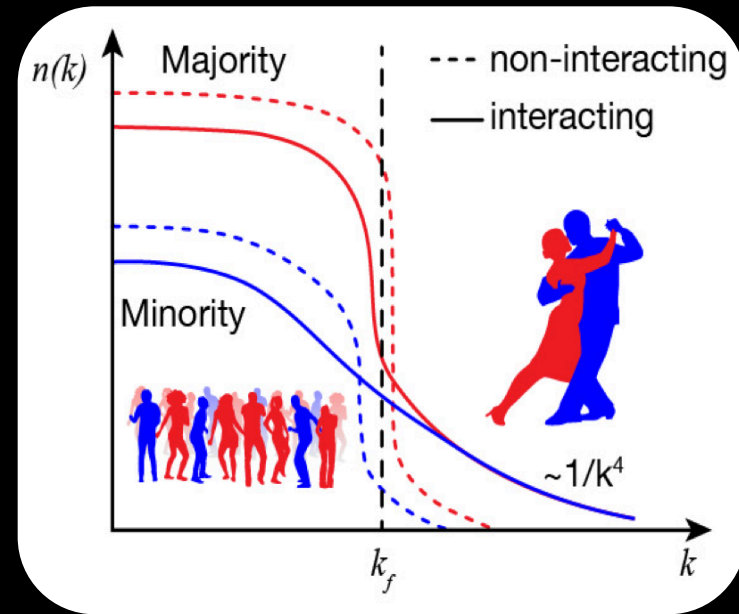


# The CaFe Experiment: Short-Range Pairing Mechanisms in Heavy Nuclei

Proposal PR12-16-004

## Spokespersons:

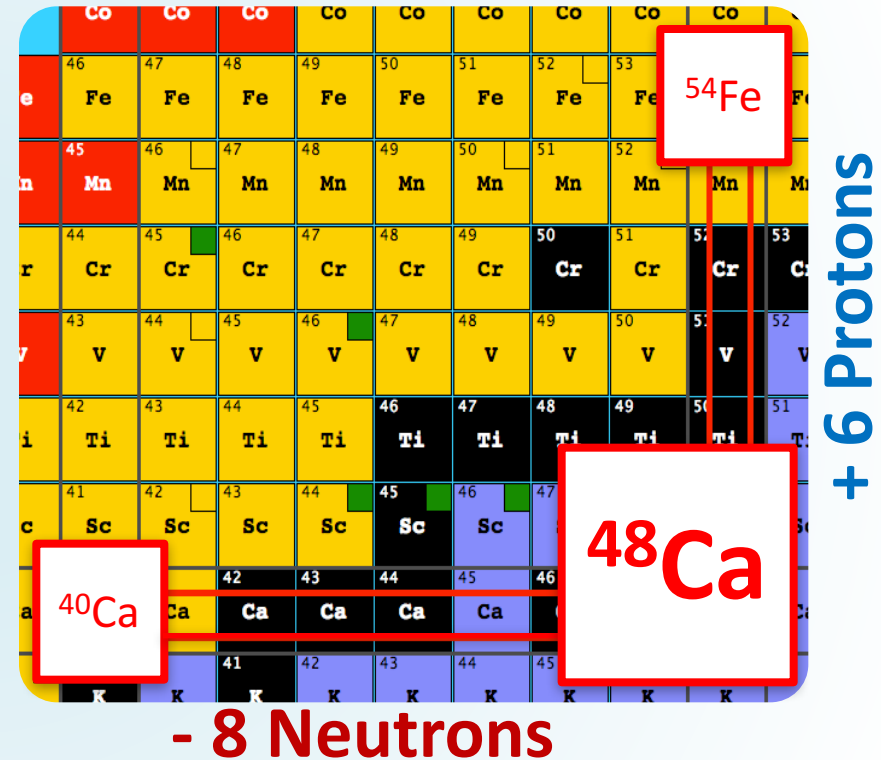
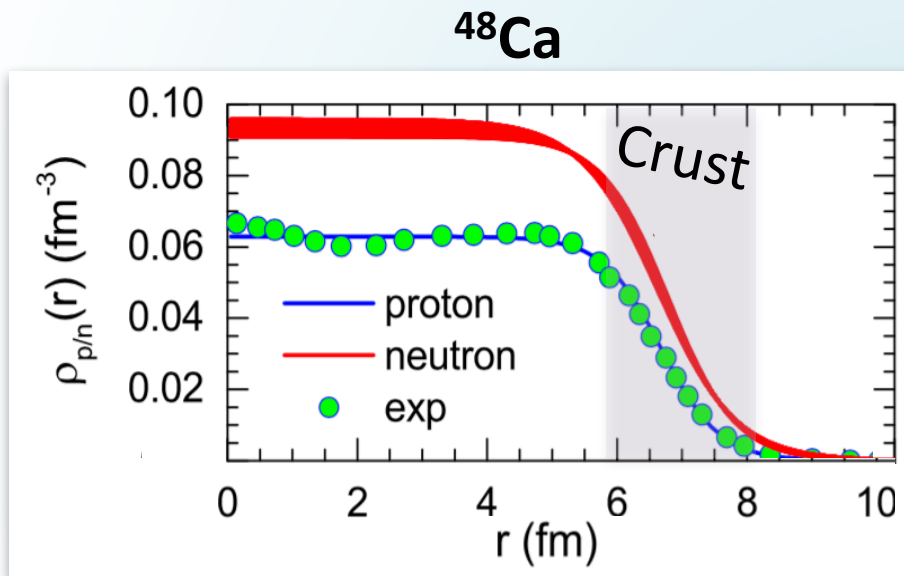
O. Hen (MIT), L. B. Weinstein (ODU),  
D. Higinbotham (JLab), E. Cohen (TAU)





# CaFe: Executive Summary

Focused measurement of short-range pairing in neutron rich nuclei and dynamics of nucleons in the  $^{48}\text{Ca}$  crust.



We will measure mean-field and correlated nucleons using the  $(e, e'p)$  reaction in  $^2\text{H}$ ,  $^{12}\text{C}$ ,  $^{40}\text{Ca}$ ,  $^{48}\text{Ca}$  and  $^{54}\text{Fe}$ .

4 days in Hall-C using standard SHMS and HMS setup.

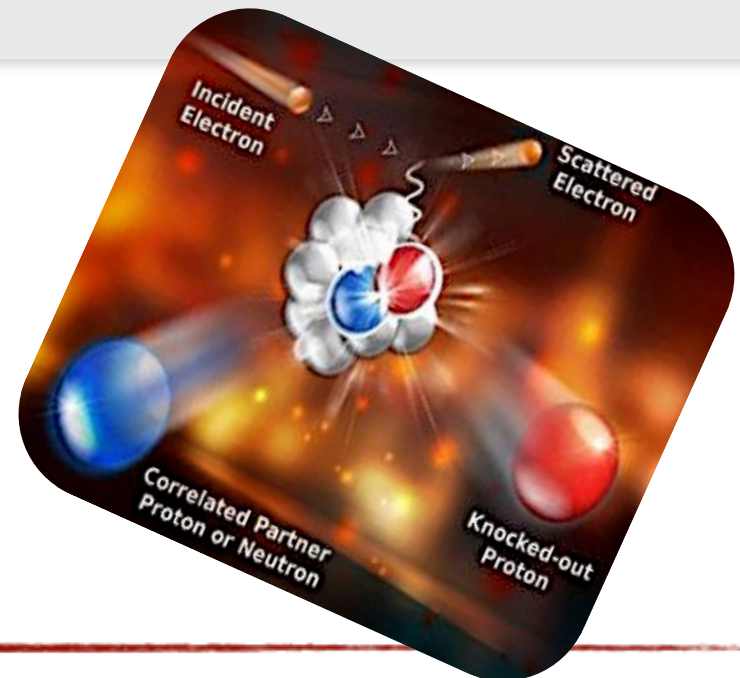
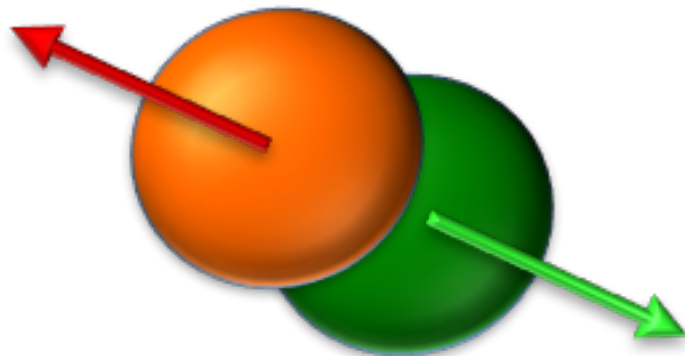


# Short-Range Correlations



Nucleon **pairs** that are close together in the nucleus (wave functions overlap)

=> Momentum space: **pairs** with *high relative momentum* and *low c.m. momentum* compared to the Fermi momentum ( $k_F$ )



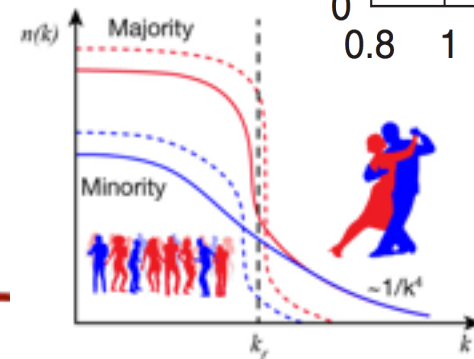
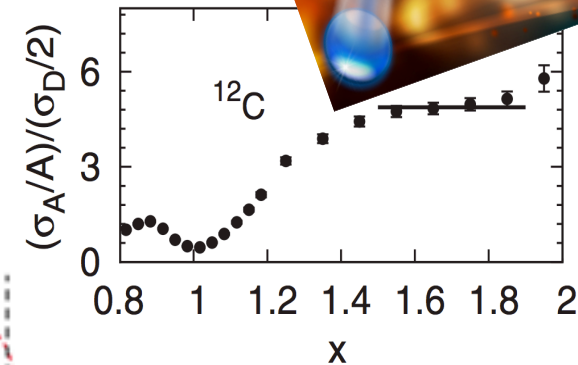
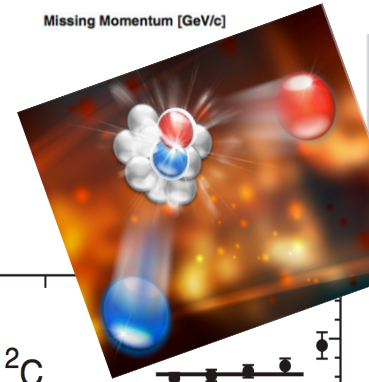
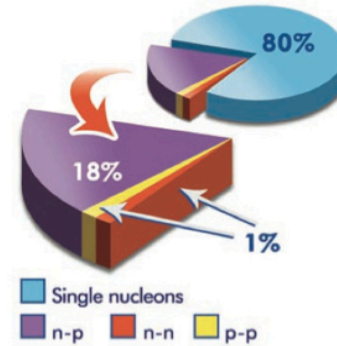
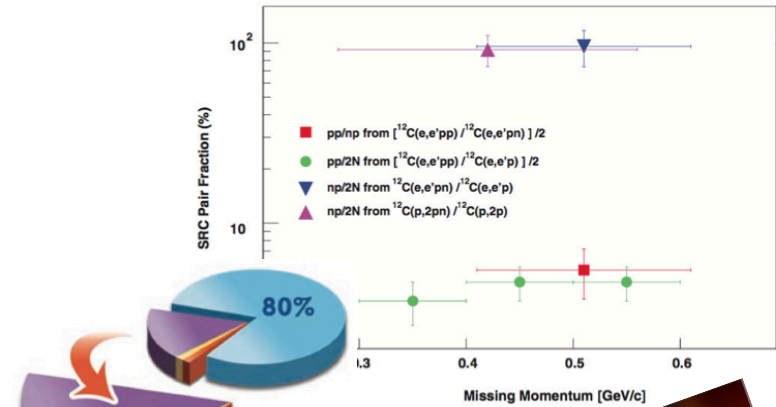


