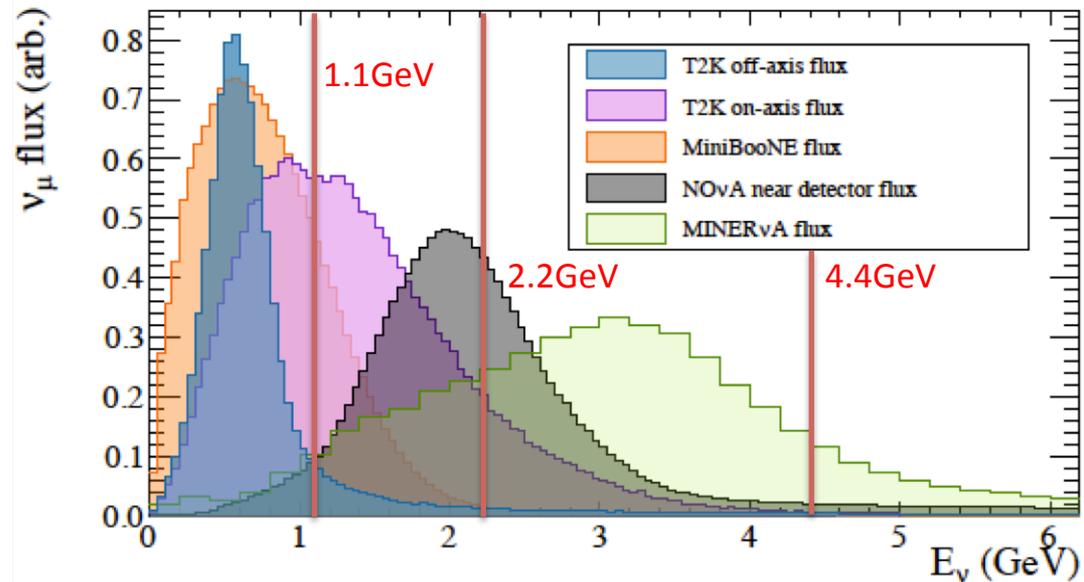
Target	length [cm]	Density * length [g/cm ²]
^3He	4 – 5	0.335
^4He	4 – 5	0.0625
^{12}C		0.221
^{56}Fe		0.118
CH_2		0.067

Good (e,e') and $(e,e'p)$ events $\times 10^6$

with e and p PID, vertex and fiducial cuts and $W < 2$

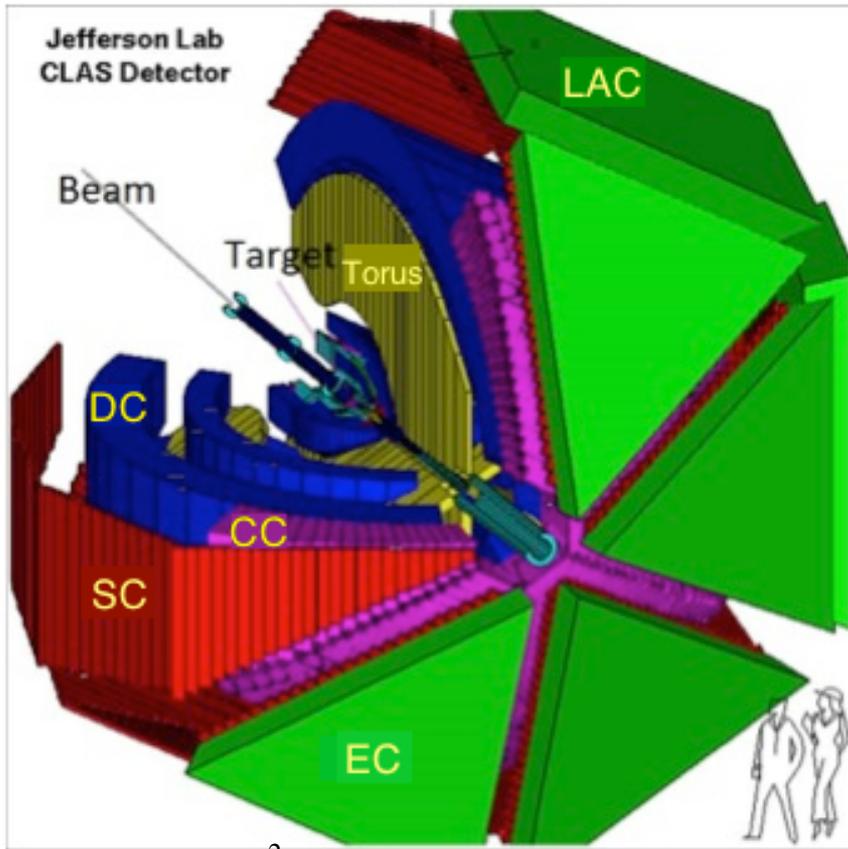
	2.2GeV (e,e')	2.2GeV $(e,e'p)$	4.4GeV (e,e')	4.4GeV $(e,e'p)$
^3He	29	12	3.9	1.4
^4He	46	17	8	2.6
^{12}C	29	11	5	1.5
^{56}Fe	1.5	0.5	0.4	0.1

Current neutrino

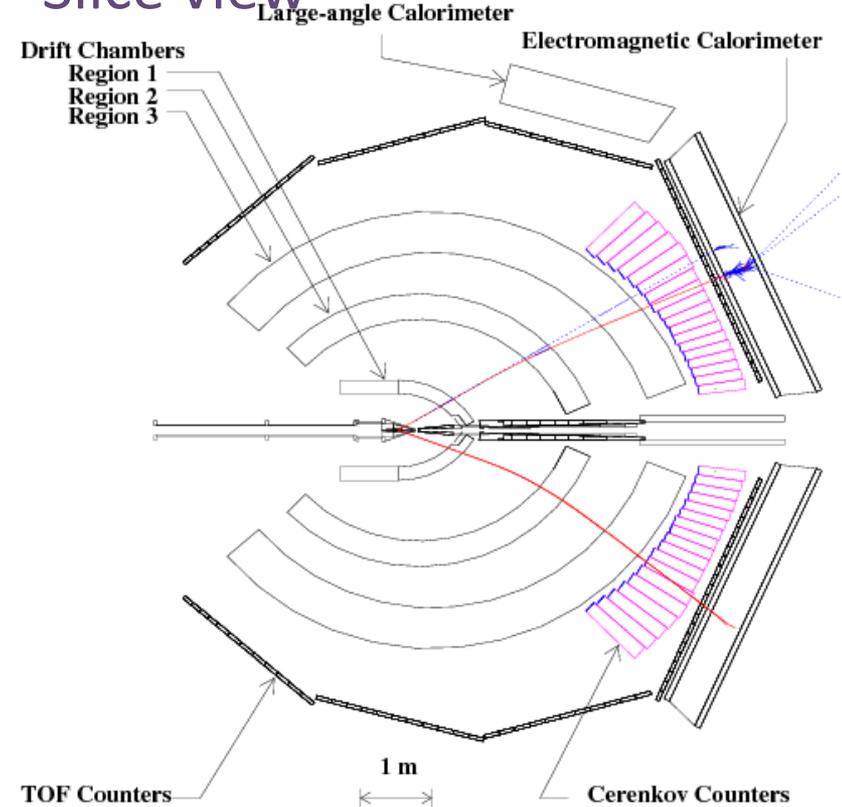


CLAS detector package

3D view



Slice view



$$\sigma \sim \left(\frac{1}{q^2 + M^2} \right)^2$$

$$M = m_{\text{photon}} \quad (\text{for } e^- \text{ scattering})$$

$$M = m_{W^\pm} \quad (\text{for CC weak interaction})$$

Scale the electron scattering data with $1/\sigma_{\text{Mott}}$ to have 'neutrino like' data!

