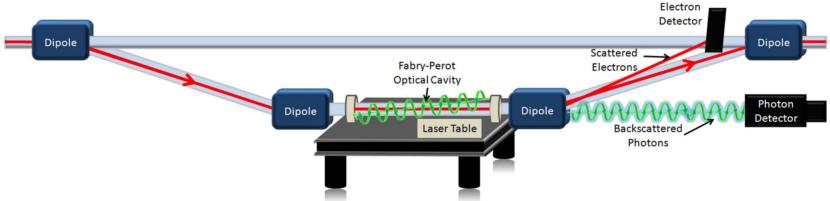
Compton Update for DVCS

Juan Carlos Cornejo DVCS Collaboration Meeting January 26, 2018

Overview of 6 GeV Era Hall A Compton Polarimeter



- Determine beam polarization by measuring asymmetry of $e + \gamma \rightarrow e + \gamma$
- Laser system: 1 W green drive laser coupled to a high gain Fabry-Perot cavity → several kW intracavity power.
- Photon Detectors: GSO ("low energy"), PbWO₄ ("high energy")
- Electron detector: silicon strip detector (evaluating other options for future)
- DAQ: integrating mode for γ-detector (restored by CMU/Gregg)
 - Under development: new counting mode DAQ for γ-detector + new electron detector DAQ (VETROC). Evaluating new integrating γ-detector DAQ options.

Compton Photon Detector 6 GeV Era

Single GSO crystal manufactured by Hitachi Chemical

- 0.5% Ce-doped Gd₂SiO₅
- 6 cm diameter x 15 cm length
- X_o~ 1.4 cm

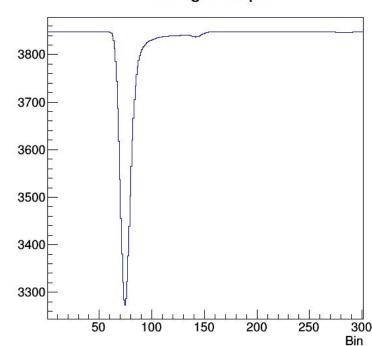


Flash ADC integrating DAQ:

- Customized Struck SIS3320 FADC @ 200 MHz
- No threshold (Accumulator 0) → Dead-timeless
- 1 Primary Data word for each helicity period (Ex: 1/30 second)
 - Additional accumulators
 - Accumulator 4 (with threshold) \rightarrow
 - "stretched-window"
 - Auxiliary diagnostic data taken simultaneously.

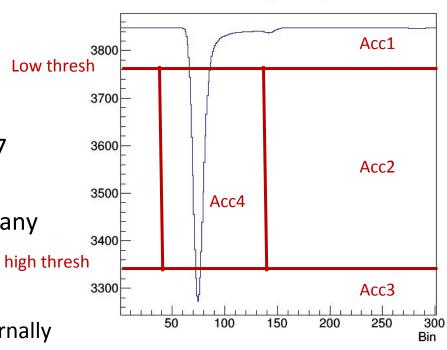
A Bit About Accumulators

- The fADC samples at 200 MHz
 - 1 sample (bin) every 4 ns
- During one helicity flip
 - \circ @ 30Hz \rightarrow ~30 ms window
- So per helicity the fADC will have ~7 million samples
- We cannot possibly readout that many samples fast enough
- So instead we read out the integral
 - Struck fADC does this integral internally
 - Has several integrating methods
- We also write few "snapshots" (300 samples) for systematic studies



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Average Snaps

Acc0 has no threshold

Thresholds are to get rid of synchrotron and high energy backgrounds.

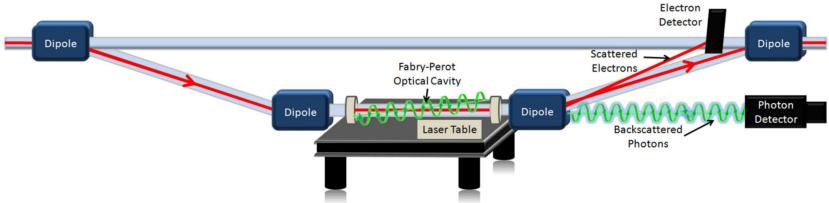
In the Past: Acc0 Prefered Method

- Working without threshold reduces systematics
- Do not need to worry about pedestal

$$A_{msr} = \frac{(S_p - P \cdot N_p) - (S_n - P \cdot N_n)}{(S_p - P \cdot N_p) + (S_n - P \cdot N_n)}$$

- Where 'S' is the integral of an accumulator for plus/minus helicities respectively.
- P is the pedestal, and N is the number of samples
- For Acc0 $N_p = N_n \rightarrow pedestal cancels out$
- For the other accumulators, we would need to precisely measure the pedestal because it does not drop out.
 - Additionally, we'd have to understand and monitor thresholds.