# Short-Range Correlations



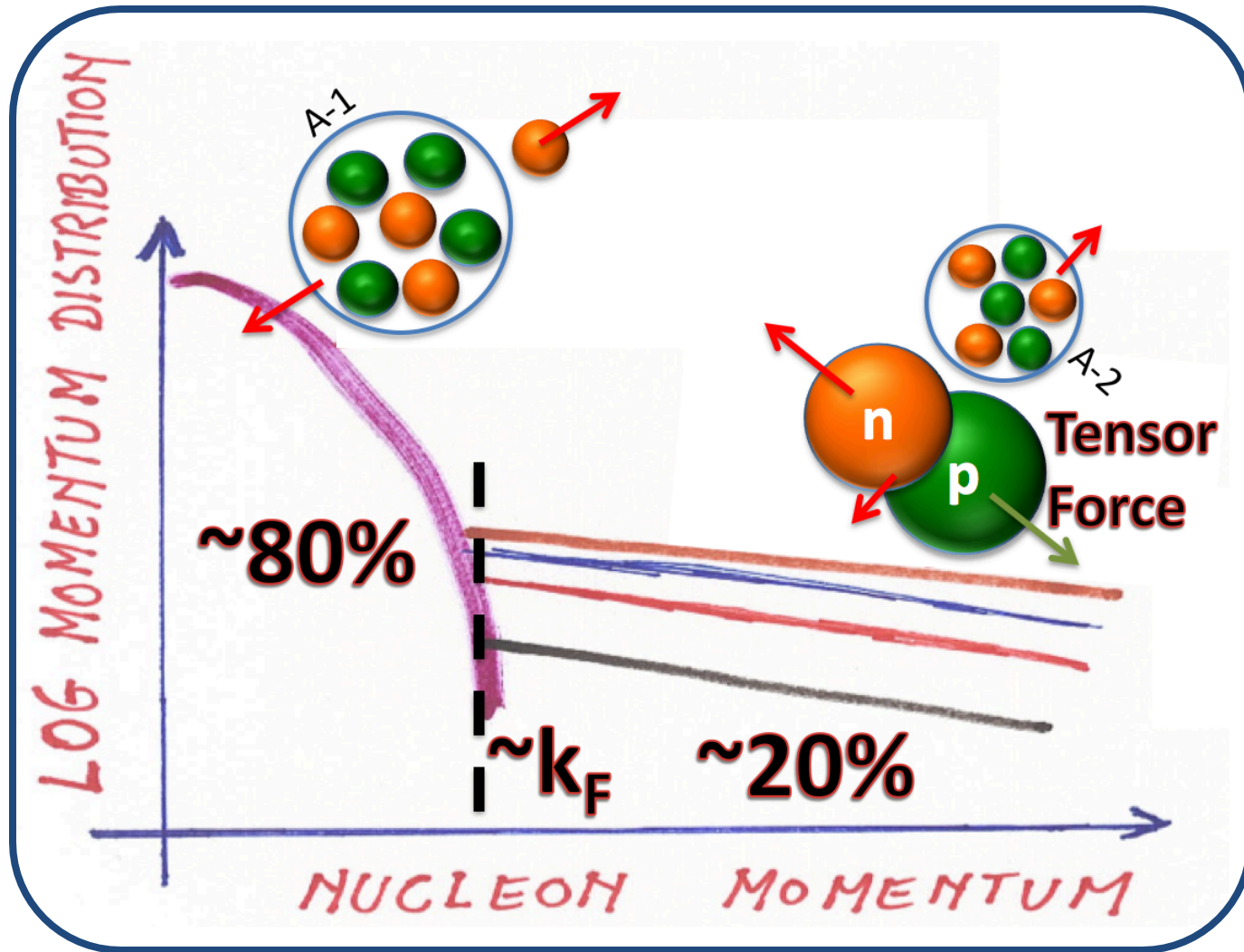
## What we've learned so far:

1. 20 - 25 % of nucleons in nuclei ( $A > 12$ ) have high-momentum ( $k > k_F \sim 250$  MeV/c).
2. High-momentum nucleons predominantly belong to 2N-SRC pairs.
3. 2N-SRC are dominated by np pairs (tensor interaction).





# SRC lead to a 'universal' structure of nuclei in momentum space





# Important to JLab!



## 6 GeV experimental SRC results:

**2 Science, 6 PRLs, RMP (forthcoming), and more...**

+ LOTS of theoretical / phenomenological work inspired by the experimental results.

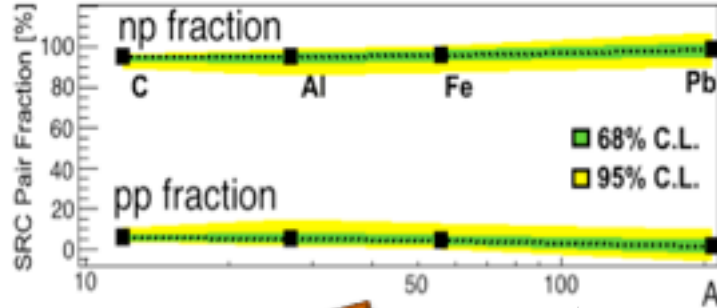
**“A highlight of the JLAB 6 GeV program was the discovery that neutron-proton pairs dominate short-range correlations, SRC, in nuclei.** This is a fundamental piece of information important for achieving a comprehensive model of nuclei. Knowledge of short-range correlations and consequently the resulting high-momentum components of the nucleon wave functions is also important information for experiments that use nuclei, for example in neutrino scattering, that rely on accurately known nuclear properties.” (JLab PAC 41)



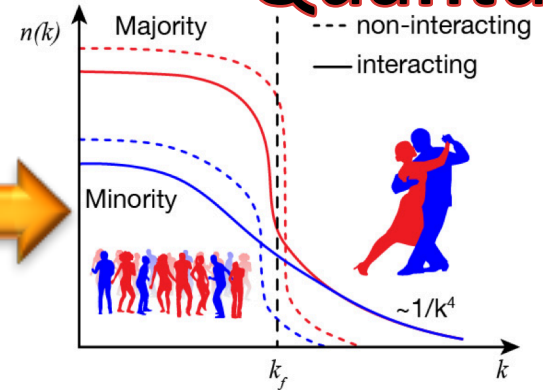
# Important to Physics! 😊



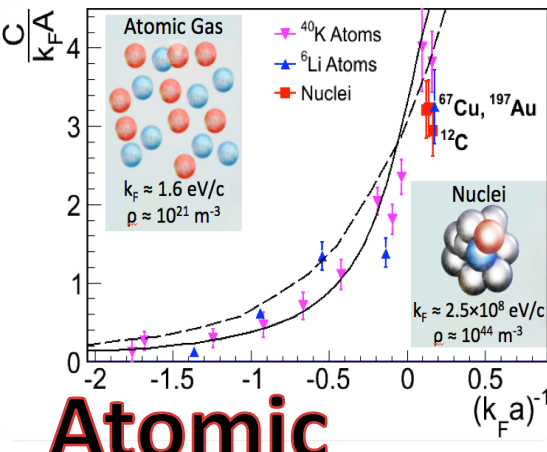
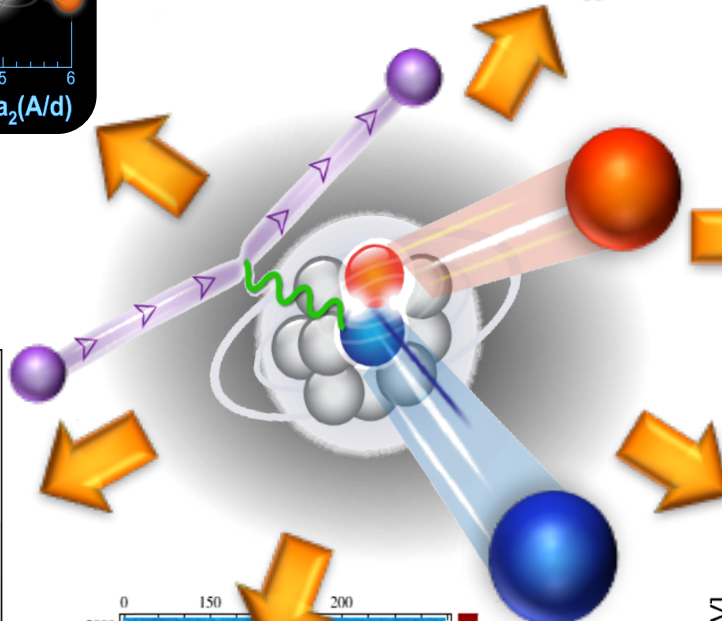
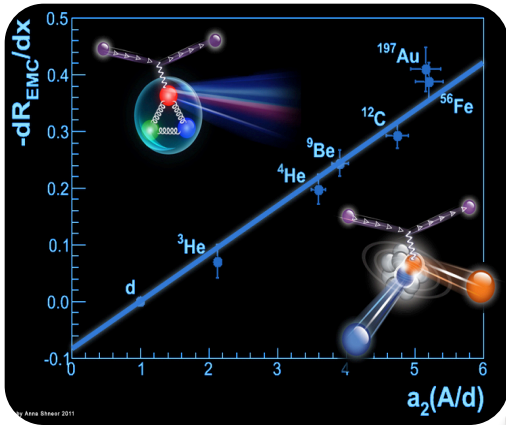
## Nuclear



## Quantum

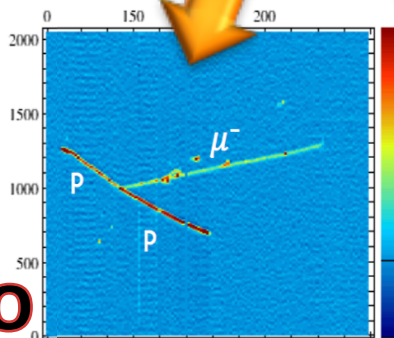


## Particle

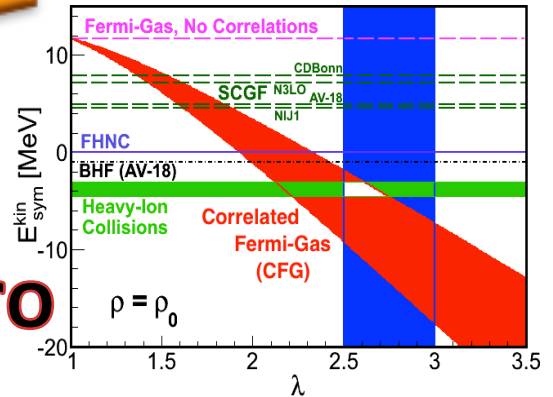


## Atomic

## Neutrino



## Astro

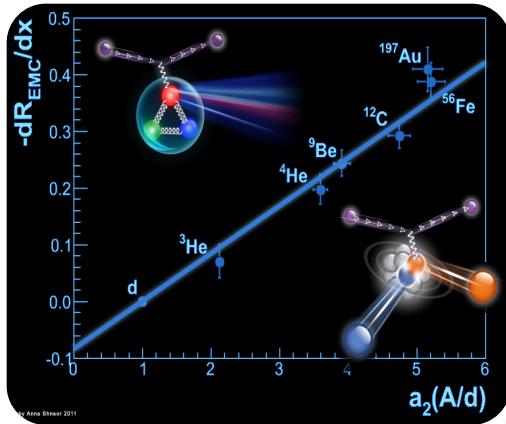




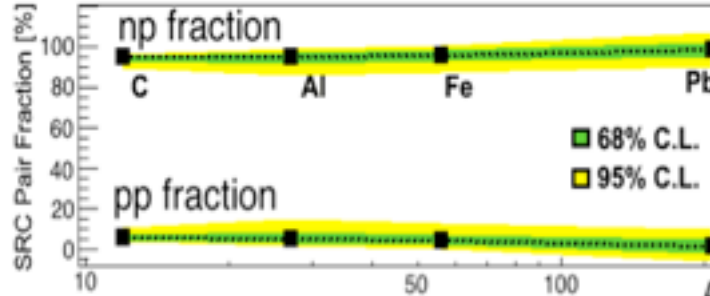
# Important to Physics! 😊



## Particle



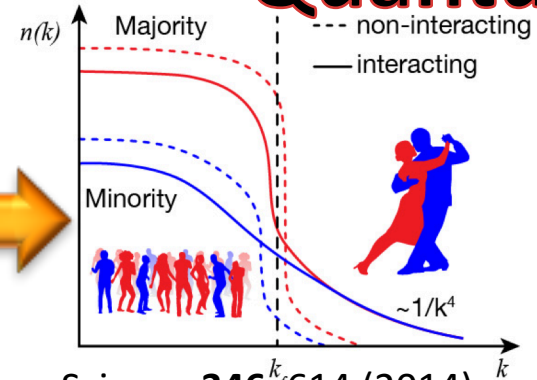
PRL **106**, 052301 (2011),  
 PRD **84**, 117501(2011),  
 PRC **85**, 047301(2012),  
 IJMPA **22**, 1330017 (2013),  
 arXiv 1607.03065 (2016).



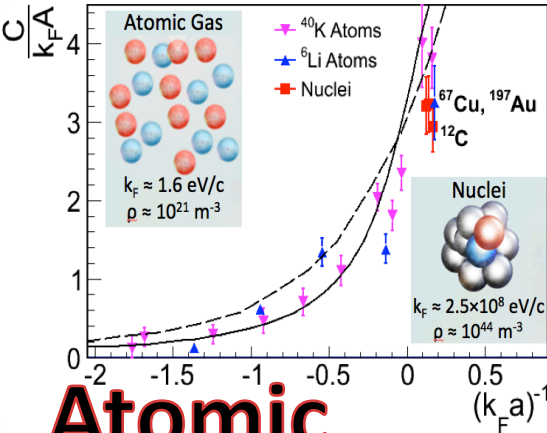
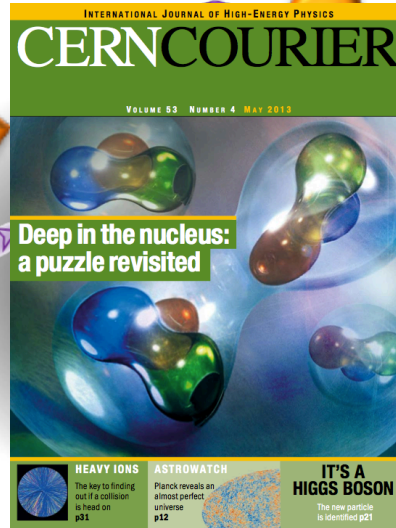
## Nuclear

Science **320**, 1476 (2008),  
 PRL **108**, 092502 (2012),  
 PLB **772**, **63** (2013),  
 PRL **113**, 022501 (2014),  
 PRC **92**, 024604 (2015).

## Quantum



Science **346**, 614 (2014).

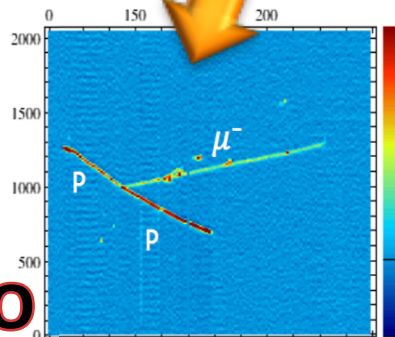


## Atomic

PRC **92**, 045205 (2015).