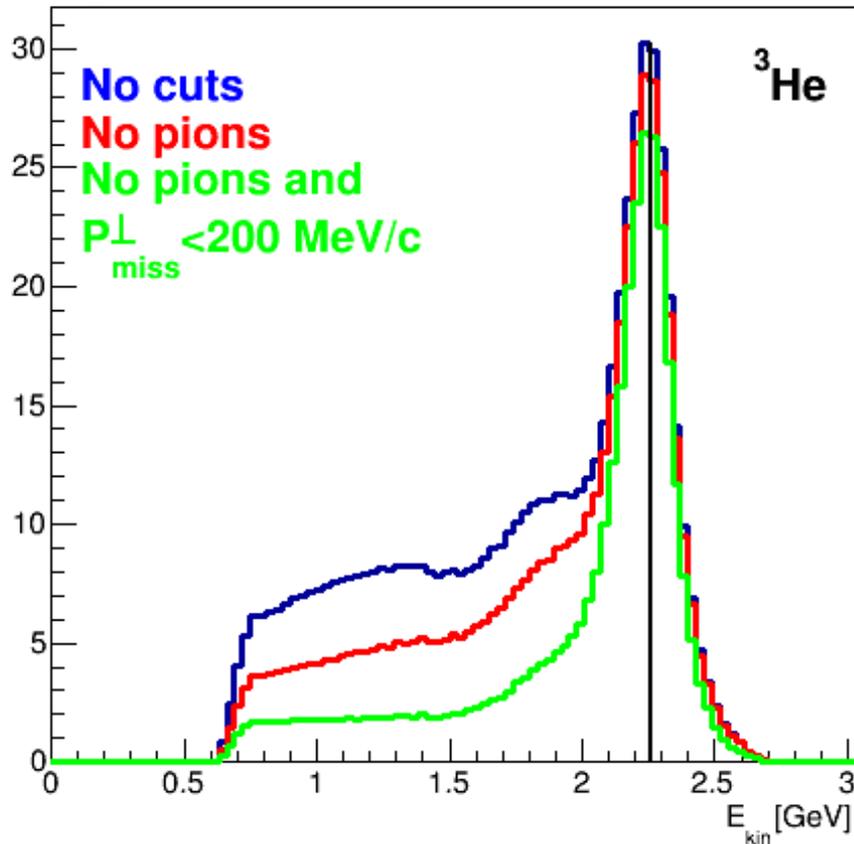
2.261 GeV analysis

Reconstructed (e,e') energy

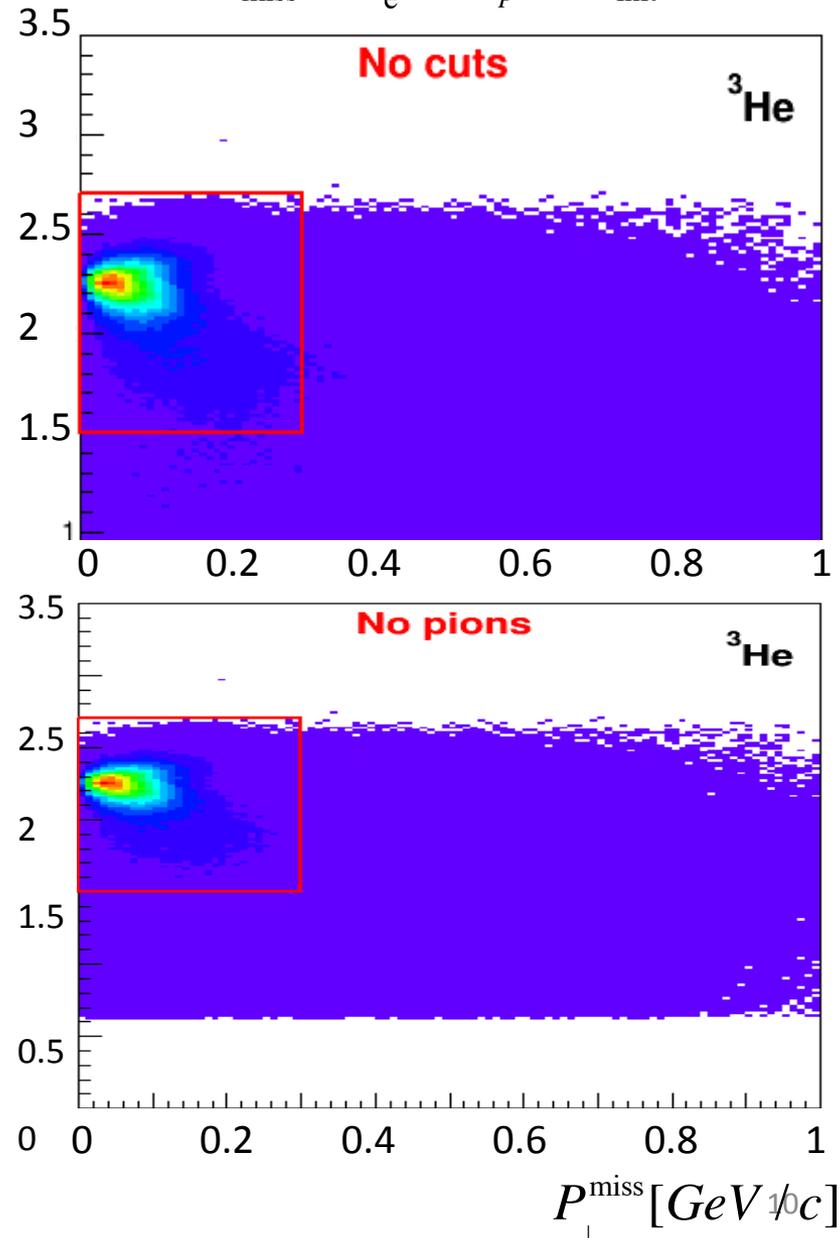
E2a ^3He 2.261 GeV

Standard CCQE event selection is 0π

e^- only

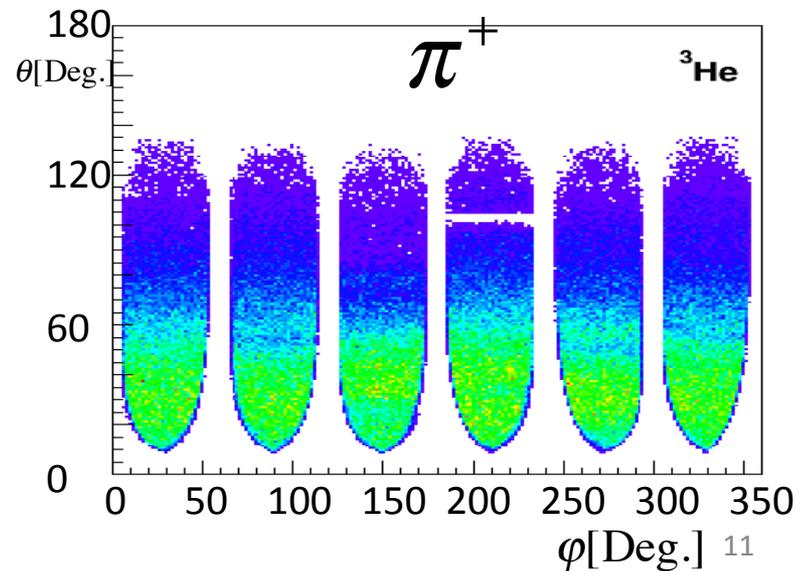
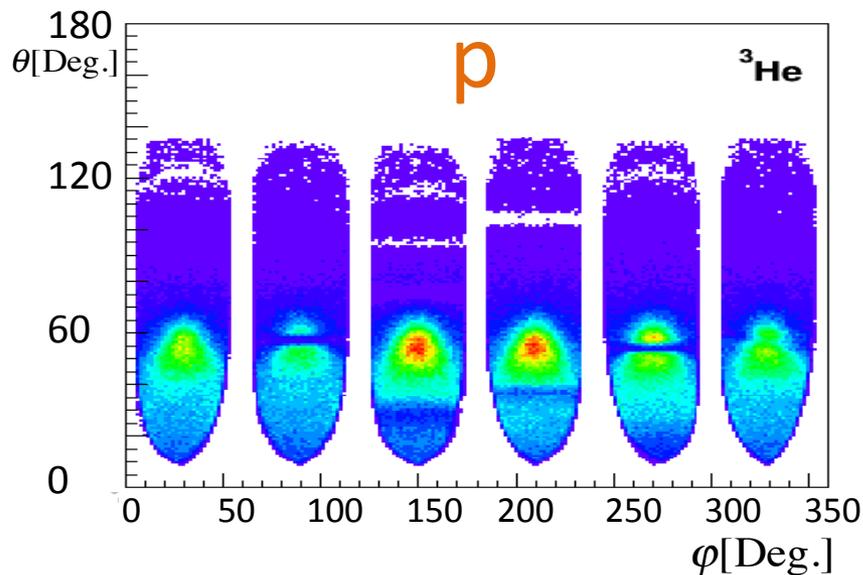
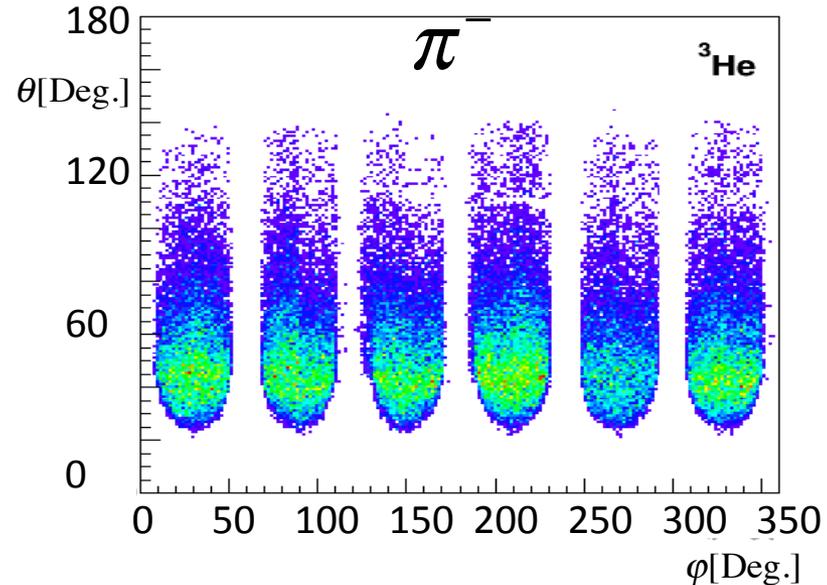
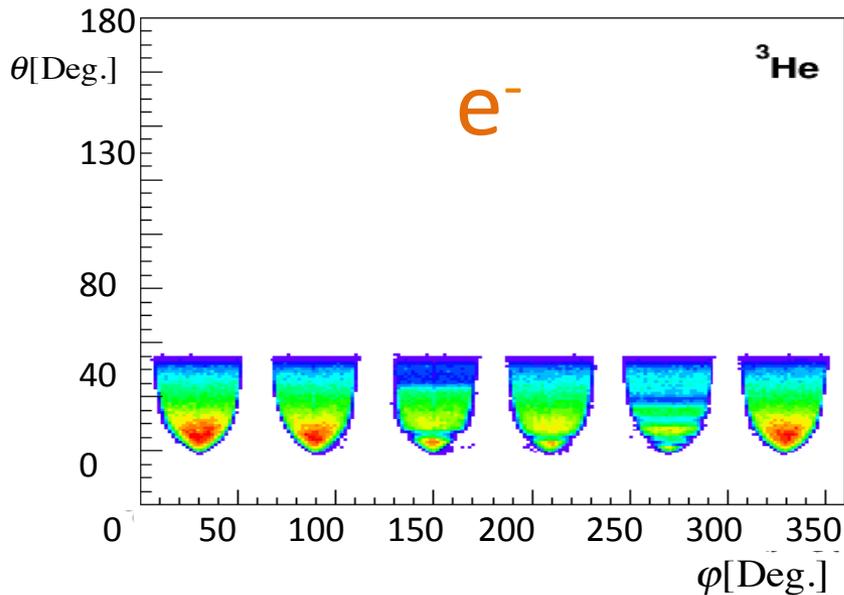


$$P_{\text{miss}}^{\perp} = P_{e^-}^{\perp} + P_p^{\perp} = P_{\text{init}}^{\perp}$$