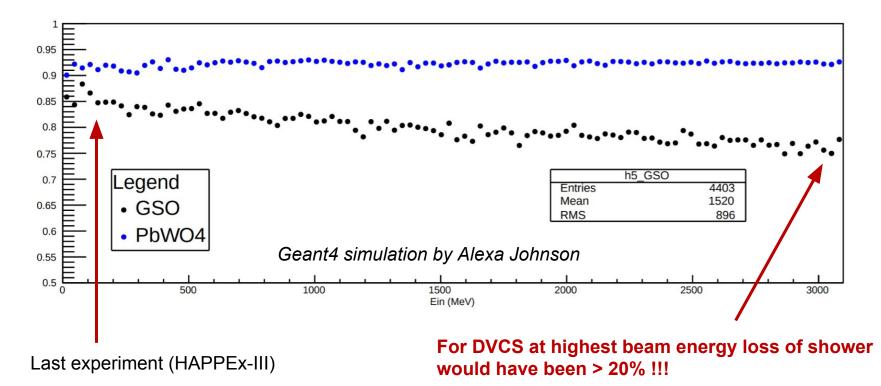
Compton Polarimeter During DVCS-III (12 GeV Era)



- Compton chicane modified to accommodate 12 GeV beam (see backups)
- Photon Detectors
 - **PbWO**₄ was used to span full range of DVCS beam energies (up to 12 GeV)
- Electron detector:
 - Not used for production running
- DAQ:
 - integrating mode for γ -detector (restored by CMU/Gregg)

Why Did We Use PbWO4 Instead of GSO?

- GSO has large radiation length (~1.4 cm)
- PbWO4 much denser so radiation length is ~0.9 cm

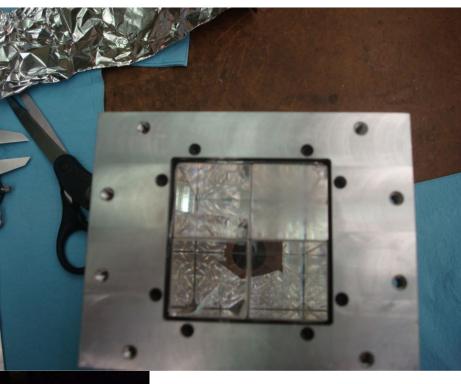


Compton Photon Detector for "High Energy" Experiments

Need denser material to contain shower of ~3 GeV photon \rightarrow PbWO₄ (2x2 stack)

- 6x6 cm x 20 cm length (total)
- On loan from Yerevan/Hall C
- Suitable for > 6 GeV beam
- Light output changes ~2%/°C
 - Wrapped in styrofoam





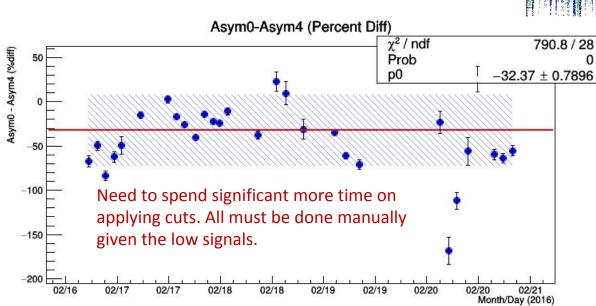
Summary of Spring 2016 Production Running

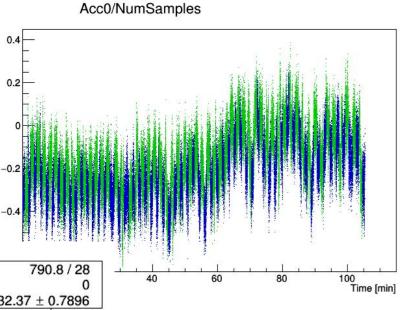
- Saw significant synchrotron radiation
 - \circ Trouble centering beam on Photon Detector \rightarrow used larger aperture collimator
 - Unexpected discrepancy between "accumulators"
 - Lots of system trouble (laser, DAQ, etc...)
- Pass 2 Data (E_{beam} = 4.4 GeV) Feb 16-23
 - 28 "good runs" (~2 hours each). 7 runs possibly recoverable
 - Accumulator data has not been fully analyzed
- Pass 4 Data (E_{beam} = 8.8 GeV) March 1 March 7
 - \circ 43 runs \rightarrow still building "good" set
 - Scaler readout in DAQ went bad, so this set of data requires lots of manual parsing through data to select good cuts.
- Pass 5 Data (E_{beam} ~ 12 GeV) March 8
 - **3 runs** (1 suspect, but very likely recoverable)
- Laser died after these three runs. No more data for spring.