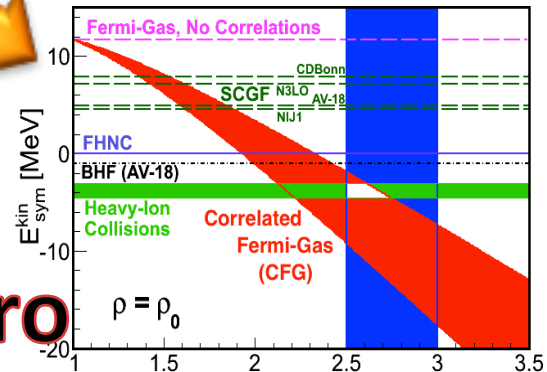
## Neutrino

PRD **90**, 012008 (2014); arXiv 1604.02482 (2016)



## Astro

PRC **91**, 025803 (2015), PRC **93**, 044610 (2016),  
 PRC **91**, 044601 (2015), PRC **93**, 014619 (2016),  
 PRC **92**, 011601 (2015), PLB **759**, 79 (2016),  
 Hen and Steiner et al., In Preparation.







# MOVING FORWARD

A large blue arrow pointing to the right, positioned below the word "FORWARD".

Our focus for today –  
Correlations in neutron rich nuclei

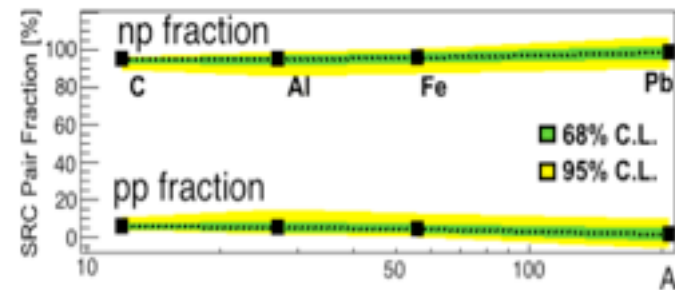
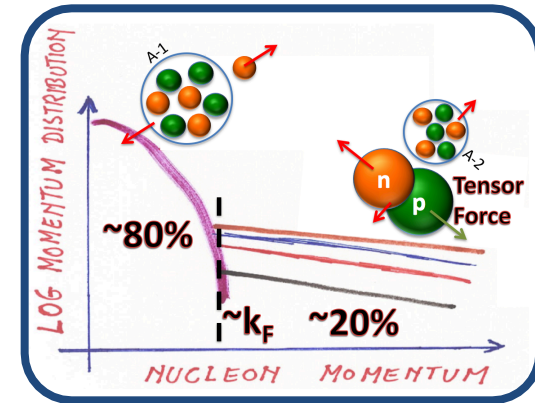
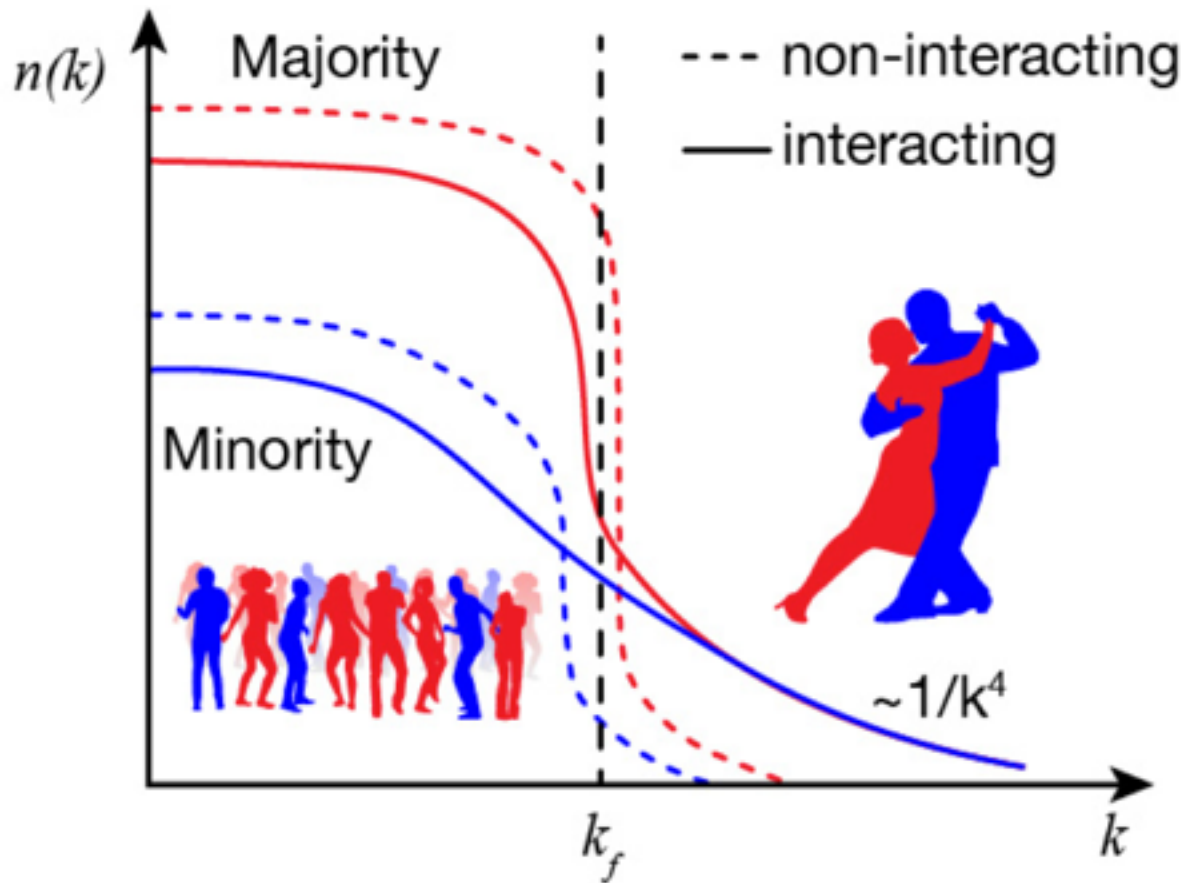




# Asymmetric nuclei



## Proton and neutron momentum distribution in imbalanced systems with short-range interactions





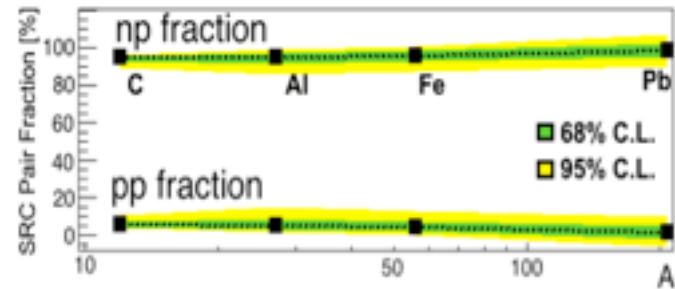
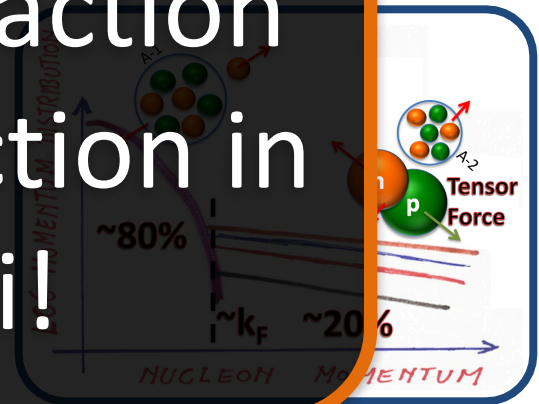
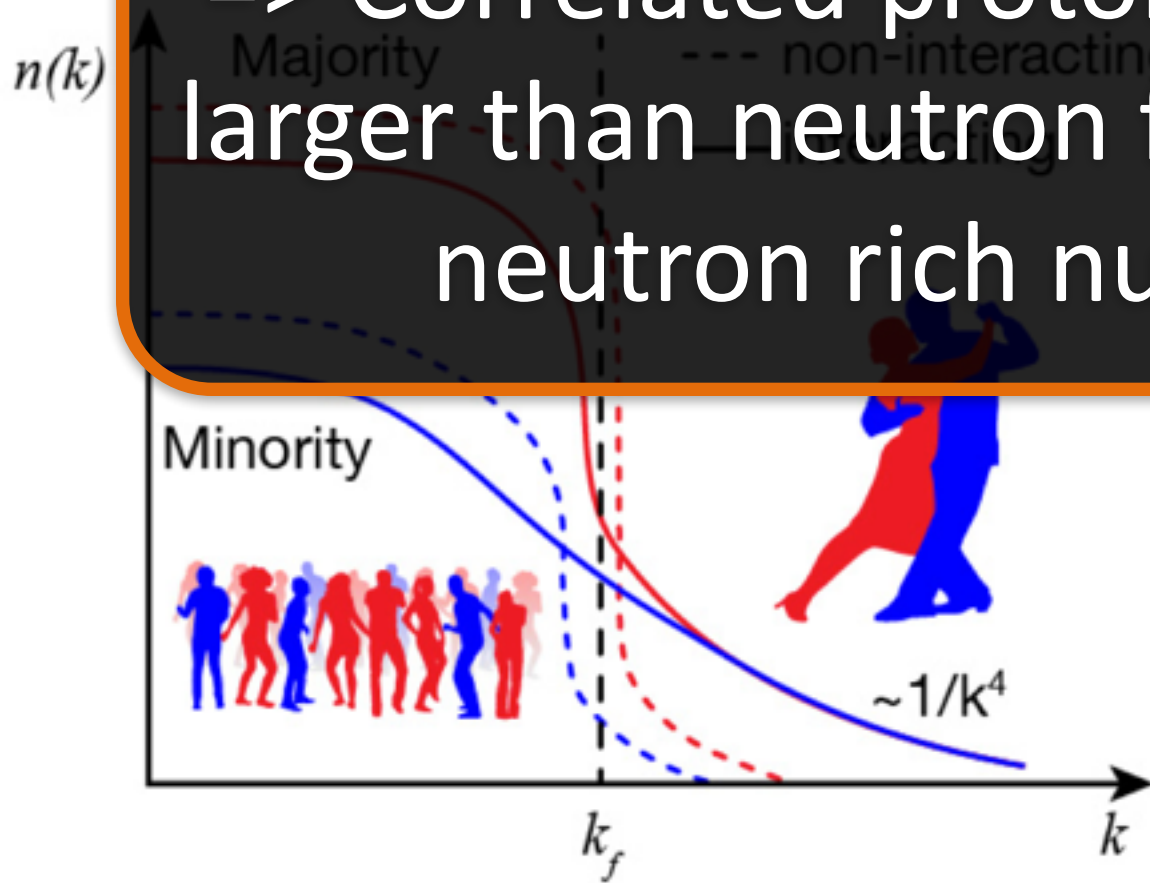
# Asymmetric nuclei



Proton and neutron momentum distribution in

imbalanced systems with short-range interactions

=> Correlated proton fraction larger than neutron fraction in neutron rich nuclei!



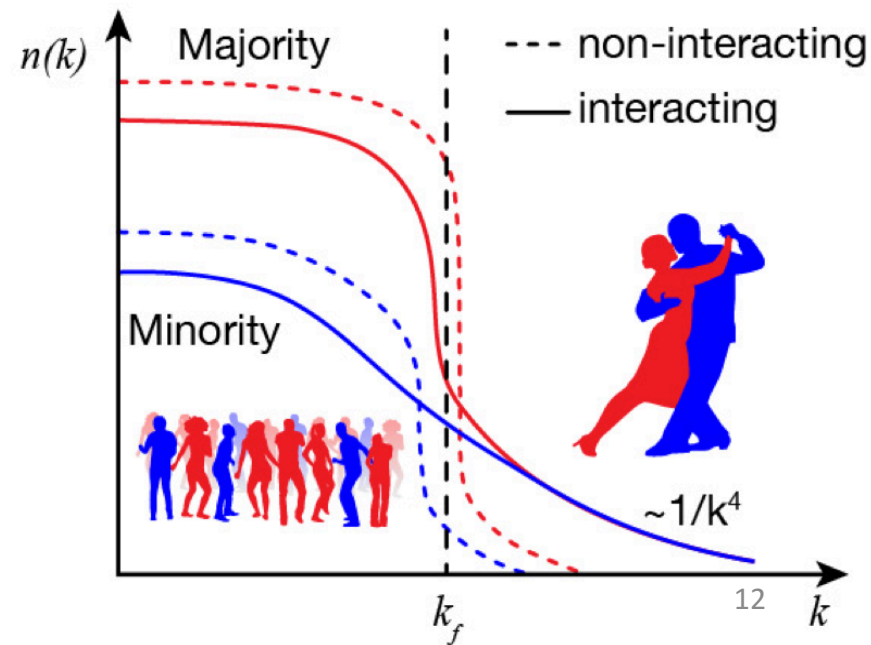
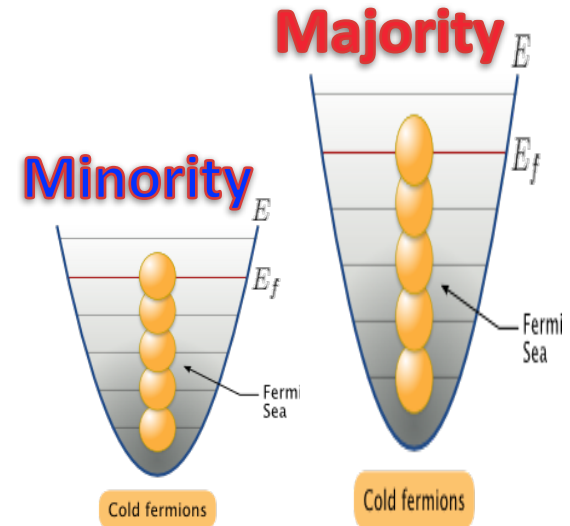
# Kinetic Energy Sharing in Asymmetric Nuclei

## Pauli Principle:

higher Fermi momentum  
=> Majority (**neutrons**)  
fermions move faster.