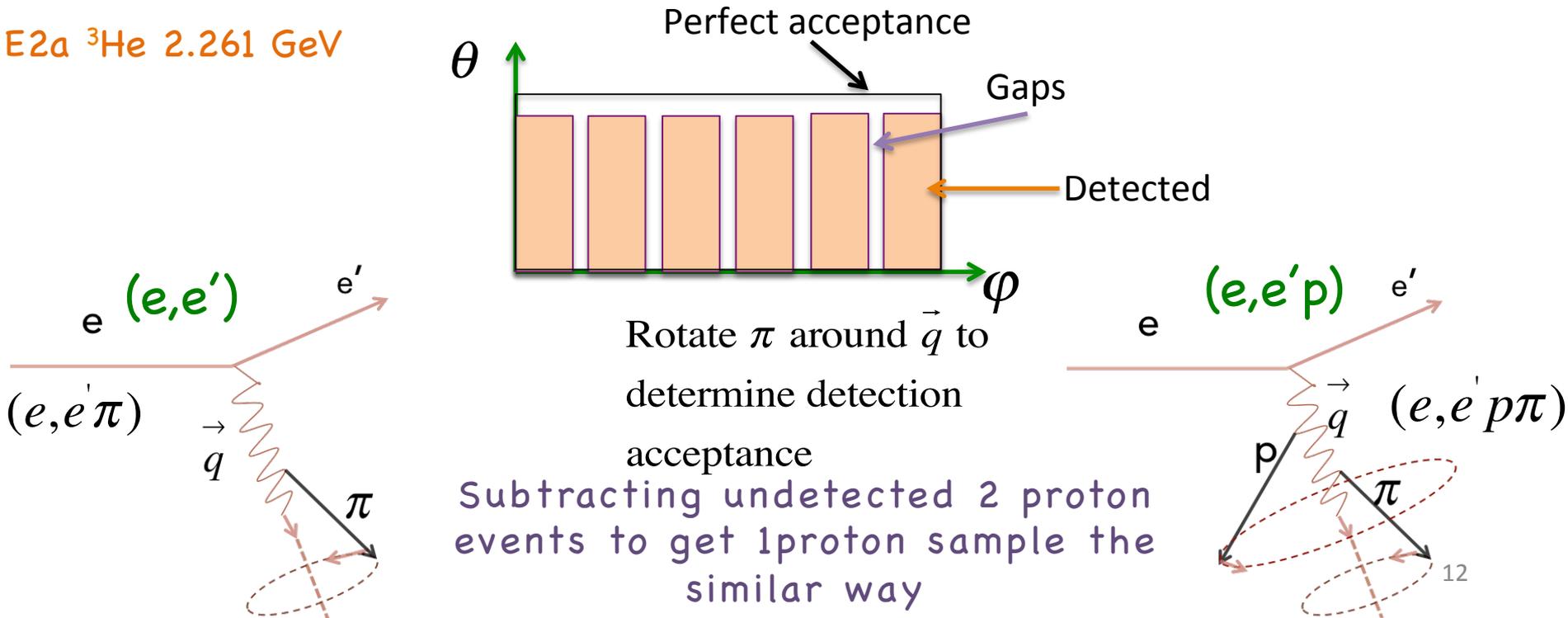
θ and φ distributions

E2a ${}^3\text{He}$ 2.261 GeV

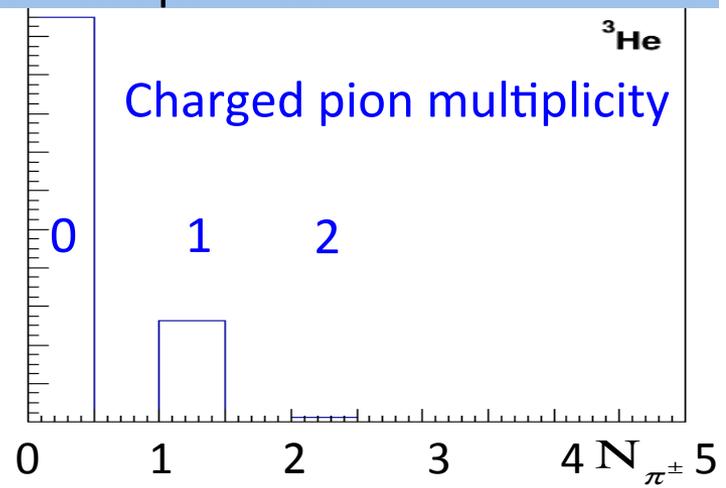
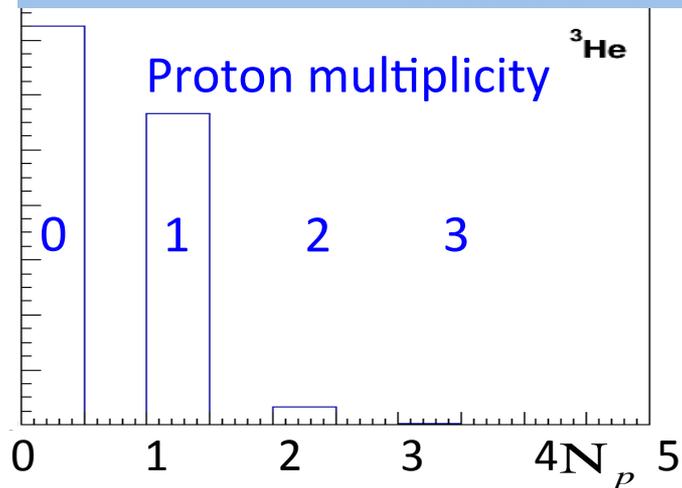


Subtracting undetected pions to get 0 pion sample

E2a ^3He 2.261 GeV



Number of events with pions and protons



Subtracting undetected pions

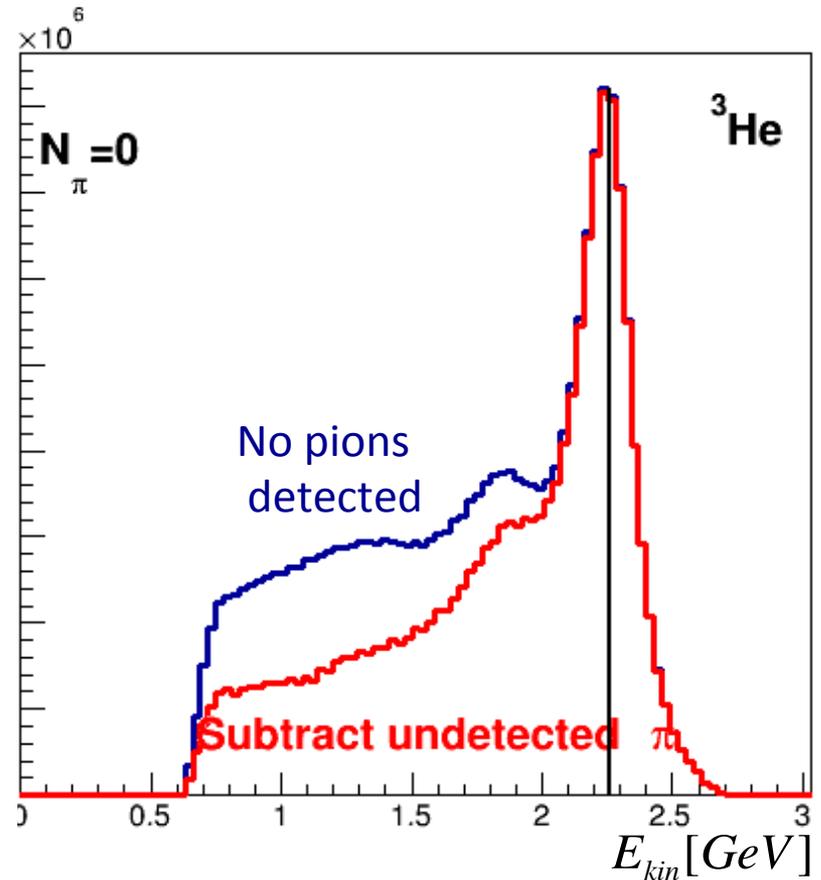
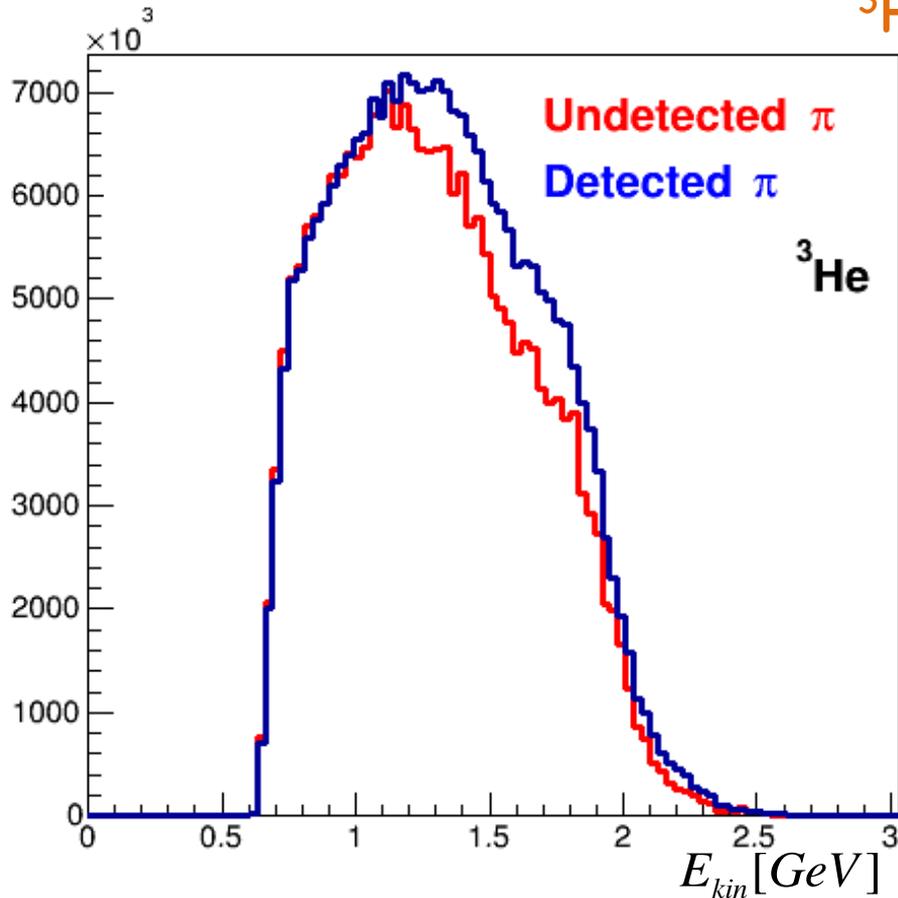
E2a ^3He 2.261 GeV