Spring 2016 Pass 2 Overview

Low energy means low signal to noise

- Must be very careful when Ο subtracting background
- Large Acc0 Acc4 discrepancy
 - Possibly helicity pickup or other effect Ο



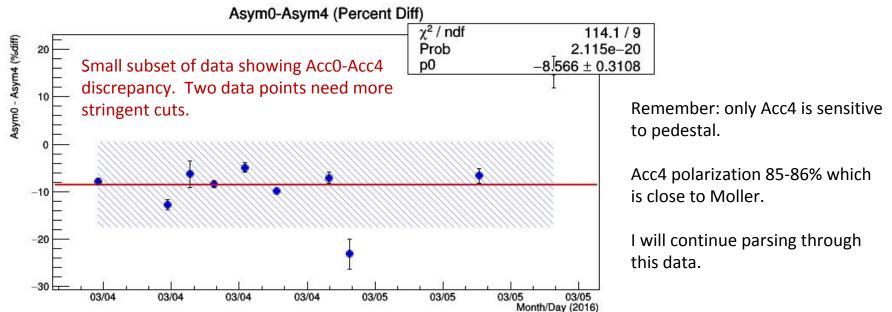


0.2

-0.4

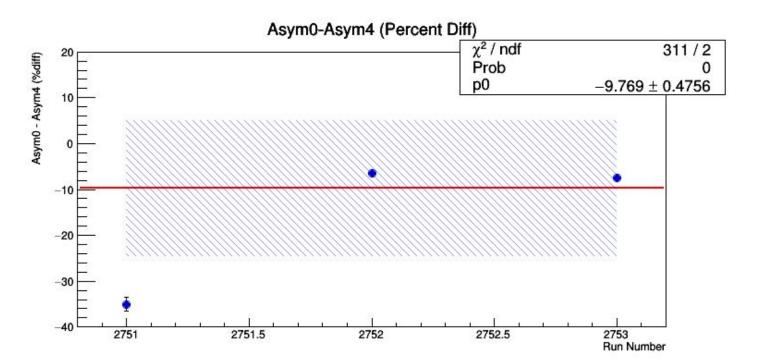
Spring 2016 Pass 4 Overview

- Still needs to be parsed through same level as Fall data
 - Acc0-Acc4 not as large as Pass 2
- Lots of bench and dedicated analysis has been performed on this data to identify discrepancy → nothing so far has explained it.



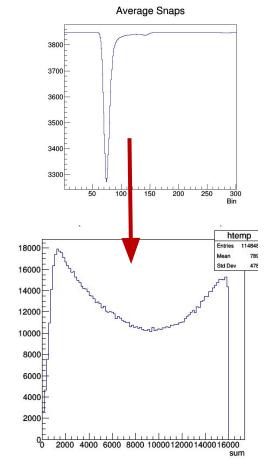
Spring 2016 Pass 5 Overview

- Only 3 data files (2 good, one suspect).
- Same Acc0-Acc4 discrepancy.
- Acc4 gives ~86 % polarization, close to Moller



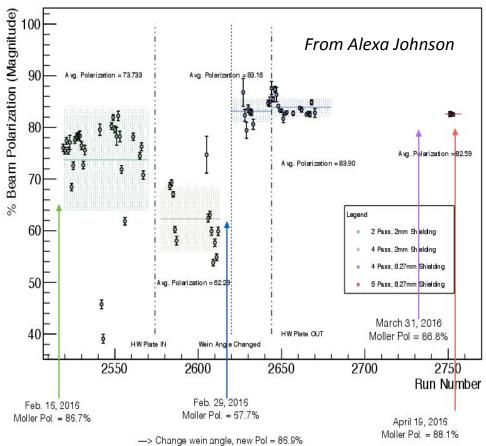
Cross Check Accumulators By Fitting Compton Spectra

- We can perform a cross check on the accumulators by taking the "triggered" data and producing a "Compton spectrum"
- Fit this spectrum directly to Monte Carlo and extract polarization.
- The systematics for this are much harder to deal with.
- Since it is secondary data we do not have very high statistics nor high precision.
- Alexa Johnson used this method to perform a cross check.



Cross Check Accumulators By Fitting Compton Spectra

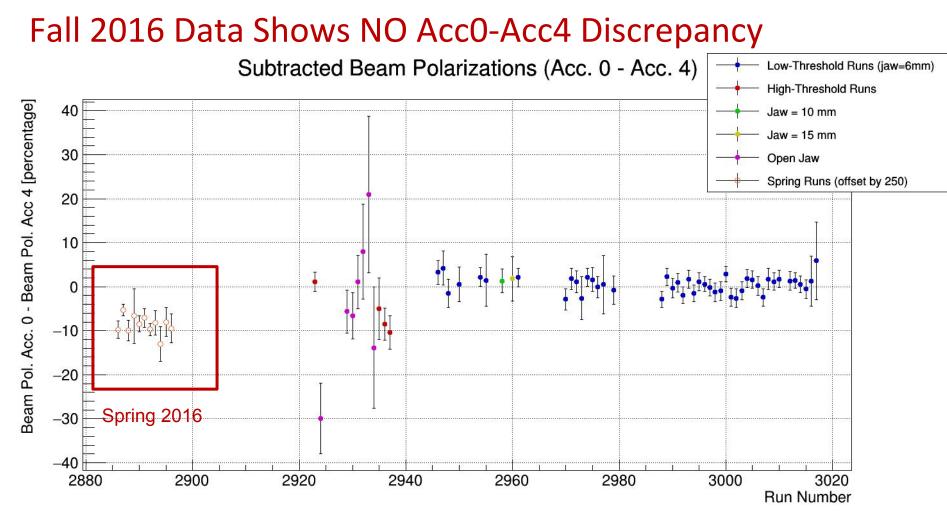
- For Pass 4 and 5 only a few percent below Moller.
- Significantly smaller at Pass2
 - Signal is very tiny at low beam energies.
 - Compton spectra has very poor resolution!
- Gives us hope that Acc4 is "ok"