## np correlations:

greater pairing probability  
=> Minority (**protons**)  
fermions move faster



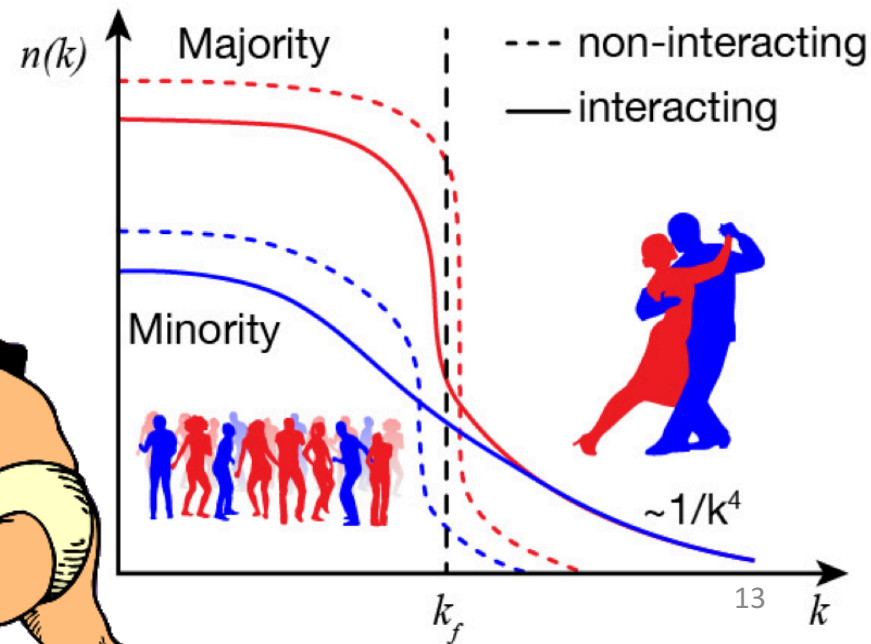
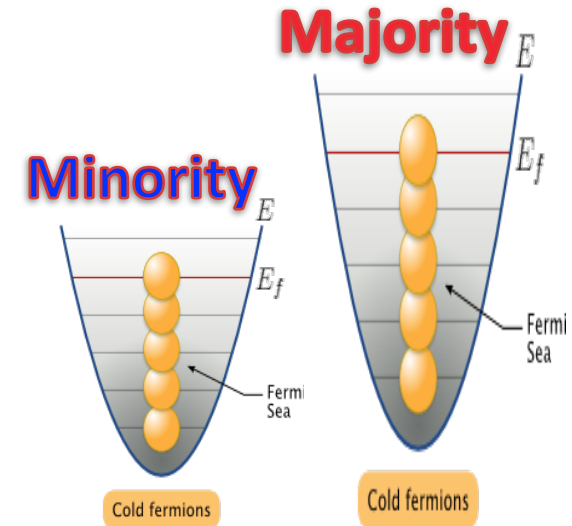
# Kinetic Energy Sharing in Asymmetric Nuclei

## Pauli Principle:

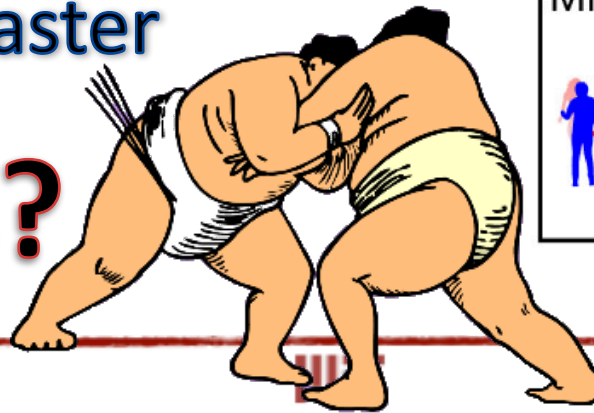
higher Fermi momentum  
=> Majority (**neutrons**)  
fermions move faster.

## np correlations:

greater pairing probability  
=> Minority (**protons**)  
fermions move faster



# Who wins?



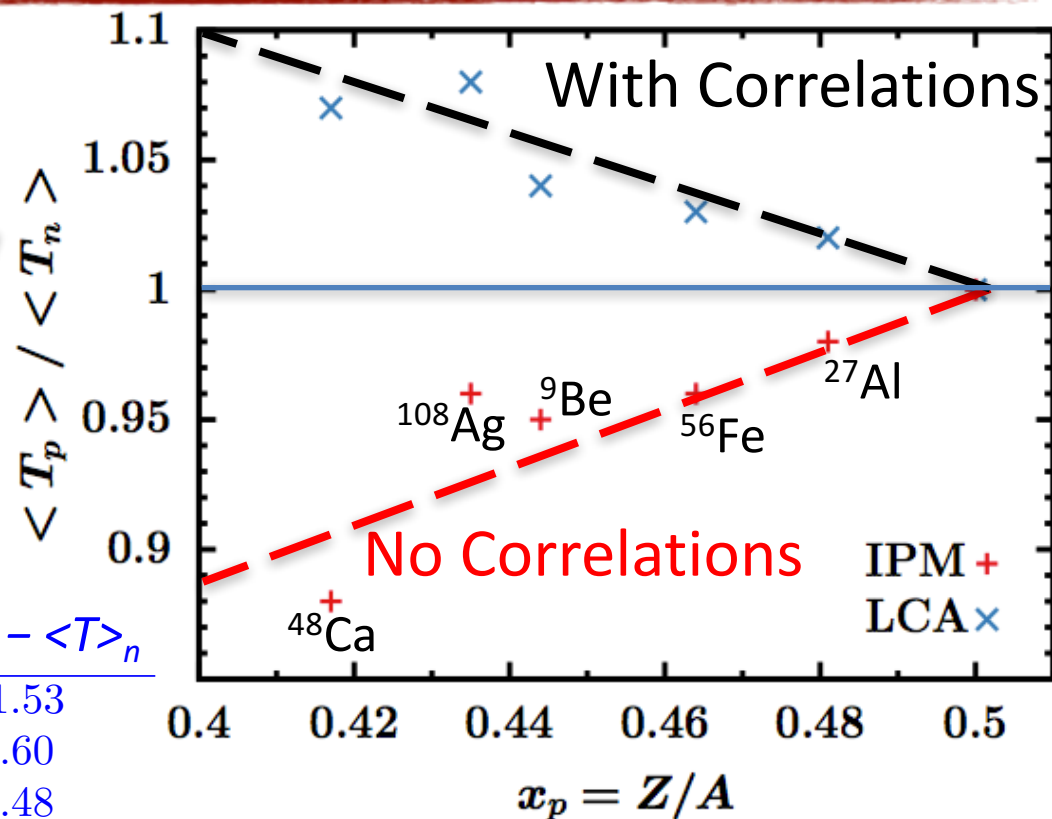


# Calculations Predict Correlations Win

$$\langle T \rangle_{\text{Minority}} \geq \langle T \rangle_{\text{Majority}}$$

Light Nuclei ( $A < 12$ )

	$\frac{ N-Z }{A}$	$\langle T \rangle_p$	$\langle T \rangle_n$	$\langle T \rangle_p - \langle T \rangle_n$
$^8\text{He}$	0.50	30.13	18.60	11.53
$^6\text{He}$	0.33	27.66	19.06	8.60
$^9\text{Li}$	0.33	31.39	24.91	6.48
$^3\text{He}$	0.33	14.71	19.35	-4.64
$^3\text{H}$	0.33	19.61	14.96	4.65
$^8\text{Li}$	0.25	28.95	23.98	4.97
$^{10}\text{Be}$	0.2	30.20	25.95	4.25
$^7\text{Li}$	0.14	26.88	24.54	2.34
$^9\text{Be}$	0.11	29.82	27.09	2.73
$^{11}\text{B}$	0.09	33.40	31.75	1.65



Heavy Nuclei ( $27 < A < 108$ ):

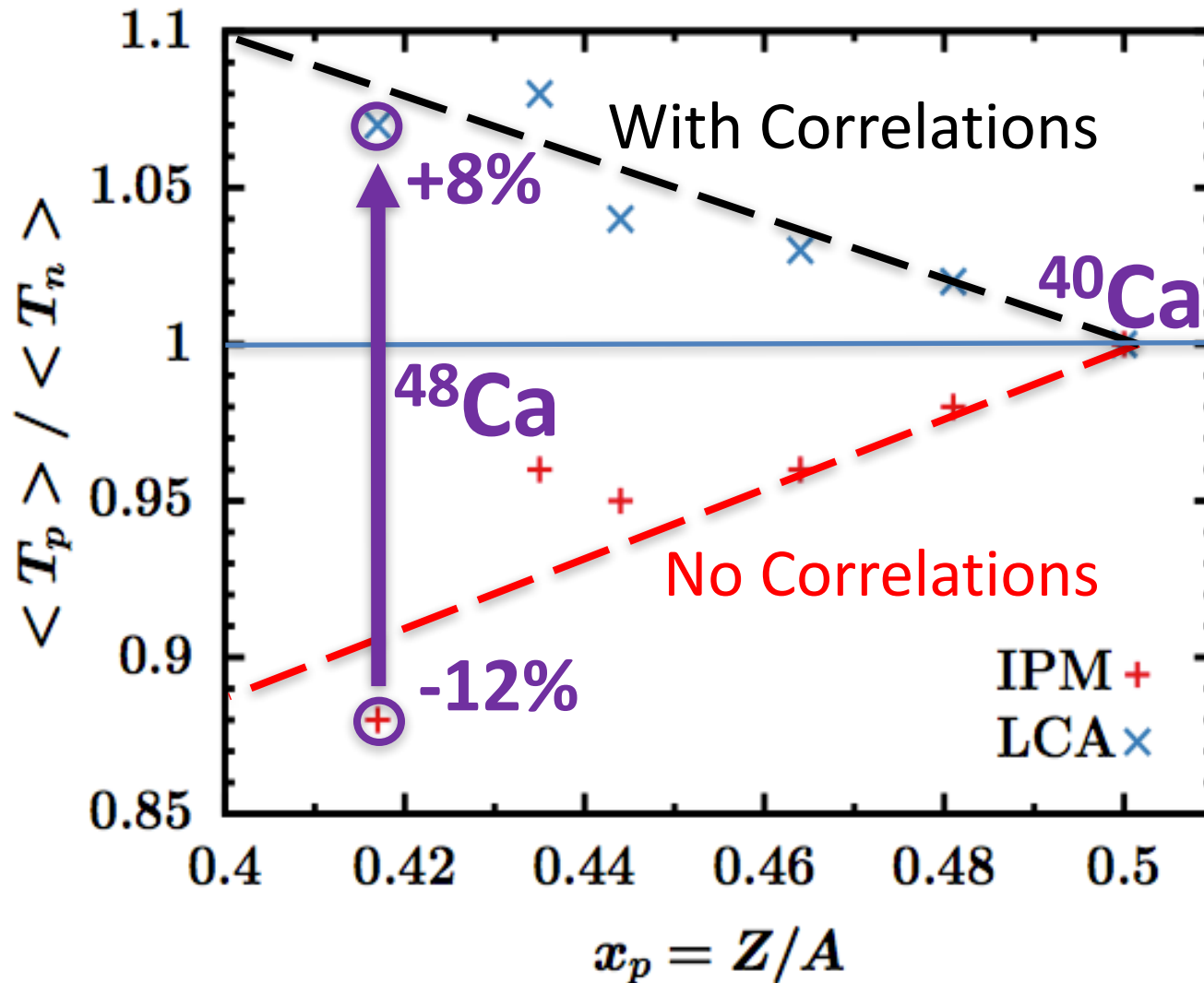
M. Vanhalst, W. Cosyn, and J.