(e,e')

Cuts

No π_+ , π_- and no photons coming from π_0 decay

^3He

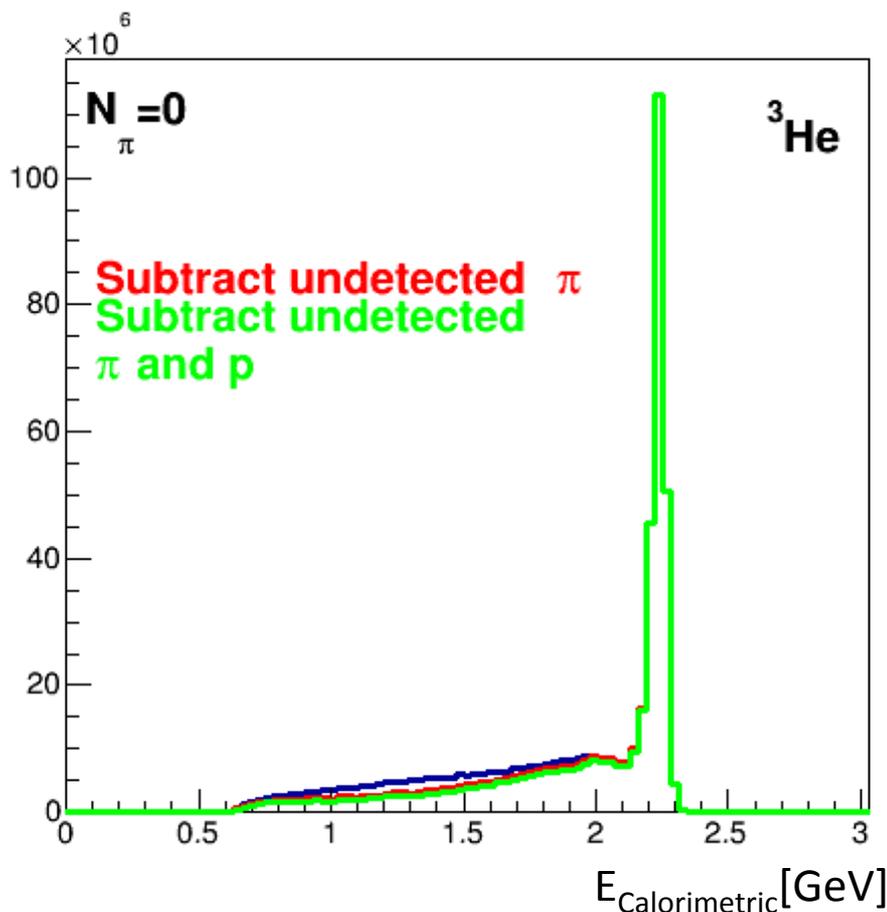
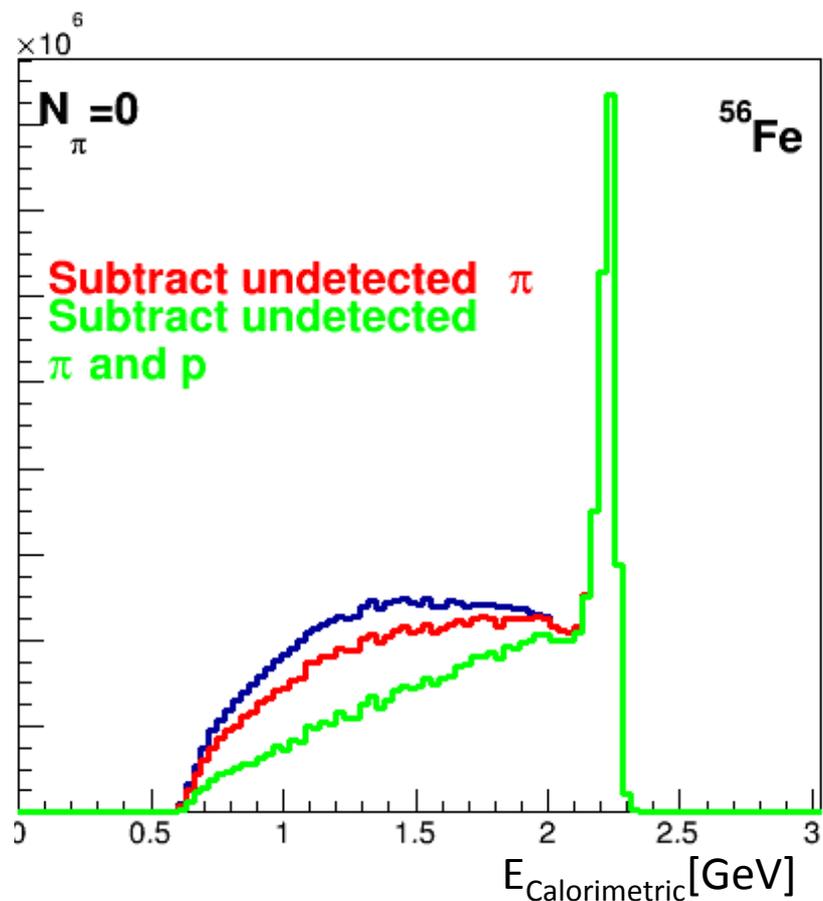


$$E_{\text{Calorimetric}} = E_{e'} + T_p + E_{\text{Binding}}$$

$E_{e'}$ - energy of scattered electron

T_p - kinetic energy of knock-out proton

E_{binding} - Difference between binding energies of A and A-1 nuclei

 ${}^3\text{He}$

 ${}^{56}\text{Fe}$


2.2GeV

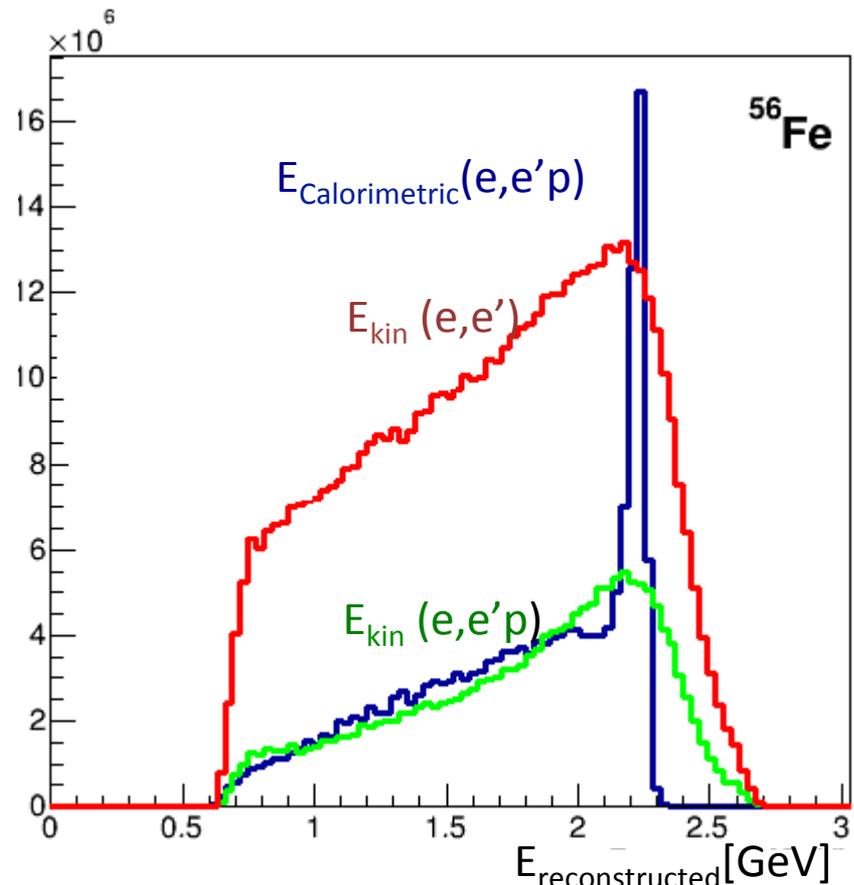
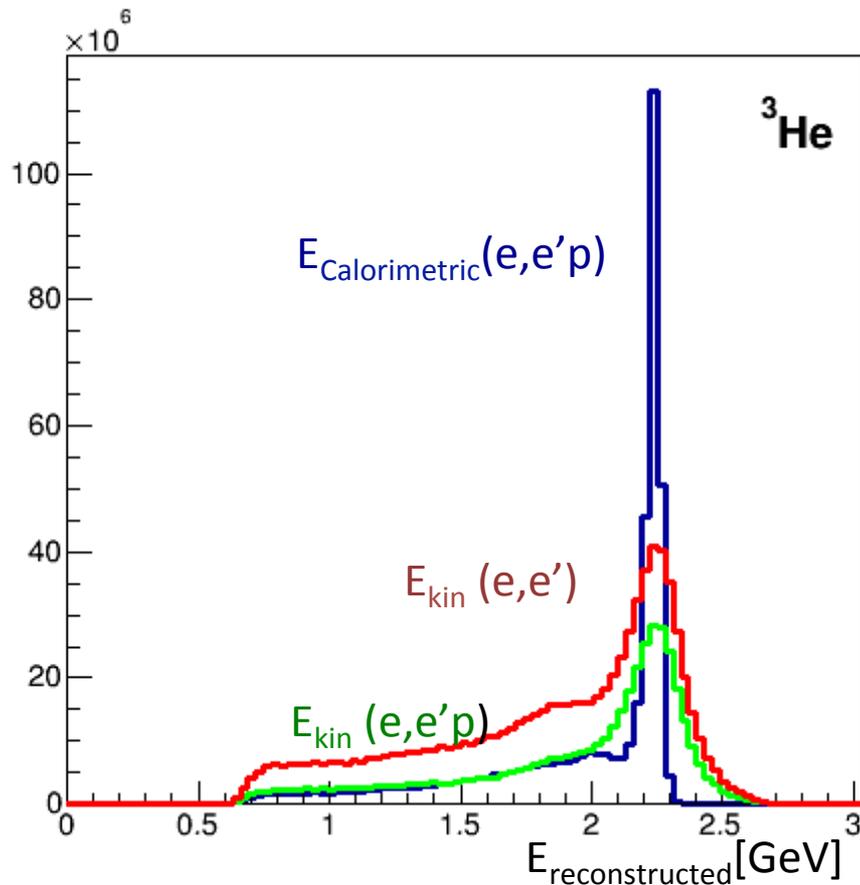
$(e,e'p) E_{\text{calorimetric}}$, $(e,e') E_{\text{rec}}$ and $(e,e'p) E_{\text{rec}}$