Beam Polarization vs Run Number

Summary of Fall 2016 Production Running

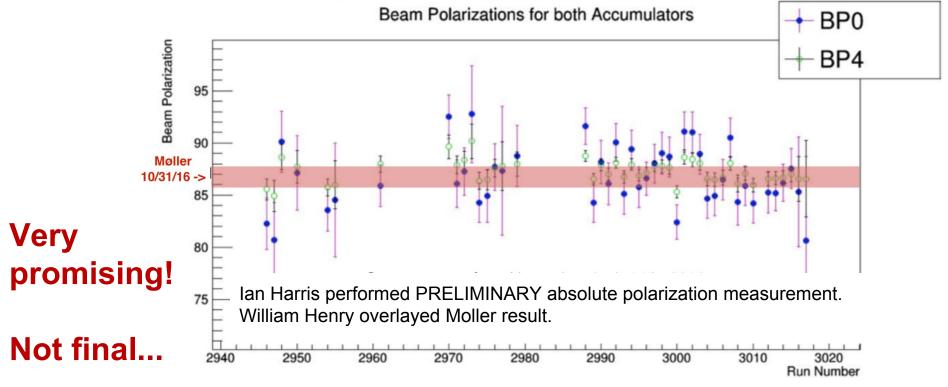
- Synchrotron radiation (8 GeV beam) issues solved!
 - Fixed a misalignment of beamline (now taken into account by MCC)
 - Installed remote controlled collimator system (Tungsten "Jaws")
 - Also a much better way of centering detector
 - \circ With synchrotron issue fixed \rightarrow lowered Accumulator Threshold
- Accumulator discrepancy has disappeared
 - But do not know why....
 - Undergraduate CMU student Ian Harris re-analyzed all of Fall 2016 data.
 - Needed to update pedestal, beam and laser calibrations etc...
- 44 "good" data runs (~2 hours each) @ 8 GeV
- No Pass5 data



High threshold shows large scatter. Low threshold shows good agreement between Acc 0 &4

Compton and Moller Finally in Agreement?

Beam polarizations given by accumulators 0 and 4 were in fair agreement (1.13 stdev)



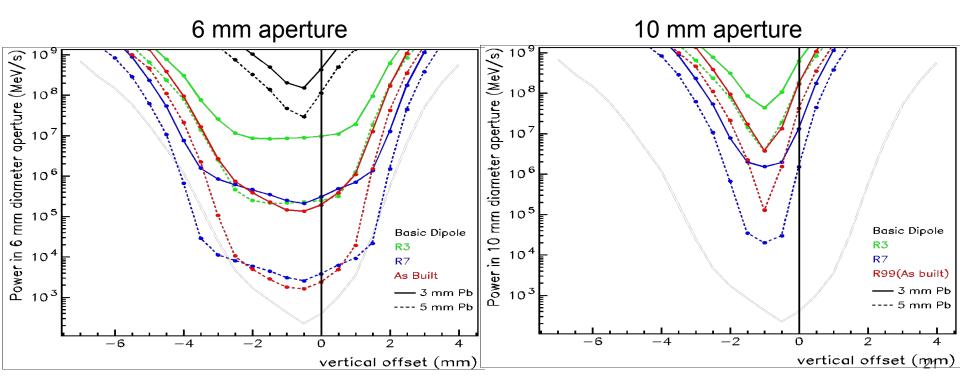
Summary

- Spring 2016 \rightarrow a lot of work needs to be done
 - No final polarizations to show you today
 - Needs lots of manual work to clean up data
 - Likely no resolution to Acc0-Acc4 discrepancy.
 - Worst case scenario: use Acc4 to monitor polarization stability, use Moller for absolute polarization value.
- Fall 2016 \rightarrow Very very promising
 - Thanks to lan's hard work, got good set of data.
 - Acc0 Acc4 discrepancy seems to have disappeared.
 - Not mentioned today: need to do some more work to explain sudden change
- No ETA on final analysis, sorry :(
- Stay tuned....