Ryckebusch, J. Phys. G **42**, 055104 (2015)





# Focusing on $^{48}\text{Ca}$



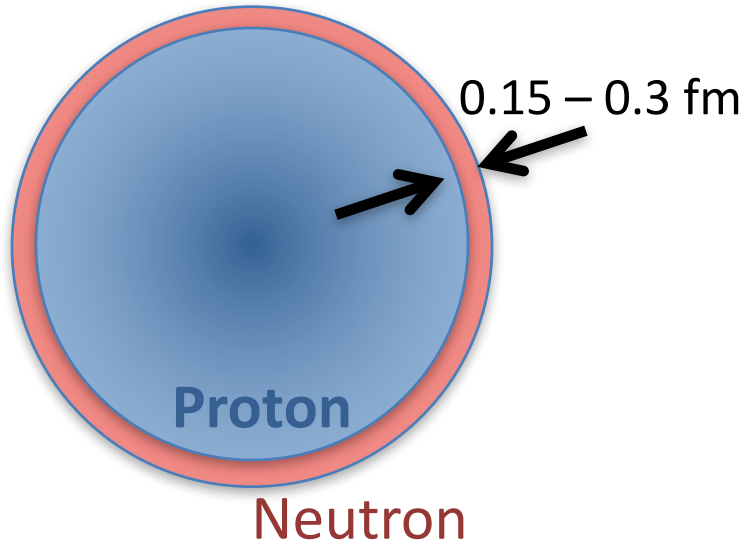


# Focusing on $^{48}\text{Ca}$



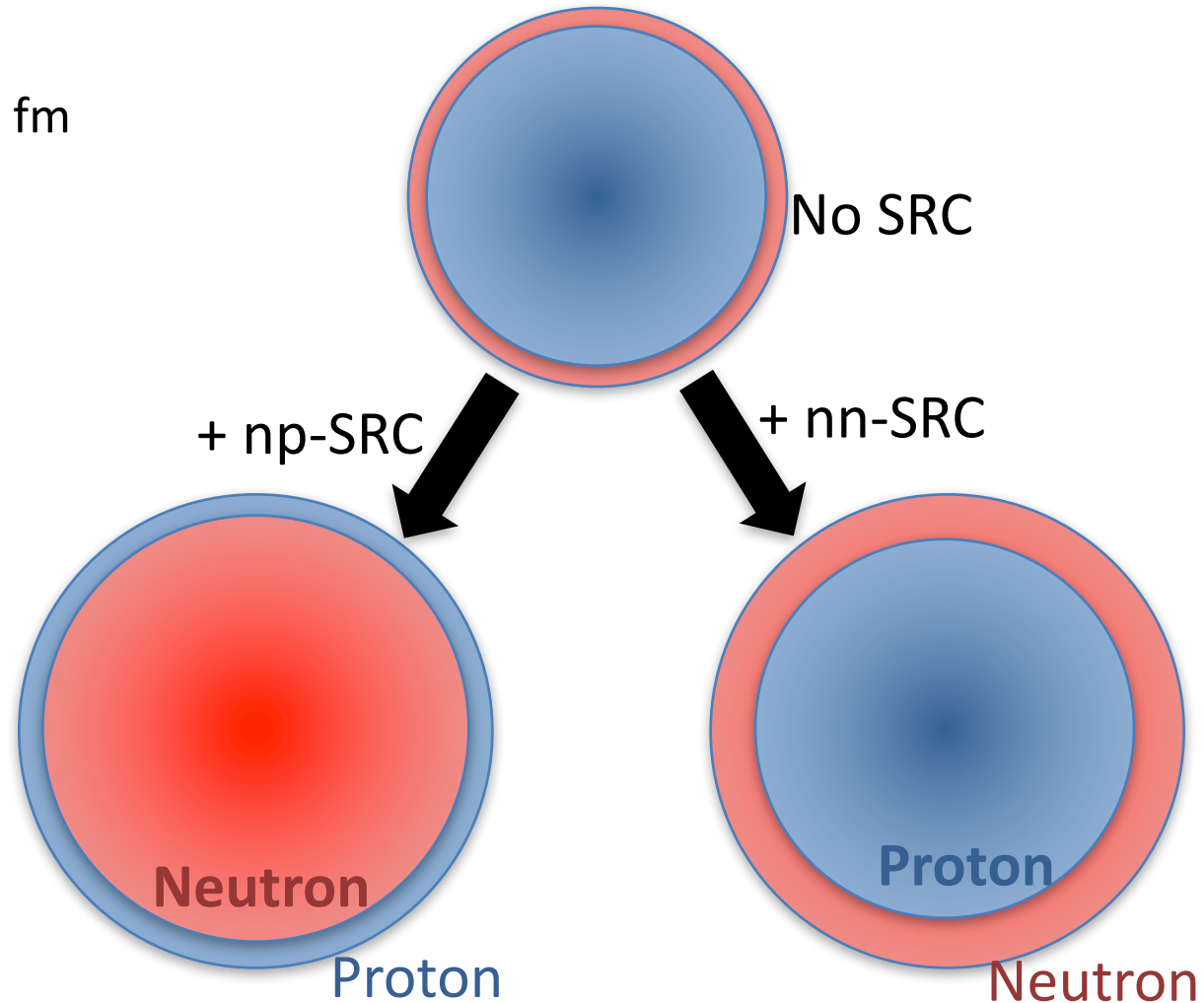
## Coordinate space:

[CREX]



## Momentum space:

[CaFe]



## Adding correlations:

- Reduce the radius.
- Inverts the momentum skin?



# Focusing on $^{48}\text{Ca}$

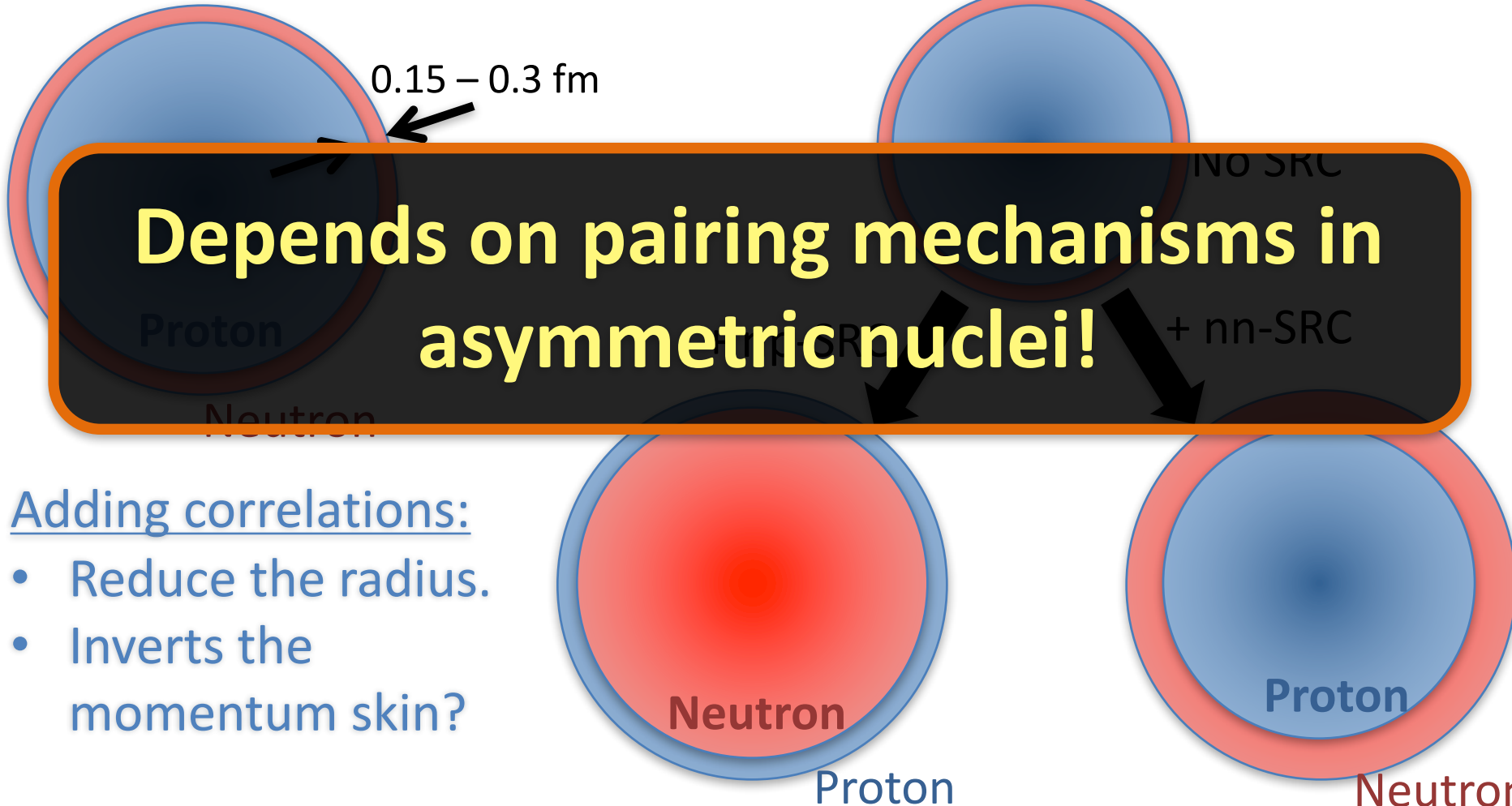


## Coordinate space:

[CREX]

## Momentum space:

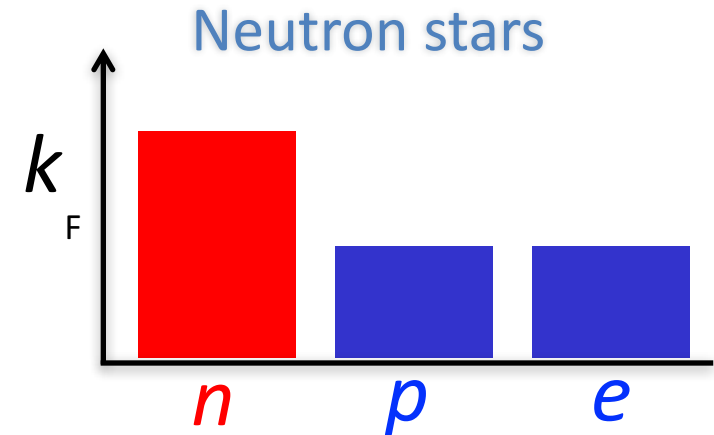
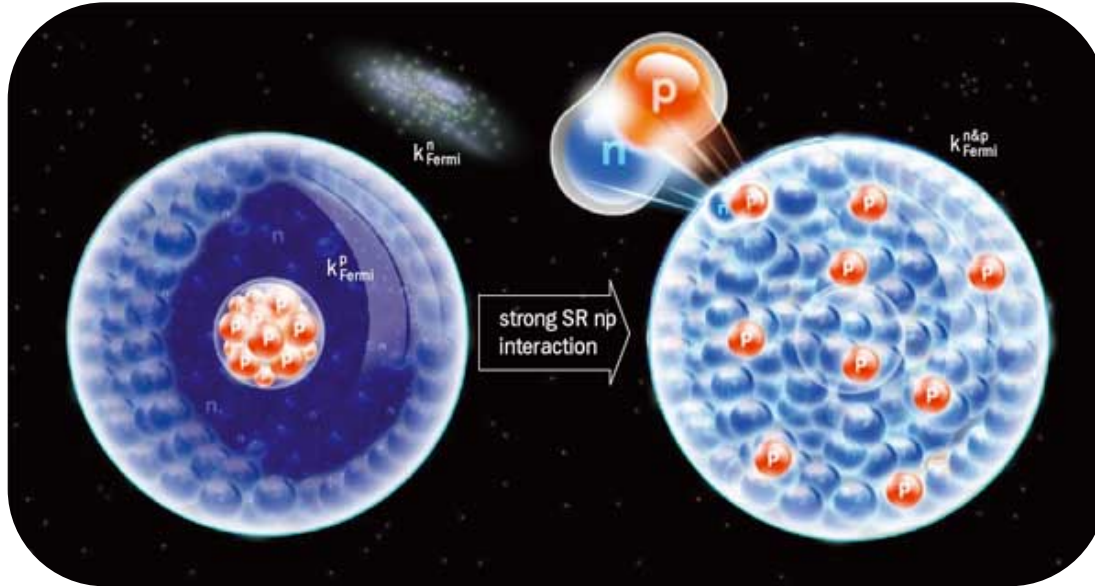
[CaFe]



**Depends on pairing mechanisms in asymmetric nuclei!**

### Adding correlations:

- Reduce the radius.
- Inverts the momentum skin?



## np-SRC pairs

- Create high-momentum protons and open 'holes' in the neutron Fermi sphere.
  - Change the URCA process cooling rate?
  - Change the spatial distribution?
- **Implementing into cooling calculations now!**



Relates to the energy change when replacing n with p.

- equation-of-state of neutron stars,
- heavy-ion collisions,
- r-process nucleosynthesis,
- core-collapse supernovae,
- more...

How can JLab help?

- **CREX / PREX** constrain the slope [  $L^\infty (dE_{\text{sym}}/d\rho) |_{\rho_0}$  ].
- **Pairing** changes the kinetic / potential balance.