Cuts

No π^+ , π^- and no photons coming from π^0 decay

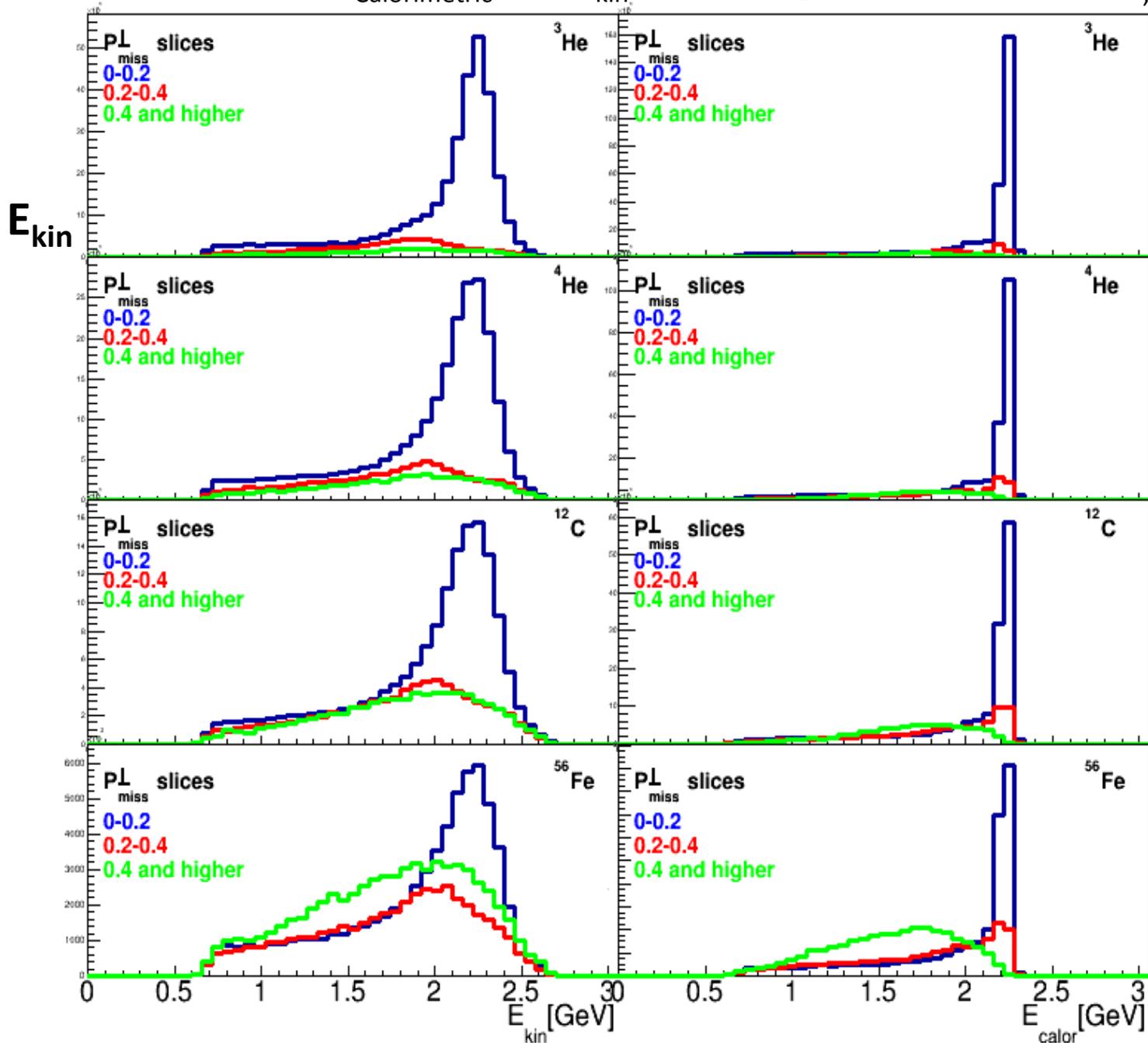
^3He

^{56}Fe



1. E_{kin} has Worse peak resolution than $E_{\text{Calorimetric}}$
2. ^{56}Fe is much worse than ^3He
3. Same tail for $E_{\text{kin}} + E_{\text{calorimetric}}$
4. ^{56}Fe predominantly tail

$E_{\text{Calorimetric}}$ and E_{kin} for all targets at 2.261 GeV in P_{\perp}^{miss} slices



$E_{\text{Calorimetric}}$

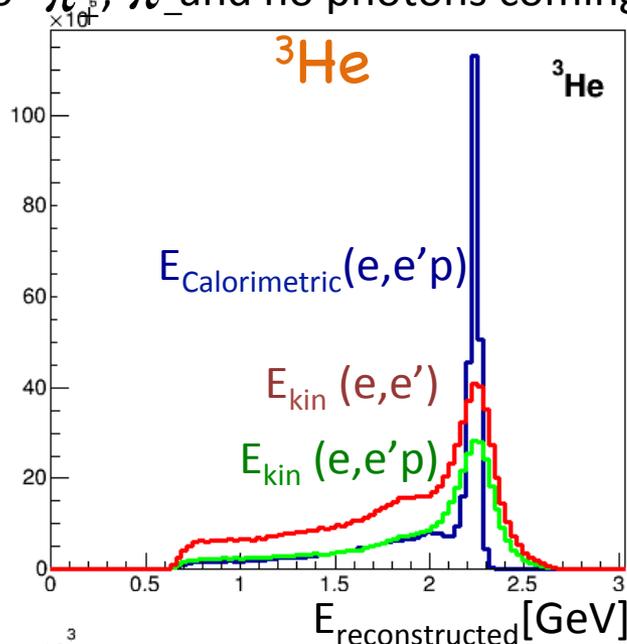
1. Increase in non-QE background with increasing P_{\perp}^{miss}
2. Radiative tail in $E_{\text{calorimetric}}$
3. Worse peak resolution for E_{kin}
4. Increase in non-QE background for heavier targets 16
5. $E_{\text{Reconstructed}}$ can be improved by cut

4.461 GeV analysis

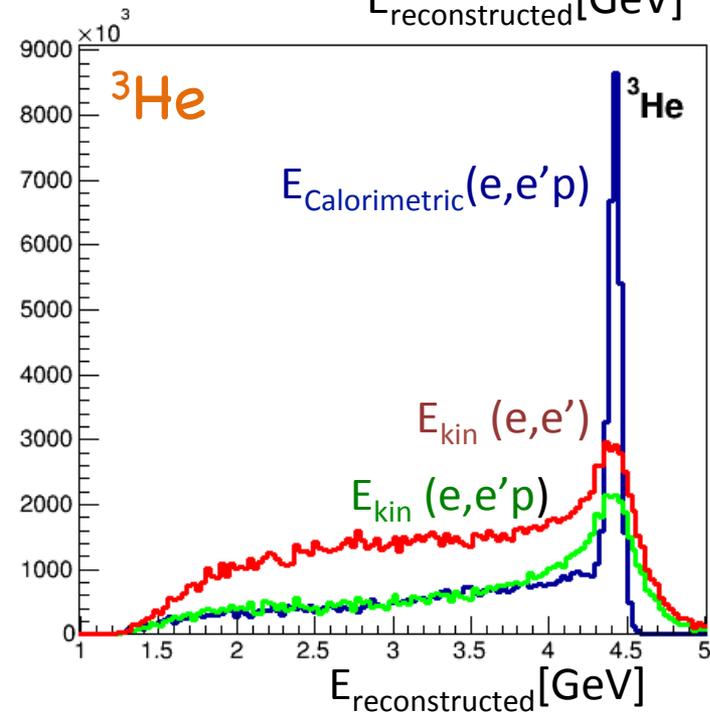
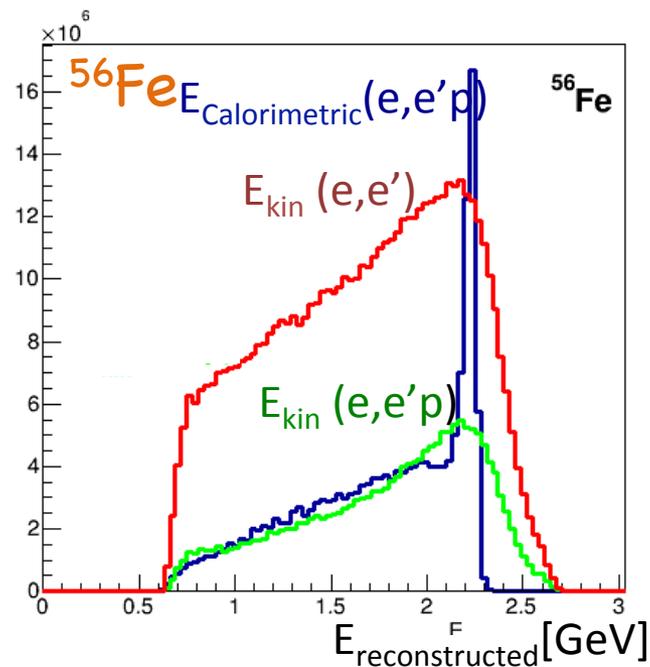
Cuts

$(e,e'p) E_{\text{calorimetric}}$, $(e,e') E_{\text{kin}}$ and $(e,e'p) E_{\text{kin}}$