Extras

Photon Detector Experience with DVCS

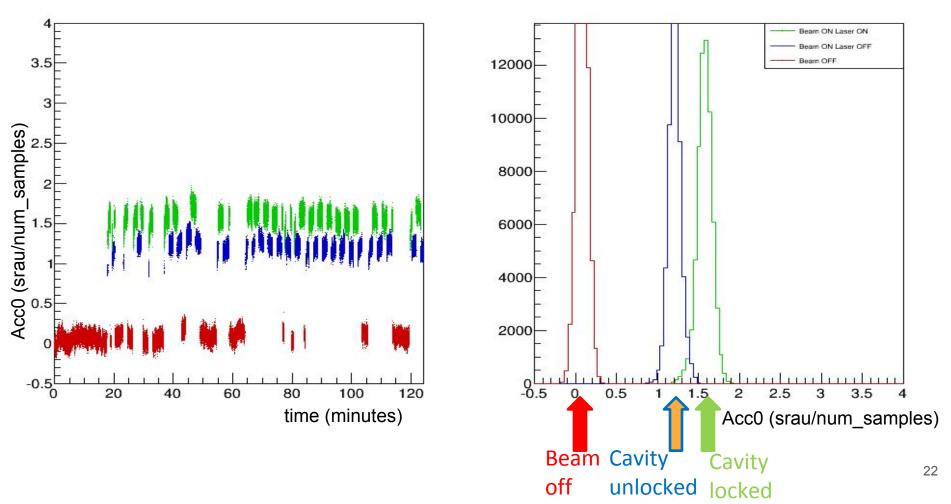
- Spring $2016 \rightarrow \text{saw significant}$ synchrotron radiation
 - \circ Trouble centering beam on Photon Detector \rightarrow used larger aperture collimator
 - Unexpected discrepancy between "accumulators"



Spring 2016 4-pass running (@15 µA)



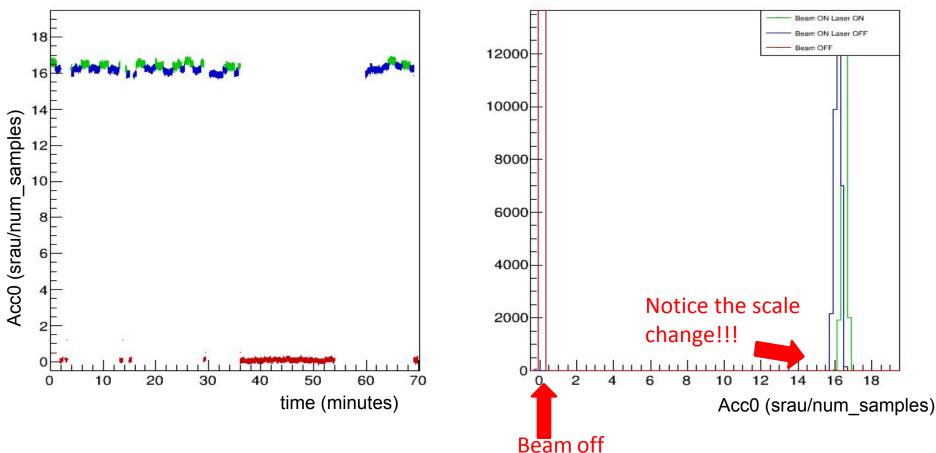
Acc0/NAcc0, Run=2631, 10mm Aperture



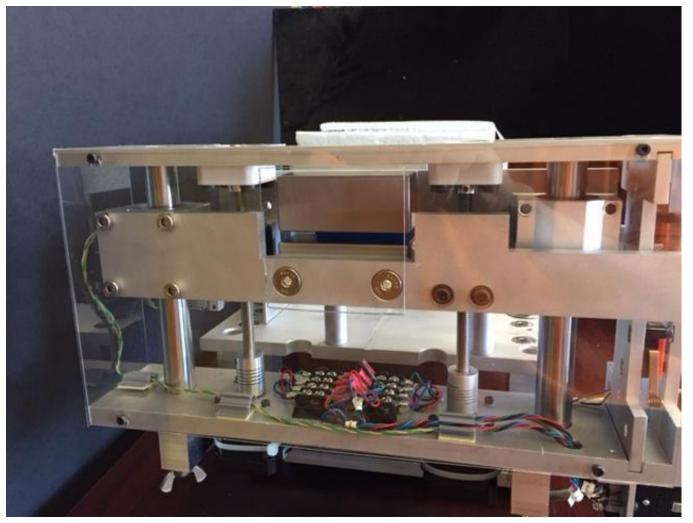
Spring 2016 5-pass running (@15 µA)

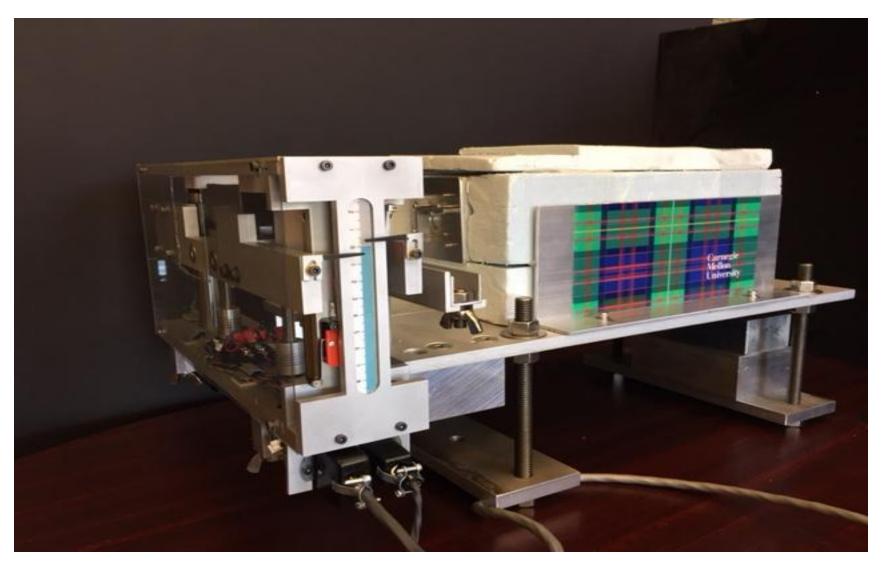
Acc0/NAcc0, Run=2751, 10mm Aperture

Acc0/NAcc0, Run=2751, 10mm Aperture



Remote (arduino) controlled Tungsten "JAWS"