## CaFe will measure the pairing mechanism

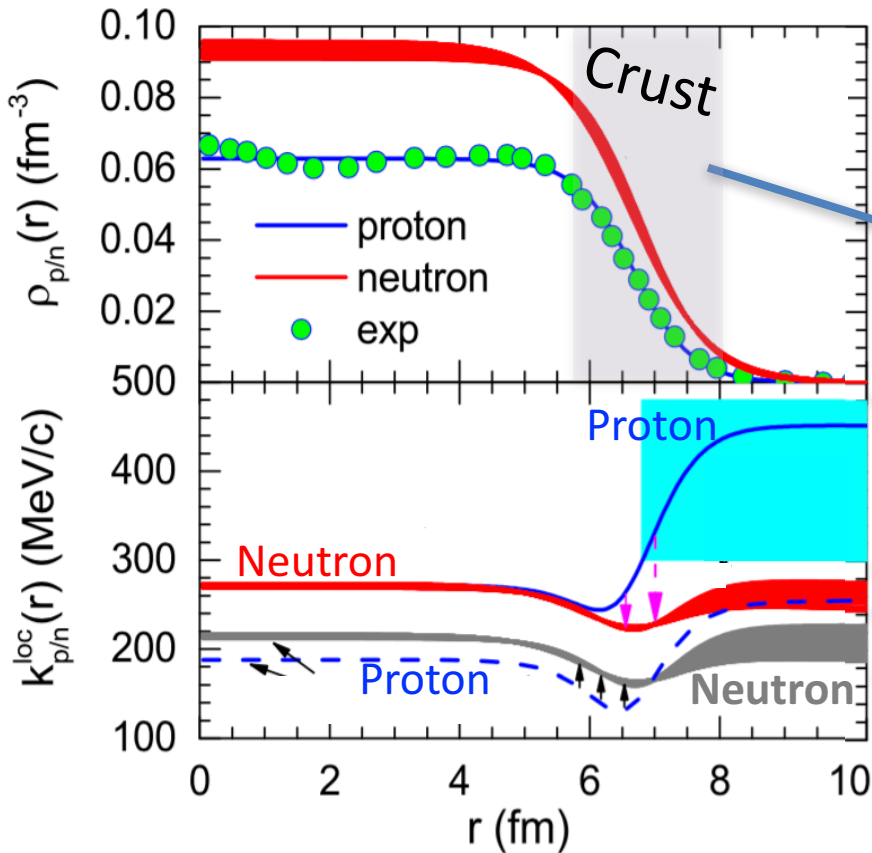




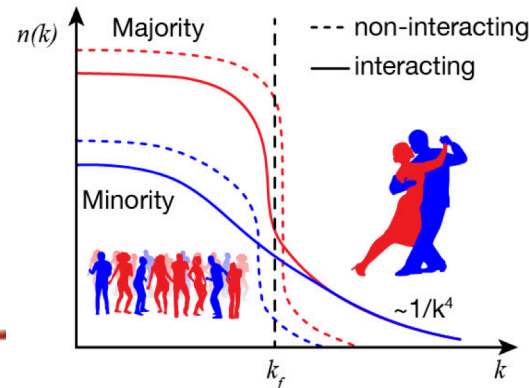
# Pairing and... Neutron Skins



## Protons move faster in the neutron skin



Protons in the **crust** 'feel' one of the largest neutron excess reachable by terrestrial experiments



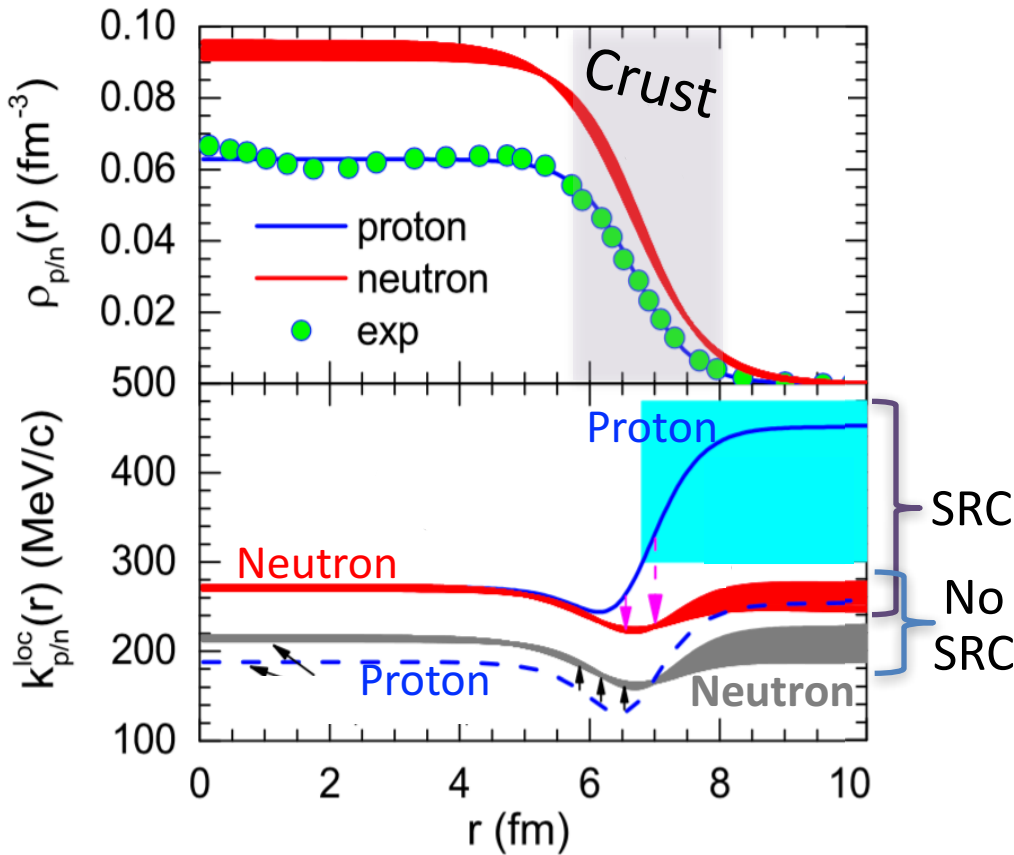
B.J. Cai et al., arXiv: 1606.08045 (2016).



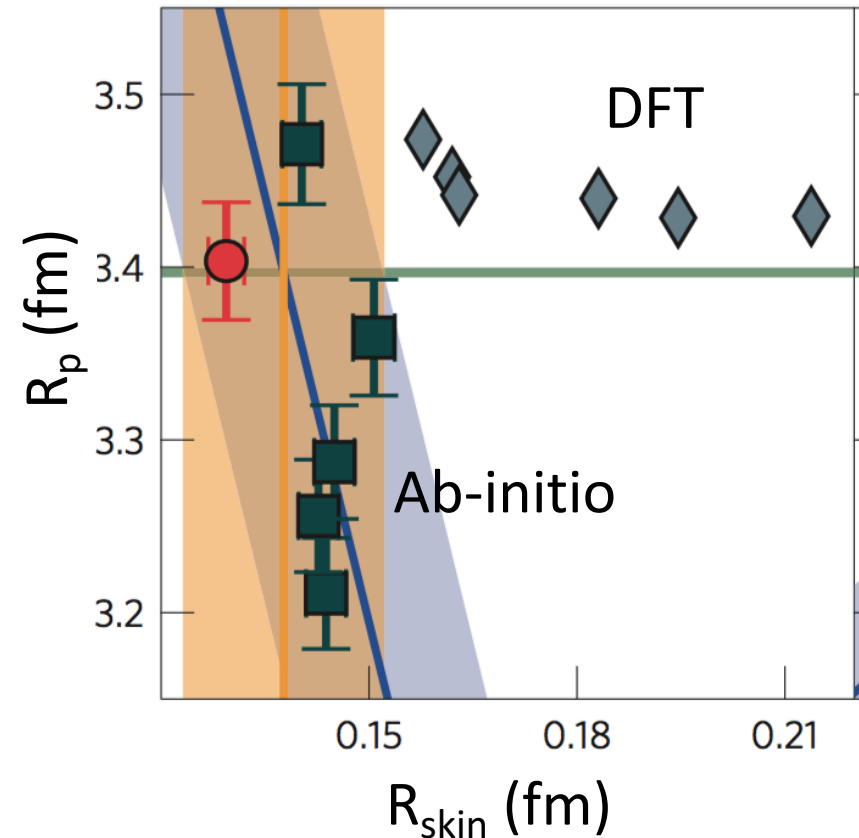
# Pairing and... Neutron Skins



## Protons move faster in the neutron skin



## Do SRC change the neutron skin?



B.J. Cai et al., arXiv: 1606.08045 (2016).

G. Hagen et al., Nature Physics 12, 186 (2016)



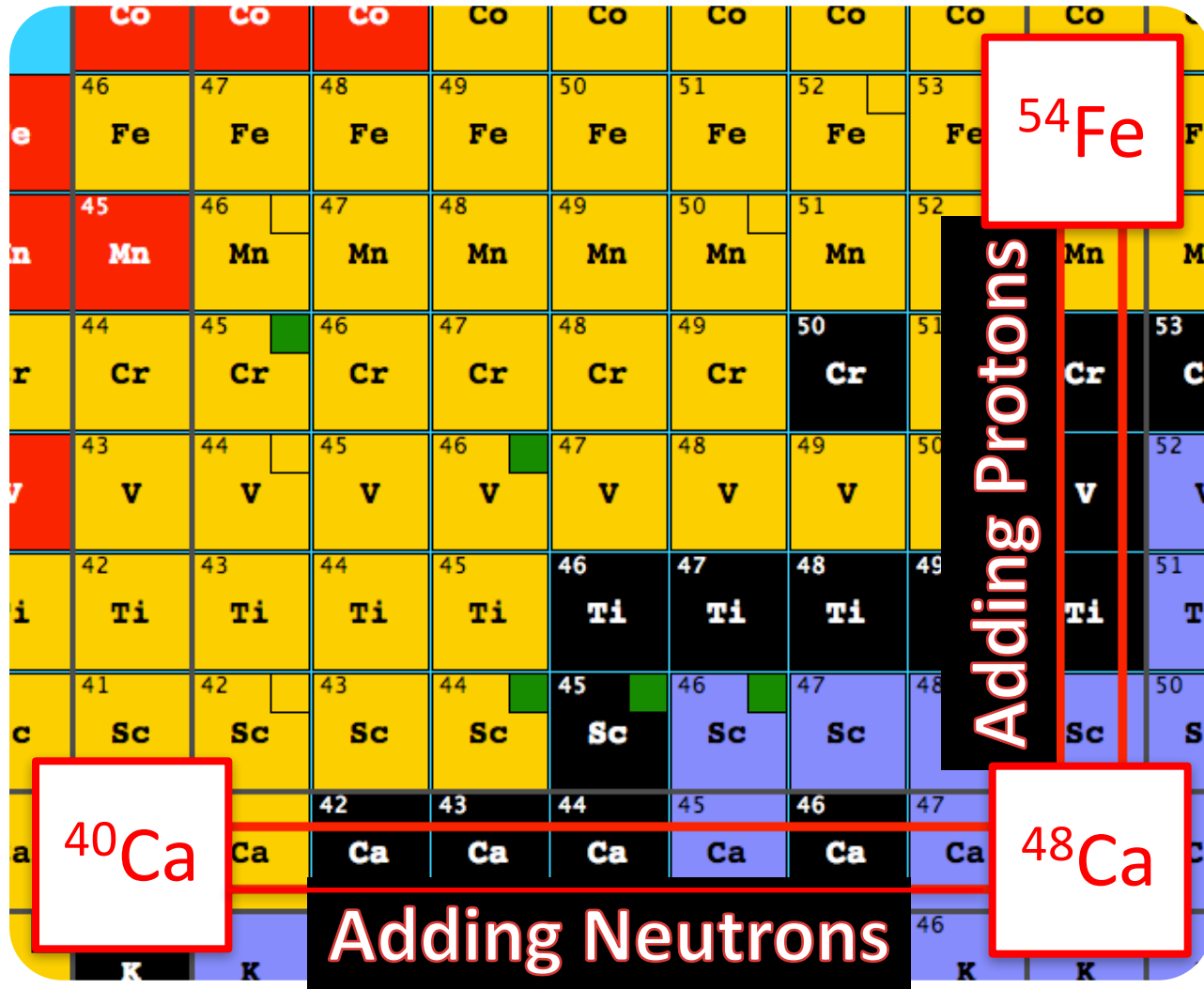
# what's the Problem?

These effects depend quantitatively on the short-range pairing mechanism in asymmetric nuclei.

**=> In comes CaFe!**



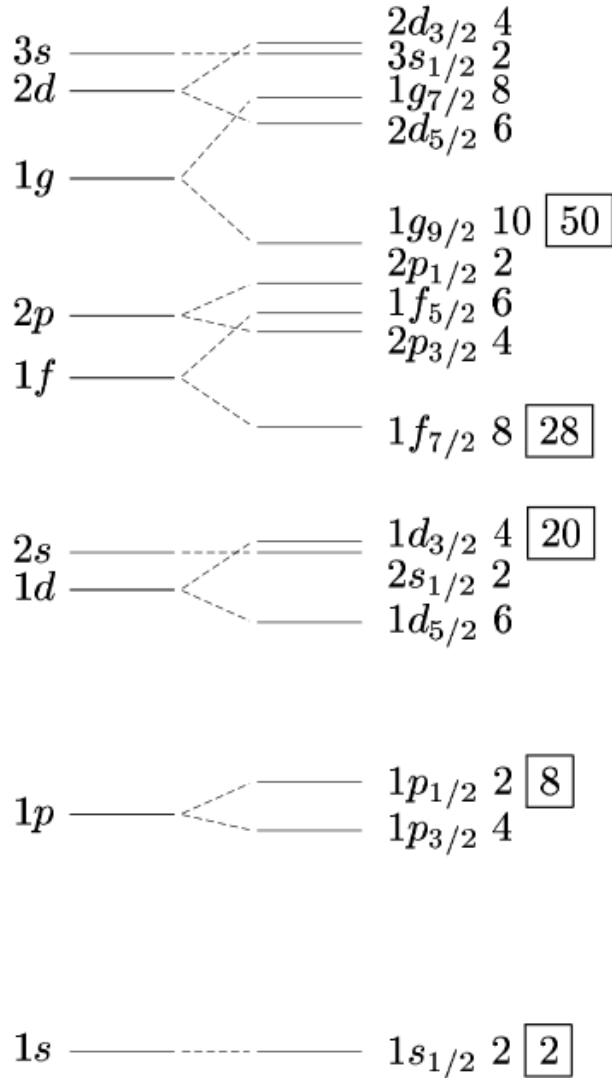
# The CaFe Triplet: A Lab for Asymmetric Nuclei



$^{48}\text{Ca}$  has a  
40% neutron  
excess!!



# The CaFe Triplet: A Lab for Asymmetric Nuclei



Nucleus	Z	N	
<sup>40</sup> Ca	20	20	Symmetric double magic
<sup>48</sup> Ca	20	28	+ Full neutron shell (1f <sub>7/2</sub> )
<sup>54</sup> Fe	26	28	Almost symmetric double magic

How do the neutrons from the outer 1f<sub>7/2</sub> shell correlate with the <sup>40</sup>Ca core?