No π_{\pm} , π_0 and no photons coming from π_0 decay

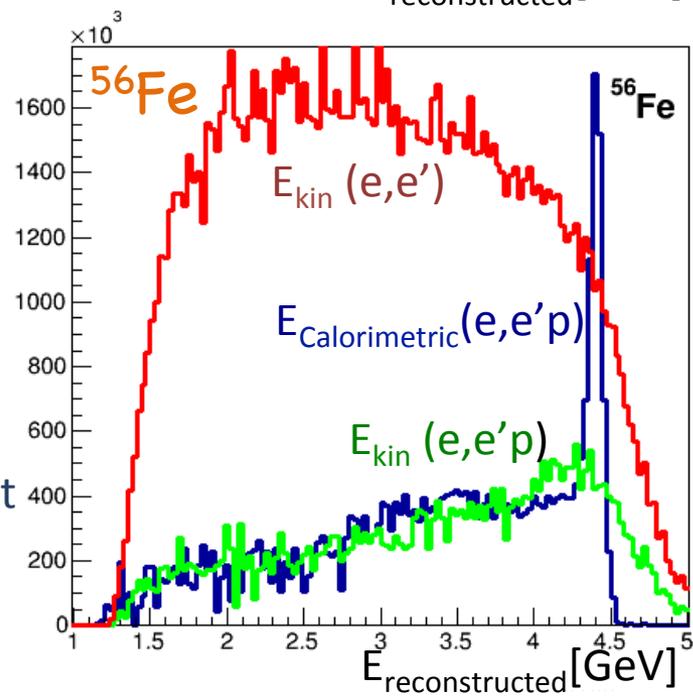


2.2 GeV



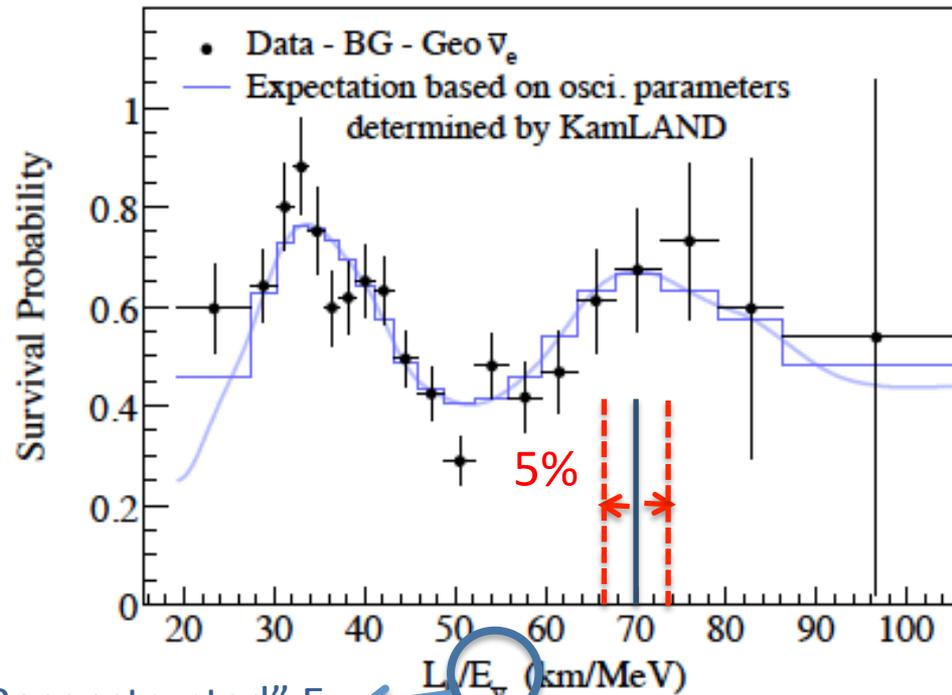
4.4 GeV

- No evident peak in E_{kin} of ${}^{56}\text{Fe}$ at 4 GeV
- Work best for ${}^3\text{He}$ at 2 GeV



Fraction of events reconstructed to within 5% of the beam energy

	2.2 GeV		4.4 GeV	
	$E_{kin}(e,e')$	$E_{calorimetric}(e,e'p)$	$E_{kin}(e,e')$	$E_{calorimetric}(e,e'p)$
^3He	0.32	0.55	0.21	0.40
^4He	0.23	0.46	0.15	0.31
^{12}C	0.2	0.39	0.12	0.29
^{56}Fe	0.16	0.26	0.09	0.22



KamLAND, PRL 100, 221803 (2008)

Error sources

- ✧ Statistical error due to the amount of the analyzed data
- ✧ Systematic error due to imperfect geometrical acceptance (to be studied)
- ✧ Errors of the weights for subtraction of undetected pions and protons
 - Statistical error due to the number of rotations is kept less than 1% with sufficient number of rotation (is not included in error calculation)
 - Systematic error due to the dependence of the cross section on the angle between $(\vec{p}_e, \vec{p}_{e'})$ and (\vec{q}, \vec{p}_π) or $(\vec{q}, \vec{p}_{\text{prot}})$ planes (is small and is being studied)

Summary

1. First use of electron data to test neutrino energy reconstruction algorithms

- use zero-pion cuts to enhance quasi-elastic event selection
- just scattered lepton (E_{kin})
 - ✧ used in Cherenkov-type neutrino detectors
- total energy of electron plus proton ($E_{Calorimetric}$)
 - ✧ used in calorimetric neutrino detectors
- improved by a transverse momentum cut to better select QE events

2. Only 0.09–0.55 fraction of events reconstruct to within 5% of the beam energy at 2 GeV

- better for lighter nuclei

3. Serious implications for neutrino oscillation Measurements

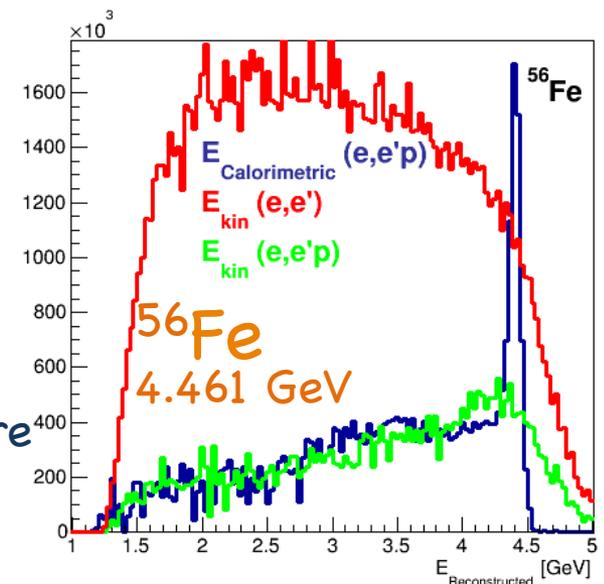
4. Tremendous interest in the neutrino community

5. Analysis note in preparation, aiming for PRL

6. Future work

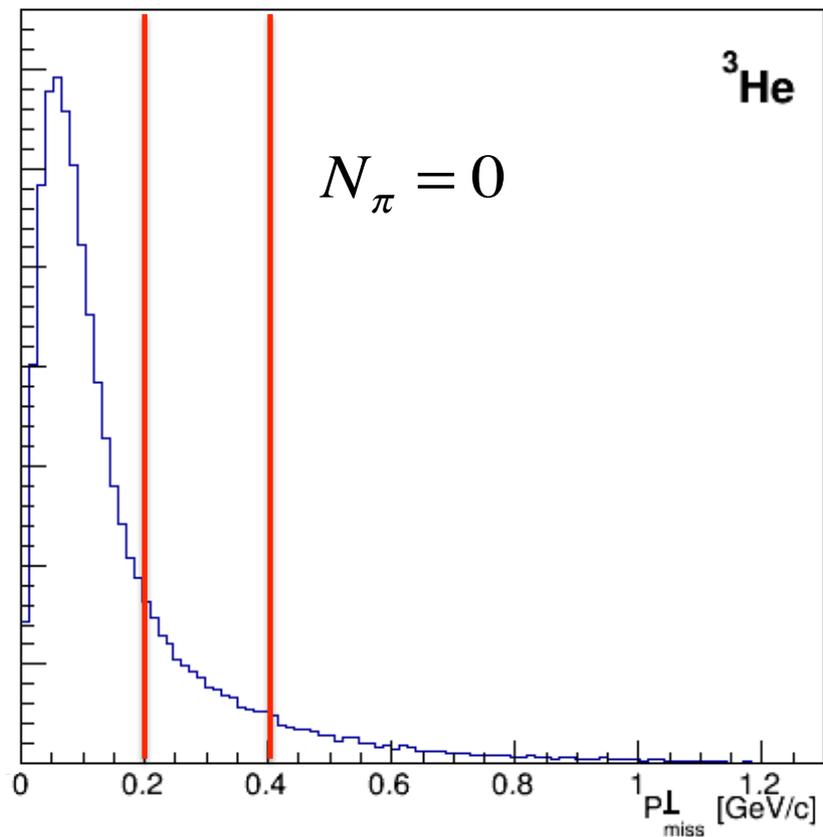
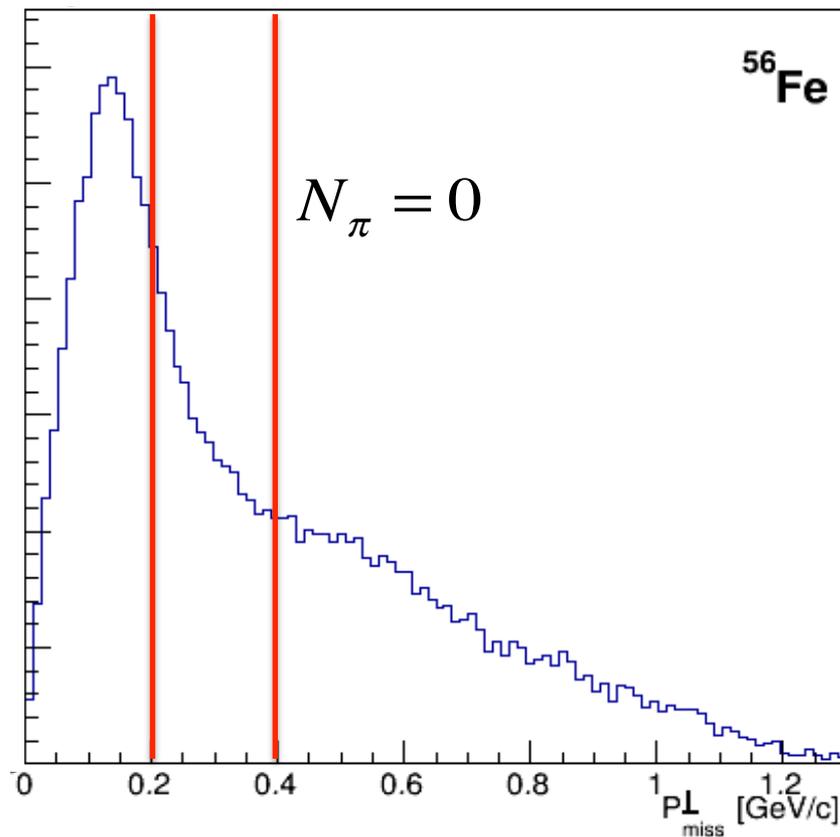
- extend analysis to other kinematic regions, more targets and energies
- Identify regions with good and bad energy reconstruction and GENIE modeling.