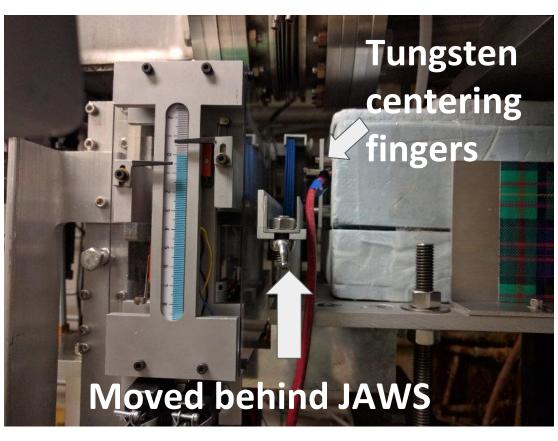




Remote (arduino) controlled Tungsten "JAWS"

Synchrotron shield in front of JAWS removed

H3ddU

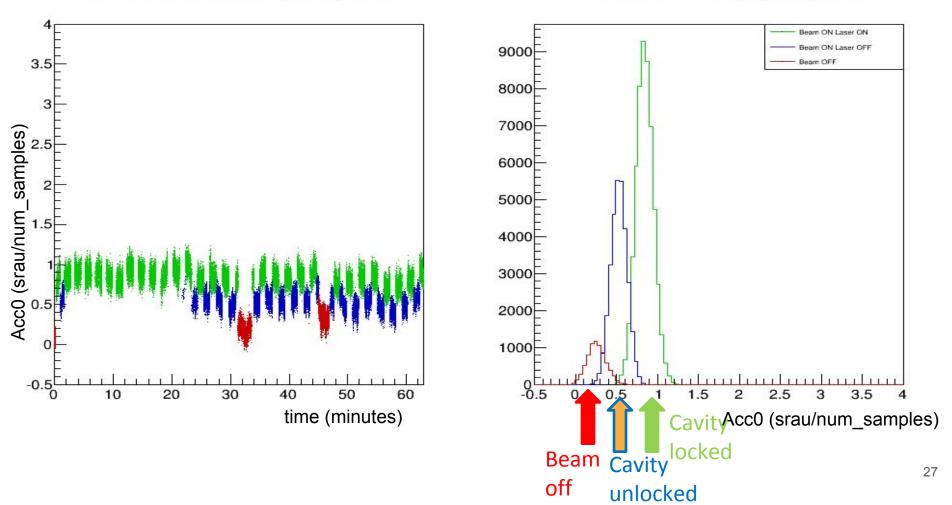


Simulations showed lead-synch shield spread photon beam and JAWS would clip Compton Spectrum.

Fall 2016 4-pass running with 10 mm aperture (@10 µA)

Acc0/NAcc0, Run=2958, 10mm Aperture

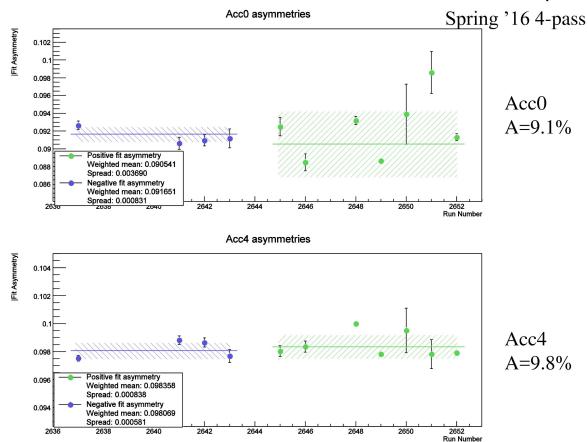
Acc0/NAcc0, Run=2958, 10mm Aperture



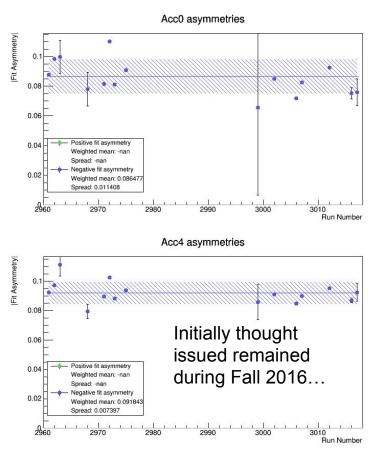
Unanticipated Discrepancy in No-Threshold vs Threshold

• Acc0 (no-threshold) is as much as 10%

lower than Acc4 (threshold)



Unanticipated Discrepancy in No-Threshold vs Threshold



- Acc0 (no-threshold) is as much as 10% lower than Acc4 (threshold)
- Many bench tests already performed, including dedicated analysis.
 - Initial primary suspect was a likely "afterglow"
 - R.E Zhu et. al. NIM A 376 (1996) found some crystals with > 15 ms afterglow.
 - Afterglow can be "small" but Acc0 integrates ~6.6 million samples → effect gets amplified.
 - See previous Compton talks for all bench tests that we ruled out.