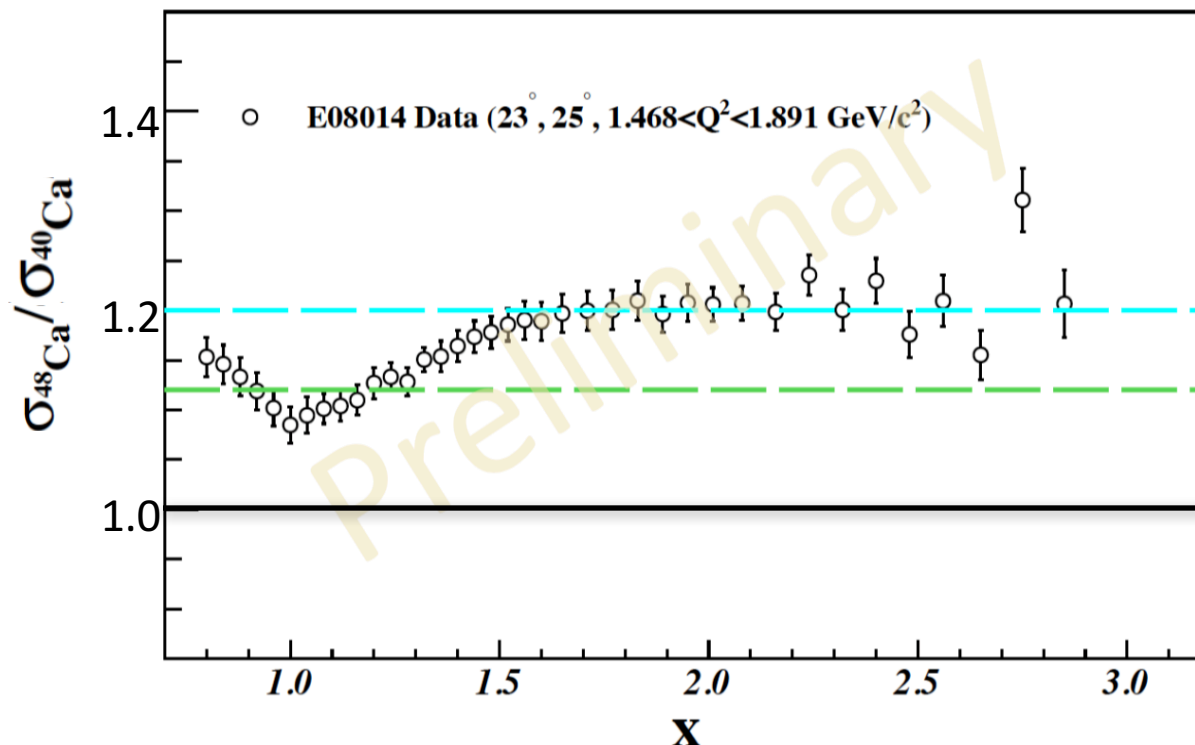




# What do we already know?



( $e, e'$ ) cross-section ratios at  $x_B > 1$  are sensitive to the TOTAL NUMBER OF SRC PAIRS:



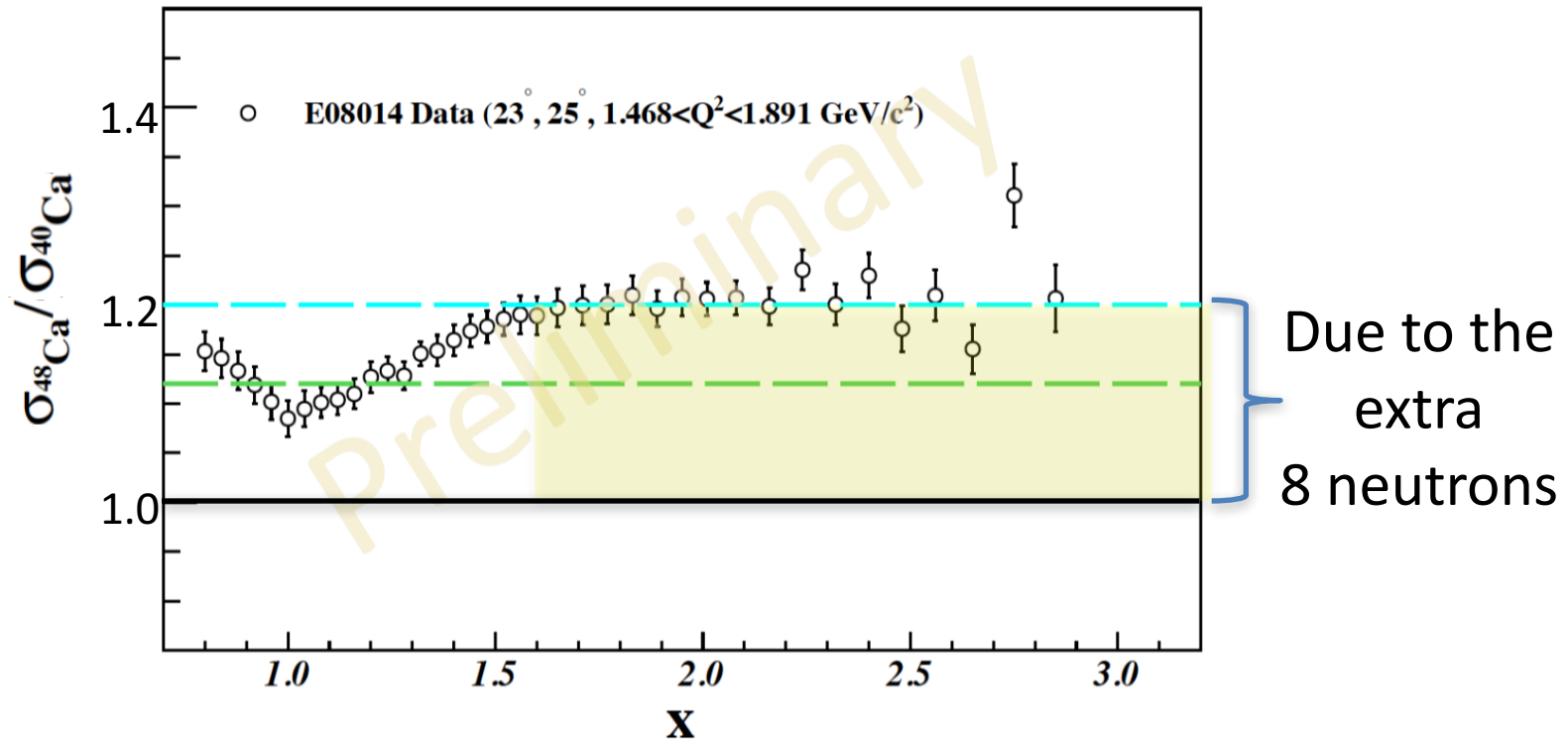
$\Rightarrow$   $^{48}\text{Ca}$ : + 20% nucleons, +20% SRC pairs!



# What do we already know?



( $e, e'$ ) cross-section ratios at  $x_B > 1$  are sensitive to the TOTAL NUMBER OF SRC PAIRS:



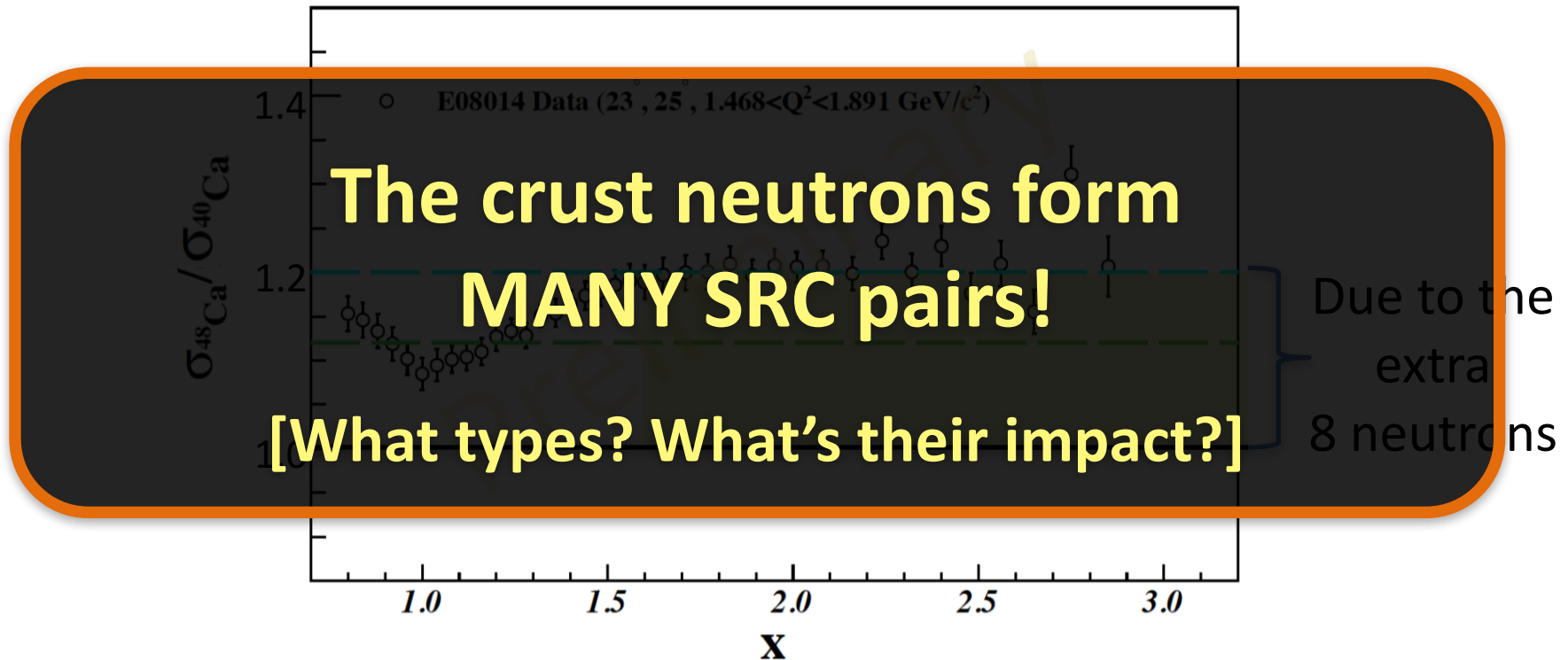
The neutrons in the outer  $1f_{7/2}$  shell (i.e. in the skin) are *equally correlated* as the nucleons in the  $^{40}\text{Ca}$  core!



# What do we already know?



( $e, e'$ ) cross-section ratios at  $x_B > 1$  are sensitive to the TOTAL NUMBER OF SRC PAIRS:



The neutrons in the outer  $1f_{7/2}$  shell (i.e. in the skin) are *equally correlated* as the nucleons in the  $^{40}\text{Ca}$  core!



# Correlations in the skin



Can  $(e, e')$  also tell us what type of pairs the  $1f_{7/2}$  neutrons create?  
Unfortunately No...

Examine two extreme hypothesis:

All pairs can form SRC pairs:

$$\sigma_{(e, e')|x>1} \propto (\sigma_p + \sigma_n)(NZ) + 2\sigma_p(Z \cdot (Z - 1)/2) + 2\sigma_n(N \cdot (N - 1)/2)$$

Only np pairs can form SRC pairs:

$$\sigma_{(e, e')|x>1} \propto (\sigma_p + \sigma_n)(NZ)$$

All-Pairs / np-Pairs = 0.94



VERY hard to discriminate



# Correlations in the skin



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Only np pairs can form SRC pairs:

$$\sigma_{(e, e')|x>1} \propto (\sigma_p + \sigma_n)(NZ)$$

All-Pairs / np-Pairs = 0.94



VERY hard to discriminate

The  $(e, e')$  reaction can not distinguish protons and neutrons  $\rightarrow$  need  $(e, e'p)$  !





