# Internal Target Experiments

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# Nucleosynthesis: $\alpha + {}^{12}C \rightarrow \gamma + {}^{16}O$

- ${}^{12}C/{}^{16}O$  ratio  $\rightarrow$  affects nucleosynthesis of heavier elements and stellar evolution
  - $\rightarrow$  massive stars: carbon, oxygen, neon burning
  - $\rightarrow$  white dwarfs: super nova type Ia
  - $\rightarrow$  end of stars: <sup>16</sup>O rich star black hole, <sup>12</sup>C rich star neutron star

T. A. Weawer and S. E. Woosley, Phys. Rep. 227 (1993) 335.

- For  $\alpha + {}^{12}C$  at T ~ 2 · 10<sup>8</sup> K, Gamow window ~ 300 keV
- Due to Coulomb barrier  $\sigma \sim 10^{-5}$  pb (direct measurement is not possible)

# Nucleosynthesis: $\alpha + {}^{12}C \rightarrow \gamma + {}^{16}O$



- Contribution to the cross section at Gamow window is complicated:
- $\rightarrow$  E1 component, 1<sup>-</sup>: subthreshold state at 7.117 MeV and broad resonance at 9.59 MeV
- $\rightarrow$  E2 component, 2<sup>+</sup>: subthreshold state at 6.917 MeV and narrow resonance at 9.85 MeV

L.R. Buchmann, C.A. Barnes, Nucl. Phys. A 777 (2006)

### Nucleosynthesis: $\alpha + {}^{12}C \rightarrow \gamma + {}^{16}O$



R. Plag et al, Phys. Rev. C 86 (2012)

### **Cross section measurements**

• Direct measurements by using:

a)  $\alpha$  beam: detection of gammas  $\rightarrow$   $S_{E1}$  and  $S_{E2}$ 

b) <sup>12</sup>C beam (inverse kinematic): detection of <sup>16</sup>O recoils  $\rightarrow$  S<sub>tot</sub>

• Indirect measurements:

a)  $\beta$  decay of <sup>16</sup>N: <sup>16</sup>O<sup>\*</sup>  $\rightarrow \alpha + {}^{12}C (S_{E1} \text{ can be obtained})$ b) inverse reaction,  $\frac{\sigma(\gamma + {}^{16}O)}{\sigma(\alpha + {}^{12}C)} = \frac{\mu c^2 E_{\alpha}}{E_{\gamma}^2} \approx 42 \text{ (for } E_{\alpha} = 1 \text{ MeV})$  $- \text{ photo-disintegration of } {}^{16}O: \gamma + {}^{16}O \rightarrow \alpha + {}^{12}C$  $- \text{ electro-disintegration of } {}^{16}O: {}^{16}O(e, e' \alpha){}^{12}C$ 

![](_page_5_Figure_1.jpeg)

• Advantage:

- a) inverse reaction: larger cross section than direct reaction
- b) JLab's Low Energy Recirculator Facility (LERF) e<sup>-</sup> beam
- c) Internal windowless gas target Jet target
- Disadvantage: cross section has more complicated structure

$$\frac{d\sigma}{d\omega d\Omega_e d\Omega_\alpha} = \frac{M_\alpha M_{12C}}{8\pi^3 W} \frac{p_\alpha^{cm} f_{rec}^{-1}}{(\hbar c)^3} \sigma_{Mott} (\nu_L R_L + \nu_T R_T + \nu_{LT} R_{LT} + \nu_{TT} R_{TT})$$

#### A. S. Raskin and T. W. Donnelly, Ann. of Phys. 191 (1989)

$$R_{T} = P_{0}(\cos\theta_{\alpha})[|t_{E1}|^{2} + |t_{E2}|^{2}] +$$

$$+ P_{1}(\cos\theta_{\alpha})\frac{6}{\sqrt{5}}\cos(\varphi_{E2} - \varphi_{E1})|t_{E1}||t_{E2}| +$$

$$+ P_{2}(\cos\theta_{\alpha})\left[-|t_{E1}|^{2} + \frac{5}{7}|t_{E2}|^{2}\right] -$$

$$- P_{3}(\cos\theta_{\alpha})\frac{6}{\sqrt{5}}\cos(\varphi_{E2} - \varphi_{E1})|t_{E1}||t_{E2}| -$$

$$- P_{4}(\cos\theta_{\alpha})\frac{12}{7}|t_{E2}|^{2}$$

• Multipole matrix elements ( $q_0 = 1.2 \text{ fm}^{-1}$ ):

$$t_{EJ} = \frac{\omega}{q} \left(\frac{q}{q_0}\right)^J \left[a'_{EJ} + \left(\frac{q}{q_0}\right)^2 b'_{EJ}(q)\right] e^{-\left(\frac{q}{q_0}\right)^2}$$
$$t_{CJ} = \left(\frac{q}{q_0}\right)^J \left[a'_{CJ} + \left(\frac{q}{q_0}\right)^2 b'_{CJ}(q)\right] e^{-\left(\frac{q}{q_0}\right)^2}$$

• Long wavelength limit  $(q \rightarrow 0)$  and continuity:

$$t_{EJ} \rightarrow -\sqrt{\frac{J+1}{J}} \left(\frac{\omega}{q}\right) t_{CJ} \qquad a'_{EJ} = -\sqrt{\frac{J+1}{J}} \left(\frac{\omega}{q}\right) a'_{CJ}$$

![](_page_8_Figure_1.jpeg)

- α-knockout experiment at NIKHEF
- $3\mu A e^{-}$  beam at 639 MeV and 615 MeV
- Energy of  $\alpha$ -particles: from 20 to 35 MeV
- Target: CO<sub>2</sub> at 1.6 bar and 300 K
- Luminosity:  $1.46 \cdot 10^{34} (\text{cm}^{-2}\text{s}^{-1})$

G. De Meyer et al., Phys. Lett. B 513 (2001)

- e<sup>-</sup> 5 mA at 100 MeV at LERF
- Mainz jet target: 132.10<sup>-6</sup> g/cm<sup>2</sup>
  (S. Aulenbacher, DarkLight meeting June 2016)
- Luminosity: 2.45 · 10<sup>36</sup> (cm<sup>-2</sup>s<sup>-1</sup>)
- Solid angle: 5 msr for  $e^-$ , 25 msr for  $\alpha$
- Data from M. Assunção et al. Phys. Rev. C 73 (2006) to estimate  $a'_{EJ}$ ,  $a'_{CJ}$  and  $\varphi_{E2} - \varphi_{E1}$
- $E_{CM} = 0.7 \text{ MeV}$

![](_page_9_Figure_7.jpeg)

![](_page_10_Figure_1.jpeg)

![](_page_11_Picture_0.jpeg)

- Perform the error analysis to estimate acceptable uncertainties
- Selection of detector set-up
- Build a simulation to study the background
- Optimize the rate

### Backup

![](_page_13_Figure_1.jpeg)