- Proposal “Electrons for Neutrinos” conditionally approved by PAC 45.

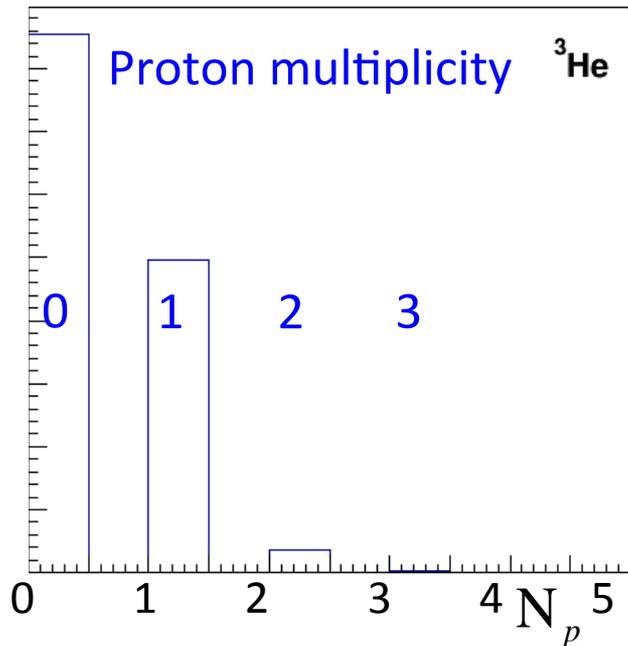


2.2 GeV

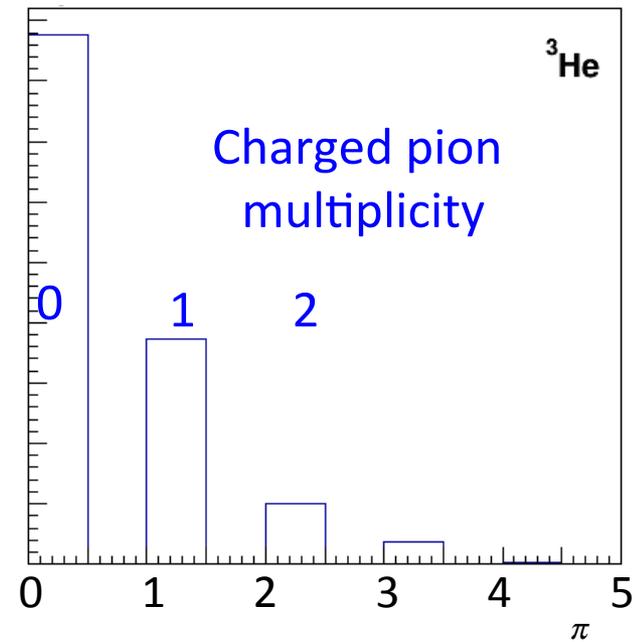
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 ${}^3\text{He}$  ${}^{56}\text{Fe}$ 

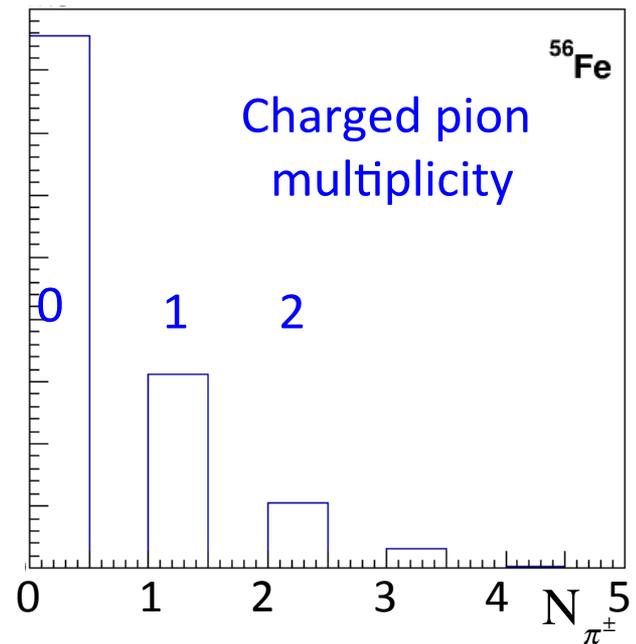
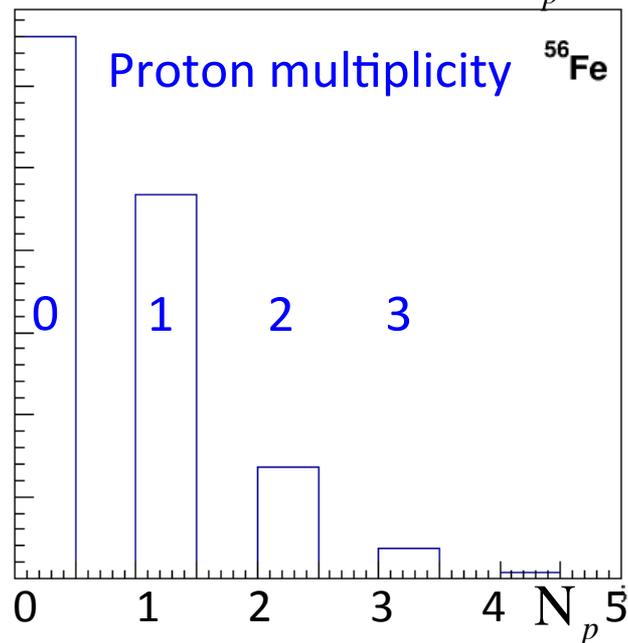
E2a 4.461 GeV



^3He



^{56}Fe



Subtracting undetected pions

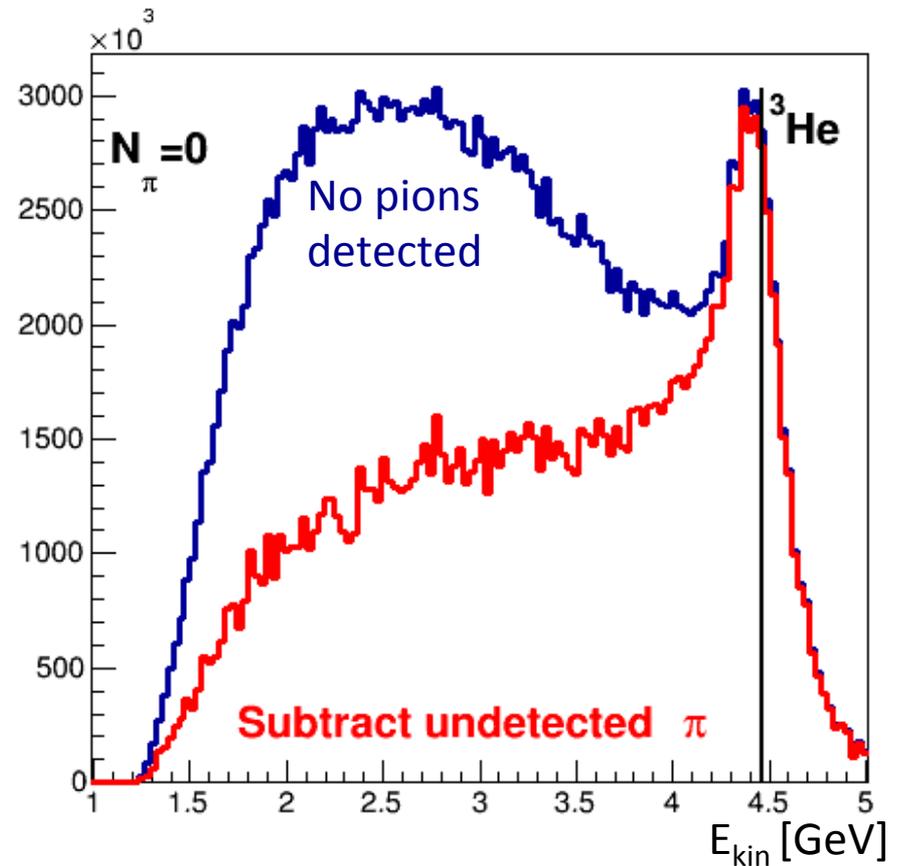
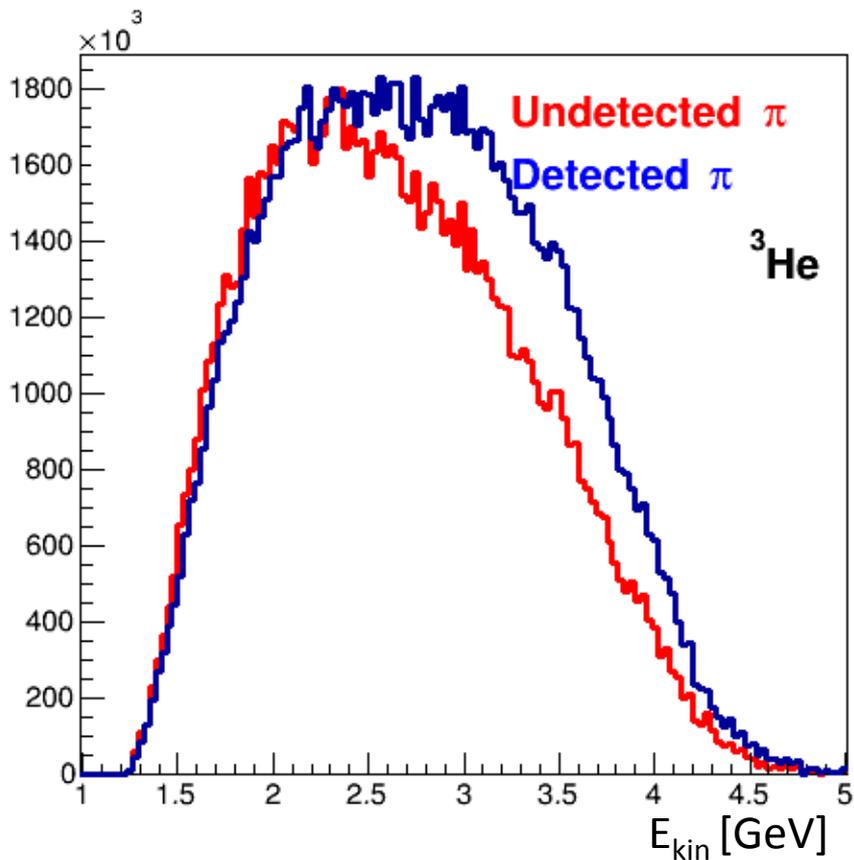
4.461 GeV