# Correlations in the skin



Can **(e,e'p)** tell us what type of pairs the  $1f_{7/2}$  neutrons create? **Yes!**

Ratios of correlated protons in different nuclei

Model	$^{40}\text{Ca}/\text{C}$	$^{48}\text{Ca}/^{40}\text{Ca}$	$^{54}\text{Fe}/^{48}\text{Ca}$	$^{54}\text{Fe}/^{40}\text{Ca}$
1 – all protons	1.7	1	1.3	1.3
2 – all pairs	11.7	1.23	1.44	1.77
3 – all $np$ pairs	10.1	1.4	1.30	1.82
4 – $S$ and $P$ $np$ pairs	10.4	1.31	1.32	1.73
5 – $l = 0, n = 0$ $np$ pairs	5.3	1.20	1.22	1.47

The sensitivity of the  $(e,e'p)$  reaction gives  
**large effects!**



# Correlations in the skin



Can **(e,e'p)** tell us what type of pairs the  $1f_{7/2}$  neutrons create? **Yes!**

Ratios of correlated protons in different nuclei

Model	$40\text{Ca}/\text{C}$	$48\text{Ca}/40\text{Ca}$	$54\text{Fe}/48\text{Ca}$	$54\text{Fe}/40\text{Ca}$
1 – all protons	1.7	1	1.3	1.3
2 – all pairs	1.17	1.25	1.44	1.77
3 – all $np$ pairs	1.0	1.0	1.0	1.82
4 – $S$ and $P$ $np$ pairs	10.4	1.31	1.32	1.73
5 – $l = 0, n = 0$ $np$ pairs	5.3	1.20	1.22	1.47

**How can (e,e'p) study the correlated protons?**

The sensitivity of the  $(e,e'p)$  reaction gives **large effects!**

C. Colle and O. Hen et al., Phys. Rev. C **92**, 024604 (2015)

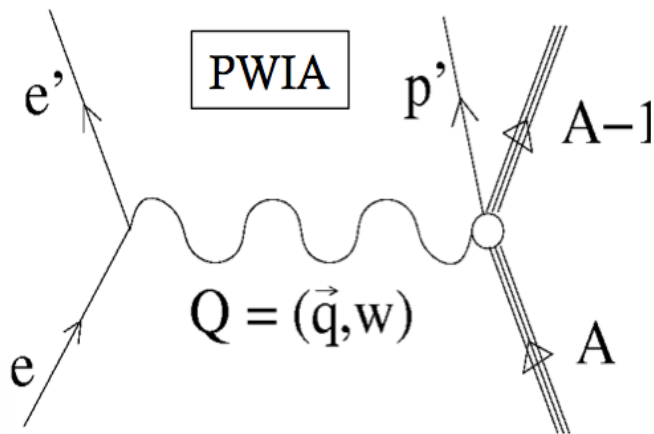


# (e,e'p) - studying the protons



$$\sigma_{(e,e'p)} = k \cdot \sigma_{ep} \cdot S_p(E_{miss}, P_{miss}) \cdot T$$

Kinematical Factor
e-p Cross Section
Spectral Function
“Transparency”



## Complications:

- Rescattering of the outgoing proton.
  - Off-shell proton cross-section.
  - Meson Exchange Currents (MEC).
  - Delta production (i.e. IC).
- => Spectral function is not an observable!**

$$E_{miss} = \omega - T_p - T_{A-1}$$

$$\vec{p}_{miss} = \vec{q} - \vec{p}' = -\vec{p}_{init}$$

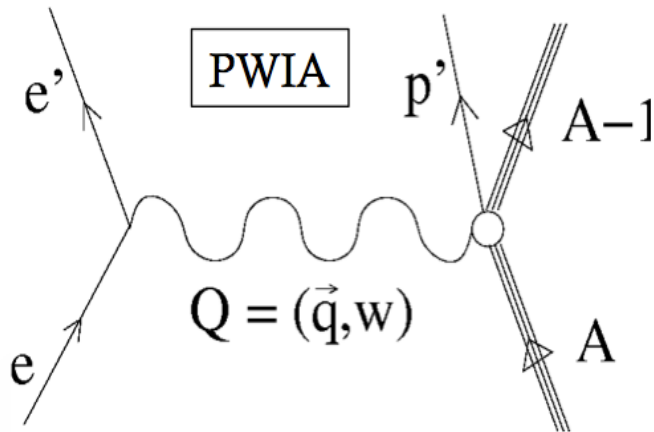


# (e,e'p) - studying the protons



$$\sigma_{(e,e'p)} = k \cdot \sigma_{ep} \cdot S_p(E_{miss}, P_{miss}) \cdot T$$

Kinematical Factor
e-p Cross Section
Spectral Function
“Transparency”



## Complications:

- Rescattering of the outgoing proton.
  - Off-shell proton cross-section.
  - Meson Exchange Currents (MEC).
  - Delta production (i.e. IC).
- => Spectral function is not an observable!**

$$E_{miss} = \omega - T_p - T_{A-1}$$

$$\vec{p}_{miss} = \vec{q} - \vec{p}' = -\vec{p}_{init}$$

## Solution:

- Choosing the ‘right’ kinematics,
- Integrate over a wide  $P_{miss}$  range,
- Extract cross-section ratios.



# Choosing Kinematics



- $E_{\text{beam}} = \mathbf{11 \text{ GeV}}$  @ 40  $\mu\text{A}$  to maximize rates.
- $^1\text{H}$ ,  $^2\text{H}$ ,  $^{12}\text{C}$ ,  $^{40}\text{Ca}$ ,  $^{48}\text{Ca}$ , and  $^{54}\text{Fe}$  targets.
- $\mathbf{Q^2 \approx 3.5 \text{ GeV}^2}$ 
  - Reduces non-nucleonic currents (MEC, IC).
  - Proton energies high enough for Glauber FSI calculations.
- $\mathbf{x_B = Q^2/2m\omega > 1.2}$  to minimize non-nucleonic currents.
- $\mathbf{\theta_{p,m,q} < 50^\circ}$  to minimize FSI.
- Two Kinematics:
  - $\mathbf{350 < p_{\text{miss}} < 600 \text{ MeV/c}}$  (“SRC”)
  - $\mathbf{p_{\text{miss}} < 250 \text{ MeV/c}}$  (“Mean-Field”)



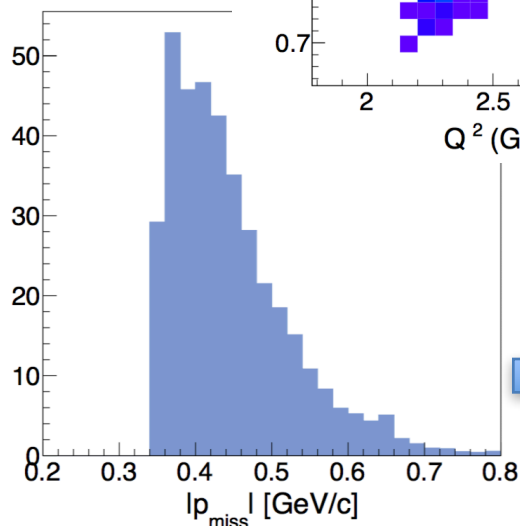
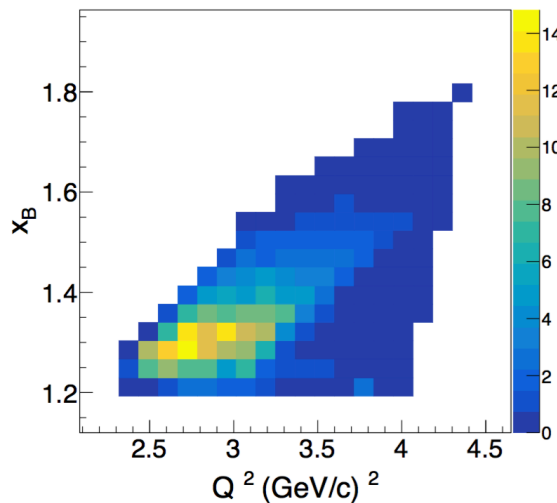
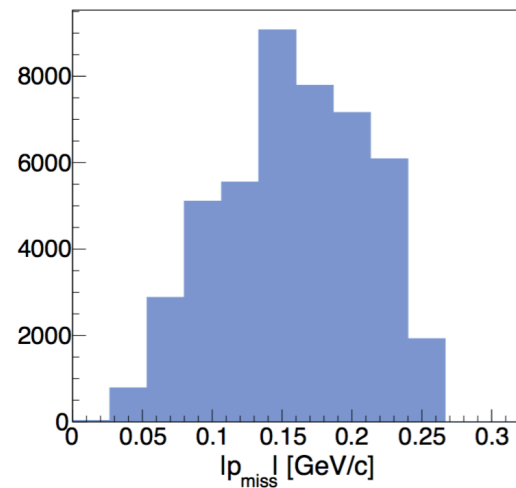
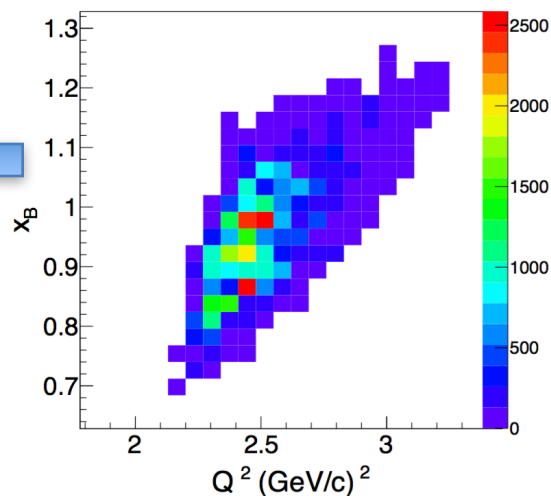
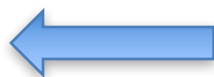


# Kinematics and Acceptance



$Q^2$	$E_{\text{Beam}}$	$E'_e$	$\theta_e$	$ \mathbf{p}_p $	$\theta_p$	$p_{\text{miss}}$
GeV <sup>2</sup>	GeV	GeV		GeV/c		GeV/c
3.5	11	9.85	10.0°	1.80	43.0°	0.45
3.5	11	9.85	10.0°	1.90	42.5°	0.15

Mean-Field Kinematics



SRC Kinematics



# The Final Observable



## (e,e'p) Cross-section Ratios for:

$$\frac{A_1(e, e'p)|_{SRC}}{A_1(e, e'p)|_{Mean-Field}} / \frac{A_2(e, e'p)|_{SRC}}{A_2(e, e'p)|_{Mean-Field}}$$

- Proportional to the relative number of correlated protons with minimal (experimental and theoretical) uncertainties.
- Tells us how many np-SRC pairs the extra  $f_{7/2}$  neutrons create!
- Reduced transparency corrections: Hen et al., Phys. Lett. B 722, 63 (2013)
  - $A^{1/3}$  dependence, T( $^{48}\text{Ca}/^{40}\text{Ca}$ )  $\sim$  6% correction.
  - Mean-field / SRC ratio further reduce the correction.
  - Well understood benchmark for symmetric nuclei ( $^2\text{H}$ ,  $^{12}\text{C}$ ,  $^{40}\text{Ca}$ ,  $^{54}\text{Fe}$ ).

# Measurement Plan and Beam-Time Request

## 4 days @ 11 GeV

kinematics	Target	Data-Taking [Hours]	Current [uA]	Expected Number of Events
high $p_{miss}$	$^2\text{H}$	12	40	4000
	$^{12}\text{C}$	10	40	4000
	$^{40}\text{Ca}$	14	40	4000
	$^{48}\text{Ca}$	14	40	4000
	$^{54}\text{Fe}$	14	40	4000
low $p_{miss}$	All Targets	16	40	10,000/Target
Commissioning and calibrations		12	20–60	
Target and Spectrometer Changes		4	N/A	
Total		96 (4 days)		



# Summary



**What?**

CaFe is a focused measurement of short-range pairing in neutron rich nuclei.

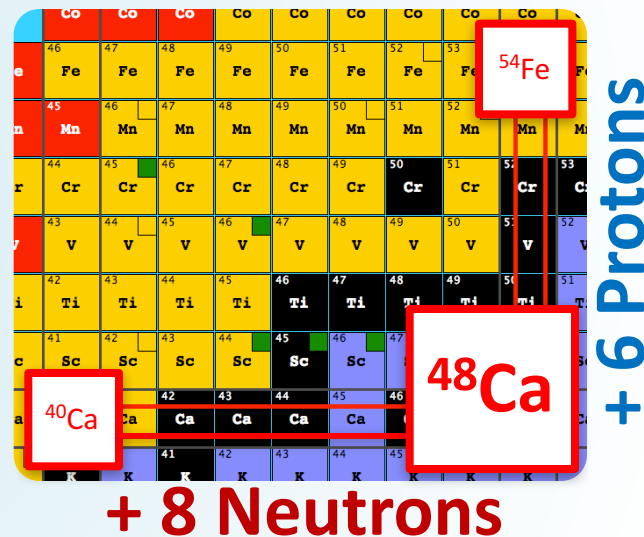
**Why?**

**What is the dynamics of the neutron skin in  $^{48}\text{Ca}$ ?**

Do correlations impact:

- Movement of nucleons in the skin?
- Extraction of static properties like the skin size?

Implications include neutron stars, the EMC effect and more...



**How?**

We will measure mean-field and correlated nucleons using the  $(e, e'p)$  reaction in  $^2\text{H}$ ,  $^{12}\text{C}$ ,  $^{40}\text{Ca}$ ,  $^{48}\text{Ca}$  and  $^{54}\text{Fe}$

We request 4 days at Hall-C using SHMS and HMS in standard configuration.







# Additional Observables (Benchmarks)

Using  $(e, e'p)$  ratios of symmetric nuclei to validate extraction procedure:

$$\frac{A_1(e, e'p)|_{Mean-Field}}{A_2(e, e'p)|_{Mean-Field}}$$

$$\frac{A_1(e, e'p)|_{SRC}}{A_2(e, e'p)|_{SRC}}$$

- ‘Traditional’ Transparency measurement.
- Well understood for symmetric nuclei ratios.
- Benchmark for Glauber calculations over a wide range of nuclei (from  $^2\text{H}$  to  $^{12}\text{C}$  to  $^{40}\text{Ca}$  and  $^{54}\text{Fe}$ ).
- ‘SRC’ Transparency measurement.
- Benchmark against  $(e, e')$  for symmetric nuclei ratios.
- Allows to extract physics in  $^{48}\text{Ca} / ^{40}\text{Ca}$  ratio where FSI corrections are small ( $\sim 6\%$ ).



## Tensor Correlations:

- Break the Fermi-Gas picture
- Reduce the kinetic symmetry energy (at  $\rho_0$ )
- Enhance the potential symmetry energy (at  $\rho_0$ )
- Softens the potential symmetry density dependence

But.... Still consistent with constrains from neutron stars observations!



# SRC and the Symmetry Energy



$$E_{sym}(\rho) = E_{sym}^{kin}(\rho_0) \cdot \left(\frac{\rho}{\rho_0}\right)^\alpha + E_{sym}^{pot}(\rho_0) \cdot \left(\frac{\rho}{\rho_0}\right)^{\gamma_i}$$