Photon Detector Experience with DVCS

- Spring $2016 \rightarrow \text{saw significant}$ synchrotron radiation
 - \circ Trouble centering beam on Photon Detector \rightarrow used larger aperture collimator
 - Unexpected discrepancy between "accumulators"
- Fall 2016 \rightarrow synchrotron radiation (8 GeV beam) issues solved!
 - ✓ Fixed a misalignment of beamline (now taken into account by MCC)
 - Installed remote controlled collimator system (Tungsten "Jaws")
 - Also a much better way of centering detector
 - \circ With synchrotron issue fixed \rightarrow lowered Accumulator Threshold
 - Accumulator discrepancy...
 - Is gone!
 - Undergraduate CMU student Ian Harris re-analyzed all of Fall 2016 data.
 - Needed to update pedestal, beam and laser calibrations etc...
 - Last nagging issue is strange Acc0 to Acc4 correlation
 - But less worrisome.

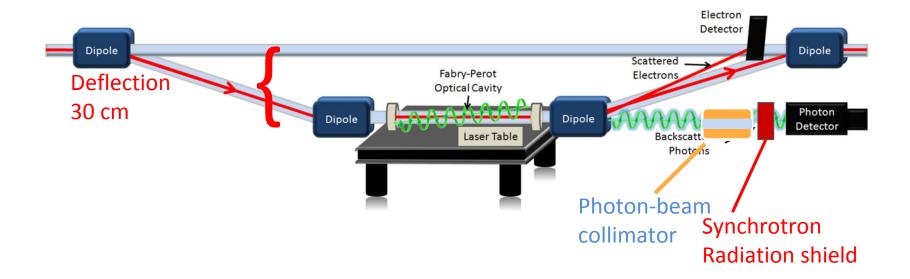
Found nothing that would indicate problem with GSO (PREX2/CREX)

Summary

- Compton had to be recommissioned for 12 GeV running.
 - Hit several bumps along the way.
- Laser system needs some work \rightarrow have plan to move forward.
- PbWO₄ Photon detector tested (will not be used for PREX/CREX2)
- Synchrotron and centering issues no longer a problem → should be smooth experience moving forward.
- Initial Acc 0 & 4 discrepancy caused us to perform lots of bench tests and dedicated analysis
 - Found nothing that indicates a problem for PREX2/CREX.
 - Resolved Acc0/4 issue now less worrisome → but will keep looking into for future experiments.

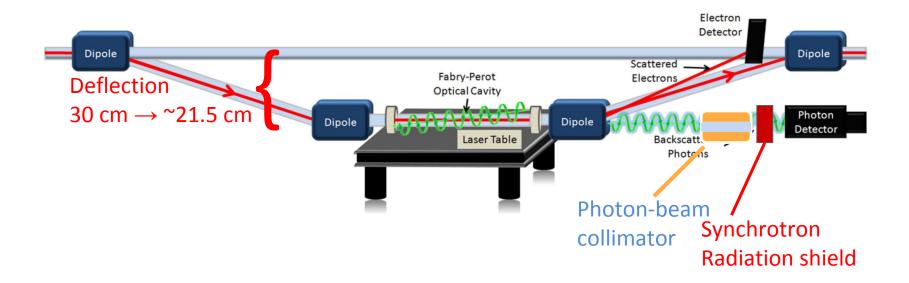
Compton Chicane during 6 GeV Era

- Vertical chicane ~30 cm
- Cylindrical collimator and thin Pb "shield" used to minimize synchrotron



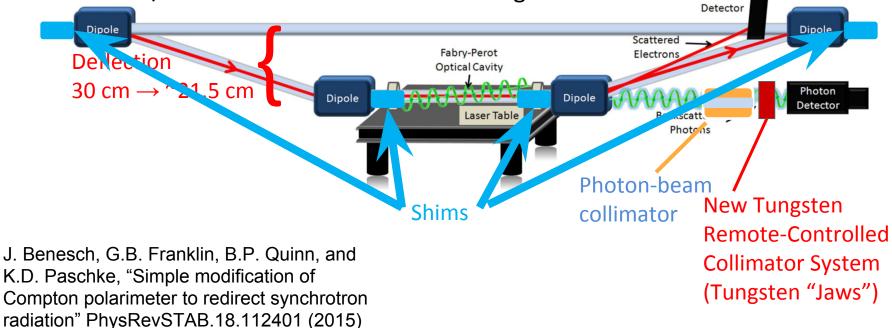
Compton Chicane at 12 GeV

- Vertical chicane deflection decreased (30 cm \rightarrow ~21.5 cm)
- Expect ~E⁴ increase in synchrotron radiation!!!