(e,e')

Cuts

No π_+ , π_- and no photons coming from π_0 decay

^3He

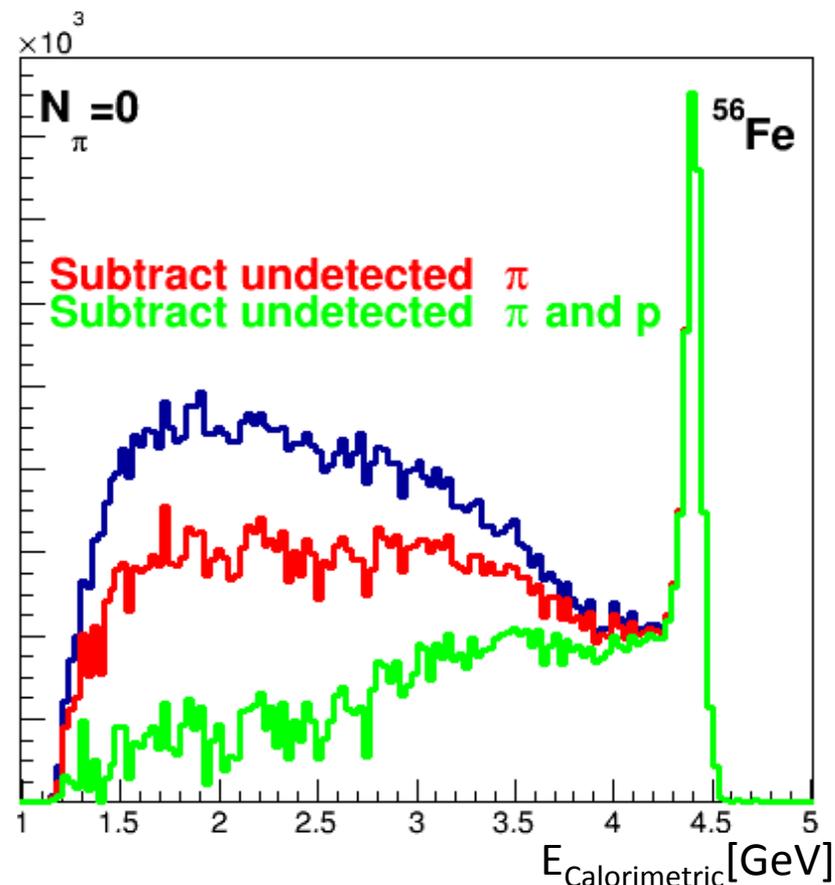
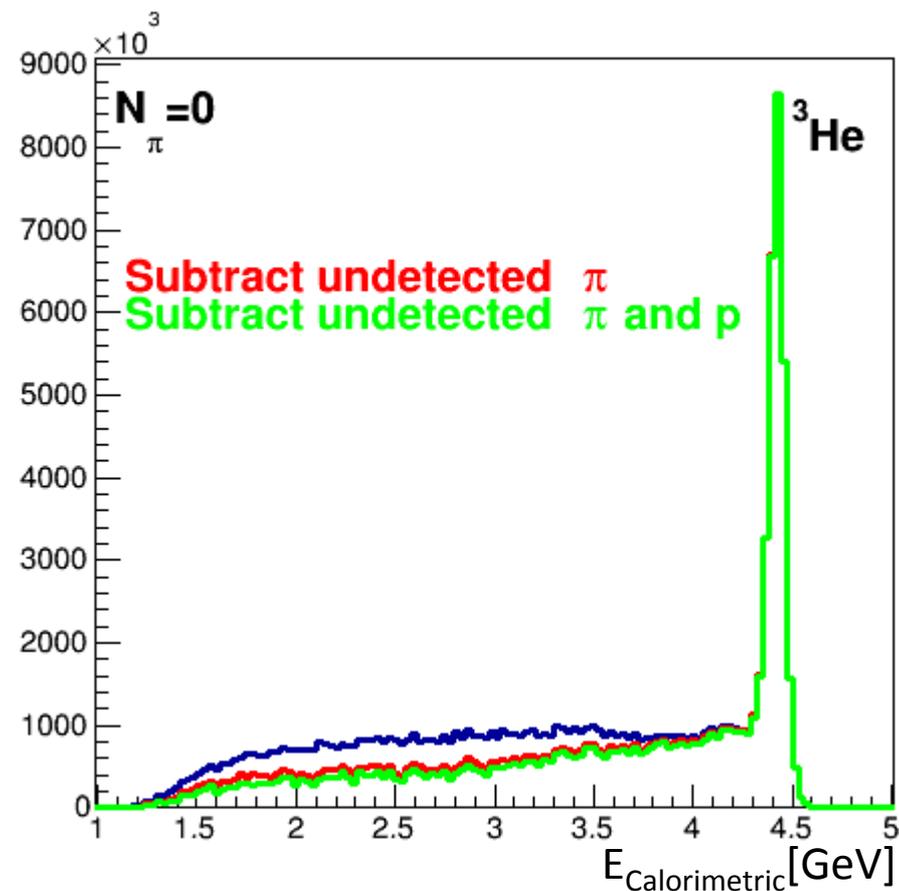


$$E_{\text{Calorimetric}} = E_{e'} + T_p + E_{\text{Binding}}$$

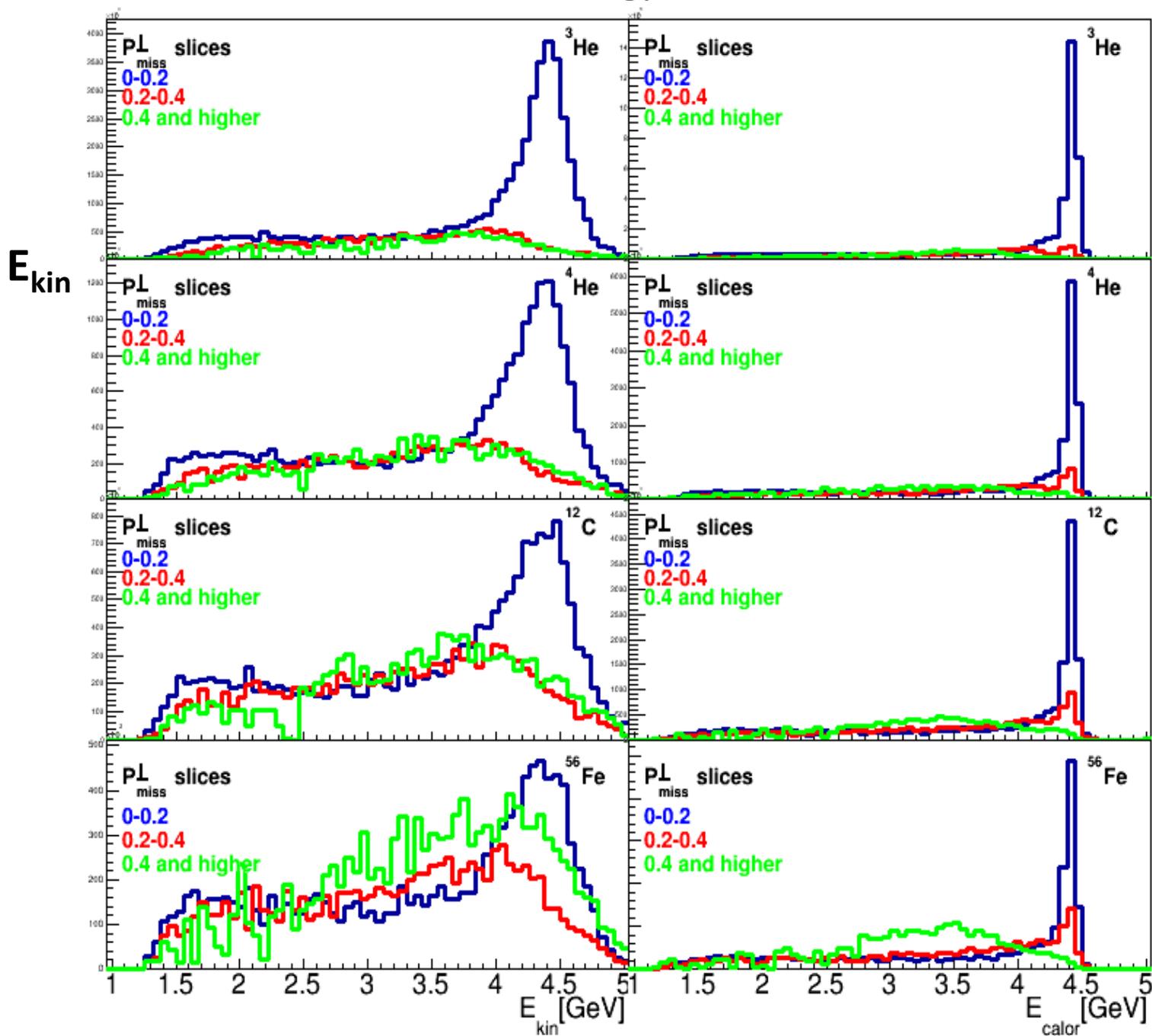
$E_{e'}$ - energy of scattered electron

T_p - kinetic energy of knock-out proton

E_{binding} - Difference between binding energies of A and A-1 nuclei

 ${}^3\text{He}$
 ${}^{56}\text{Fe}$


Energy reconstruction in P_{\perp}^{miss} slices



$E_{\text{Calorimetric}}$

1. The peak at beam energy broader than at 2GeV
2. Background tail is the smallest with respect to the peak for Ecalorimetric at 2 GeV and $P_{\text{miss}}^{\perp} < 200\text{MeV}/c$
3. $E_{\text{Reconstructed}}$ can be improved by cut $P_{\text{miss}}^{\perp} < 26$