Compton Chicane at 12 GeV

- Vertical chicane deflection decreased (30 cm \rightarrow ~21.5 cm)
- Expect ~E⁴ increase in synchrotron radiation!!!
- Modify fringe field with added 'shims'.
- PREX2/CREX will **also** use this new configuration.

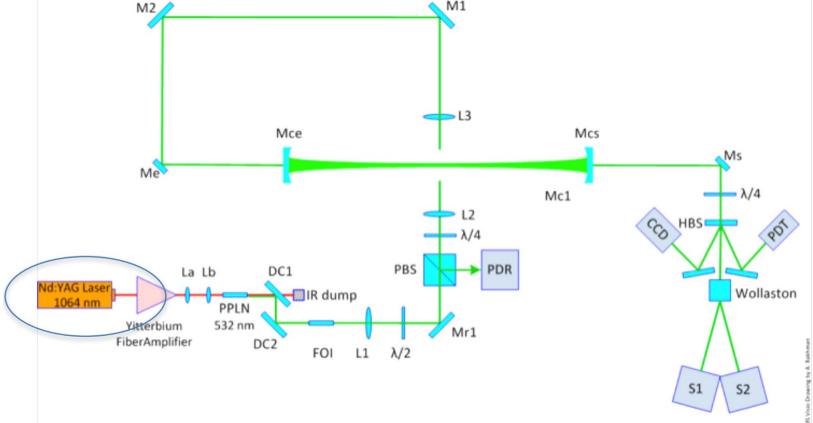


Electron

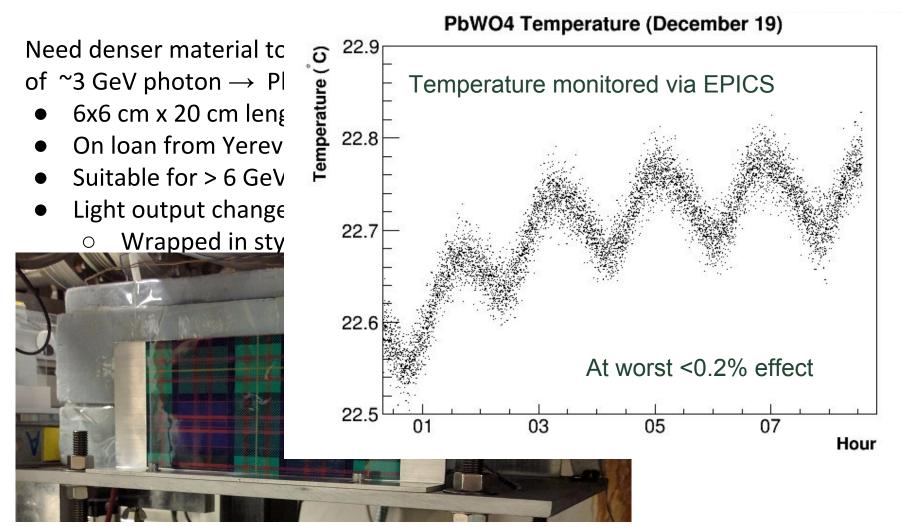
Compton Laser System Experience with DVCS/GMp

• Compton laser system was fully functional at the start of Fall 2016

Replaced a seed laser + fiber amplifier during Summer 2016



Compton Photon Detectors: "High Energy"

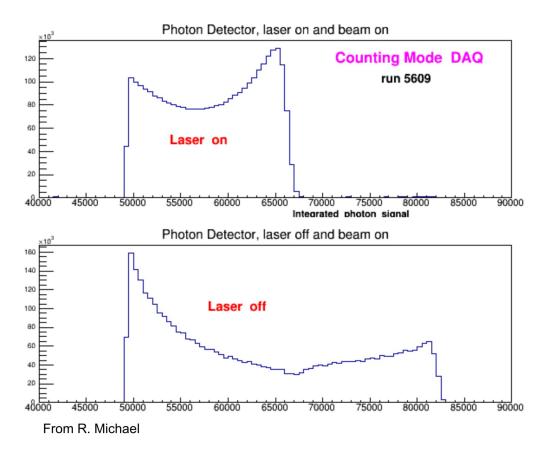


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Helicity pickup

- Helicity flip rate (@ 30 Hz) is **not delayed** \rightarrow significant pickup
 - Helicity correlated difference with no-delay: 0.04 rau/num_samples
 - Modified helicity bit and manually delayed it till end of MPS window
 - Helicity correlated difference with manual delay: max 0.001 rau/num_samples
 - Reminds us of why we **need** delayed helicity reporting

Progress on "counting mode" Photon Detector DAQ



- Progress made during Fall 2016
 - Simultaneously ran integrating and counting DAQ (S. Liu, R. Michaels & A. Camsonne)
- Present status:
 - Readout of individual photons via a single JLab FADC-250
 - Still missing helicity info in order to extract asymmetries.
- Counting mode can be better integrated with electron DAQ (VETROC)